

Resolving the High Energy Universe with Strong Gravitational Lensing

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Mentor: Margaret Geller

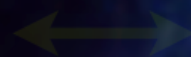
HST - 0.01" 0.5" 0.10"

Reso
with

60 pc



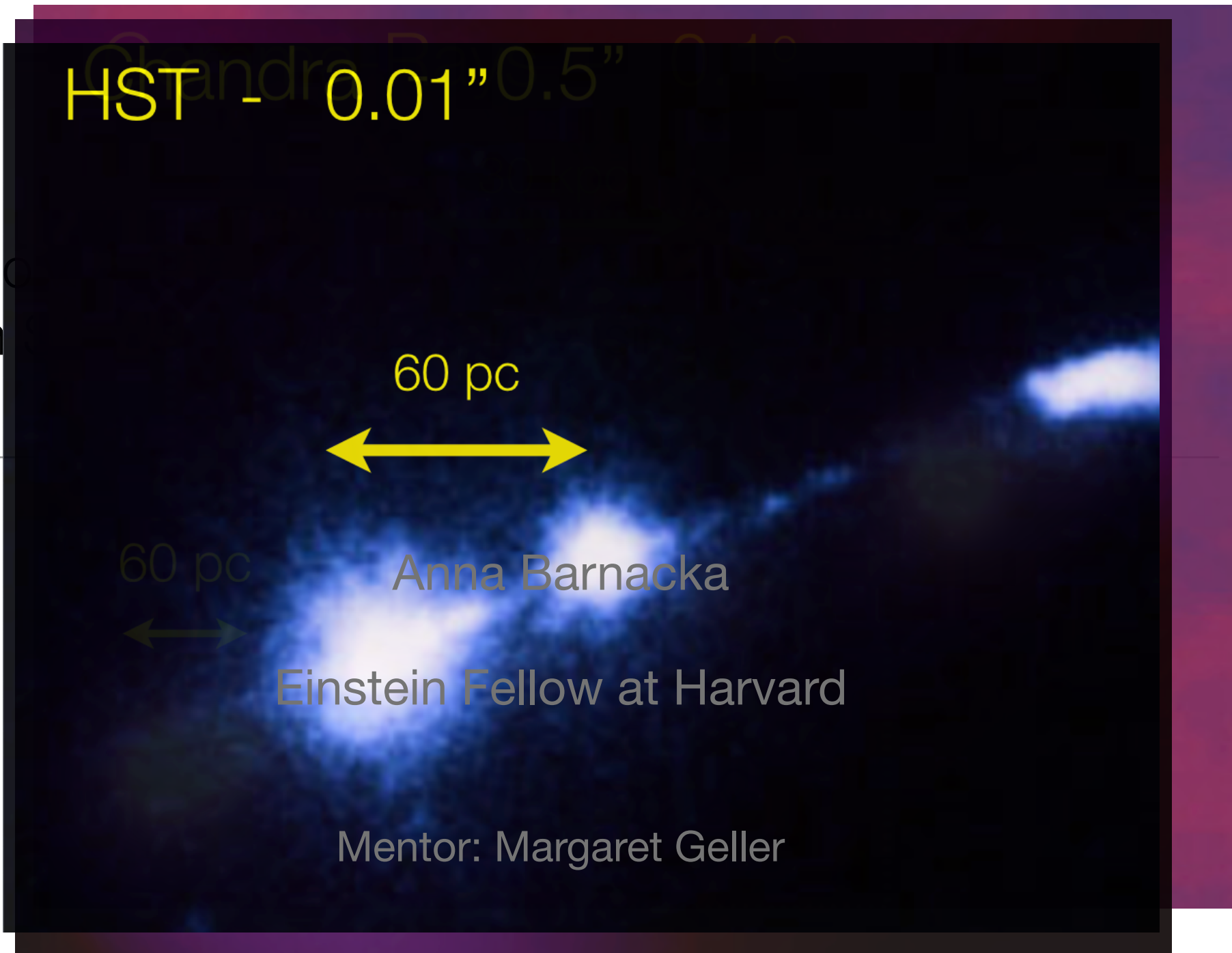
60 pc



Anna Barnacka

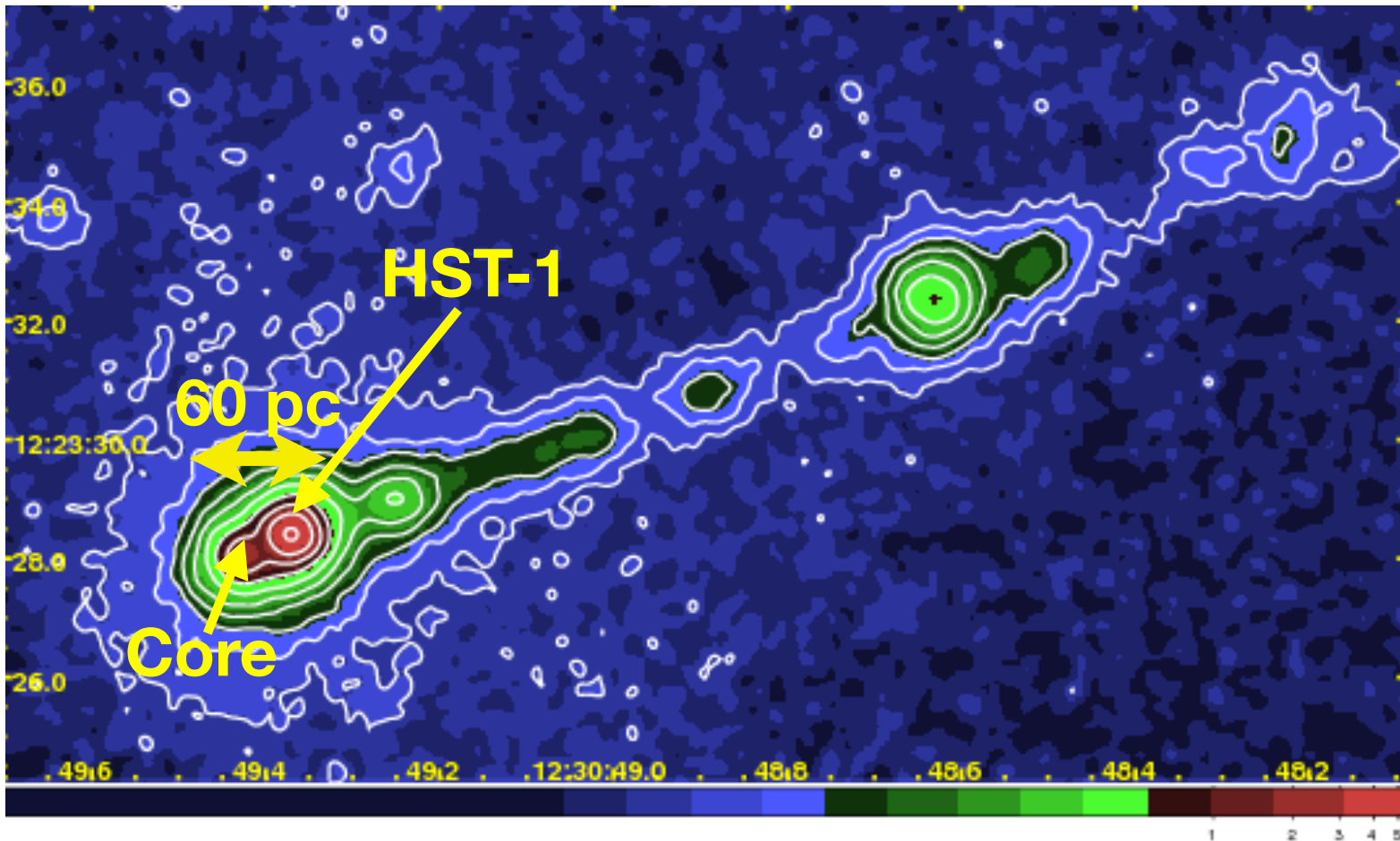
