



# Fermi Mission Status and Plans for the Future

Judy Racusin (NASA/GSFC) on behalf of the Fermi Mission Team



### Fermi Status



### Fermi Spacecraft & Operations

- Continues to operate as expected
- Closely monitoring performance of all observatory subsystems, no degradation of observatory performance
- Observation Modes
  - Dec 2013 Dec 2014 in Galactic center biased survey mode
  - Currently in 50 degree rock sky survey
  - In last year:
    - 3 Target of Opportunity (ToO) observations (~20 days)
    - 25 Autonomous Repoint Requests (~2.6 days)

### Large Area Telescope (LAT)

- Major analysis upgrade with Pass 8 event reconstruction pipeline
- New catalogs

### **Gamma-ray Burst Monitor (GBM)**

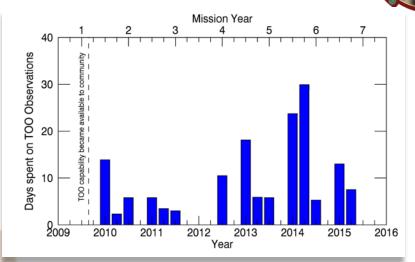
- New localization contours
- Ongoing work to improve automation (RoboBA)



### Fermi Status: Observations



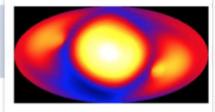
- Continues to operate as expected
- Closely monitoring performance of all observatory subsystems, no degradation of observatory performance
- Observation Modes
  - Dec 2013 Dec 2014 in Galactic center biased survey mode
  - Currently in 50 degree rock sky survey
  - In last year:
    - 3 Target of Opportunity (ToO) observations (~20 days)
    - 25 Autonomous Repoint Requests (~2.6 days)



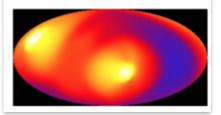
**Galactic Center Pointing** 

TOOs since last Symposium:

- 3C279
- Nova SGR 2015 No.2
- GRB 150201A



50 degree rocking survey mode



More information on Fermi TOOs: http://fermi.gsfc.nasa.gov/ssc/observations/too/



# Fermi Status: Large Area Telescope (LAT)



# Pass 8 improvements in:

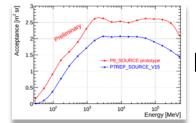
- acceptance
- effective area
- energy resolution
- PSF
- sensitivity
- field of view

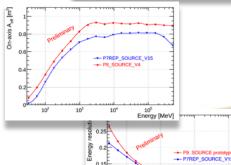
P8 public release June 24, 2015

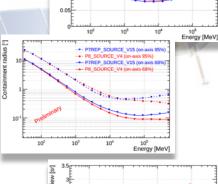
### More Pass 8 details in:

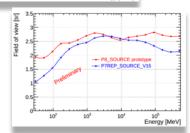
- talk by M. Wood
- posters by L. Baldini,
   E. Bloom, M. Testa,
   M. Wood

P8 Results throughout the Symposium



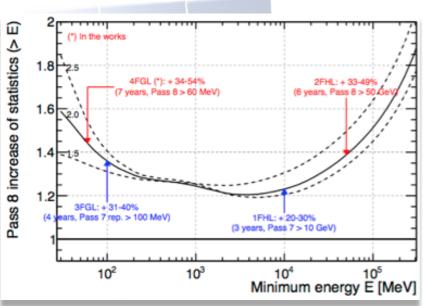






### Large Area Telescope (LAT)

- Major analysis upgrade with Pass 8 event reconstruction pipeline
- New catalogs



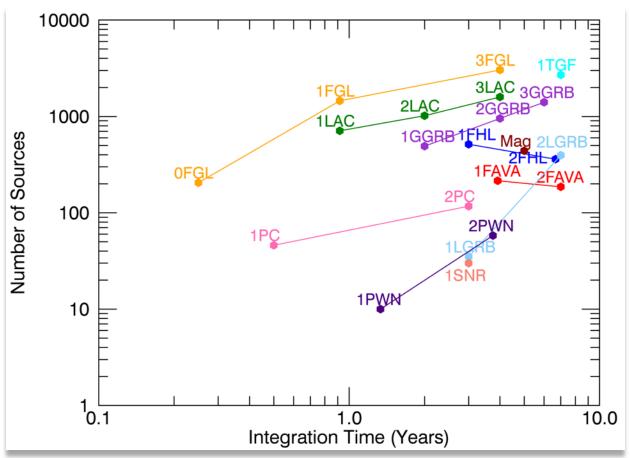


### Fermi Status: Catalogs



### Fermi Catalogs

- LAT
  - FGL (General)
  - FHL (High-energy)
  - LAC (AGN)
  - PC (Pulsars)
  - LGRB (GRBs)
  - FAVA (Flaring sources)
  - SNR (supernova remnants)
  - Solar flares (upcoming)
- GBM
  - GGRB (GRBs)
  - Mag (Magnetar bursts)
  - TGF

