



Fermi  
Gamma-ray Space Telescope

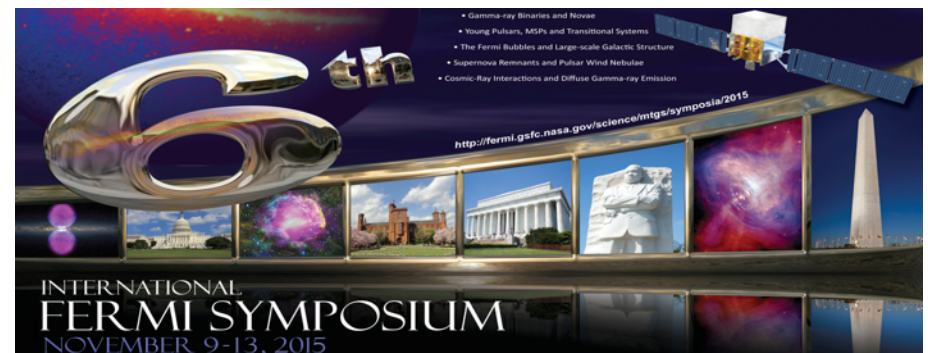
# Cosmic Ray Electron Spectrum with the Fermi-LAT

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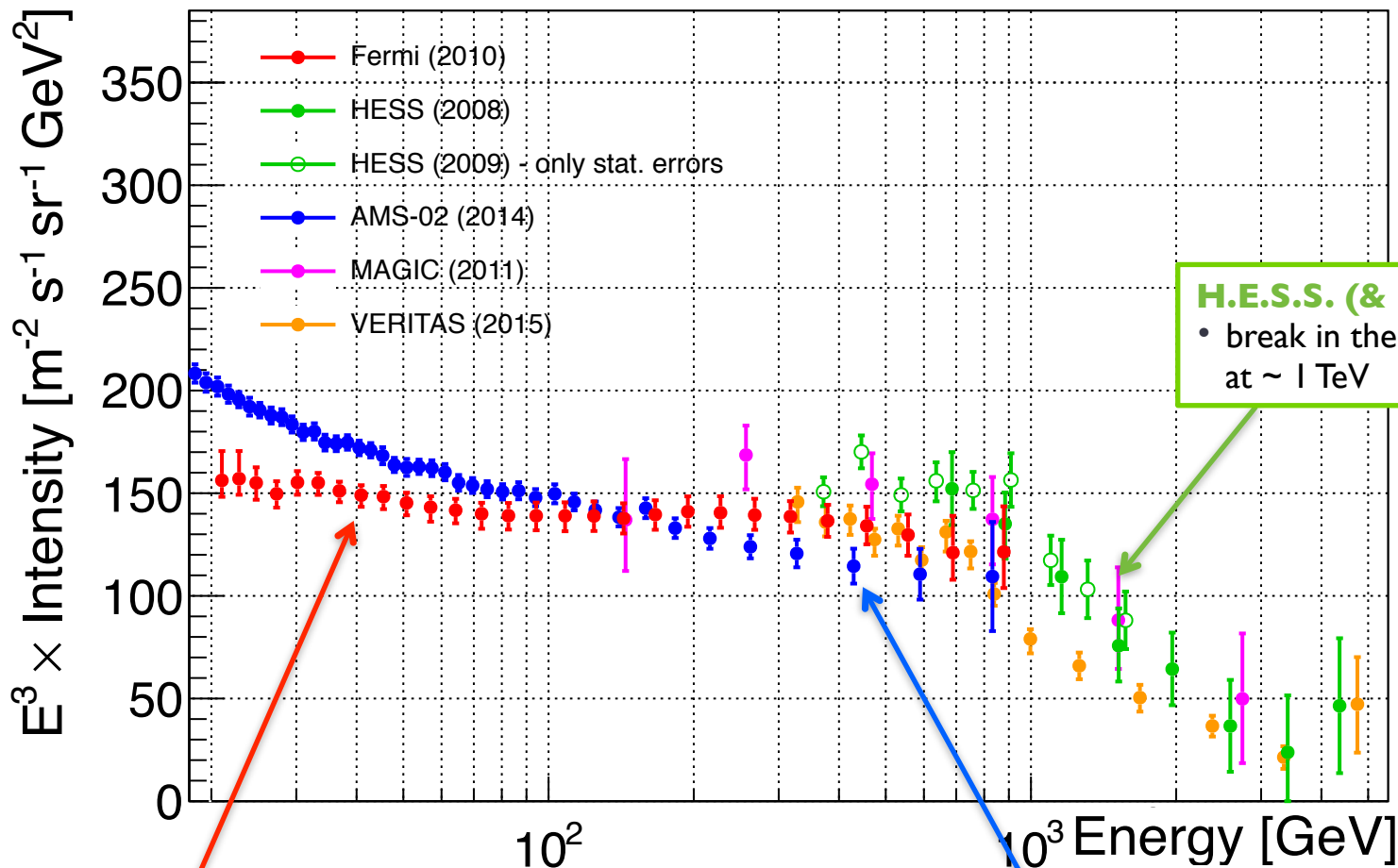
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on behalf of the **Fermi-LAT Collaboration**



# CRE( $e^+e^-$ ) experiments



## H.E.S.S. (& Cherenkov)

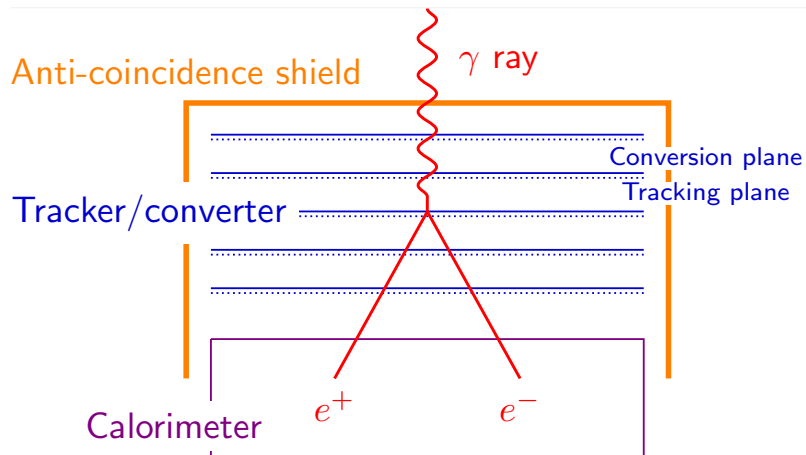
- break in the spectrum at  $\sim 1$  TeV



