



Fermi

Gamma-ray Space Telescope

Fermi LAT observations of high- and intermediate-velocity clouds: implications for cosmic rays

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Seth Digel

on behalf of
the *Fermi*-LAT collaboration

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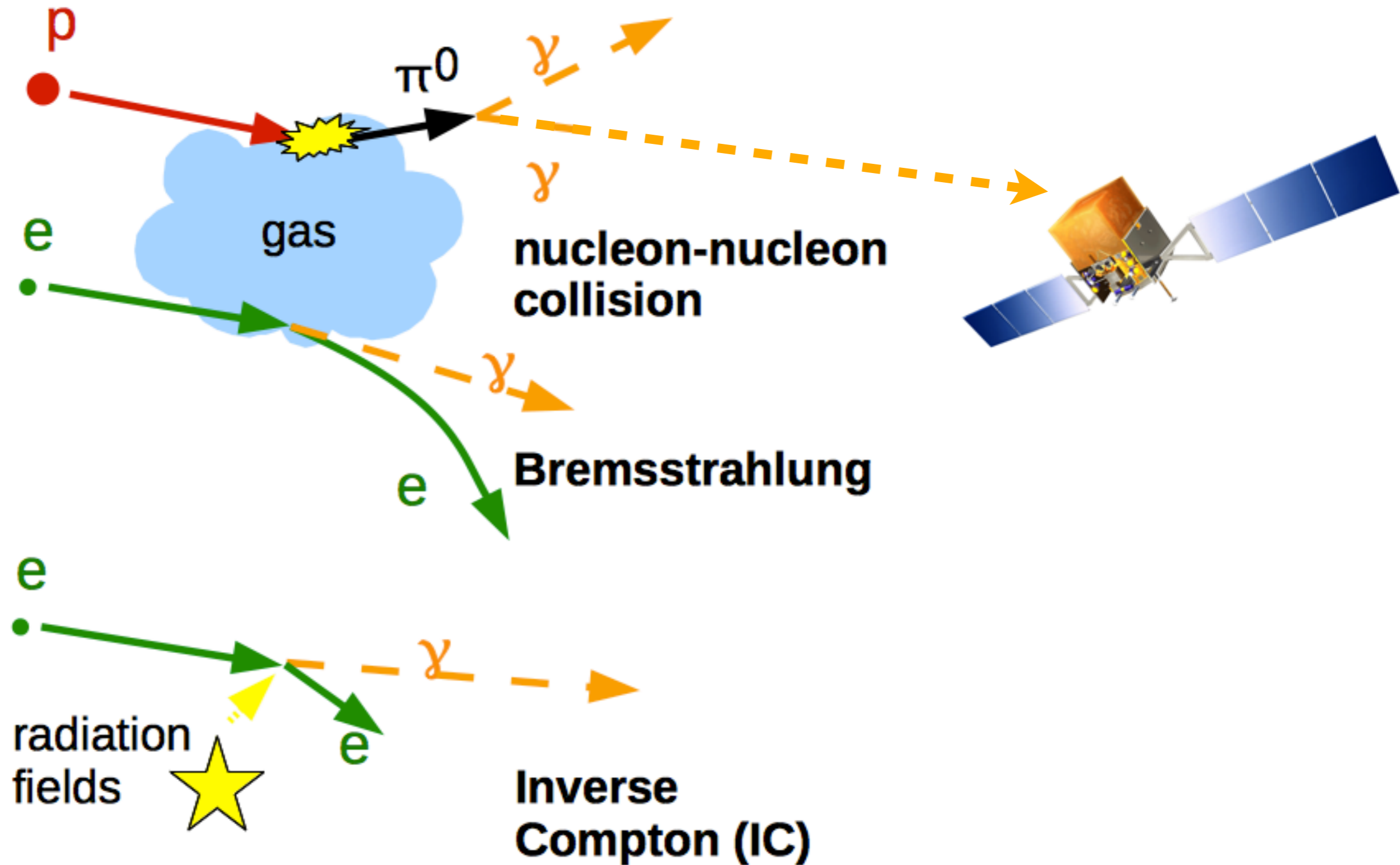
The Galactic cosmic-ray paradigm

below 10^{15} eV

- origin in the disk of the Milky Way
- propagate in $>$ kpc halo (indirect evidence)