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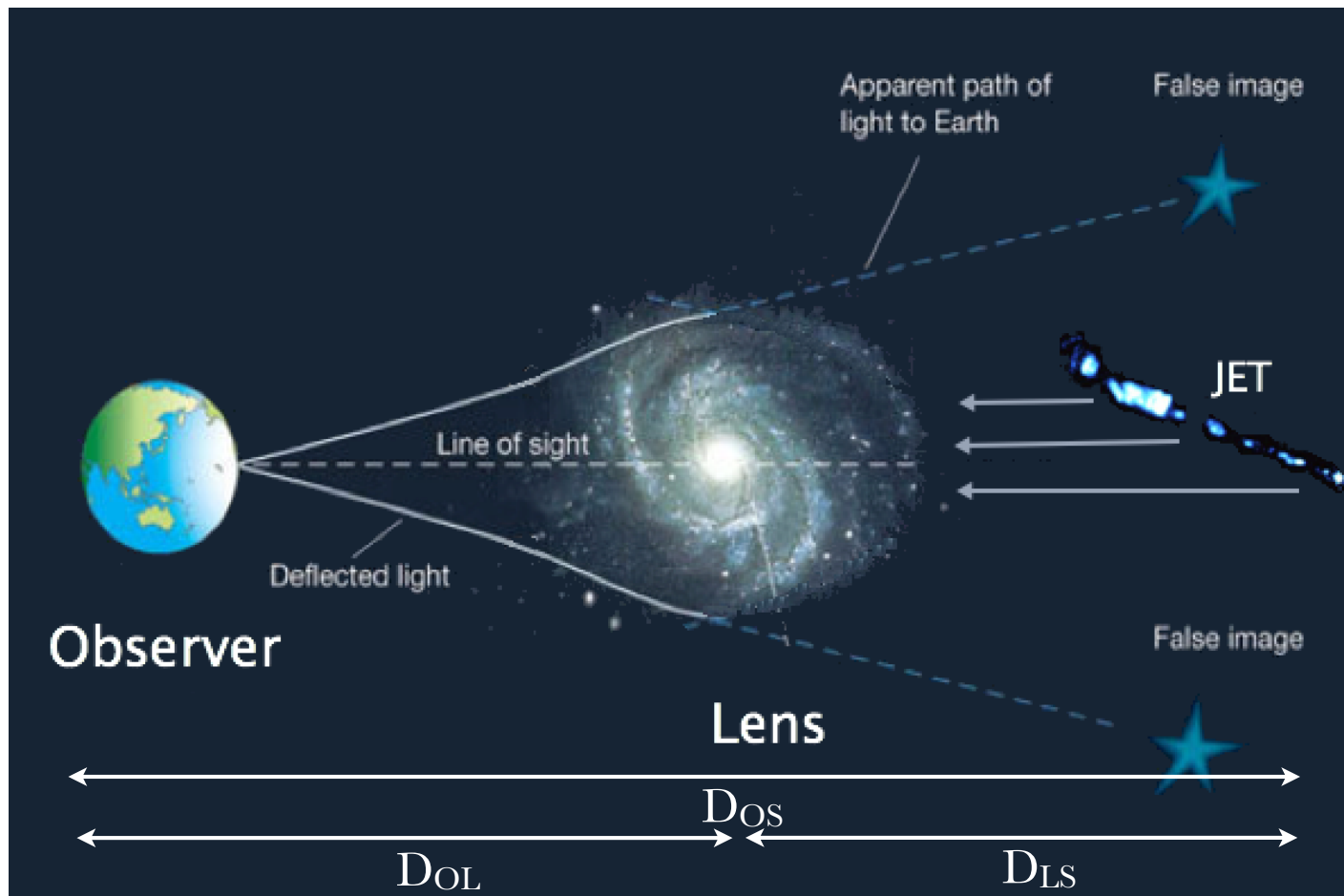
X-Ray Jets - Lessons from Chandra

Increased x-ray emission by a factor of 50 from the HST-1 knot (Harris et al. 2006,2009)
Core and HST-1: Separation ~ 60 pc



Flares from knots along the jets

M87 Gravitationally Lensed?