## Fermi

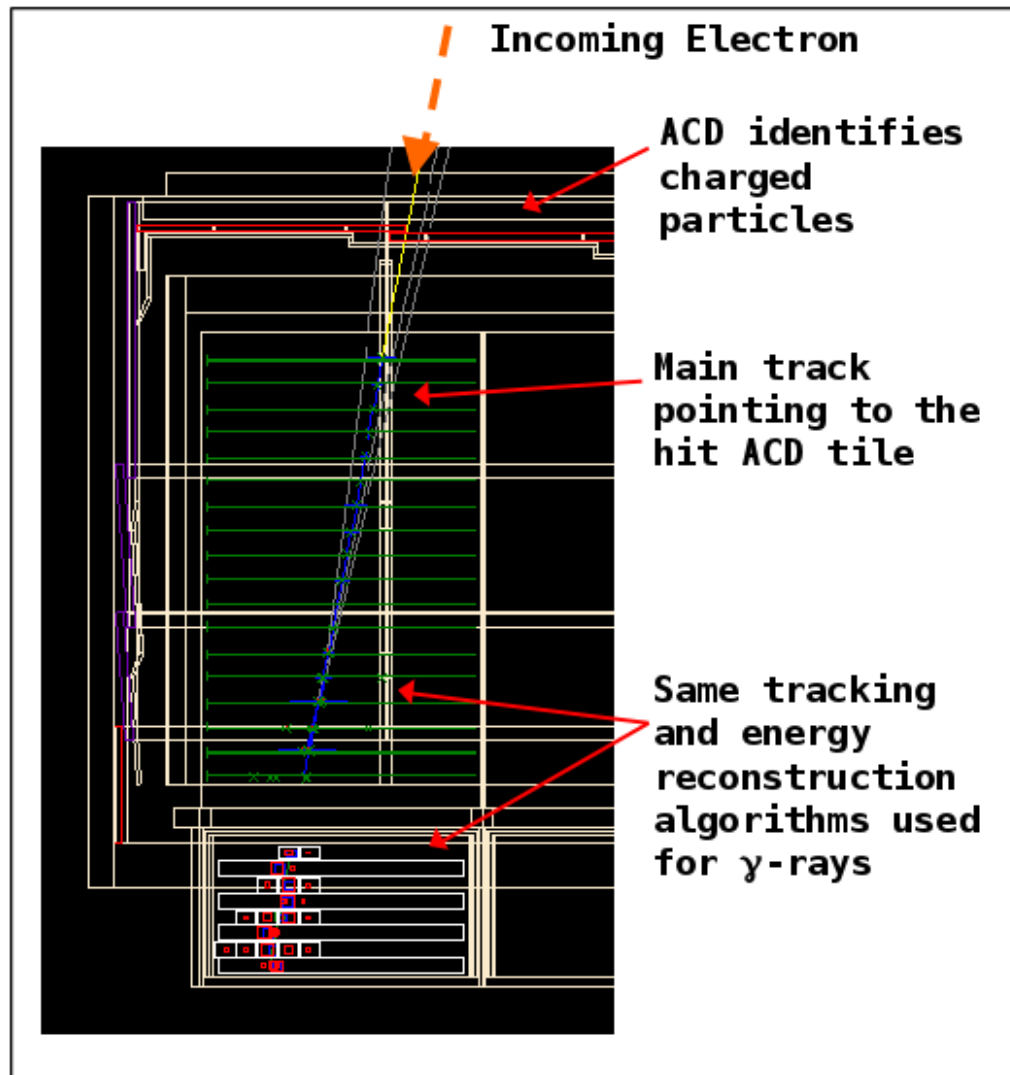
- First high-statistics measurement of inclusive spectrum between 70 GeV and 1 TeV
- Measurements compatible with single power-law over the entire energy range  $\propto E^{-3.08 \pm 0.05}$

## AMS-02

- Range: 0.5 GeV to 1 TeV.
- No structures observed.
- From 30 GeV to 1 TeV described by a single power law with  $\gamma = -3.170 \pm 0.008(\text{stat+syst}) \pm 0.008(\text{E scale})$



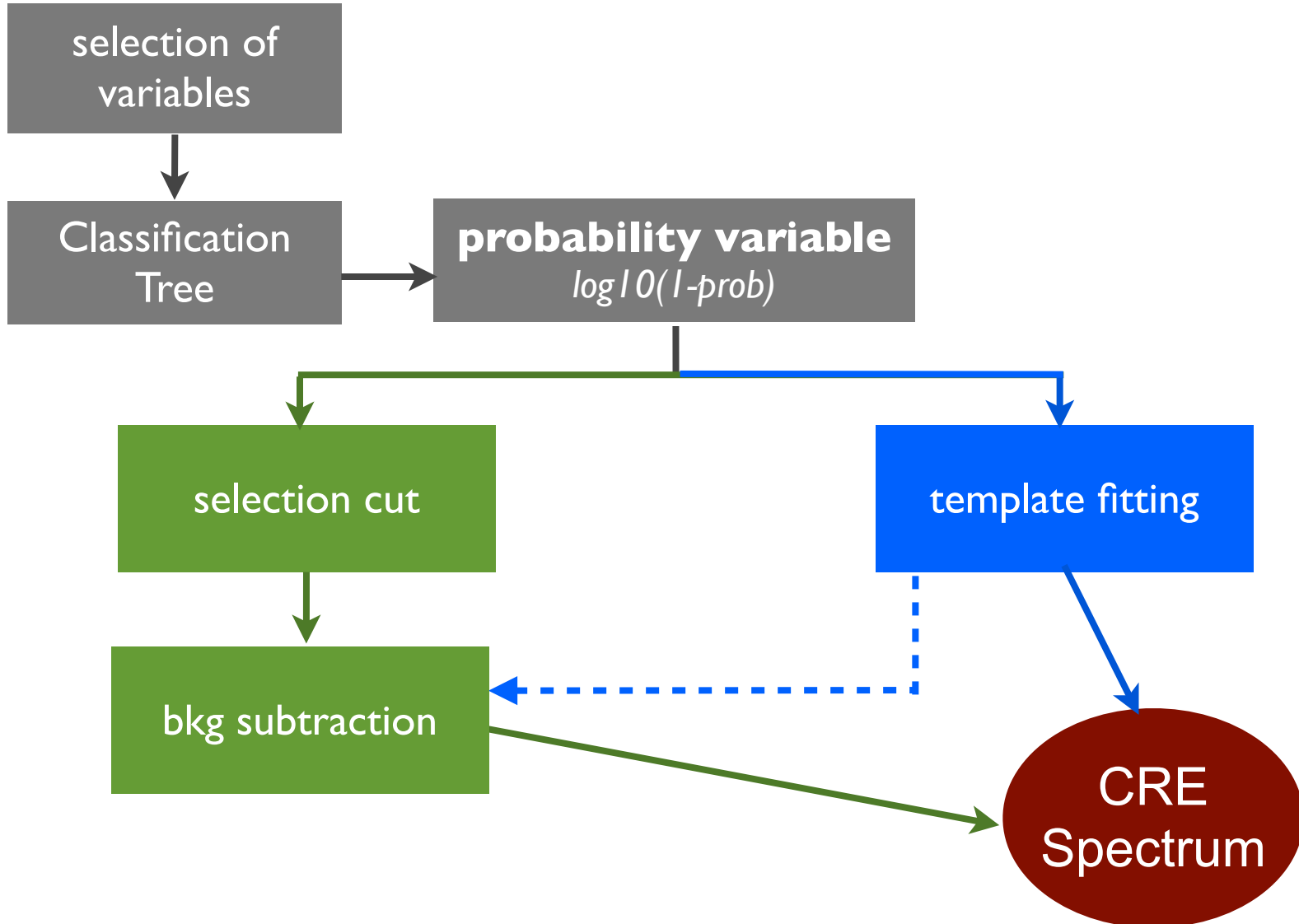
- 
 Detector is designed for E. M. showers:
  - ➔ naturally including electrons
  - ➔ event reconstruction works also for electrons
  
- 
 Electron identification requires dedicated **event selection**





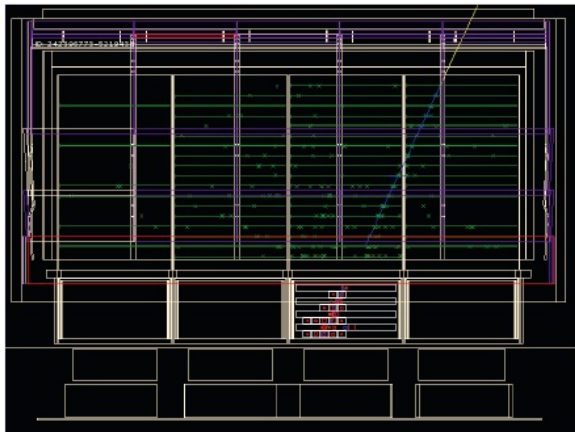
- 🔧 **Objective:** discriminate signal ( $e^+/e^-$ ) from background (mainly  $p$ ) and compute the CRE energy spectrum
  
- 🔧 **Data set:** ~ 7 years
  
- 🔧 **Energy range:** 50 GeV - 2.5 TeV
  
- 🔧 **Event reconstruction:** Pass 8
  
- 🔧 **Event selection:**
  - **PRECUTS**
  - **Classification Tree:** TMVA with the Boosted Decision Tree method
    - trained on MC data samples (signal=MC electron, bkg=MC proton)
    - in 0.25 wide log energy bins → training optimized for the whole energy range
  
