

# Afterglow Observations of *Fermi*-LAT Events and the Emerging Class of Hyper- Energetic Gamma-Ray Bursts

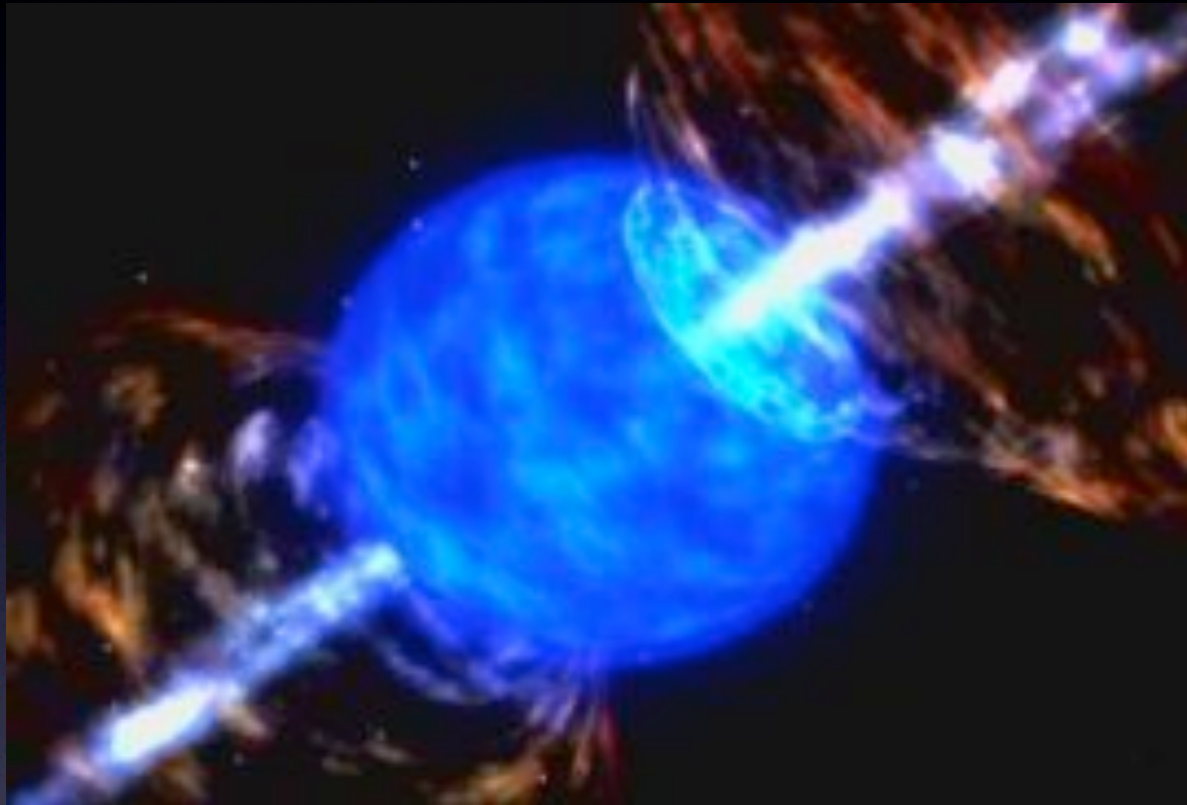
S. Bradley Cenko

