The TANAMI Program Southern-Hemisphere AGN on (Sub-)Parsec Scales



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The TANAMI Program

Tracking Active Galactic Nuclei with Austral Milliarcsecond Interferomerty



- bimonthly VLBI monitoring of extragalactic jets south of $\delta = -30^{\circ}$ since 2007
- simultaneous dual-frequency observations at 8.4 & 22.3 GHz

The TANAMI Array



- dual-freq. observations with LBA, NASA's DSN, Hartebeesthoek
- additional 8.4 GHz monitoring with GARS, TIGO & Warkworth

Source Selection



- initially: hybrid radio & γ -ray selected sample of southern extragalactic jets
 - γ -ray loud sub-sample based on EGRET results
 - flux limited radio loud subsample
- \rightarrow new Fermi/LAT detected sources continuously added
- \rightarrow initially 43, currently 79 sources

Multiwavelength Approach

 contemporaneous high resolution VLBI monitoring at 8 & 22 GHz

 \rightarrow evolution of simultaneous spectral index maps at pc scales



TANAMI images: simultaneous 8.4 & 22.3 GHz, spectral index map (Kadler et al. in prep.)

more about PKS 0208-512 in J. Blanchard's talk!

Multiwavelength Approach

in addition to dual-frequency VLBI monitoring...

- Fermi/LAT
- pointed observations with RXTE
- Swift survey program
- optical program with Rapid Eye Mount (REM, INAF)
- flux density monitoring with ATCA
- Ceduna-Hobart Interferometer (CHI)
- ightarrow time evolution of simultaneous SEDs
- → SED modeling by NRL & University of Würzburg





http://pulsar.sternwarte.uni-erlangen.de/tanami/pubs



- Ojha et al. 2010: first 8.4 GHz-epoch paper
- first-ever VLBI images for some of newly added Fermi-bright sources
- contributions to simultaneous broadband SEDs of several sources

First Results

- contributions to LAT-publications: PKS 1454-354, SED paper, Cen A core,
- TANAMI-1FGL-analysis (led by M. Böck):
 - 55/75 sources LAT-detected
 - all 8 BL Lacs but only 24/32 Quasars (75%)
 - similar result as for MOJAVE
 - upper limits on γ-ray fluxes for TANAMI sources not detected by LAT
 - 2 new detections beyond 1FGL
- high resolution observations of the γ-ray bright galaxy Centaurus A ...

 γ -ray spectral index distribution of TANAMI sources (preliminary, M. Böck)

More information:

http://pulsar.sternwarte.uni-erlangen.de/tanami Ojha et al. 2010, A&A, 519, A45

TANAMI observations of Centaurus A

- four 8 4 GHz observations
- one simultaneous 22.3 GHz epoch
- closest AGN: d ~ 3.8 Mpc
 ⇒ 1 mas ≅ 0.018 pc



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C. Müller et al. 2011, A&A, 530, L11

First dual-frequency images of Cen A:

- \rightarrow Resolve innermost mas-scale jet into discrete components at both frequencies
- $\rightarrow\,$ Well collimated jet at P.A $\sim 50^\circ$ with opening angle $\lesssim 12^\circ$
- \rightarrow Study spectral changes at sub-parsec scales

Spectral Index Map of Cen A's Sub-pc Scale Jet



- High resolution spectral index map
- Core shift of $\Delta \alpha_{rel} = -0.25\,{\rm mas}$ $\Delta \delta_{rel} = -0.2\,{\rm mas}$
- Inverted spectrum in core region
- Remarkable flat spectrum over inner few mas of jet
- Multiple optically thick emission regions

C. Müller et al. 2011, A&A, 530, L11



(Abdo et al. 2010)

SED of Cen A core emission

- quasi-simultaneous + archival data
- LAT accuracy ∼ 0.°1

What are the productions sites of the γ -rays?



(Ab do et al. 2010)

SED of Cen A core emission

- quasi-simultaneous + archival data
- LAT accuracy $\sim 0.^{\circ}1$
- → Constraints on emission models of broadband SEDs



 $\rightarrow \ \ \, \mbox{Multiple possible regions of} \\ high energy emission$

Cen A Jet Kinematics at Sub-parsec Scales



C. Müller et al. 2011, in prep.

Cen A Jet Kinematics at Sub-parsec Scales



C. Müller et al. 2011, in prep.

Cen A Jet Kinematics at Sub-parsec Scales

ore 6.**B** $\operatorname{ore}^{\operatorname{at}}$ ore J4J5 $0\overline{.1~\rm pc}$ 20 mas

- complex substructure
- stationary component at $\sim 3.5\,{\rm mas}$
- jet widening & flux decrease at $\sim 23\,{\rm mas}$

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Jet Kinematics



- mean apparent jet speed $v_{\rm app,mean} \approx 2.7 \, {\rm mas/yr} \approx 0.16 c$
- moderate peak-flux variability
- differential motion: fastest component with $v_{\rm app} \approx 4 \, {\rm mas/yr}$

For whole TANAMI sample:

- * first spectral index maps and kinematics for all sources
- * joint *Fermi* analysis
- studies on individual sources
- new telescopes: Katherine (Northern Territory), ASKAP & Yarragadee (Western Australia)

For Cen A:

- * Proper motion analysis for jet and counterjet
- * Evolution of spectral index
- * Provide key parameters for broadband emission models

- TANAMI is the only large VLBI monitoring program of southern AGN
- bimonthly, simultaneous dual-frequency observations
- complementary multiwavelength observations
- Cen A: best-ever image of an AGN jet
- sub-parsec scale spectral index map
- multiple possible production sites of γ -rays