

National Aeronautics and Space Administration



Fermi
Gamma-ray Space Telescope

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Detection of GeV Emission from Starburst Galaxies with the Fermi LAT

Keith Bechtol on behalf of the
Fermi LAT Collaboration

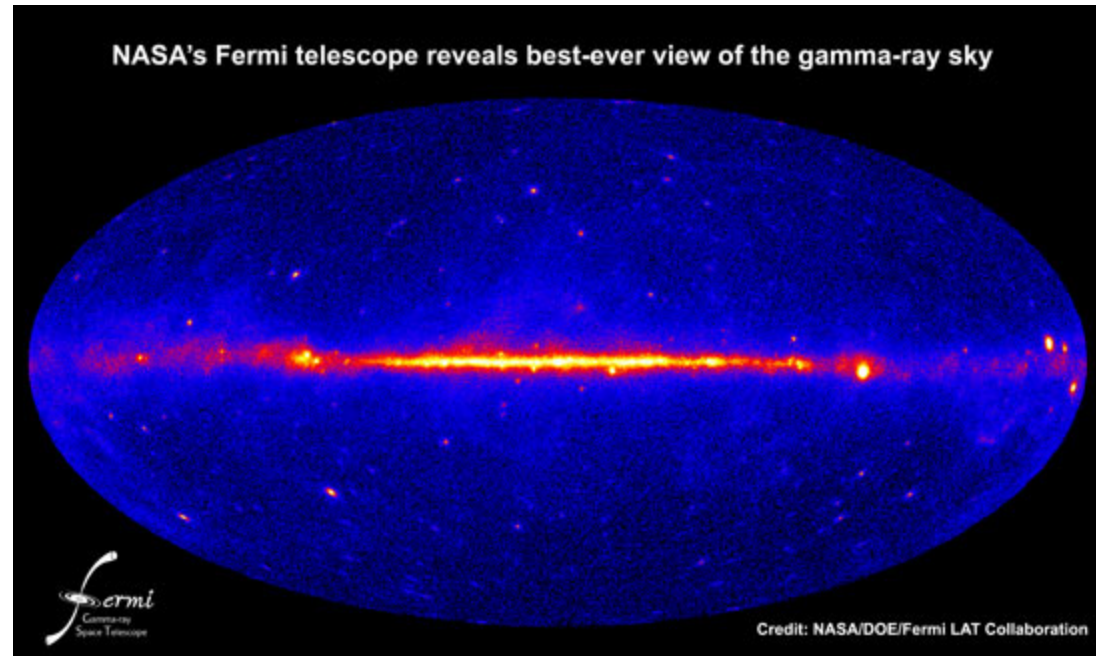
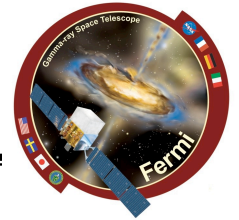
4 November 2009





- **Fermi LAT has detected steady, point-like, emission above 200 MeV from two starburst galaxies**
 - **M82 (6.8σ)**
 - **NGC 253 (4.8σ)**
- **Diffuse gamma-ray emission from star-forming galaxies**
 - **Starburst galaxies M82 and NGC 253**
- **Observations and results**
 - **Detection significance maps**
 - **Point-like and steady**
 - **Integral fluxes consistent with galactic diffuse emission**
- **Interpretation**
 - **Correlate star-formation with enhanced cosmic-ray intensity**

Diffuse Galactic Gamma-ray Emission

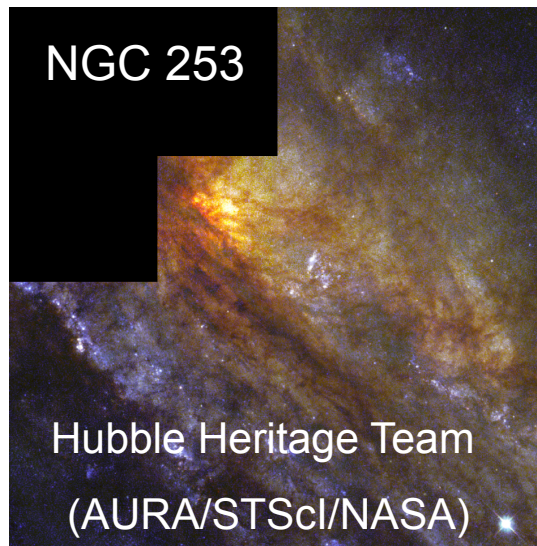
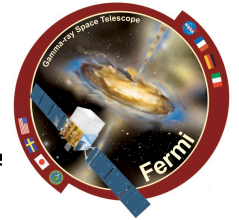