2 November 2010

Gamma-Ray Bursts 2010 - Annapolis

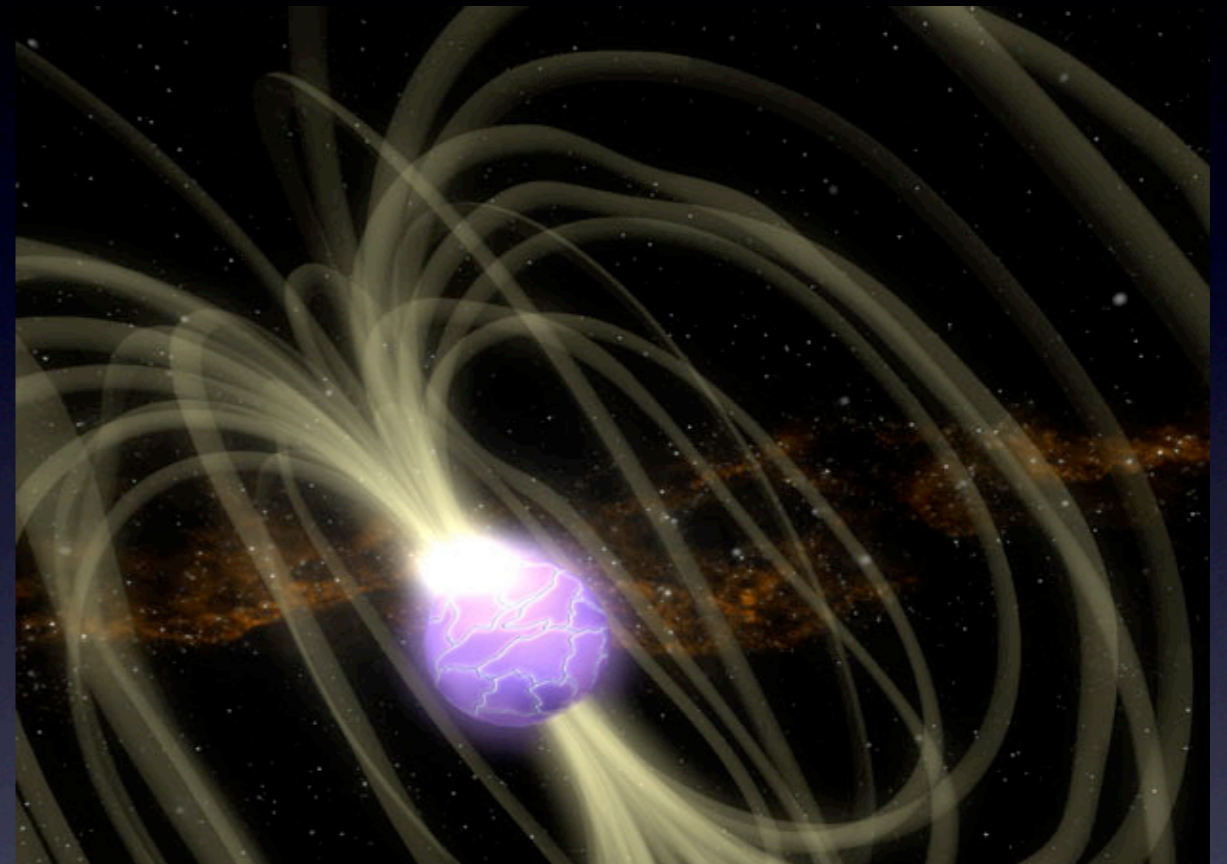
[astro-ph/1004.2900](https://arxiv.org/abs/astro-ph/1004.2900)