- 🔧 **CRE spectrum**

# Analysis steps



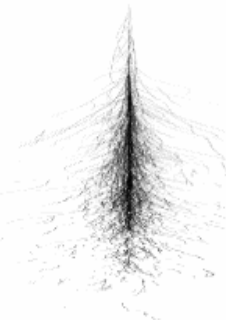


# Selection of variables

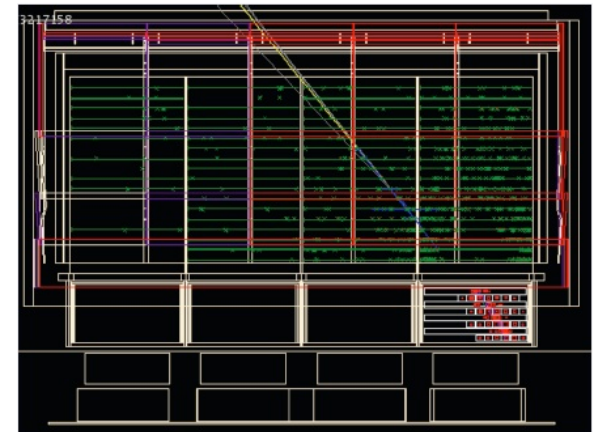
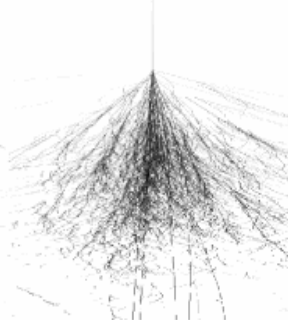


1 TeV electron candidate

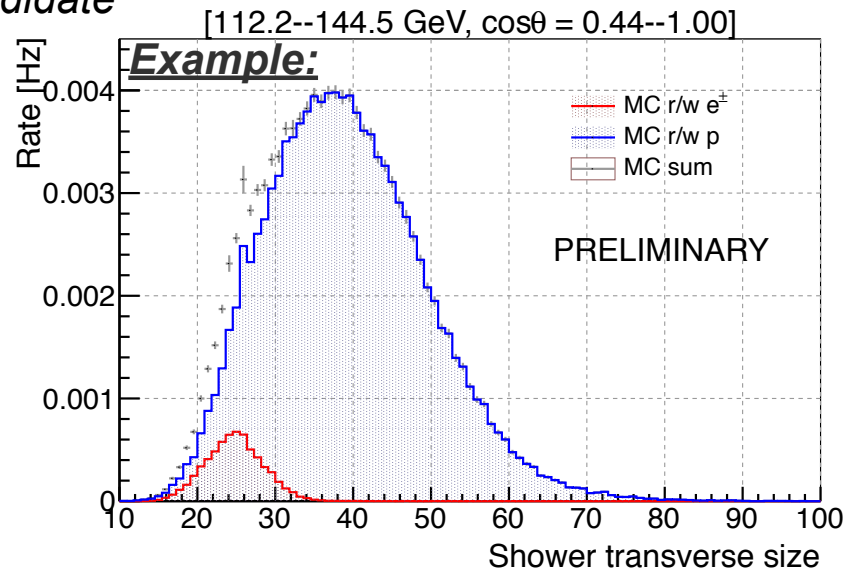
EM shower



Hadronic shower



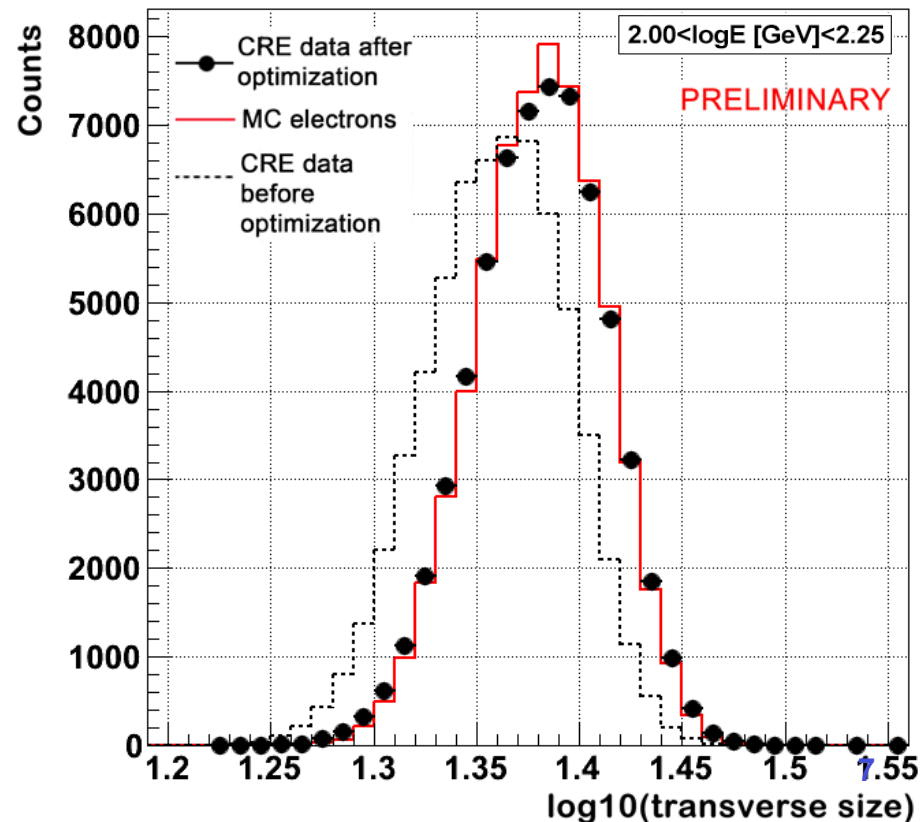
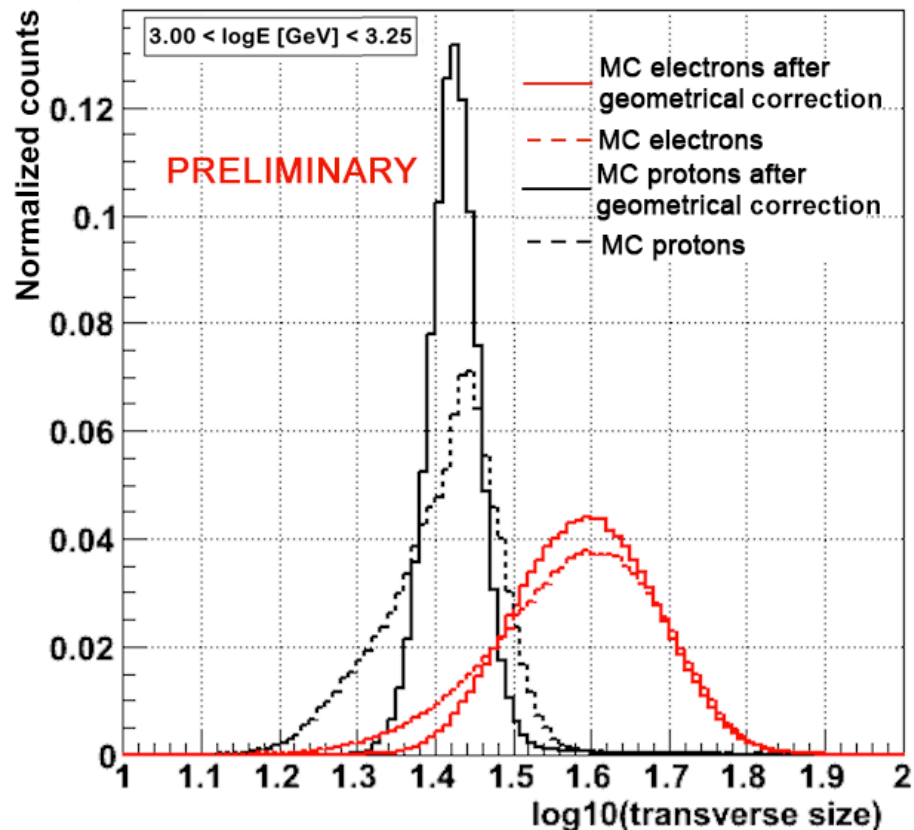
1 TeV proton candidate