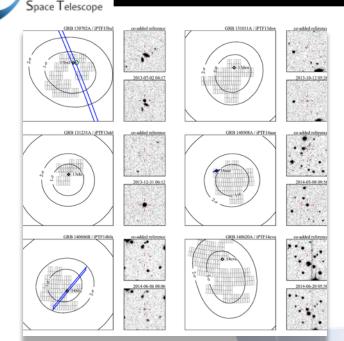


See talks by M. Ajello, D. Kocevski See posters by G. Vianello, T. Brandt, G. Fitzpatrick, A. Allafort, D. Yu

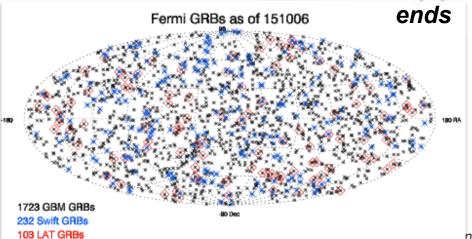
### Sermi Gamma-ray

# <sub>mໍເ</sub> Fermi Status: Gamma-ray Burst Monitor (GBM)





#### Singer et al. (2015)



#### **New Localization contours**

- include statistical and asymmetric systematic errors (Connaughton et al. 2015)
- Automatically generated and distributed via GCN
- Useful for follow-up with wide-FoV optical telescopes (e.g. iPTF, MASTER)
- Especially important for LIGO/Virgo RoboBA (coming soon)
- Ground automated positions to ~4.5° radius + contours ~1 minute after trigger

### **Gamma-ray Burst Monitor (GBM)**

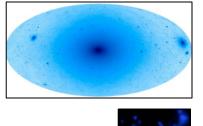
- New localization contours
- Ongoing work to improve automation (RoboBA)

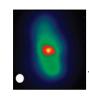


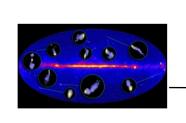
### Fermi Science Menu

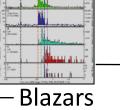








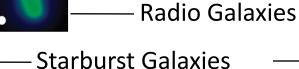




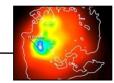


**GRBs** 





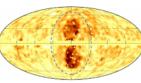




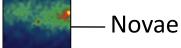


**Globular Clusters** 

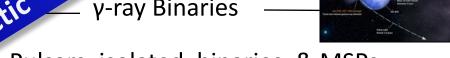








Galactic γ-ray Binaries



Pulsars: isolated, binaries, & MSPs

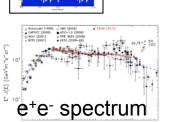


Terrestrial γ-ray Flashes

**Unidentified Sources** 



6th International Fermi Symposium - Arlington, \





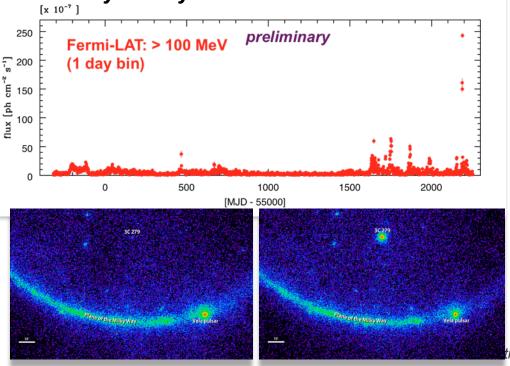
### Science Highlights: Time Variable Sources



### Blazar 3C279

- Most dynamic blazar flare ever seen (x10 in 1 day) in June 2015
- Fermi Target of Opportunity Observation allowed measure of short-timescale variability
- Triggered multiwavelength campaign