The most striking feature of the GeV gamma-ray sky is the diffuse emission from our own galaxy

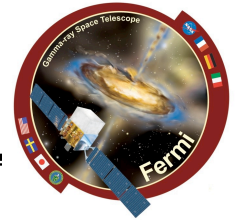
Cosmic rays interacting with interstellar medium

- $\text{CR}_{\text{protons}} + \text{gas} \rightarrow \text{neutral pion-decay}$
- $\text{CR}_{\text{electrons}} + \text{radiation fields} \rightarrow \text{inverse Compton}$
- $\text{CR}_{\text{electrons}} + \text{ambient protons} \rightarrow \text{bremsstrahlung}$

Starburst Galaxies

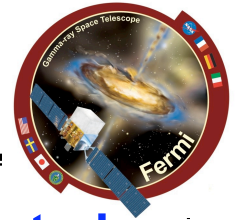


- **Starburst galaxies distinguished by regions of rapid star formation, 10-1000 × Milky Way rate**
 - Correspondingly high supernovae rates
 - Dense clumps of molecular gas
 - Highly luminous at infrared wavelengths, radio correlation
- **M82 and NGC 253**
 - Two closest starburst galaxies (~3 Mpc)
 - Edge-on viewing angles
 - Small (~100 pc scale) starburst regions
 - Star formation rate ~10 × Milky Way rate
 - Lack active nuclei
 - Extensively studied in multiple wavebands, detailed modeling/predictions