Deflection angle:

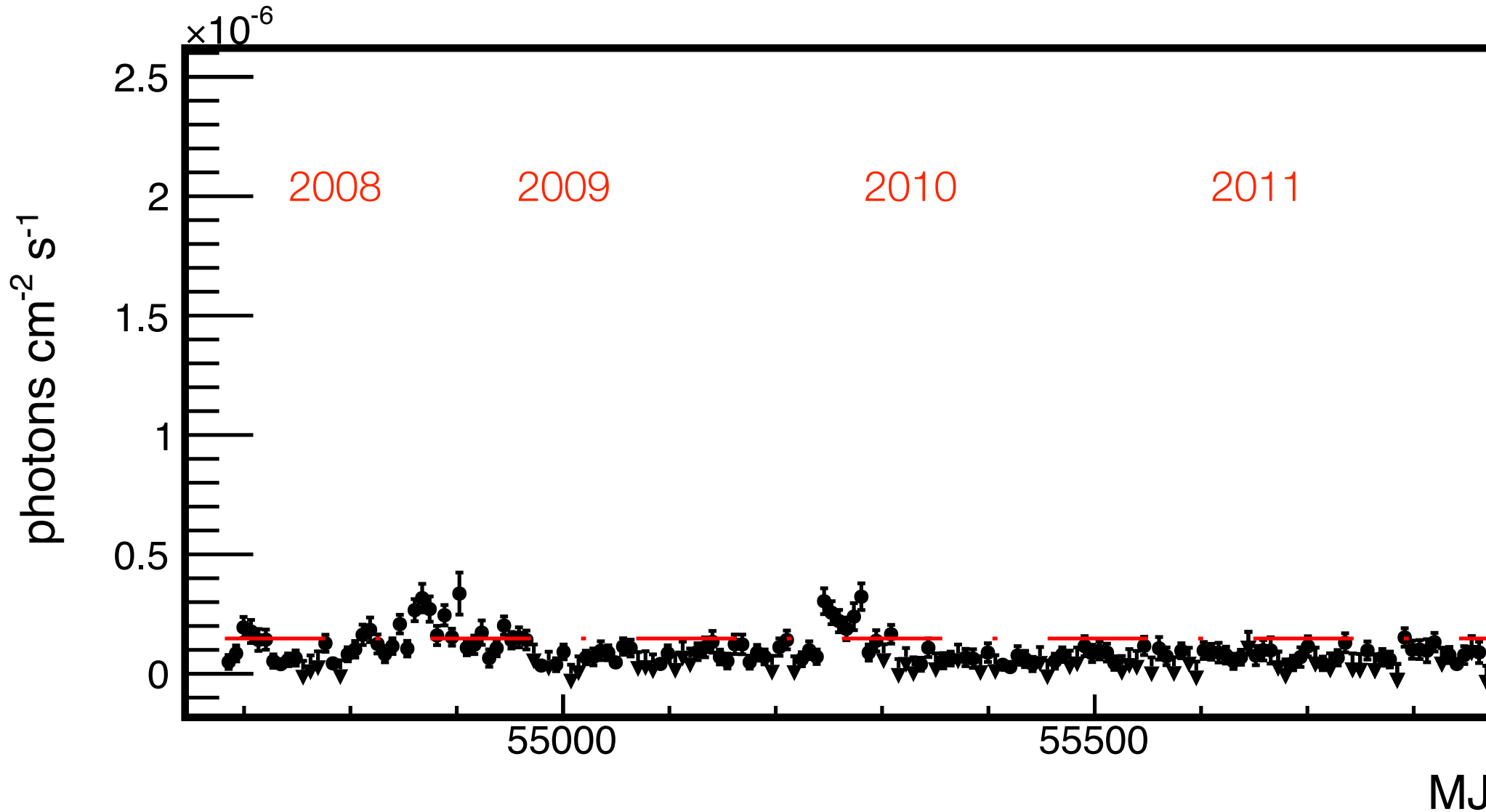
$$\alpha = \frac{4GM(r)}{c^2} \frac{1}{r}$$