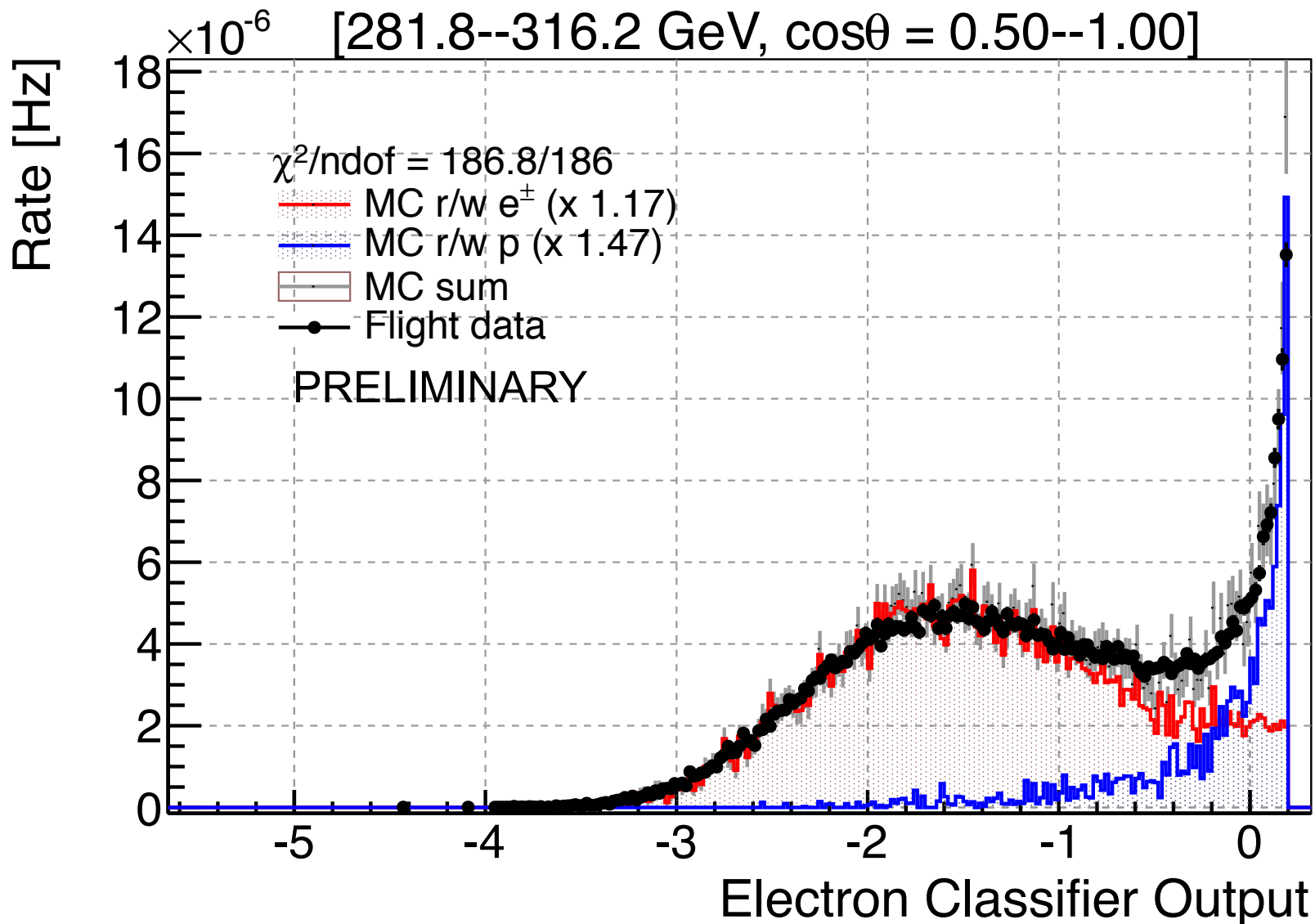
The LAT uses shower topology information to separate the electron signal from the hadronic background



- Variables chosen according to: **good MC-data agreement** and **high separation efficiency**
- Separation efficiency improved after correcting some observables for **geometry** dependence
- MC-data agreement optimized after correcting for **energy** and **incidence angle** dependences