# Motivation: Central Engines



Collapsar: Woosley 93, MacFadyen & Woosley 99

$$E_{\max} \sim 0.1 Mc^2 \sim 10^{54} \text{ erg}$$

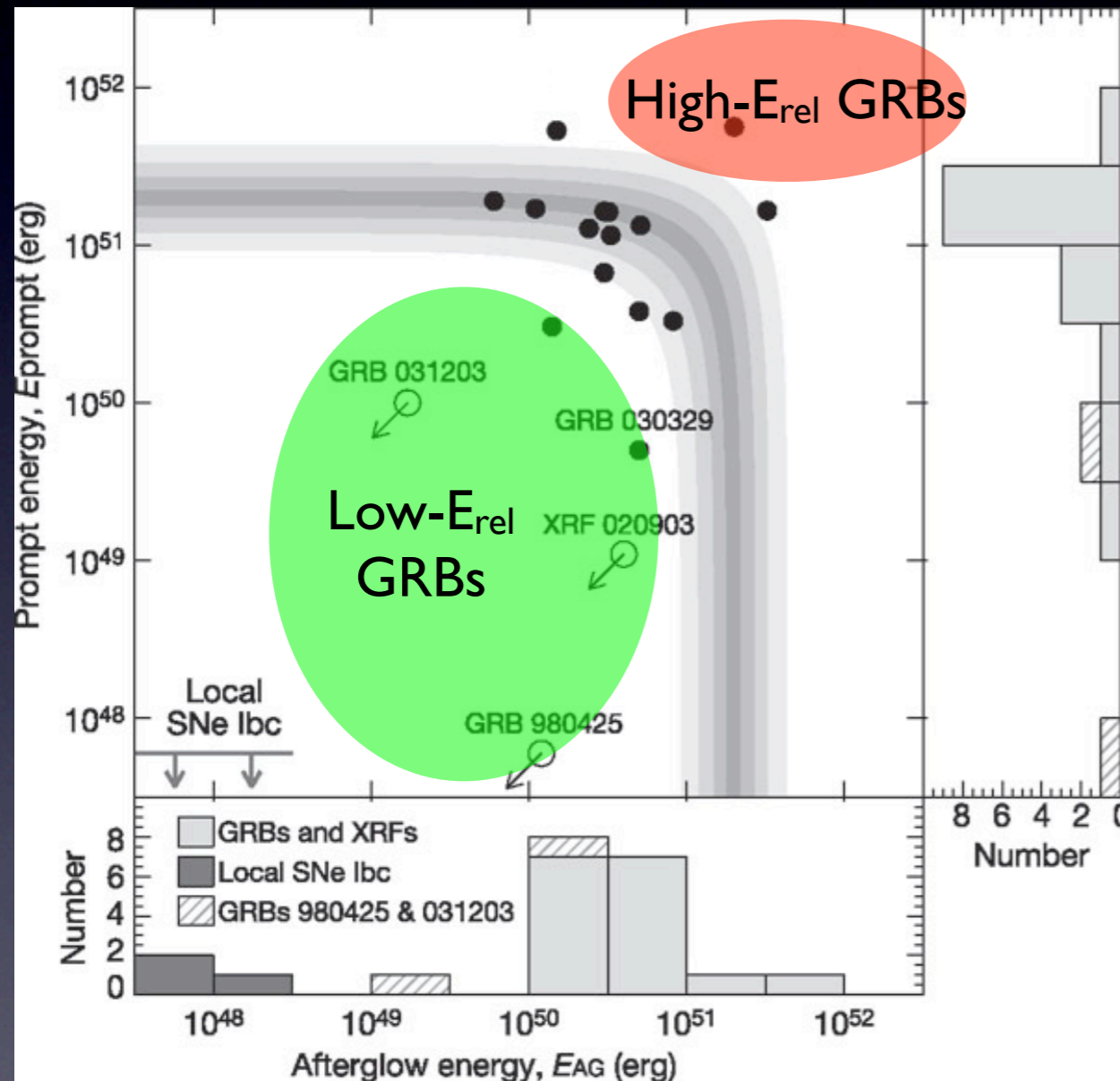


Magnetar: Usov 92, Duncan & Thompson 92

$$E_{\max} \sim I \Omega^2 / 2 \sim 3 \times 10^{52} \text{ erg}$$

Central engine models highly constrained by  
*geometry* ( $\gamma\theta$ ) and *energetics* ( $E_{\max}$ )