Images separation - a few arcseconds
time delay
magnification ratio

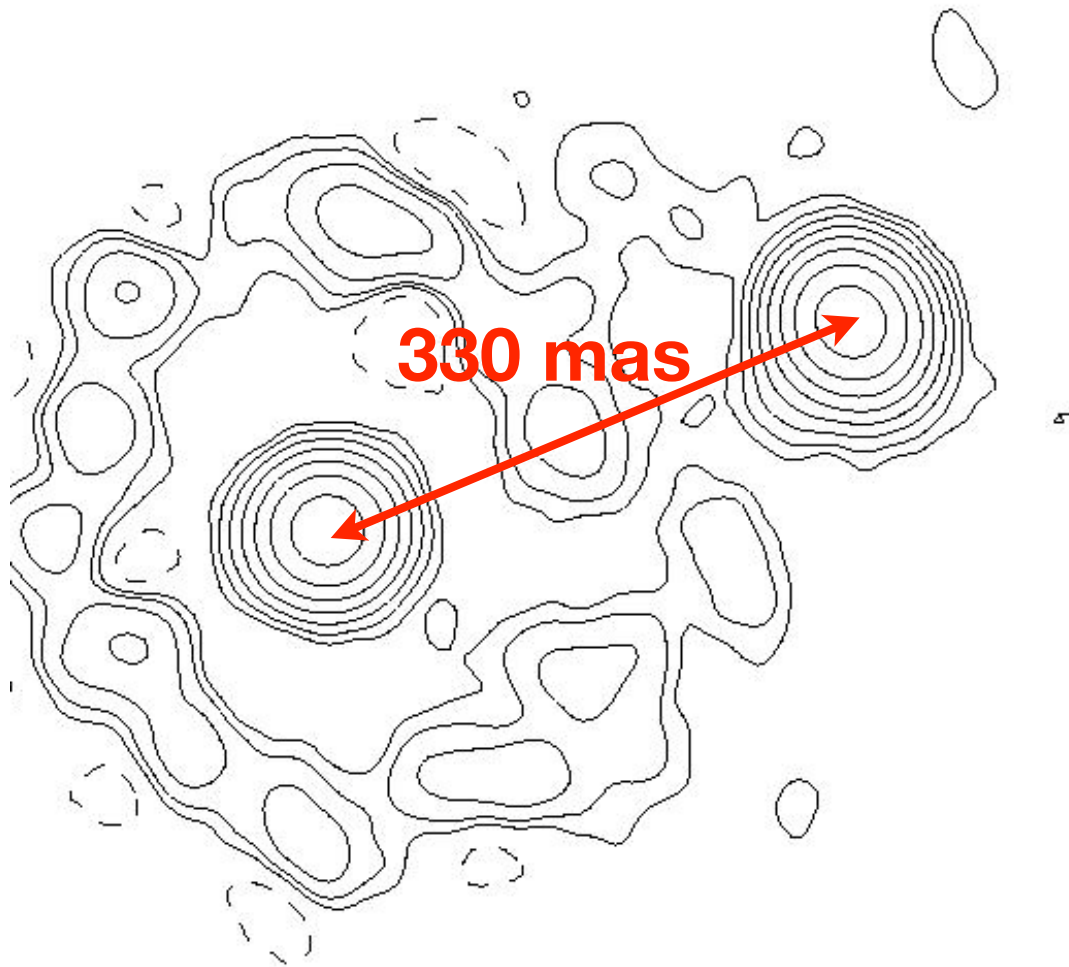
M87 as a Toy Model

- $z_s=1, z_l = 0.6$
- Einstein radius ~ 2.2 kpc (0.45")
- 60 pc $\sim 0.01'' \sim 3\%$ Einstein radius
- Differences between the **core** and the **HST-1**:
 - **difference in time delay: ~ 2 days**
 - **difference in magnification ratio: ~ 0.2**