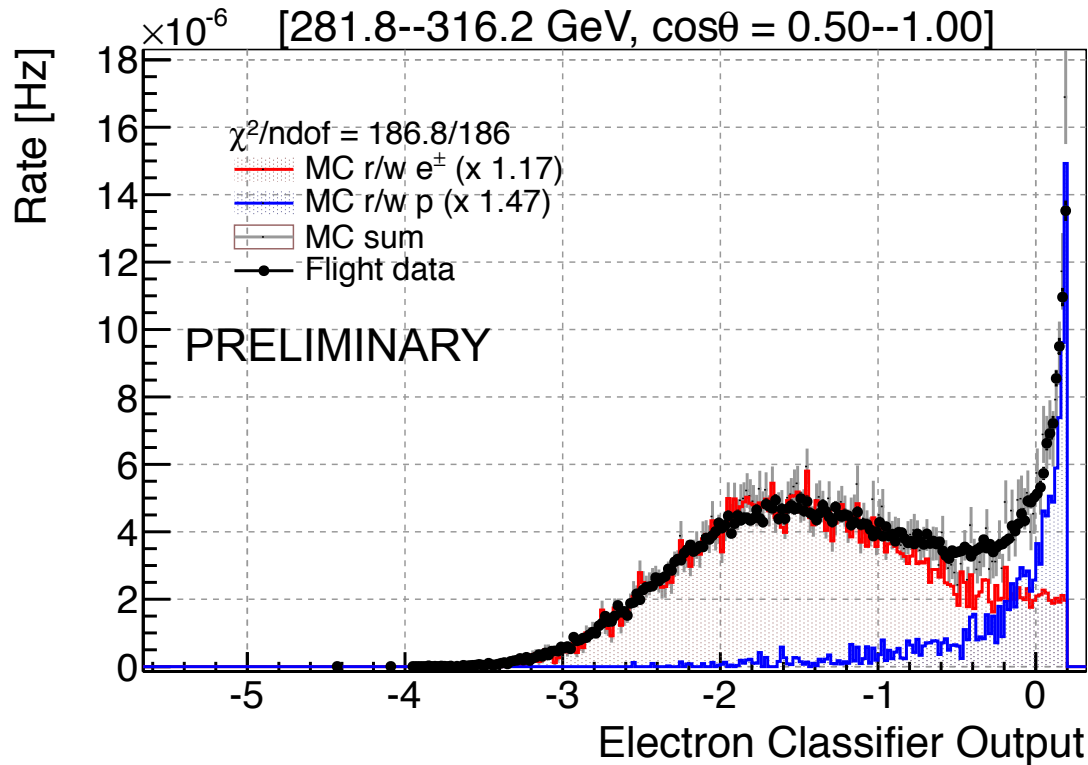


# Probability variable





# a. Template fitting

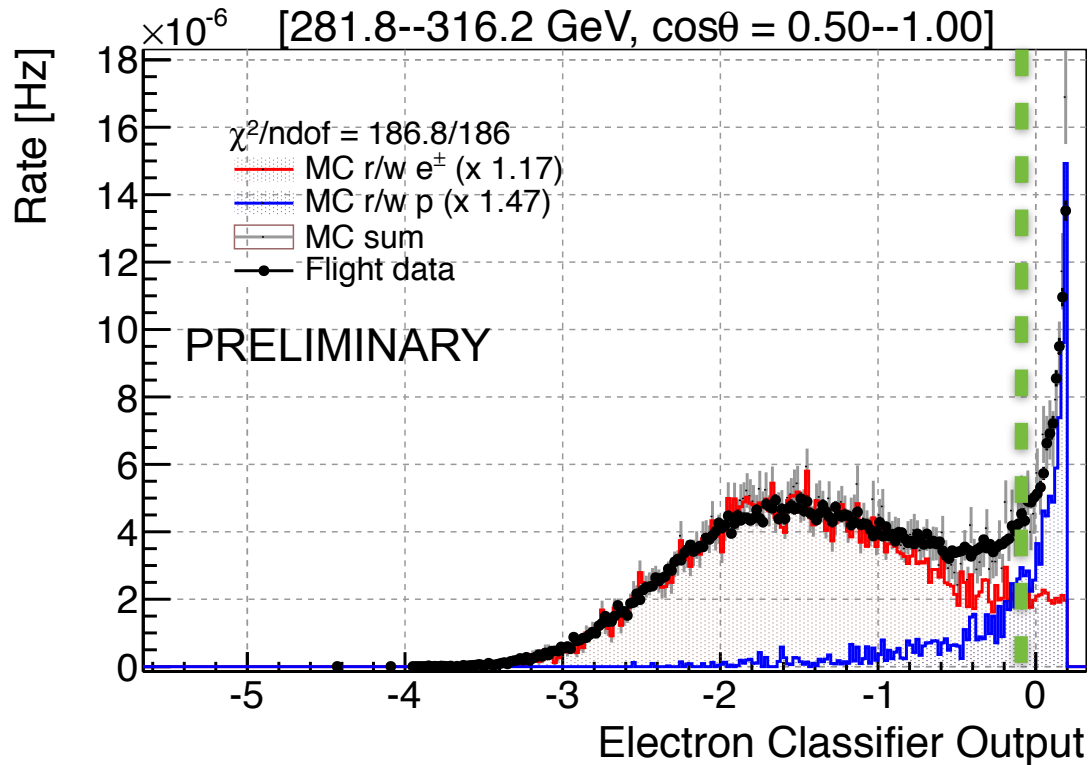


## Fit the data with the MC electron and proton template

- signal rate = (rate from flight data) x (fraction of electrons from MC)
- use that rate for the spectrum directly!

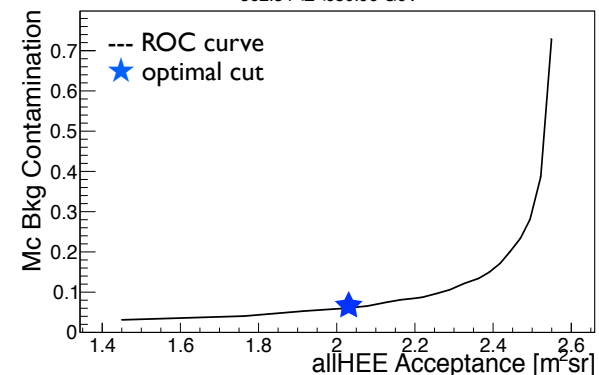
MC proton template (i.e. parameterization of AMS-01) is renormalized and fit to the data in each energy bin