γ rays as a charged particle tracer

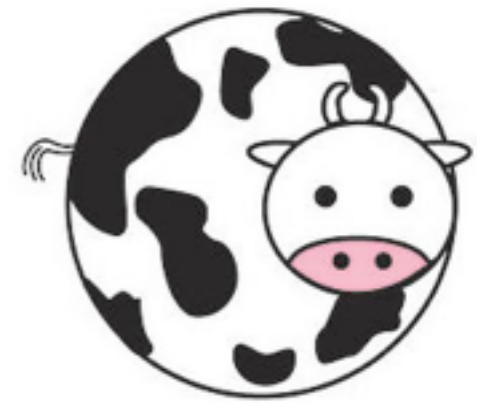


Chasing cosmic rays in the halo

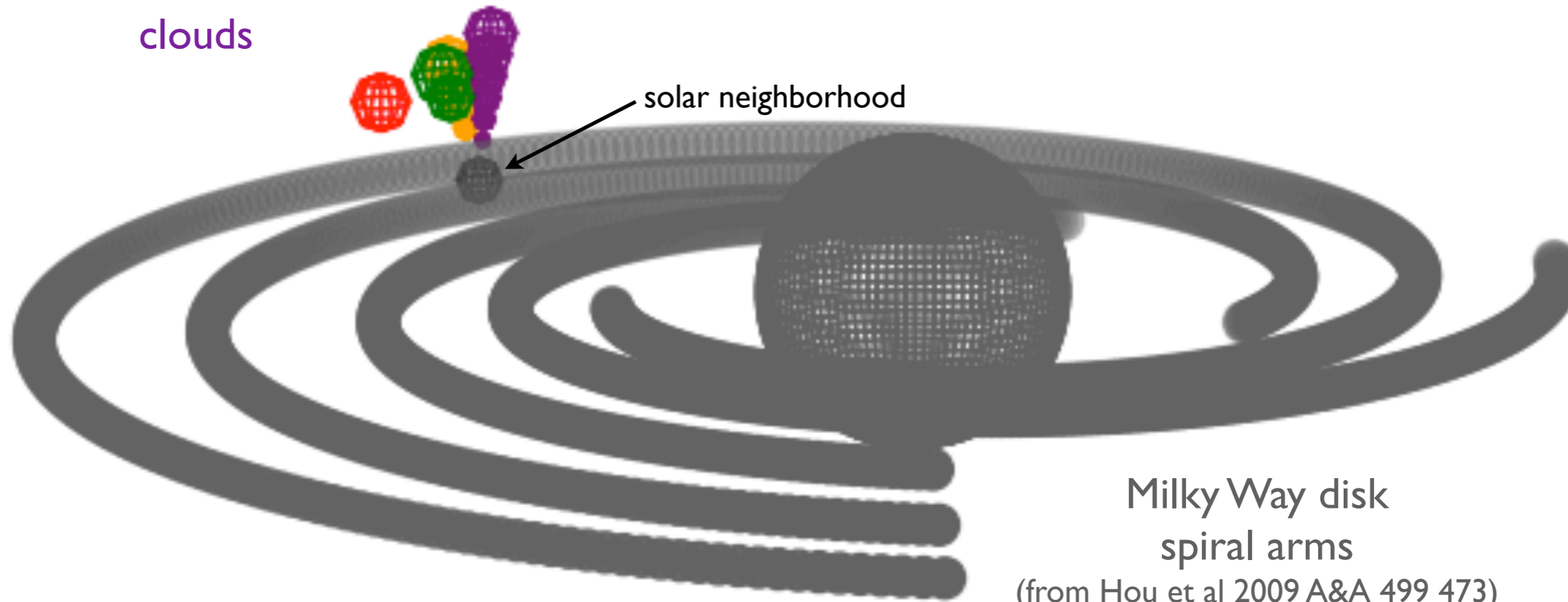
High-Velocity (HV) Cloud
infalling fresh gas