Temporal Resolution at Gamma Rays



Lensed Gamma-Ray Jets: B2 0218+35



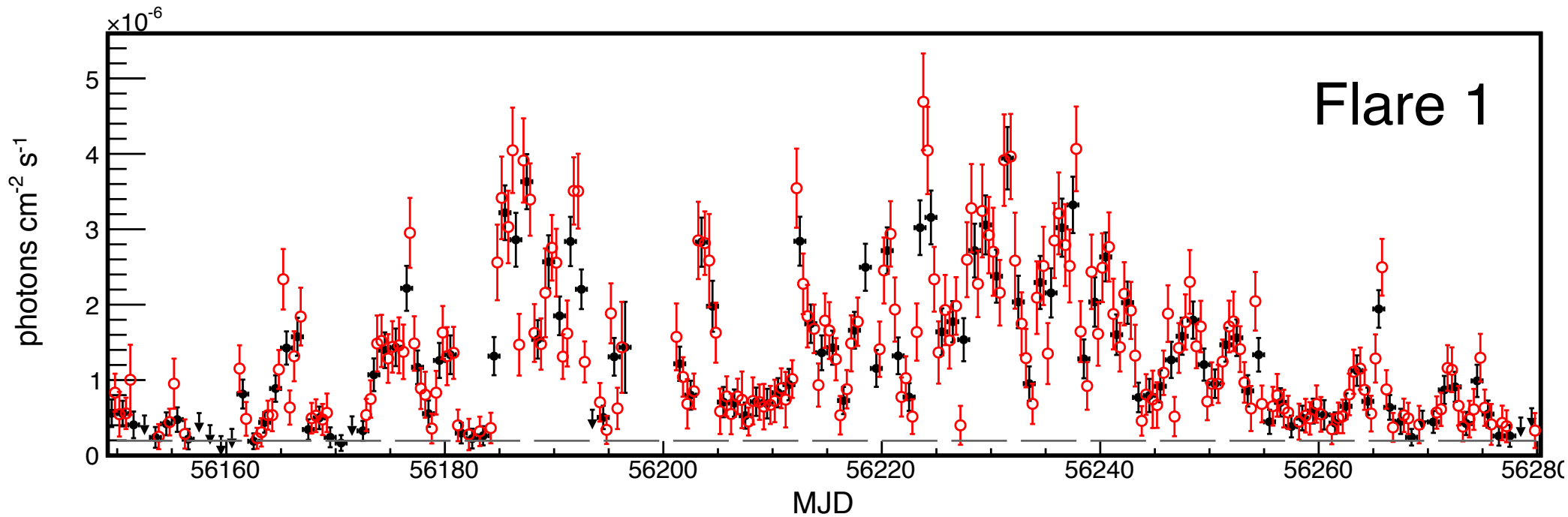
Source $z = 0.944$,
Lens $z = 0.6847$

Radio Time Delay
 10.5 ± 0.5 days

Magnification Ratio
 3.62 ± 0.06

1.687 GHz, Patnaik et al. (1992)

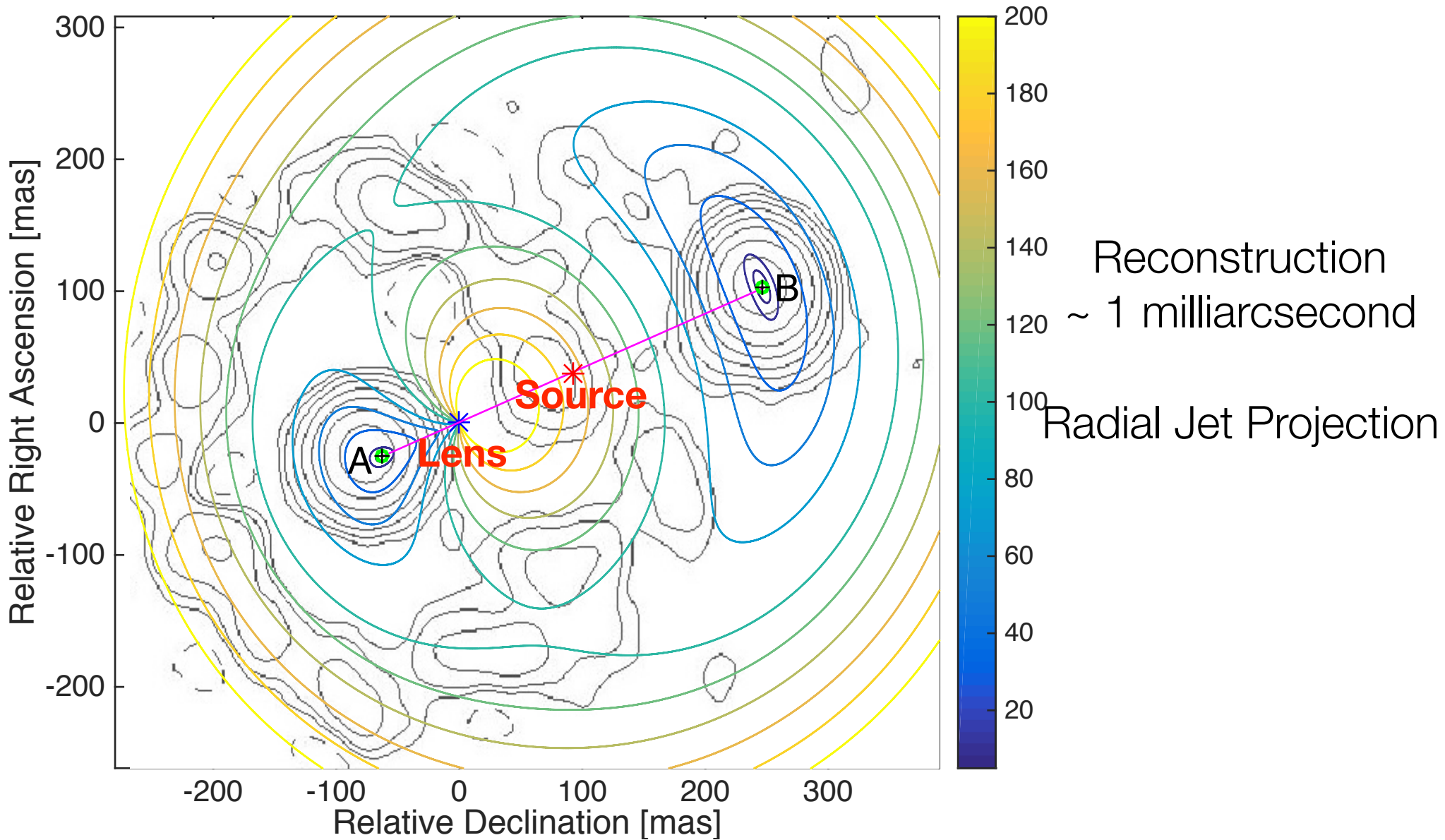
Gamma-Ray Time Delay



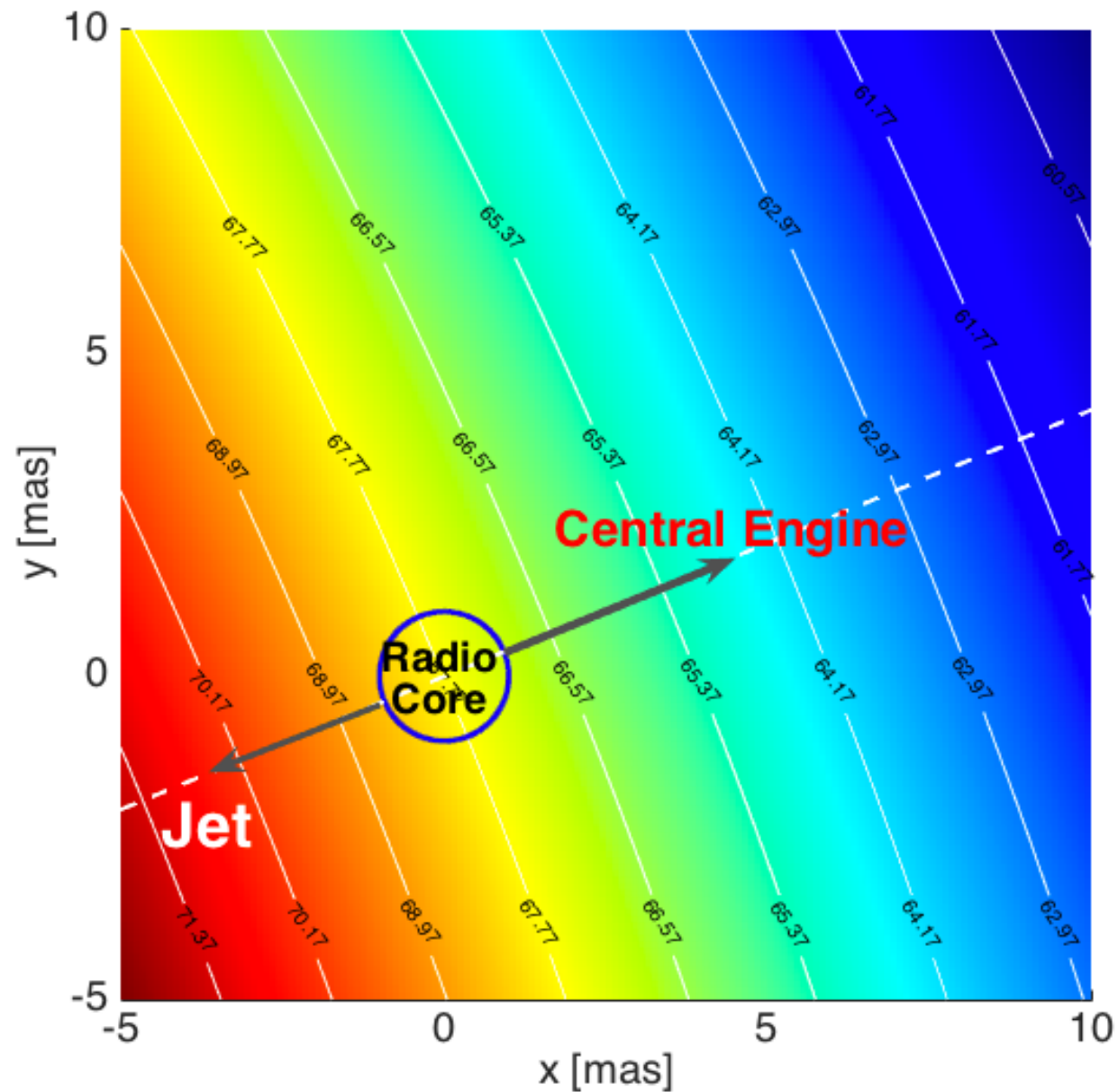
Time Delay = 11.38 ± 0.13 days (Barnacka et al., submitted)

Time Delay = 11.46 ± 0.16 days (Cheung et al. 2014)