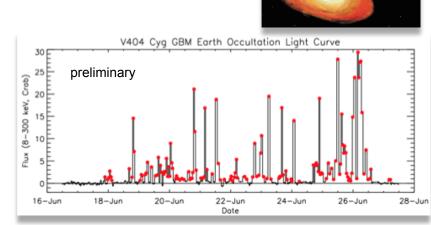
### See talk by M. Hayashida



# **Galactic Binary V404 Cyg**

- First outburst of this low mass Xray binary since 1989
- 169 GBM triggers June 15-27
- Also detected by MAXI, Swift, INTEGRAL, & many others in multiwavelength campaign

See talks by D. Huppenkothen & P. Jenke





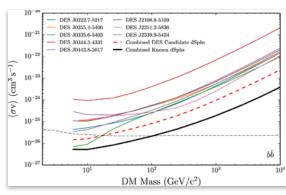
# **Science Highlights: Dark Matter**



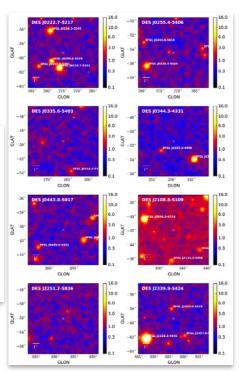
### dSph galaxies limits

- DES discovered 8 new dSph galaxies (DES collab, arXiv:1508.03622)
- LAT limits are most constraining yet (Drlica-Wagner et al. 2015)

### See talk by A. Drlica-Wagner, R. Caputo See posters by A. Geringer-Sameth, M. Mazziotta



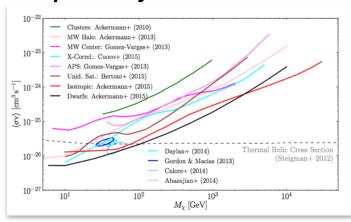
Drlica-Wagner et al. (2015)



### **Galactic Center Excess**

- 1-10 GeV excess within 10° of Galactic center
  - ~40 GeV DM annihilation?
  - or unresolved astrophysical sources?

See talks by A. Albert, D. Nieto, F. Donato, C. Weniger, B. Safdi, D. Malyshev, A. Viana
See poster by T. Linden



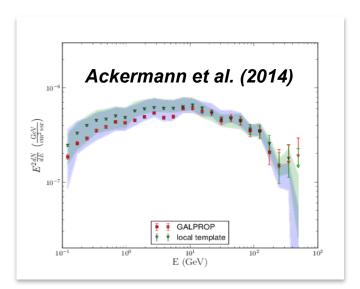


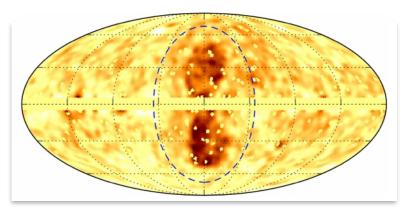
### Science Highlights: The Fermi Bubbles



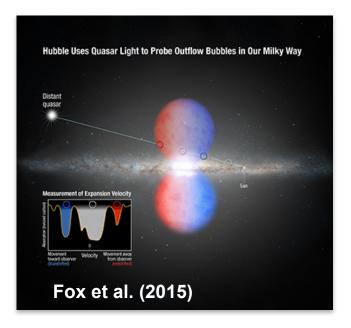
- High-energy cutoff at ~100 GeV
- Significant enhancement in southeastern region
- Evidence >900 km s<sup>-1</sup> wind via HST UV spectroscopy of quasar behind bubbles indicating Galactic Center activity in last ~2.5-4 Myr

#### See talks by M. Su & V. Dogiel





Ackermann et al. (2014)





# **Guest Investigator Program**



# Cycle 9 deadline: Jan 22, 2016 GI Program Details:

Funding for analysis of *Fermi* LAT and/or GBM data and/or correlative observations
Funding for theoretical studies related to *Fermi* 

Pointed mode or ToO observations NRAO, NOAO, Arecibo, VERITAS, INTEGRAL observations related to Fermi science