Intermediate-Velocity (IV) Cloud
gas ejected from disk

Target selection



target
HV and IV
clouds



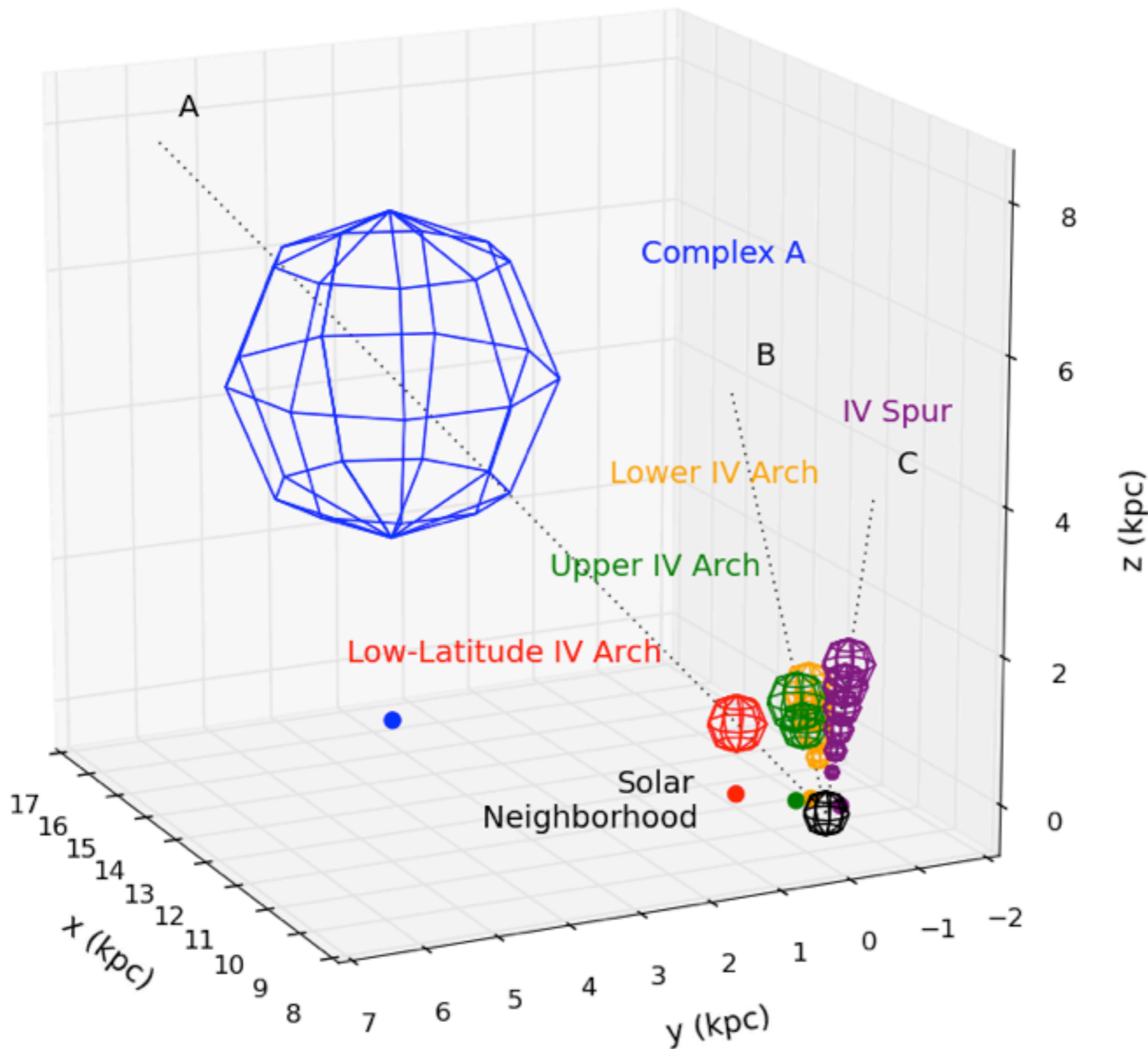
Milky Way disk
spiral arms
(from Hou et al 2009 A&A 499 473)