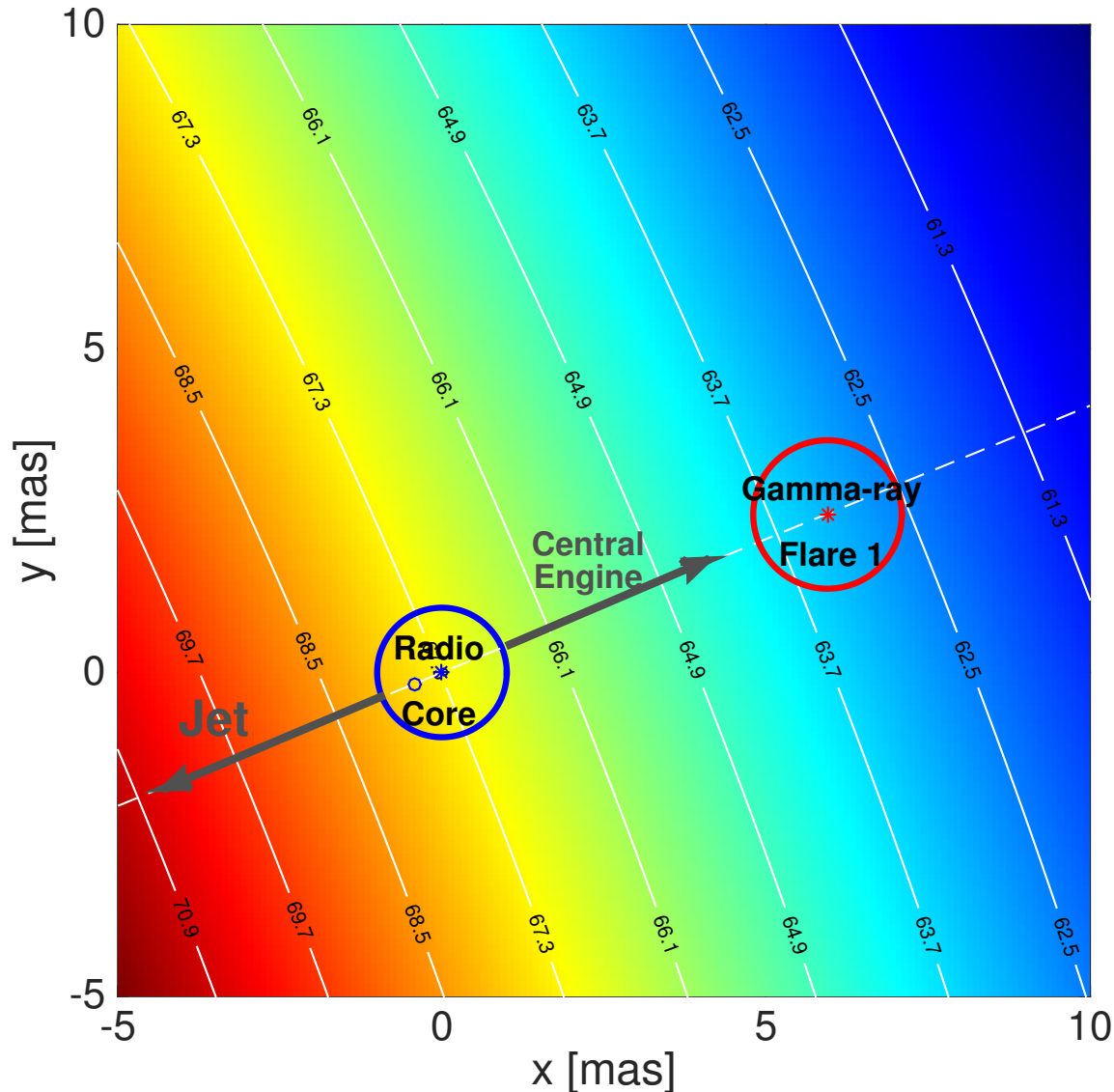
Lens Modeling



The Hubble Parameter Space



The Origin of Gamma-ray Flare



Gamma-ray flare
occurred

51 ± 8 pc

from the 15 GHz core
toward the central
engine.

~ 3 sigma effect
(Barnacka et al.
submitted)

Summary

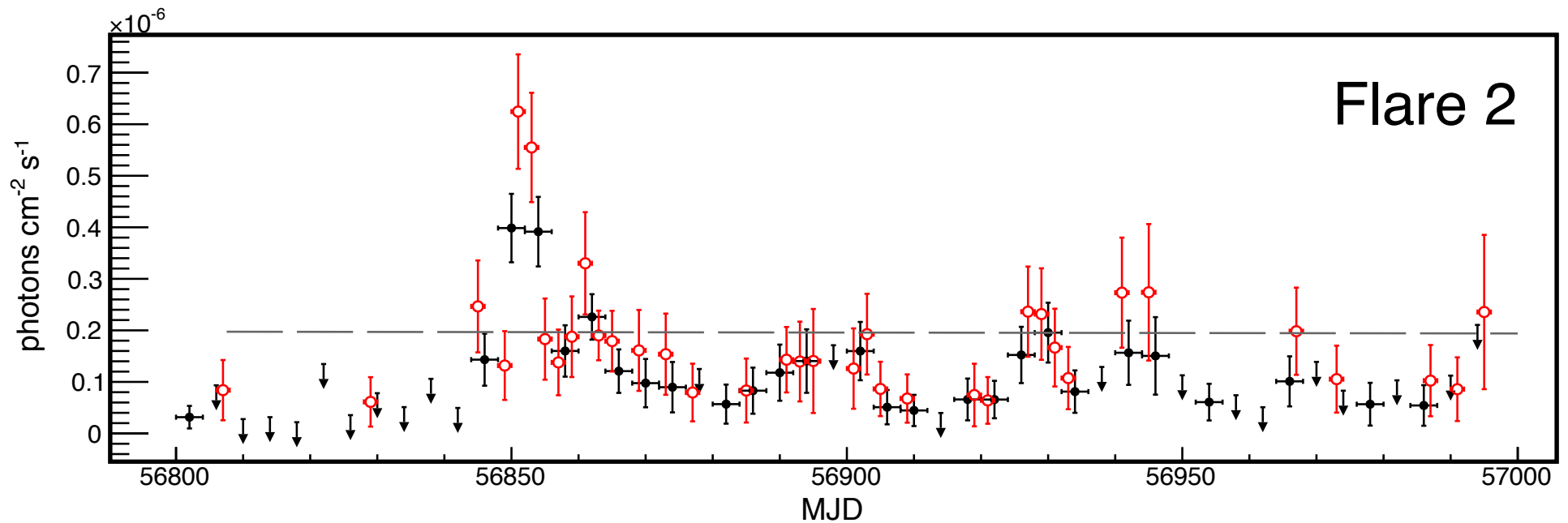
- **Strong Lensing:**
 - **Powerful Tool to Resolve High Energy Universe**
 - **Effective Spatial Resolution ~ 1 = miliarcsecond**
 - improvement $\times 10,000,000$