# Engine Diagnostics: $E_{\text{rel}}$

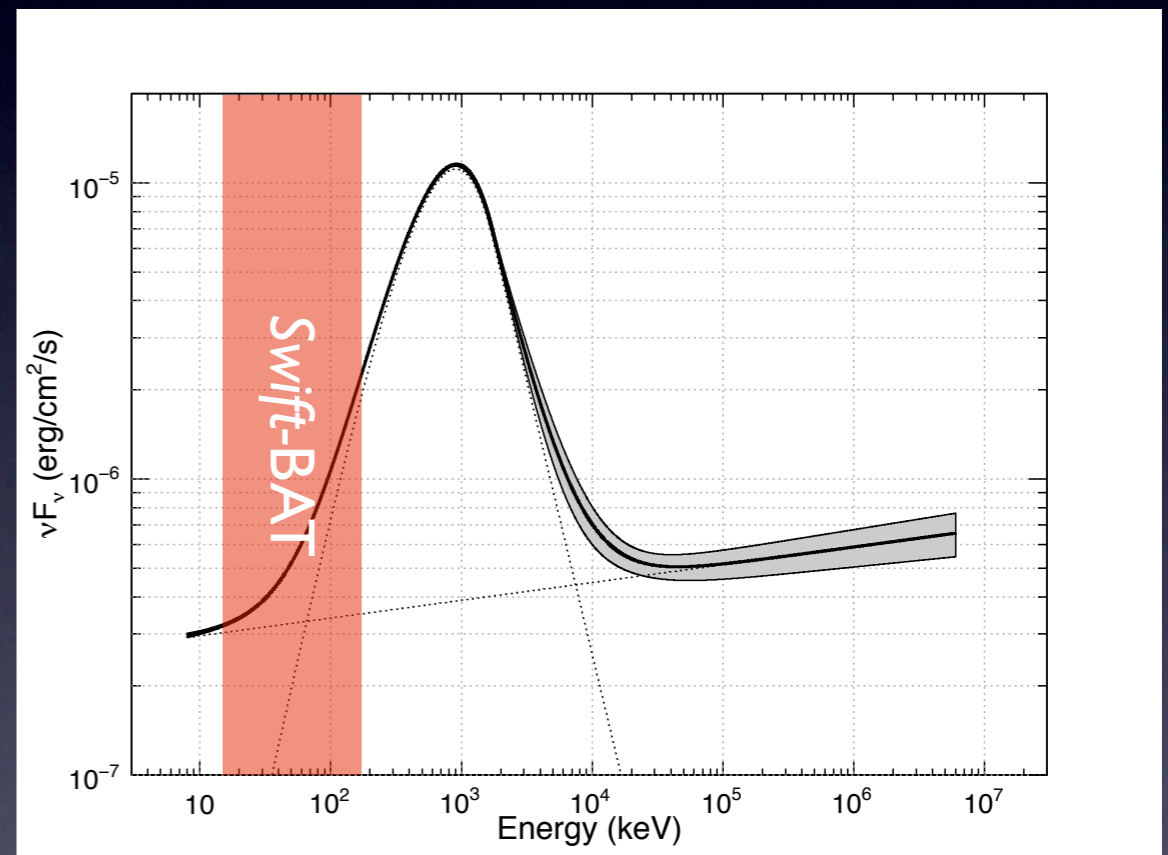


- Existence of a large population of uncollimated, sub-luminous GRBs suggests a diversity of engine mechanisms
- More recently, growing evidence for a substantial population of over-luminous GRBs with  $E_{\text{rel}} > 10^{52}$  erg
- Difficult to accommodate within magnetar framework