# b. Bkg subtraction



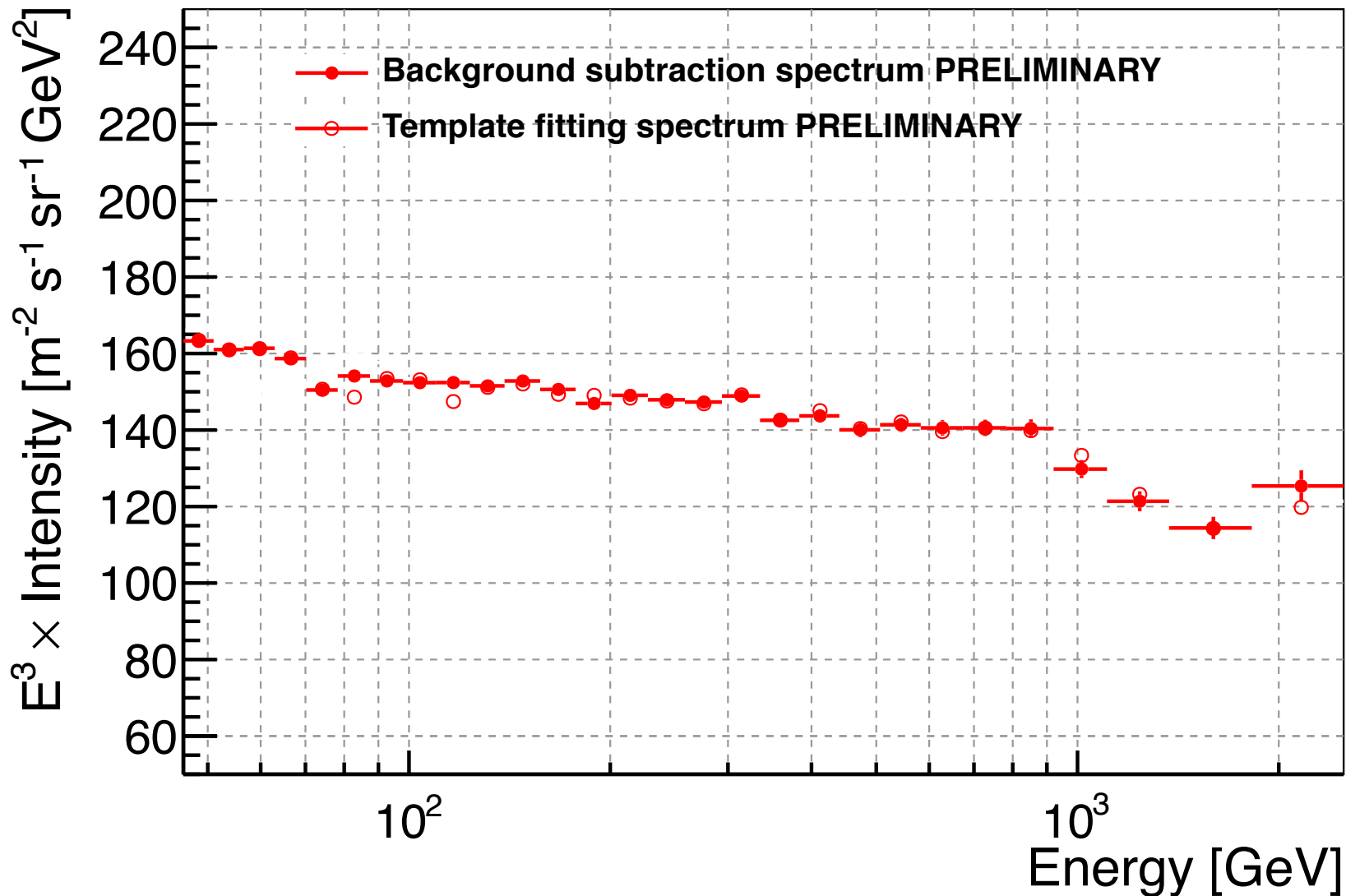
562.34 < E < 630.96 GeV

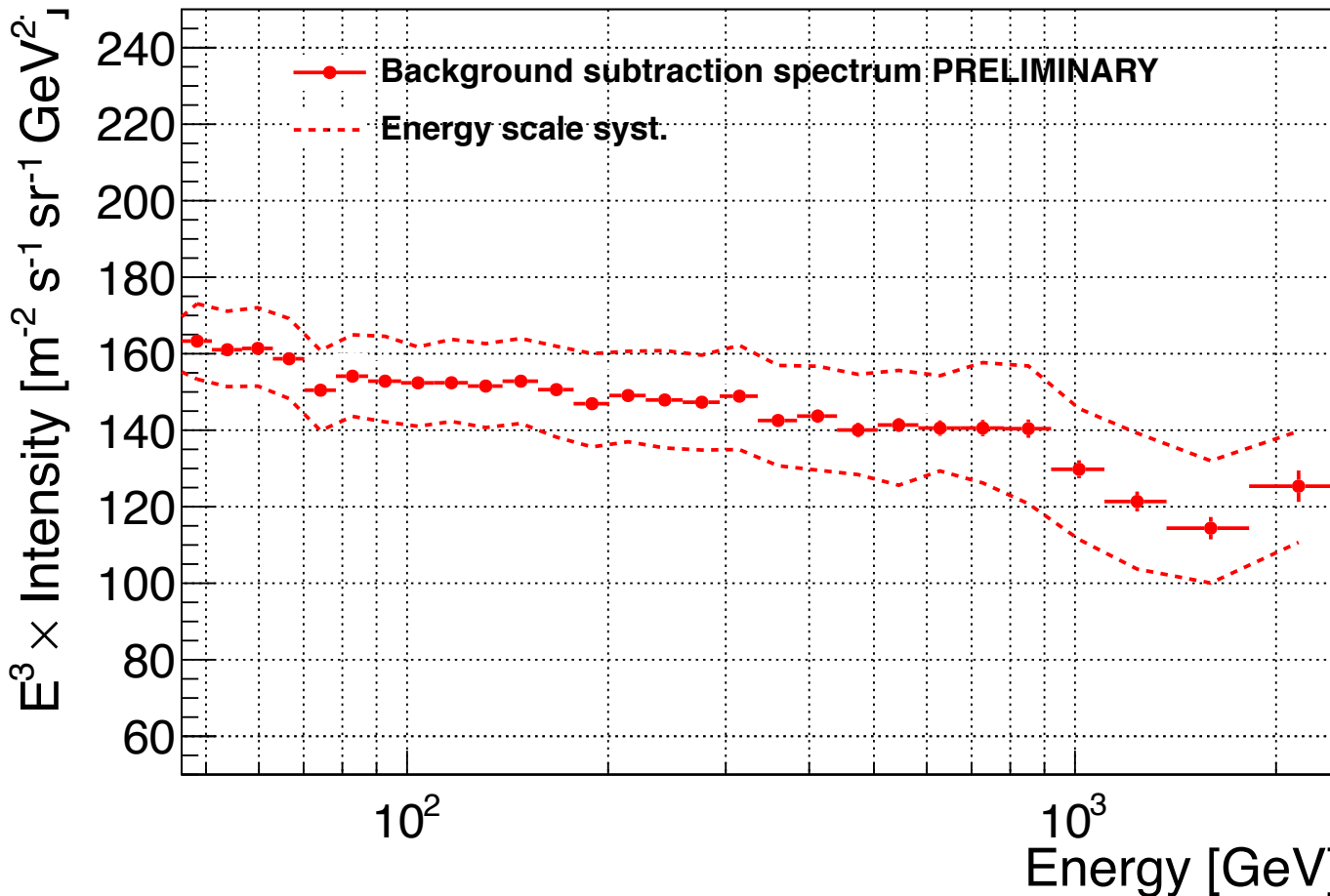
- Cutting on the CT probability variable:**
  - “**optimal cut**” = point on the performance curve (ROC) in which the slope becomes greater than a defined threshold
  - fit the cut values as a function of energy  
→ **analytical function of  $\text{Prob}(E)$**



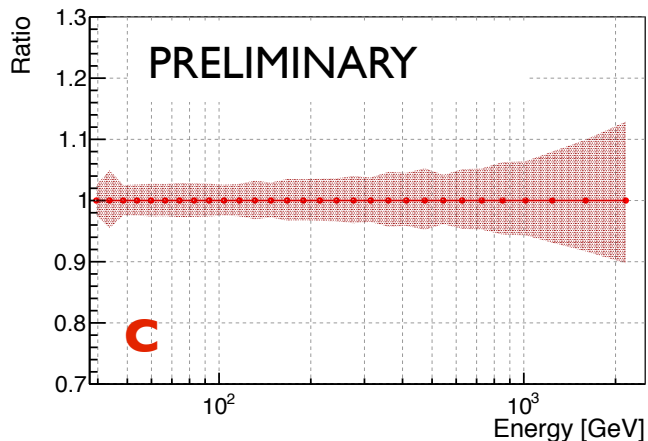
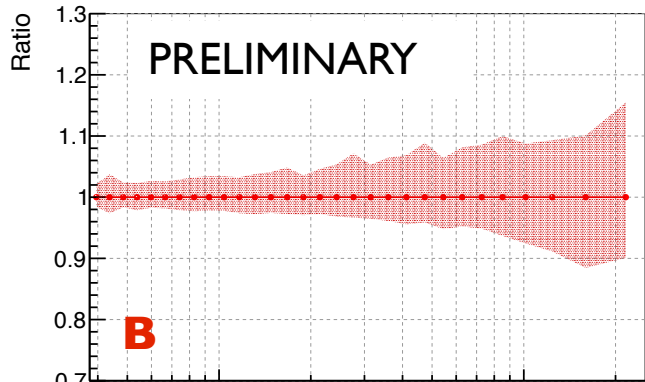
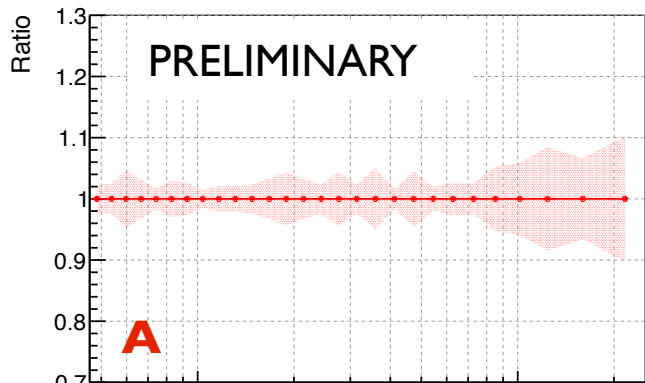
- Bkg contamination estimated after applying “correction” from template fitting 10

# Resulting CRE spectra





- The uncertainty on the absolute energy scale is the **largest source of systematics**
- Pass 8 in-flight measurement of the absolute energy scale via geomagnetic cutoff study  
→ **3.7% offset** around 10 GeV
- We have rescaled the whole spectrum by 3.7% and we have estimated the error on this scaling factor to be **2% at 10 GeV** and increasing up to **5% at ~ 1 TeV**



## A) Event selection:

- main uncertainty at this level is due to the estimation of the effective area → we take it into account by varying the signal efficiency between 30% and 90%