Backup Slides

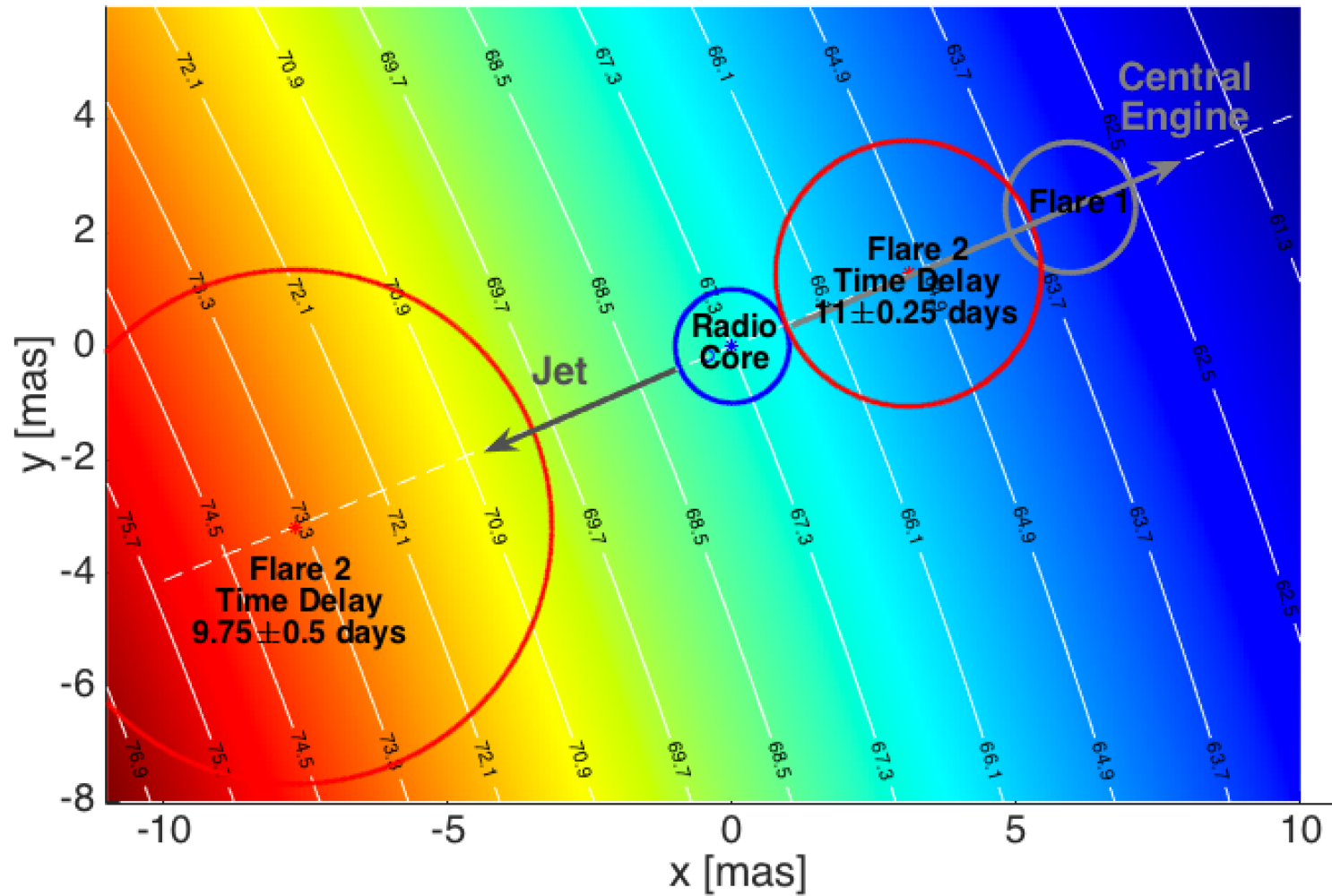
Flare 2

Gamma-ray Flare 2: Time Delays

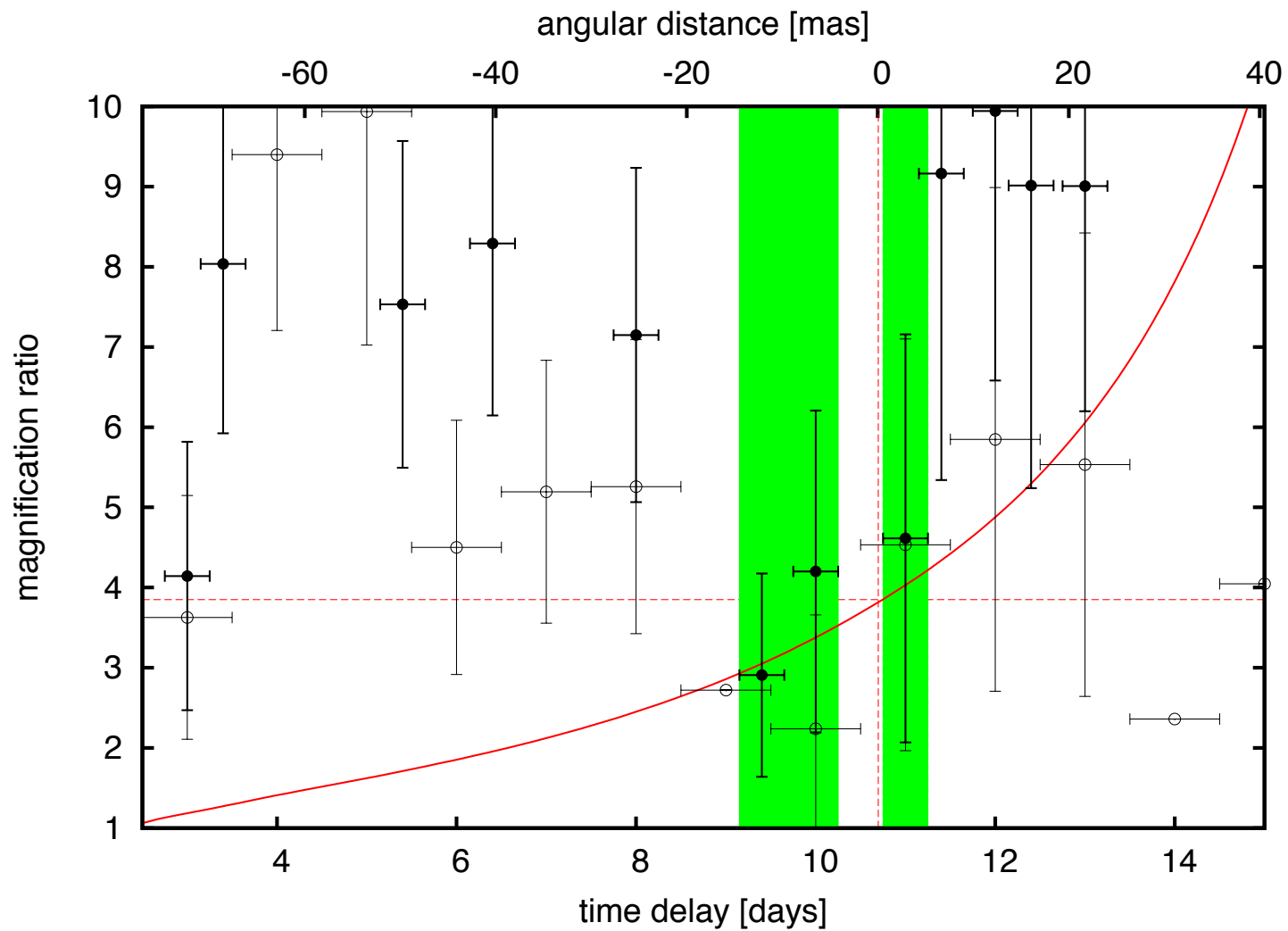


Time delay: 9.75 ± 0.5 or 11.0 ± 0.25 days