# Puzzling Energetics of *Swift* GRBs

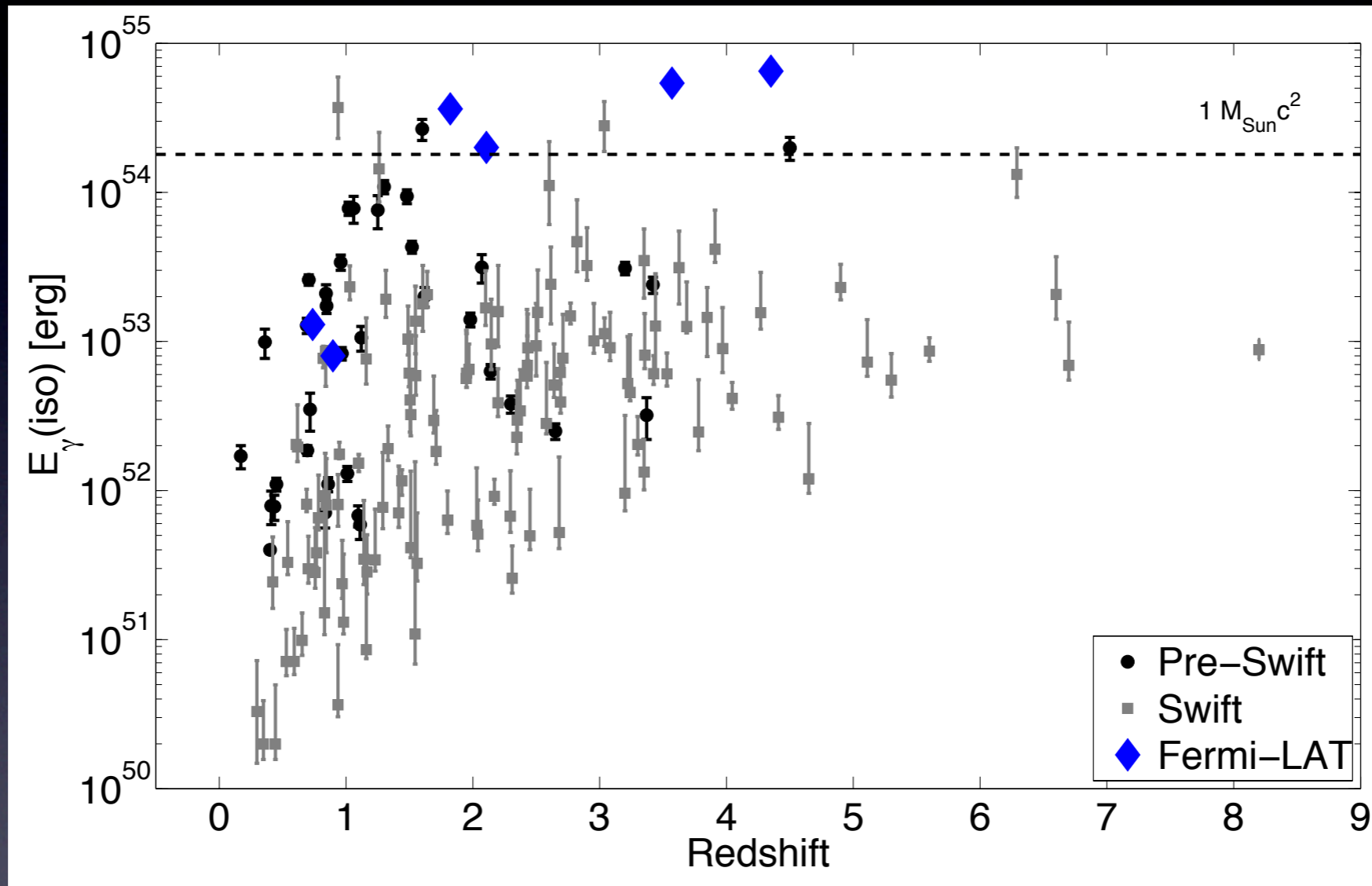
- Despite extensive searches, only a small fraction of *Swift* events with jet breaks ( $\theta$ ) measurements
- Faint, distant afterglows difficult for detailed follow-up studies
- Poor constraints on  $E_\gamma$  from limited bandpass of *Swift*-BAT