selection criteria

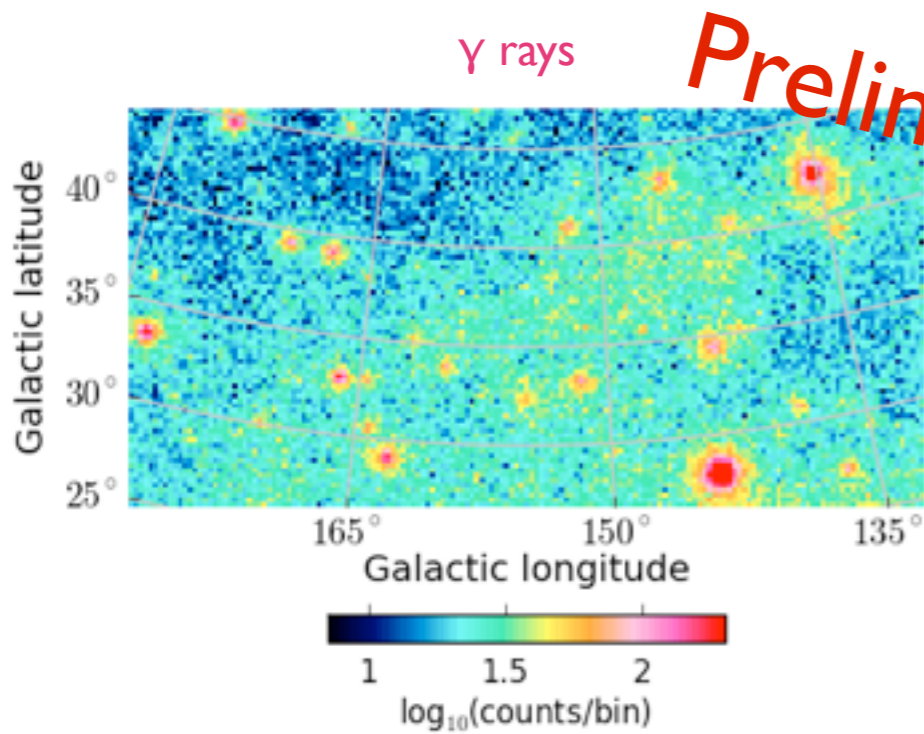
- distance/altitude bracket (stars)
- mass - distance detectable by the LAT

Meet the targets

distance/altitude brackets
Wakker 2001 ApJS 136 463



Analysis method



$$= \text{LAT PSF} \otimes \text{exp} \cdot$$

interstellar medium tracer data
 LAB HI survey
 CfA CO survey
 Planck dust thermal emission model

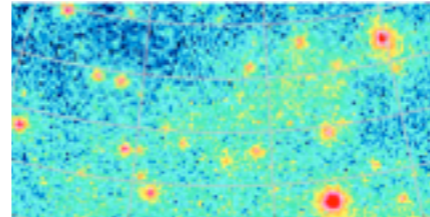
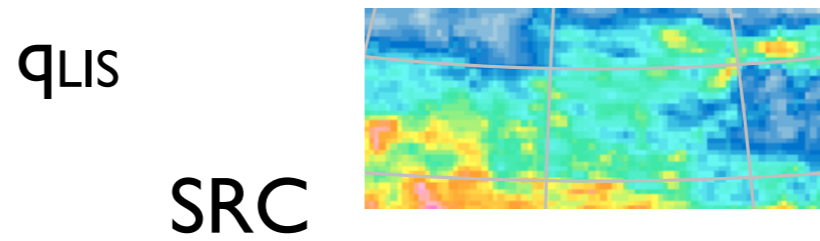
$$\begin{aligned}
 & \left[X_{\text{HI},1} \cdot \text{local HI} + X_{\text{HI},2} \cdot \text{HI Low-Latitude IV Arch} + X_{\text{HI},3} \cdot \text{HI Complex A} \right. \\
 & \left. + X_{\text{CO}} \cdot \text{CO (traces local H}_2\text{)} + X_{\text{DNM}} \cdot \text{dust excess (traces dark neutral medium)} \right] \cdot q_{\text{LIS}}(E) + \\
 & + \text{IC}(l,b,E) + \text{Isotropic}(E) + \text{3FGL}
 \end{aligned}$$