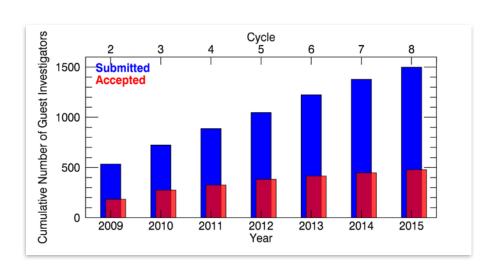
Funds are dispersed to GIs as soon as they are available

http://fermi.gsfc.nasa.gov/ssc/proposals/

GI program continuing to expand to new users

Average award has decreased to reconcile shrinking budget

Oversubscription rate 5:1 highest in astrophysics division





### Fermi Analysis Tools

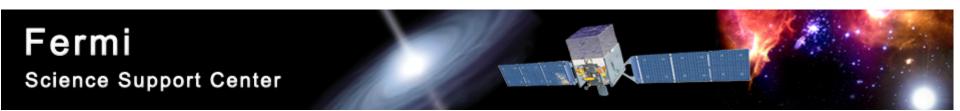


With Pass 8 data release, FSSC/LAT team released a major revision to the LAT Science Tools, documentation, and analysis threads (revised >300 individual files)

- changes/additions to event classes + Instrument Response Functions (quality, front/back, PSF, EDISP)
- new P8 diffuse models

#### **New User Contributed Software**

- GBM orbital background subtraction tool
- LAT XML manipulation tools
- http://fermi.gsfc.nasa.gov/ssc/data/analysis/user/







Goal: Utilize the full potential of Pass 8 and experience of 7 years of Fermi operations to efficiently search for short (<hours) and medium (~days) timescale transients

- Reduce data latency
- Transient search pipelines
- Streamline Target of Opportunity (ToO) process
- Expedite follow-up observations





Goal: Utilize the full potential of Pass 8 and experience of 7 years of Fermi operations to efficiently search for short (<hours) and medium (~days) timescale transients

- Reduce data latency
- Transient search pipelines
- Streamline Target of Opportunity (ToO) process
- Expedite follow-up observations

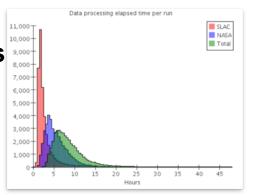
All data go public immediately

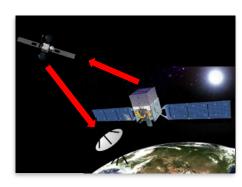
### **Data Latency Steps**

- spacecraft TDRSS
- ground network MOC
- MOC Instrument Processing Centers -FSSC

### **Improvements**

- FOT already implementing greater frequency of shorter TDRSS passes and new algorithm to chose passes
- reorder data subsets downlink order
- faster transfer from ground network to MOC





See poster by D. Thompson





Goal: Utilize the full potential of Pass 8 and experience of 7 years of Fermi operations to efficiently search for short (<hours) and medium (~days) timescale transients

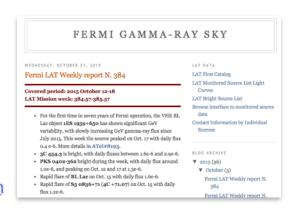
**Automatic Science Processing (ASP)** 

- Reduce data latency
- Transient search pipelines
- Streamline Target of Opportunity (ToO) process
- Expedite follow-up observations

6 hr & 1 day search timescales

Used by *Fermi* Flare Advocates for many transient detections (*Fermi* Sky Blog, ATels)

The Astronomer's Telegram



#### Fermi All Sky Variability Analysis (FAVA)

1 week search timescale (+3 day)

aperture photometry technique

Weeks Analyzed: 359
FAVA Detections (>5x): 2705

Associated: 2440
Unassociated: 265

See talk by D. Kocevski





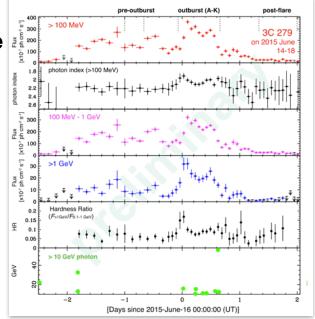
Goal: Utilize the full potential of Pass 8 and experience of 7 years of Fermi operations to efficiently search for short (<hours) and medium (~days) timescale transients

- Reduce data latency
- Transient search pipelines
- Streamline Target of Opportunity (ToO) process
- Expedite follow-up observations

Short-term (days-weeks) pointed observations

- increase exposure ~x2 above survey mode
- better quality data to measure short timescale variability and higher S/N spectra
- at expense of even exposure on the rest of sky