GRB 090902B



Abdo *et al.*, 2009

# *Fermi*-LAT and the Brightest GRBs

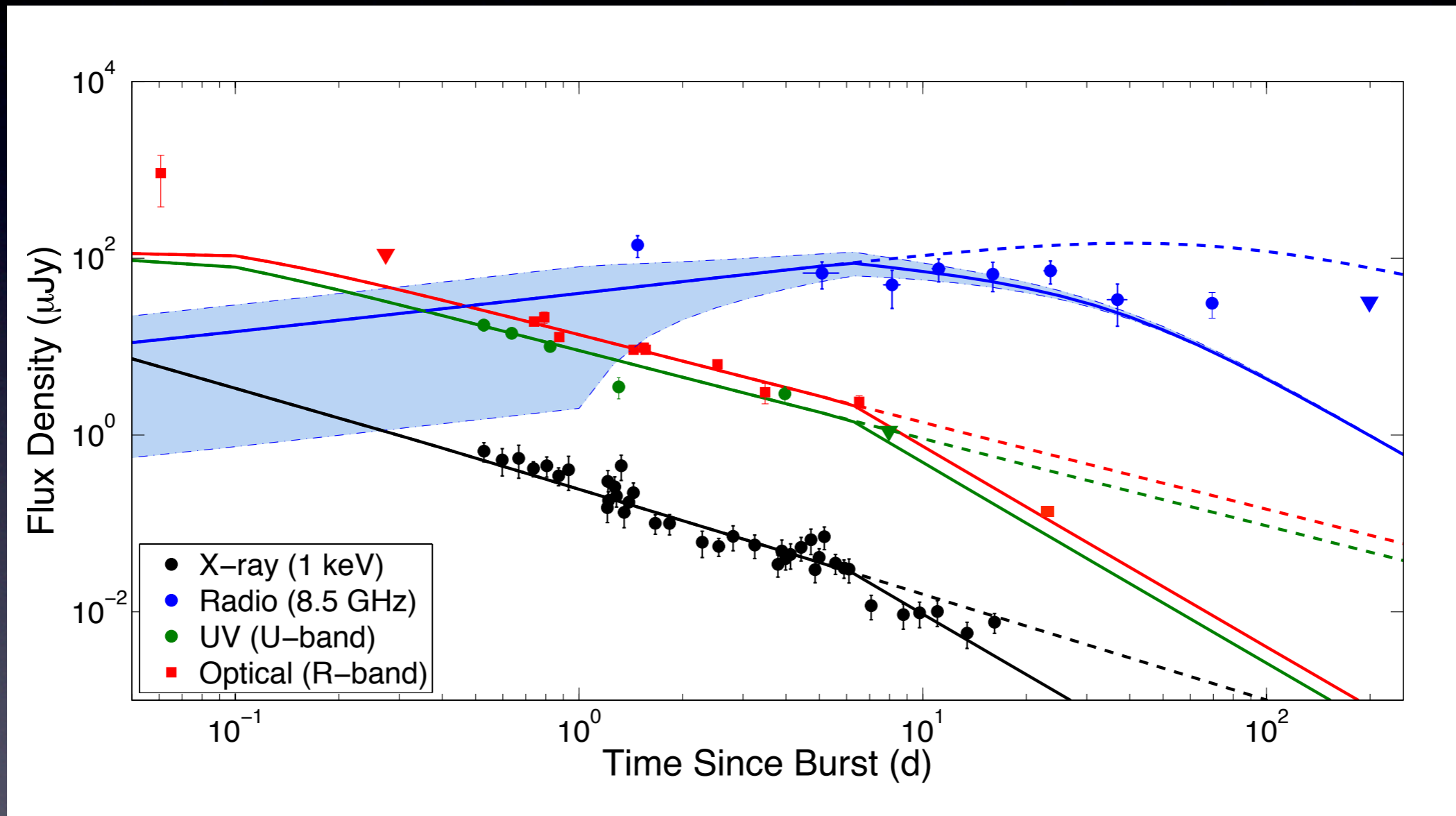


The *Fermi*-LAT offers an efficient way to target large  $E_{\gamma, \text{iso}}$  events