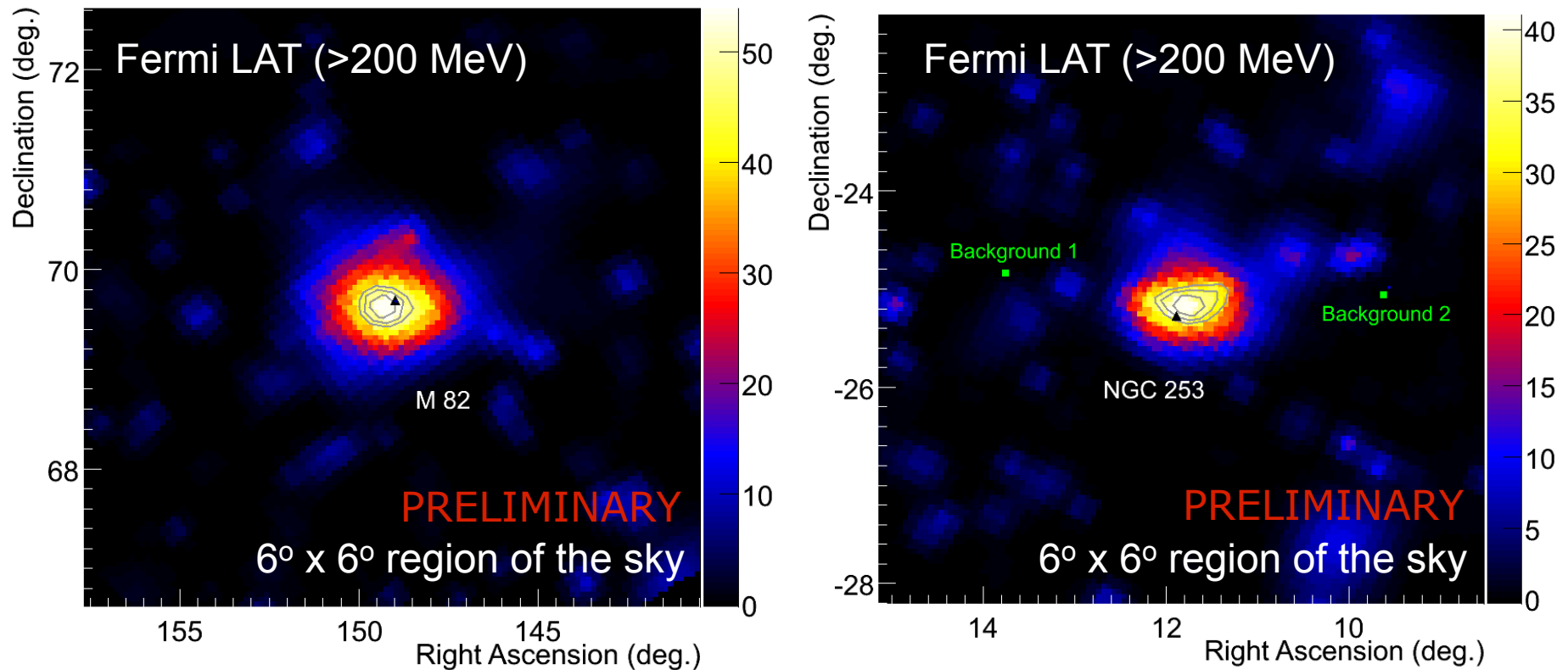


- **Observation periods**
 - **11 months (August 2008 to July 2009)**
 - Exclude time periods when Earth limb enters field of view (rocking angle cut at 43°)
- **Event selection**
 - **Energy > 200 MeV**
 - **Diffuse class (cosmic-ray background contamination < 10 %)**
 - **Zenith angle < 105°**
 - **Post launch instrument response functions (P6_V3)**
- **Model the region**
 - **10° radius region of interest**
 - **Galactic diffuse (template / hybrid / ring) + isotropic component**
 - **Include all significantly detected LAT sources within region of interest (1st year Catalog)**

Detection Significance Maps



Galactic diffuse, isotropic diffuse, and point sources subtracted

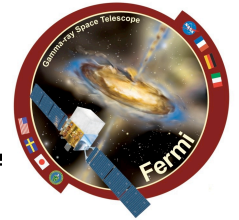


$$\text{Test Statistic (TS)} = -2 \log(L_{\text{source}} - L_{\text{no source}})$$

0.68, 0.95, 0.99 confidence level localization contours

Appear as LAT point sources, starburst regions unresolved

Point-like and steady emission

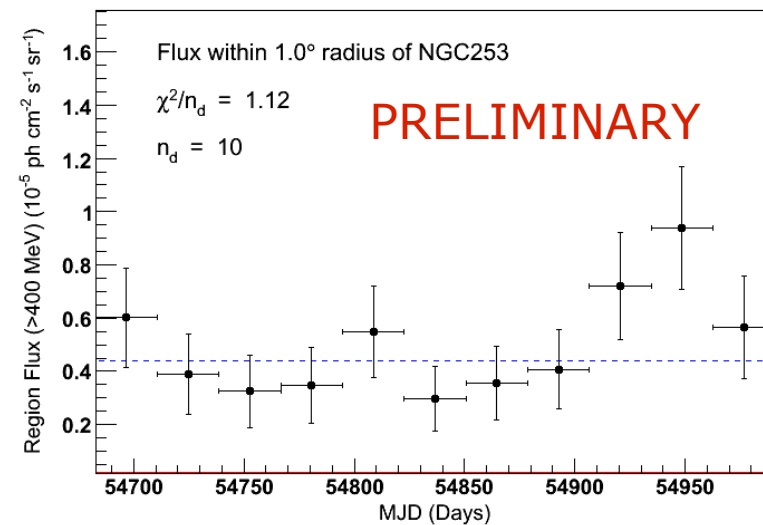
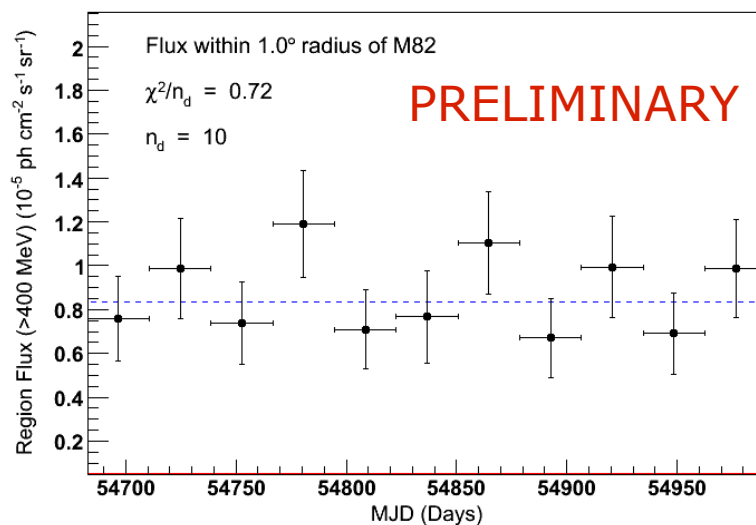


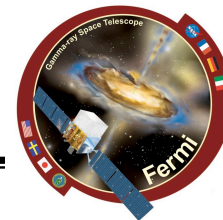
- **Spatial extension upper limits**
 - **Gaussian model, radius which contains 68 percent of flux (R_{68})**

| | R_{68} 95% confidence upper limit | Disk extent (optical) |
|----------------|-------------------------------------|-----------------------|
| M82 | 0.18° | 0.19° |
| NGC 253 | 0.30° | 0.45° |