See talk by M. Hayashida for more on 3C279



http://fermi.gsfc.nasa.gov/ssc/observations/too/





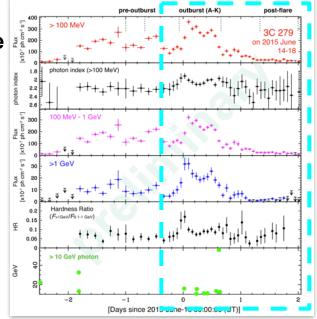
Goal: Utilize the full potential of Pass 8 and experience of 7 years of Fermi operations to efficiently search for short (<hours) and medium (~days) timescale transients

- Reduce data latency
- Transient search pipelines
- Streamline Target of Opportunity (ToO) process
- Expedite follow-up observations

Short-term (days-weeks) pointed observations

- increase exposure ~x2 above survey mode
- better quality data to measure short timescale variability and higher S/N spectra
- at expense of even exposure on the rest of sky

See talk by M. Hayashida for more on 3C279



http://fermi.gsfc.nasa.gov/ssc/observations/too/





Goal: Utilize the full potential of Pass 8 and experience of 7 years of Fermi operations to efficiently search for short (<hours) and medium (~days) timescale transients

- Reduce data latency
- Transient search pipelines
- Streamline Target of Opportunity (ToO) process
- Expedite follow-up observations

Fermi mission will promptly update community via Ferminews, ATel, GCN, direct communication with observers, etc. with ToO details to encourage multiwavelength coordination and follow-up

... Science!

see also Multiwavelength workshop on Friday

Fermi mailing lists: http://fermi.gsfc.nasa.gov/ssc/library/newsletter/ Fermi MW coordination: http://fermi.gsfc.nasa.gov/ssc/observations/multi/



# Optimizing the High Energy End of the LAT Energy Range



Goal: Utilize the full potential of Pass 8 and maximize the science return from the high energy end of the LAT

### **Sensitivity increases**

- faster at high energies (~t, photon limited)
- relative to low energies (~t<sup>1/2</sup>, background limited)

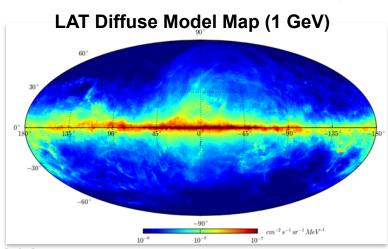
#### Diffuse emission model

- needed for all medium/long timescale analyses
- built from surveys of interstellar gas, Fermi data
- especially difficult at high energies where no templates exist and unique features (e. g. Fermi bubbles)

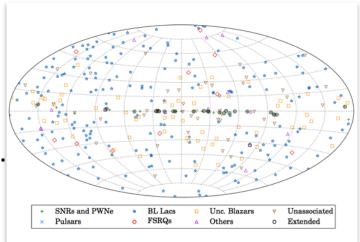
### **Updated catalogs**

needed for source populations at high energies (e.g. 2FHL)

See talks by J. Perkins, M. Ajello See poster by S. Bonnefoy



2FHL Catalog, Ackermann et al. 2015, arXiv: 1508.04449





# Maximizing the Scientific Potential of Long Baseline Observations

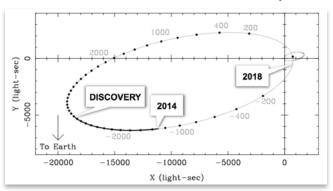


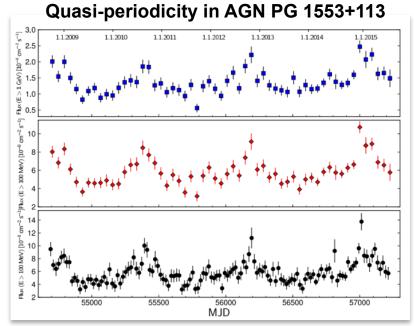
Fermi's unique ability to monitor the whole sky over the last 7 years has yielded rich datasets of variable sources

- Periodic
  - Binaries with periods of years
  - Solar flares
- Variable
  - AGN variability to correlate with multiwavelength studies
- Rare Events
  - Pulsar state transitions
  - Outbursts (e.g. Crab)

Deep stares require updated diffuse & sources catalogs
Long baseline variability requires regular calibration & understanding of the instrument stability

Pulsar/Be-star binary systems: PSR B1259-63 periastron outburst in 2011/2014, J2032+4127 may show similar in 2018