scaling factor
 γ -ray emissivity
 cosmic rays

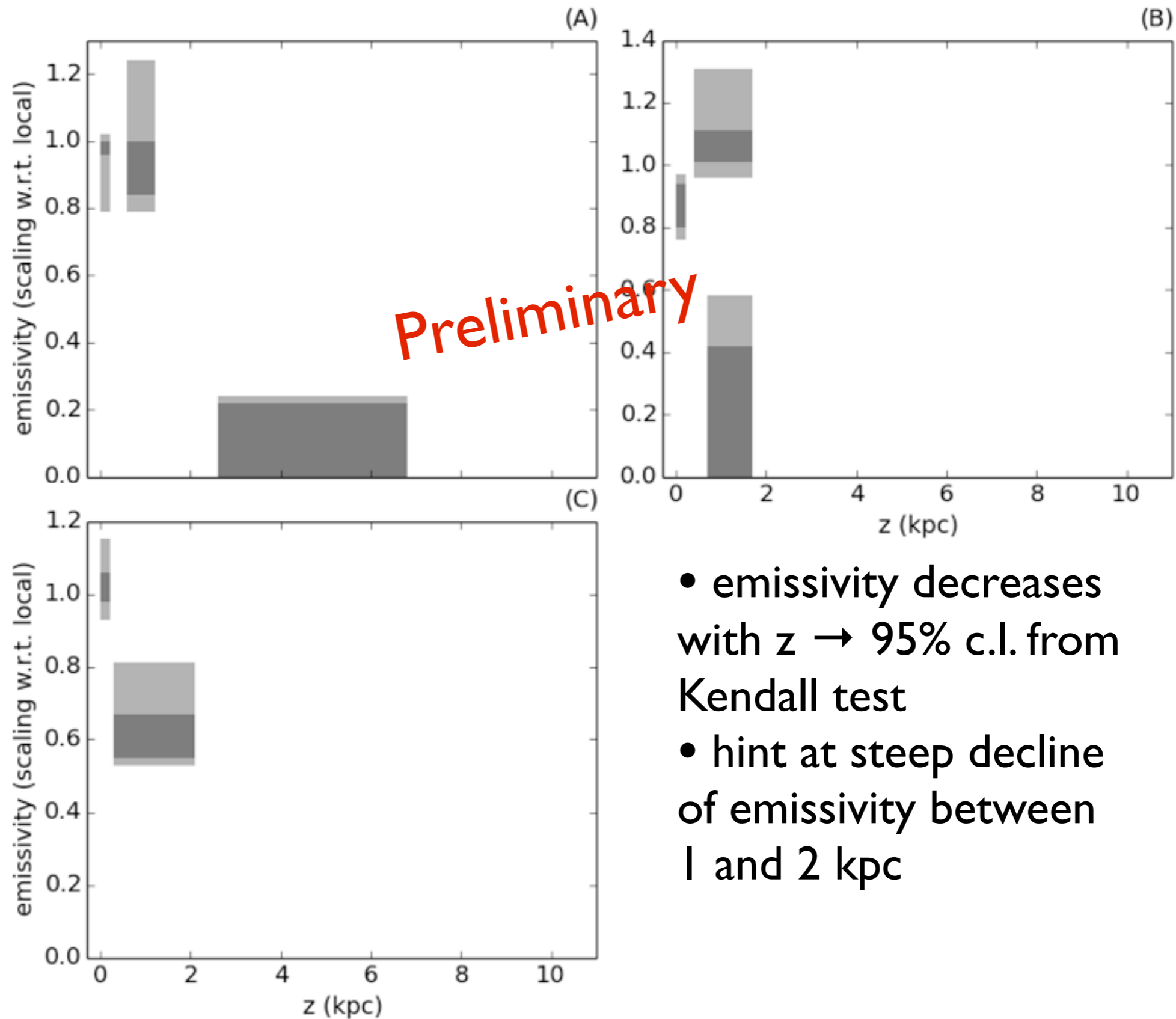
γ -ray emission rate (emissivity) per H atom per dE in Local Interstellar Space Casandjian (2014)

LAT analysis

- 73 months of P7REP data
- binned likelihood analysis
300 MeV - 10 GeV
- systematics:
 - uncertainties in model input
 - jackknife
 - LAT instrument response