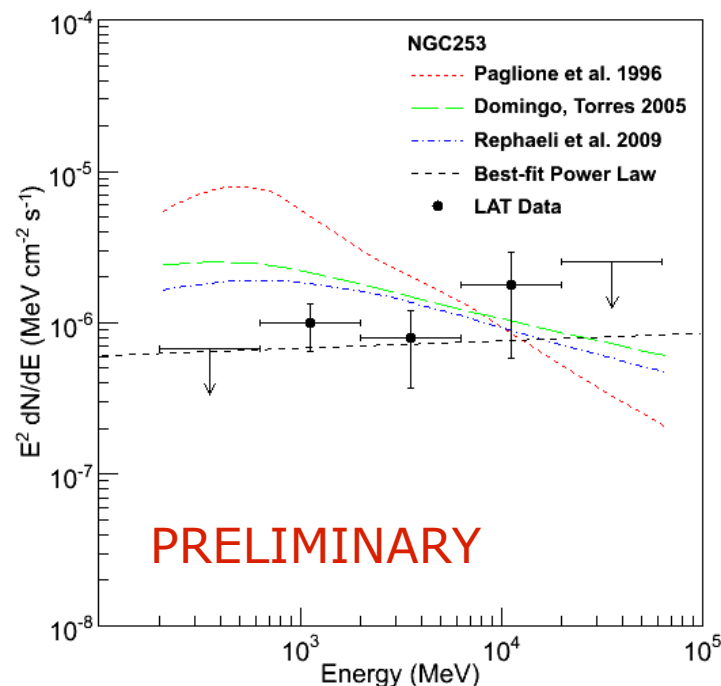
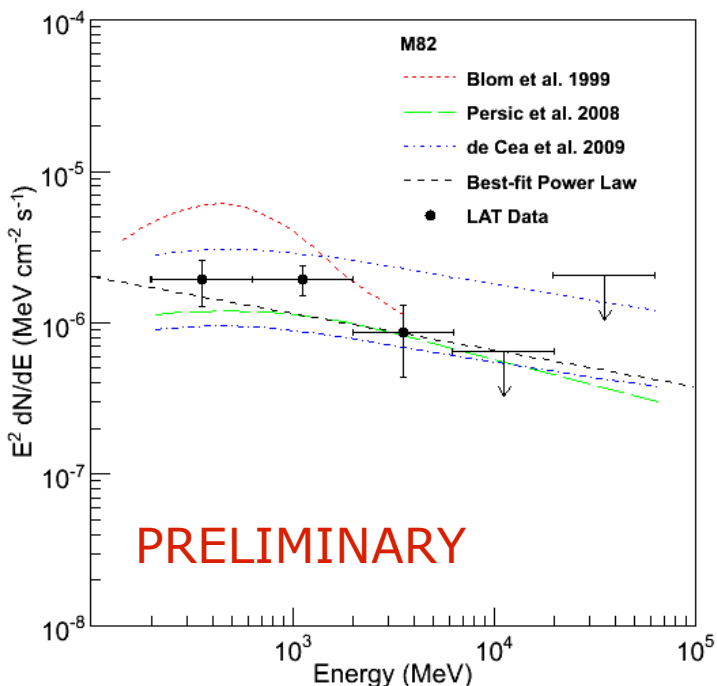
- **Both sources consistent with constant flux level**

Monthly count of the photons in the direction of each source (Aug 08 - Jul 09)