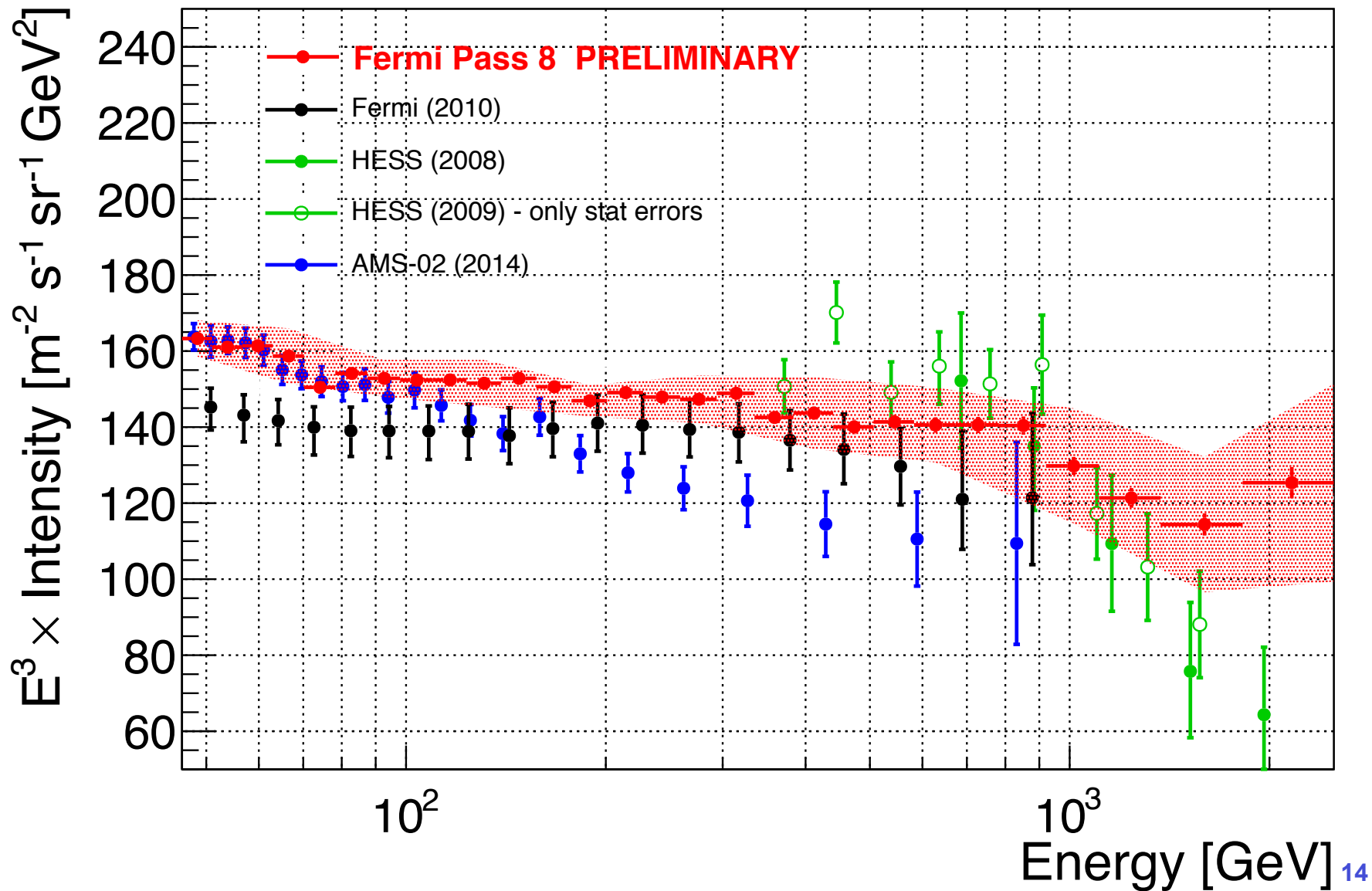
## B) Correction factor:

- the band is calculated by moving the correction magnitude by one sigma

## C) Bkg simulation reliability:

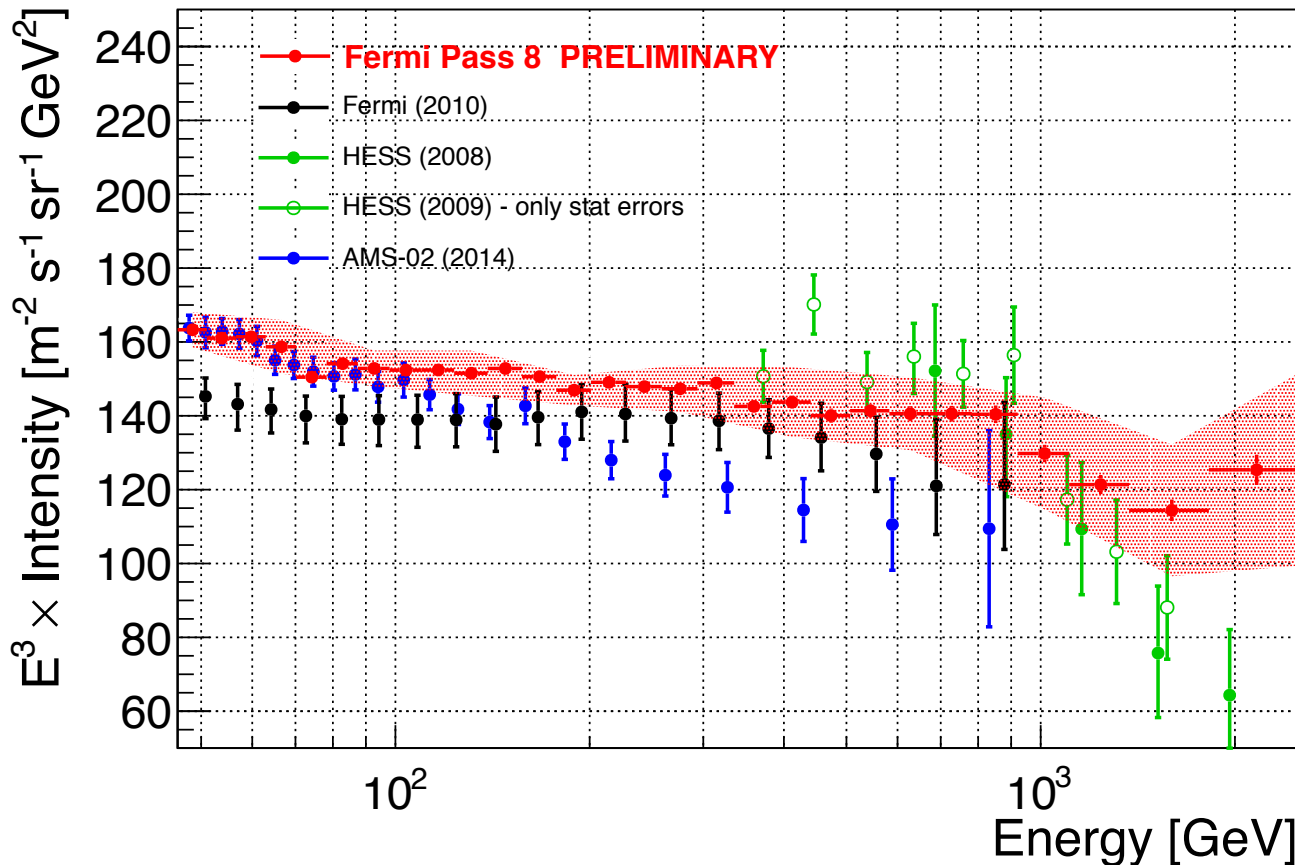
- the uncertainties related to the MC simulation of hadronic interactions could produce a 30% uncertainty on the residual contamination

# CRE inclusive spectrum



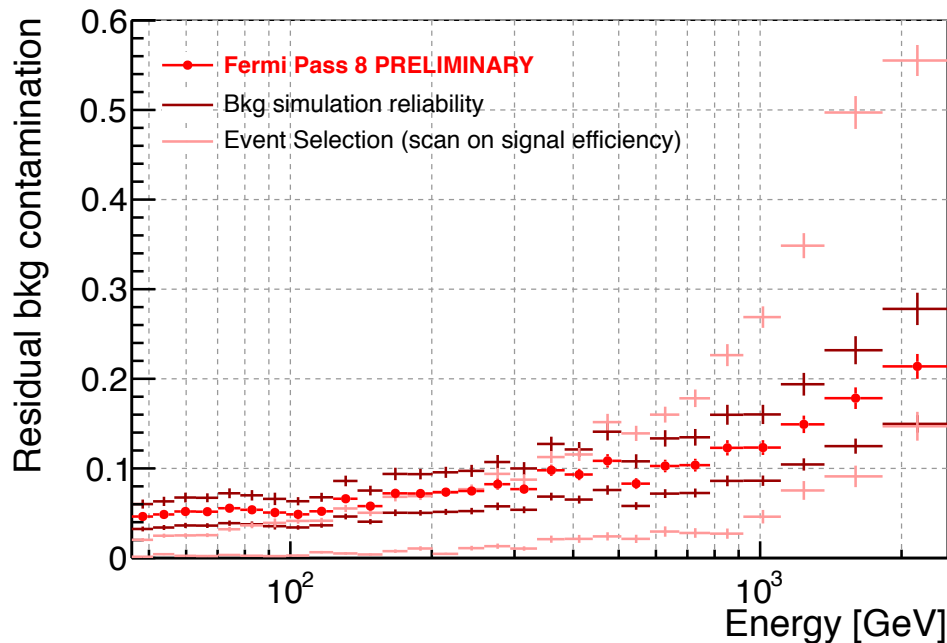
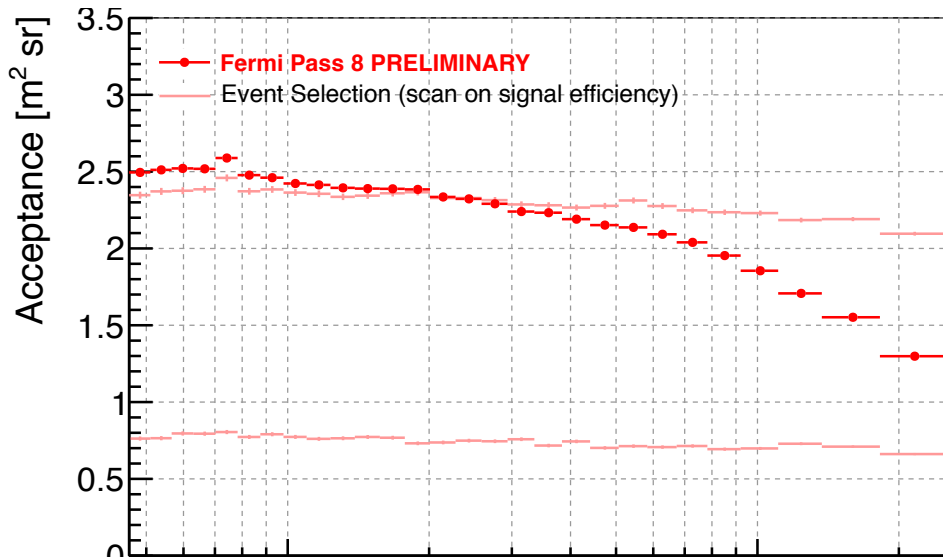