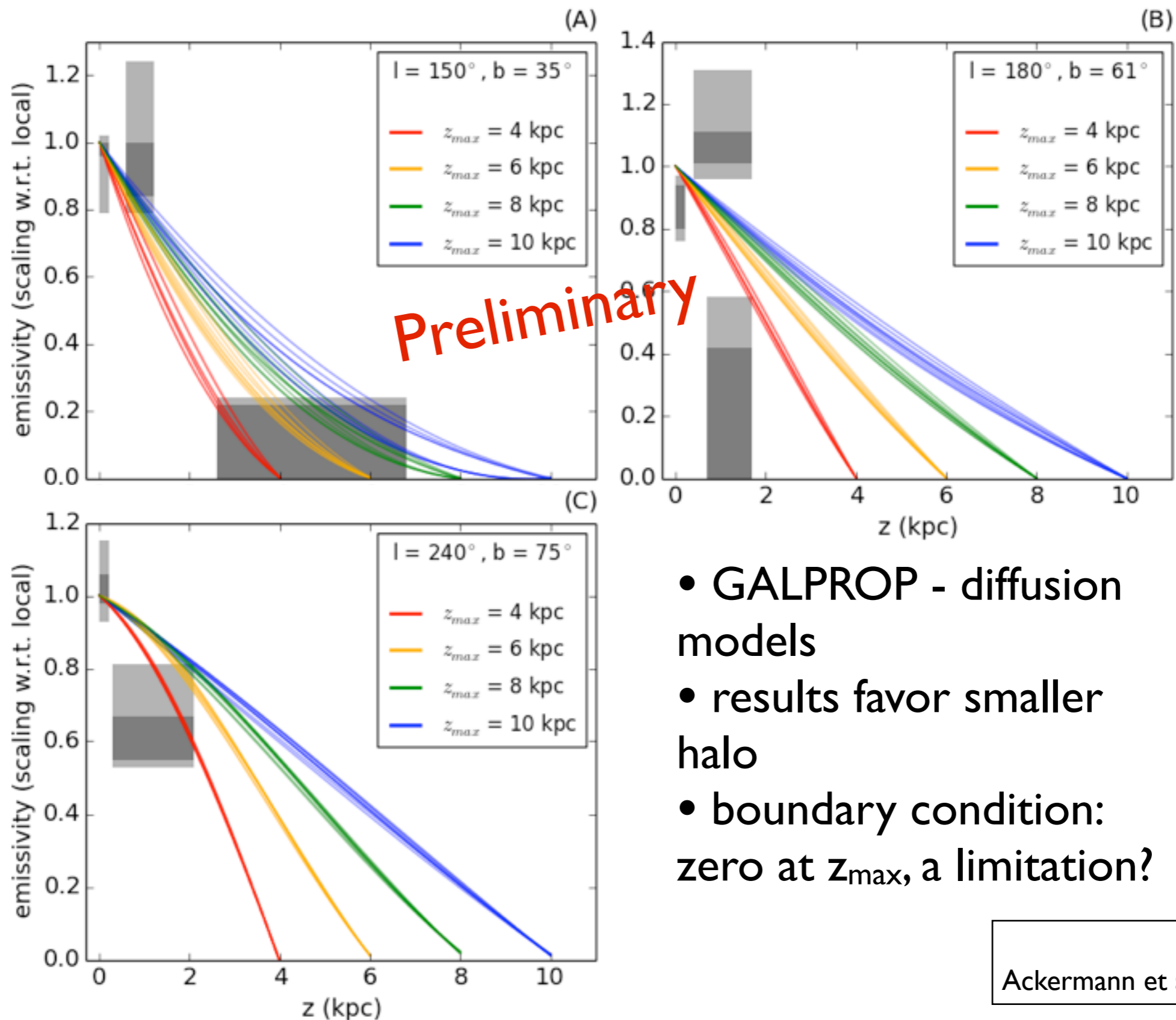


Comparison to propagation models



- emissivity decreases with $z \rightarrow$ 95% c.l. from Kendall test
- hint at steep decline of emissivity between 1 and 2 kpc

Comparison to propagation models



- GALPROP - diffusion models
- results favor smaller halo
- boundary condition: zero at z_{max} , a limitation?

models from
Ackermann et al 2012 ApJ 750 3

Final remarks

- HV and IV clouds are suitable targets to measure cosmic rays in the Milky Way halo with LAT data
- decrease of cosmic-ray densities with distance from the disk at 95% c.l. → corroborates origin in the disk
- hint at steep decline of densities between 1 kpc and 2 kpc above the disk
- can constrain cosmic-ray propagation and outflow into extragalactic space