Ackermann et al. 2015, arXiv:1509.02063

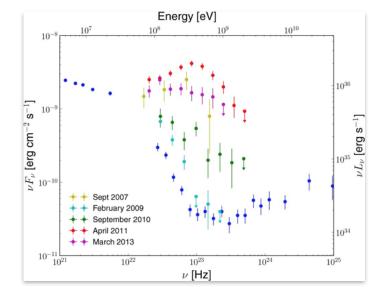
More on long baseline observations in talk by D. Thompson



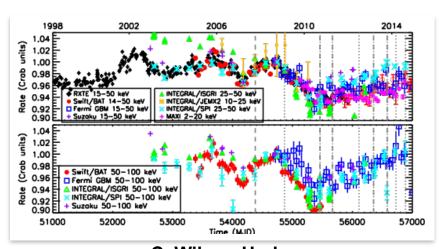
### The Future: The Crab



- Gamma-ray pulsations up to ~400 GeV (MAGIC Collab., arXiv: 1510.07048)
- Nebula has shown bright flares in high-energy gamma-rays lasting ~days with short-timescale variability (LAT, AGILE)
- Long-term Variability in hard X-ray "standard candle" (GBM)
- Emission regions and acceleration mechanisms not well understood
- Rapid response to flares from Crab or discovery of similar phenomena in other PWNe will be important to initiate multiwavelength follow-up



**Buehler & Blandford (2014)** 



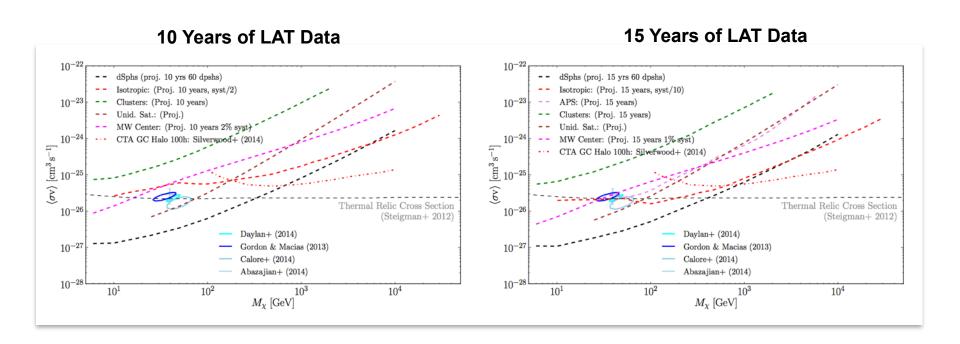
C. Wilson-Hodge



### The Future: Dark Matter



- Additional dwarf spheroidal galaxies will likely be discovered by large-scale optical surveys (e.g. DES)
- Mass range excluded by LAT will reach ~350 GeV over next 4 years







### Advanced LIGO/Virgo

- GBM is most likely instrument to detect and localize an electromagnetic counterpart to a binary neutron star merger (on axis)
- LAT all sky monitoring could also provide coincident transient source
- talks by P. Shawhan & V. Connaughton

#### IceCube PeV Neutrinos

- GRB and/or Blazar origin?
- talks by E. Waxman, M. Kadler & M. Santander

### **Pulsar Timing Arrays**

- Fermi continues to provide additional pulsars, and putative gravitational wave sources like possible SMBH binary PG 1553+113
- talk by S. Ciprini & P. Shawhan





### Advanced LIGO/Virgo

- GBM is most likely instrument to detect and localize an electromagnetic counterpart to a binary neutron star merger (on axis)
- LAT all sky monitoring could also provide coincident transient source
- · talks by P. Shawhan & V. Connaughton

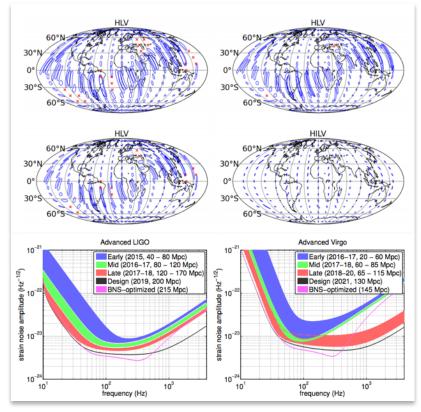
#### IceCube PeV Neutrinos

- extragalactic?
- talks by E. Waxman, M. Kadler & M. Santander

### **Pulsar Timing Arrays**

- Fermi continues to provide additional pulsars, and putative gravitational wave sources like possible SMBH binary PG 1553+113
- talk by S. Ciprini & P. Shawhan

GBM localizations will drastically reduce follow-up area, and will help to identify the host galaxy, redshift, environment, etc.