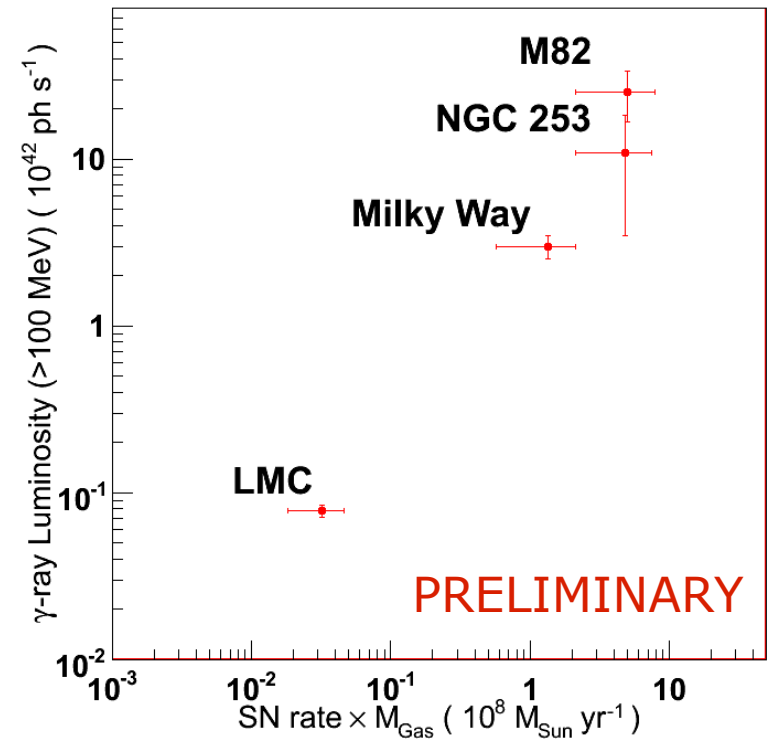
Observed integral fluxes consistent with models of diffuse galactic gamma-ray emission, but data do not yet tightly constrain spectral shapes



| | Flux (>100 MeV) (10^{-8} ph cm^{-2} s^{-1}) | Photon Index |
|----------------|---|--|
| M82 | $1.6 \pm 0.5_{\text{stat}} \pm 0.3_{\text{sys}}$ | $2.2 \pm 0.2_{\text{stat}} \pm 0.05_{\text{sys}}$ |
| NGC 253 | $0.6 \pm 0.4_{\text{stat}} \pm 0.4_{\text{sys}}$ | $1.95 \pm 0.4_{\text{stat}} \pm 0.05_{\text{sys}}$ |



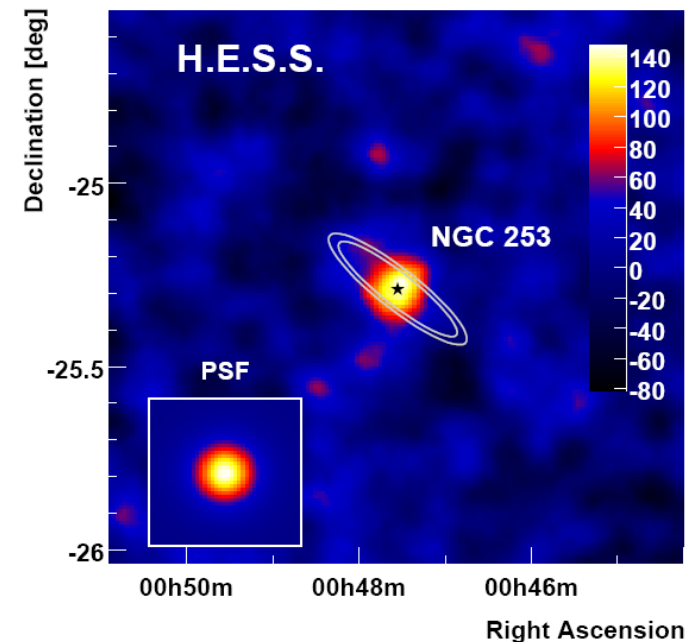
- **Enhanced cosmic-ray intensity required to explain the observed starburst gamma-ray fluxes**
- **Supernovae trace massive star formation**
 - **Supernovae remnants are probable source of galactic cosmic rays, although data not yet conclusive**
- **Beware!**
 - **Star-formation rate and gas density *non-uniform* throughout galaxies (Resolved LMC gamma-ray image)**
 - **Large uncertainty in distance measurements**



Compare gamma-ray luminosity and product of supernovae rate and gas mass in each galaxy



- **TeV detections summer 2009**
 - **NGC 253 (H.E.S.S.)**
 - **M82 (VERITAS)**
- **Fermi LAT results combined with the TeV data will fill in spectral energy distributions**
 - **Discern spectral shapes with greater certainty, constrain emission mechanisms**
- **TeV observations confirm steady sources**
- **Starbursts unresolved, TeV emission predominantly in central region**
- **LAT all-sky survey can point out additional candidates for TeV observatories**



Summary and Outlook



- **Two starburst galaxies, M82 and NGC 253, detected in 1st year of the Fermi mission**
 - **New class of gamma-ray sources**
 - **Diffuse emission arising from cosmic-ray interactions**
- **Searching for the source of galactic cosmic rays**
 - **Cosmic-ray intensity linked to massive star formation**
- **Ongoing Fermi LAT all-sky survey will look for additional starbursts and other star-forming galaxies**
 - **Normal star-forming galaxies are much more common than active galaxies but most often are fainter and unresolved**