# Plan of Attack

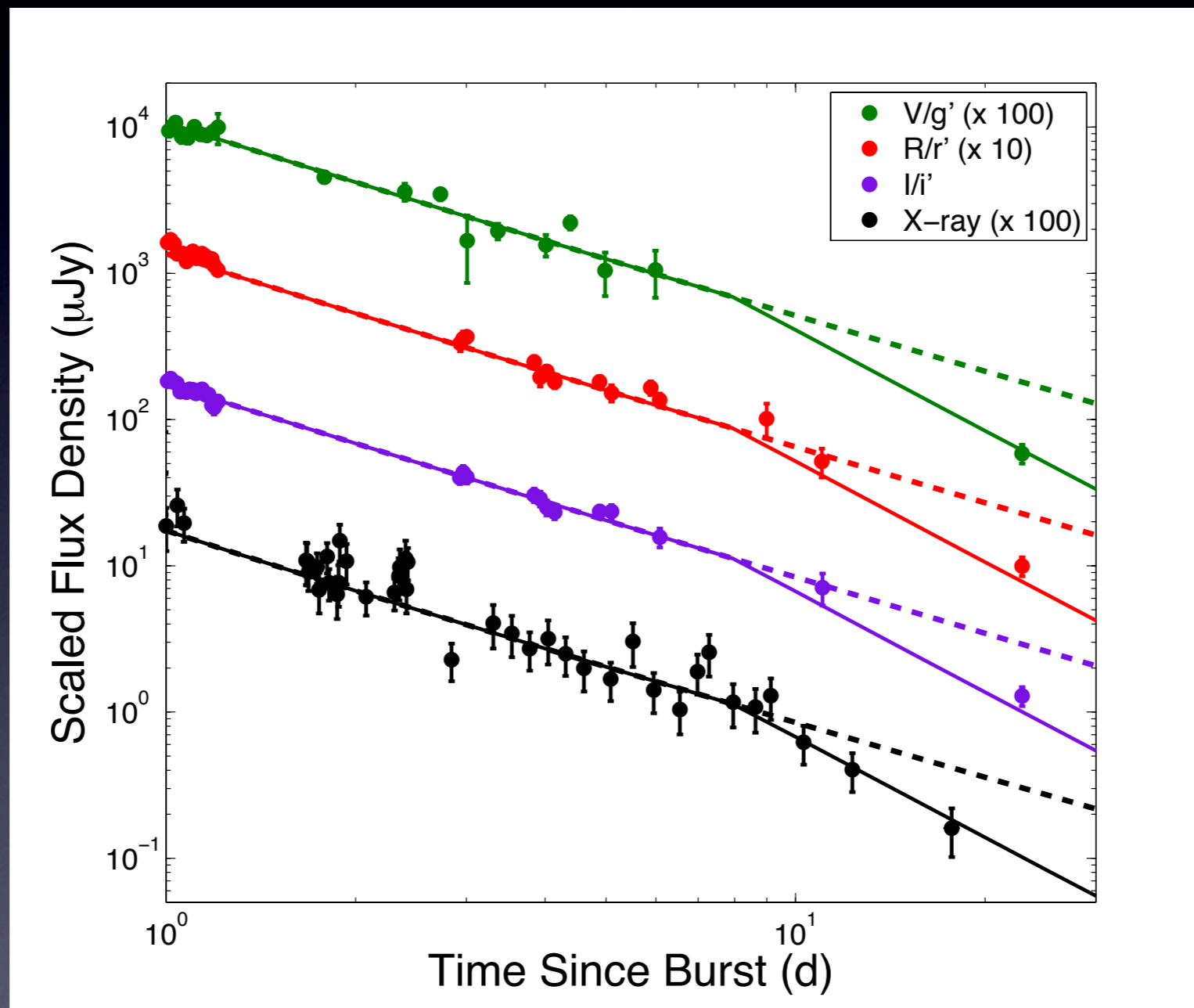
- Focus on brightest ( $\uparrow E_{\gamma,iso}$ ) GRBs, brightest afterglows, easiest jet breaks and hence opening angles to measure
- Fermi-*LAT* efficiently targets these GRBs, while also providing robust  $E_{\gamma,iso}$  and  $\gamma$  measurements
- Detailed afterglow follow-up in X-ray, optical, and radio out to late times
- Broadband modeling to infer opening angles, energetics, and circumburst density (this work)
- Host galaxy spectroscopy and photometry to correlate with larger scale environment (in prep)

# Light Curve Modeling: I



GRB 090902B:  $\theta \sim 3^\circ$ ;  $E_{\text{rel}} \sim 7 \times 10^{51}$  erg