# CRE inclusive spectrum



- Central points from the “optimal cut”, the shaded band has been obtained by summing in quadrature all the studied systematics (except for the energy scale)
- Compatible with **AMS** up to  $\sim 100$  GeV but different spectral index
- Disagreement wrt published spectrum (**Fermi 2010**):
  - likely due to “ghost” signal not taken into account in the acceptance in our first analysis

# Instrument response



**Acceptance** resulting after applying the “optimal cut”

+

● **acceptances** resulting by varying the signal efficiency between 30% and 90% also shown

**Contamination** below 20% when applying the “optimal cut”

+

● **contamination** resulting by varying the signal efficiency between 30% and 90%

● **contamination** resulting by taking into account the bkg simulation reliability



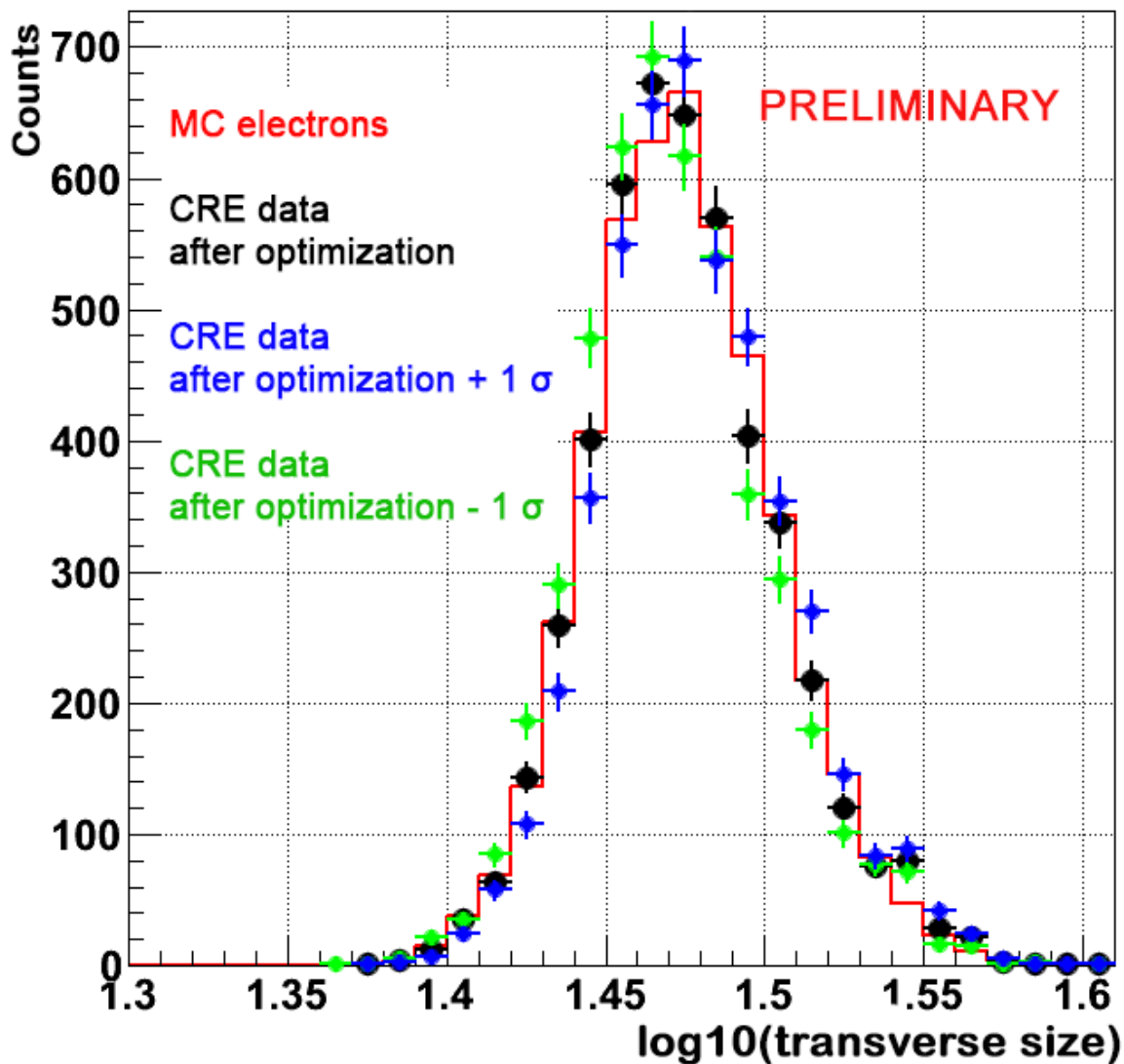
- **We performed a new measurement of CRE spectrum with Pass 8**
- **Improvements in the new analysis:**
  - almost 7 times the PRD **data set**
  - new event reconstruction & selection (**Pass 8**)
  - new **multi-variate analysis** tool
  - new selection of variables → **variables** are now “calibrated”
  - new **CTs** trained in energy bins → training optimized for the whole energy range
  - ➔ **new CRE spectrum** and associated **systematics**
- See also the posters:
  - More **technical details** on the event selection and systematics ([M.Negro et al - Diff.6](#))
  - Analysis extended down to **7 GeV** by including the DGN filter effect ([A. Manfreda et al - In/An.10](#)) → in the overlapping energy region it's in agreement with this one
  - Cross-check the **energy scale** at TeV energies by measuring the Earth Limb spectrum ([F. Spada et al - In/An.7](#))
  - Study CRE **anisotropies** to validate possible astrophys. interpretations ([N. Mazziotta et al - Diff.5](#))
  - Viable **interpretations** of our CRE spectrum ([F. Donato et al - Diff.3](#))



# BACKUP SLIDES



$3.00 < \log E \text{ [GeV]} < 3.25$





**PRECUTS** = TRIGGER FILTER + QUALITY CUT + ALPHA CUT

- TRIGGER FILTER:** the event triggers the LAT and passes the on-board gamma filter  
`'(GltGemSummary&0x20)==0 && (GltGemSummary&0x40)==0 && FswGamState == 0'`
- QUALITY CUT:** the event has at least a reconstructed track, a minimal PSF quality and the path length in the Cal is larger than the Cal on-axis thickness  
`'EvtCalCsIRLn>8 && CalIRawEnergySum>5000 && TkrNumTracks>0 && WP8CTPSFTail>0.05'`
- ALPHA CUT:** MC doesn't reproduce accurately interactions of  $\alpha$  and heavy ions in the LAT  $\rightarrow$  cut removing the majority of  $\alpha$  and heavies