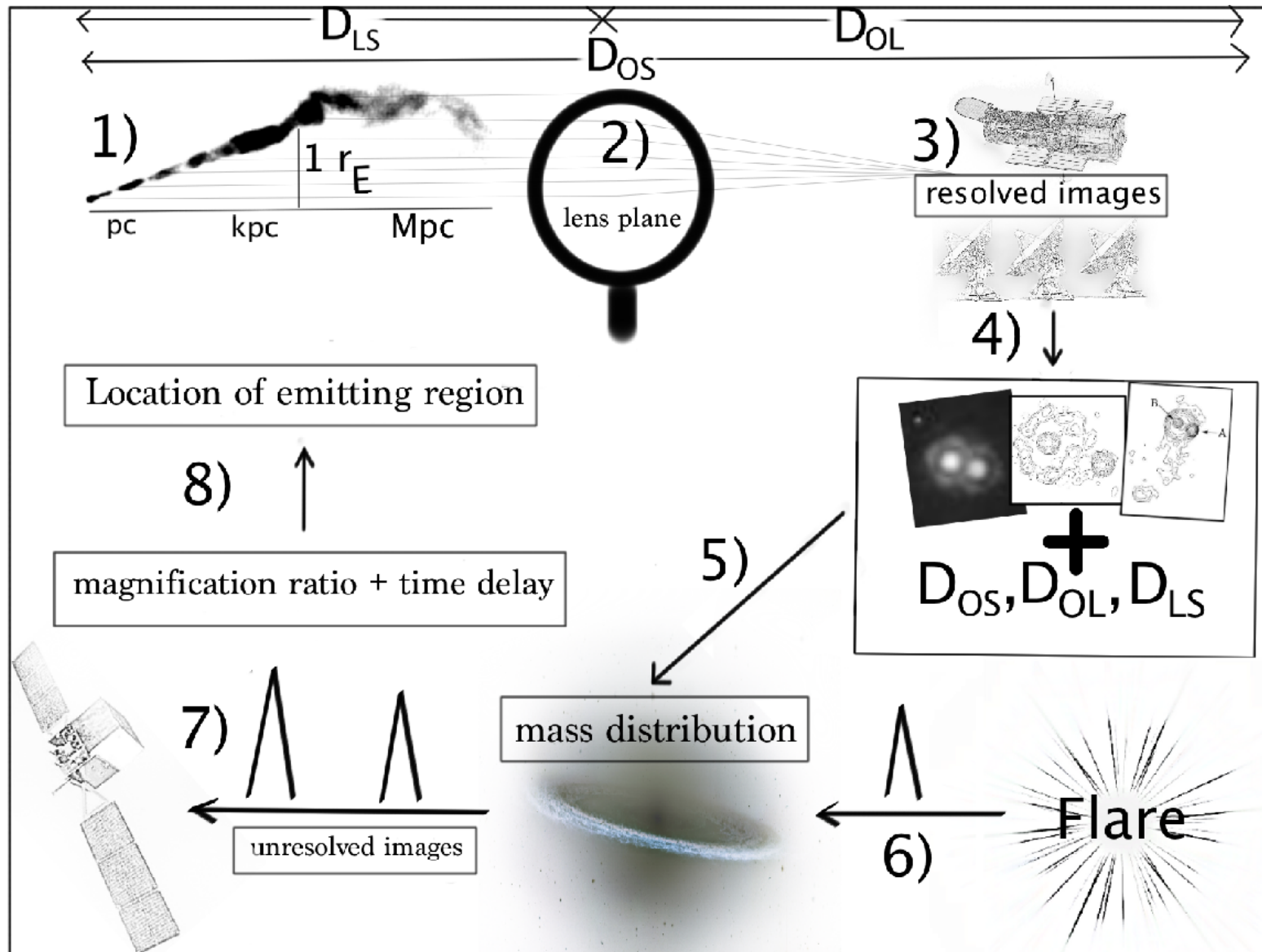
Spatial Origin of Flare 2



Gamma-ray Flare 2: The Maximum Peak Method



Application of strong lensing



Ambiguity of Gamma-Ray Origin