LIGO/Virgo Collaboration (2013)





### Advanced LIGO/Virgo

- GBM is most likely instrument to detect and localize an electromagnetic counterpart to a binary neutron star merger (on axis)
- LAT all sky monitoring could also provide coincident transient source
- talks by P. Shawhan & V. Connaughton

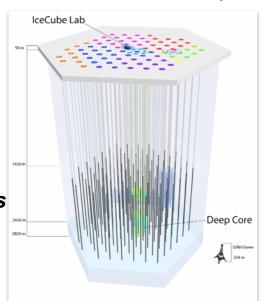
#### IceCube PeV Neutrinos

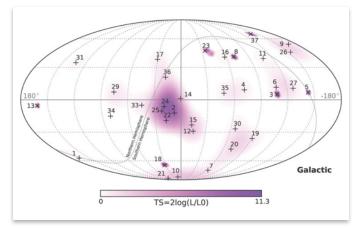
- extragalactic?
- talks by E. Waxman, M. Kadler & M. Santander

### **Pulsar Timing Arrays**

- Fermi continues to provide additional pulsars, and putative gravitational wave sources like possible SMBH binary PG 1553+113
- talk by S. Ciprini & P. Shawhan

Fermi's wide sky coverage provides unique capability to search for contemporaneous flaring in photon data and neutrinos





IceCube Collaboration (2013)





### Advanced LIGO/Virgo

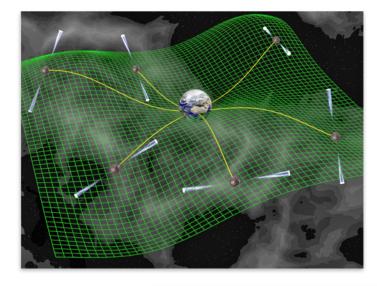
- GBM is most likely instrument to detect and localize an electromagnetic counterpart to a binary neutron star merger (on axis)
- LAT all sky monitoring could also provide coincident transient source
- talks by P. Shawhan & V. Connaughton

#### IceCube PeV Neutrinos

- extragalactic?
- talks by E. Waxman, M. Kadler & M. Santander

### **Pulsar Timing Arrays**

- Fermi continues to provide additional pulsars, and putative gravitational wave sources like possible SMBH binary PG 1553+113
- talk by S. Ciprini & P. Shawhan



binary periods of ~years are in frequency range of PTAs

evenly sampled all-sky data is ideal for searching for these periodicities



Ackermann et al. 2015, arXiv:1509.02063



### **NASA Astrophysics Senior Review**



- Every 2 years all operating missions in their extended phase compete for funding to continue operations
  - Missions in 2016 Senior Review: Fermi, Kepler (K2),
     NuSTAR, Spitzer, Swift, XMM
  - Chandra & Hubble separate process
- 2014 Panel Report
  - "The Fermi Observatory ... is a unique asset to the NASA portfolio"
  - "The Fermi GI program has been very successful, and has directly led to several important science discoveries."
  - "The SRP recommends continuation of the Fermi extended mission through FY18"

http://science.nasa.gov/astrophysics/2014-senior-review-operating-missions/



### **Conclusions**



- 2016 Astrophysics Senior Review Proposal
  - Propose to extend the *Fermi* mission through to 2020
  - Draft proposal in preparation due Jan 22, 2016
  - The Fermi Mission welcomes input from the community, especially throughout the Fermi Symposium
- Please continue to think of new and innovative ways to use the Fermi instruments and data
- Looking forward to many interesting results this week!



### **Advertisements**



# **Tooning the Extreme Cosmos**

### Free tickets still available!

http://fermi.gsfc.nasa.gov/science/mtgs/tooning/

# Take a wander into DC for a special mixture of *Fermi* and Art



# Next Huntsville GRB Workshop

October 24-28, 2016 in Huntsville, Alabama

Organizers: Valerie Connaughton, Neil Gehrels, Adam Goldstein

**Details soon!**