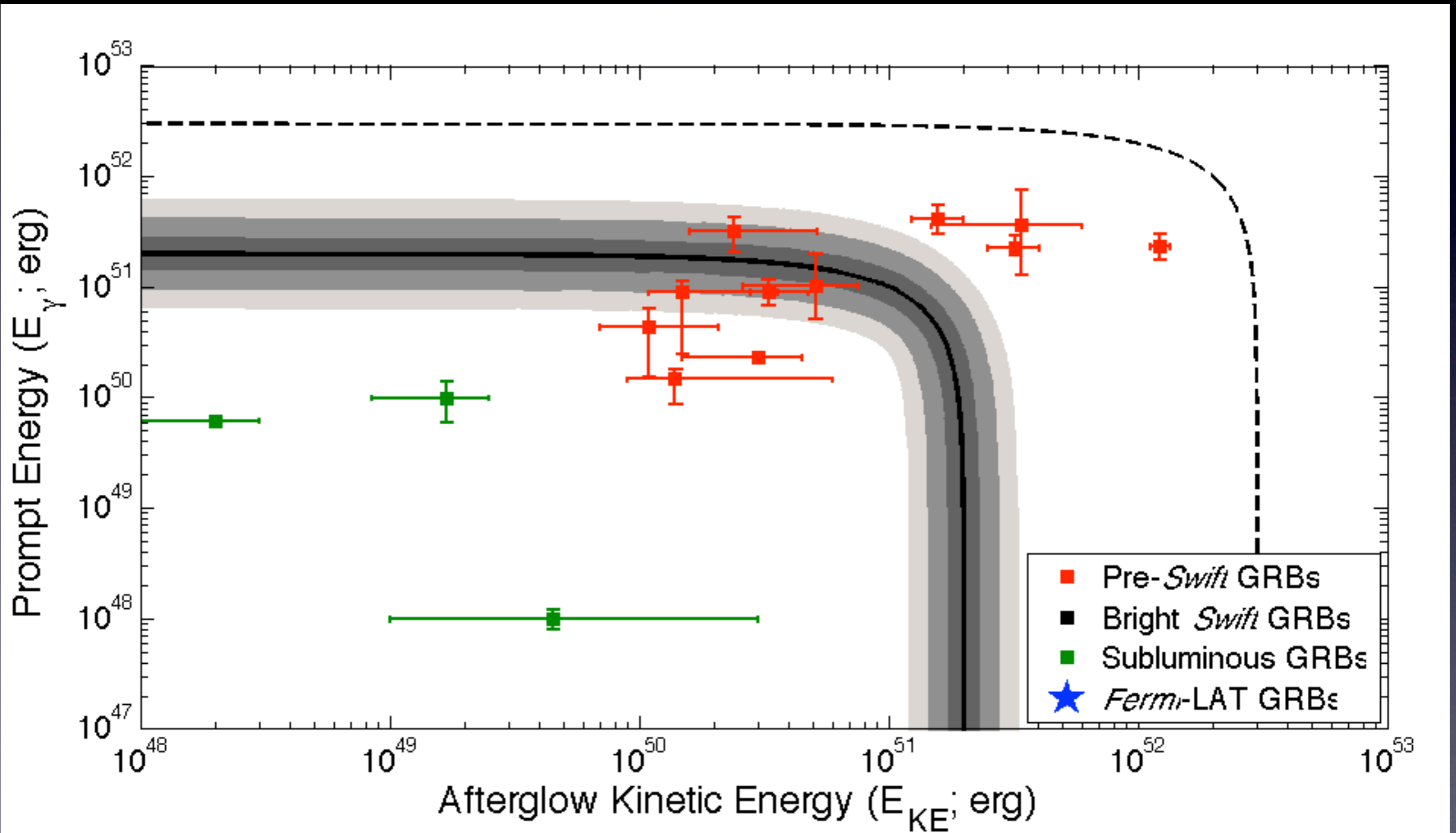
# Light Curve Modeling: II



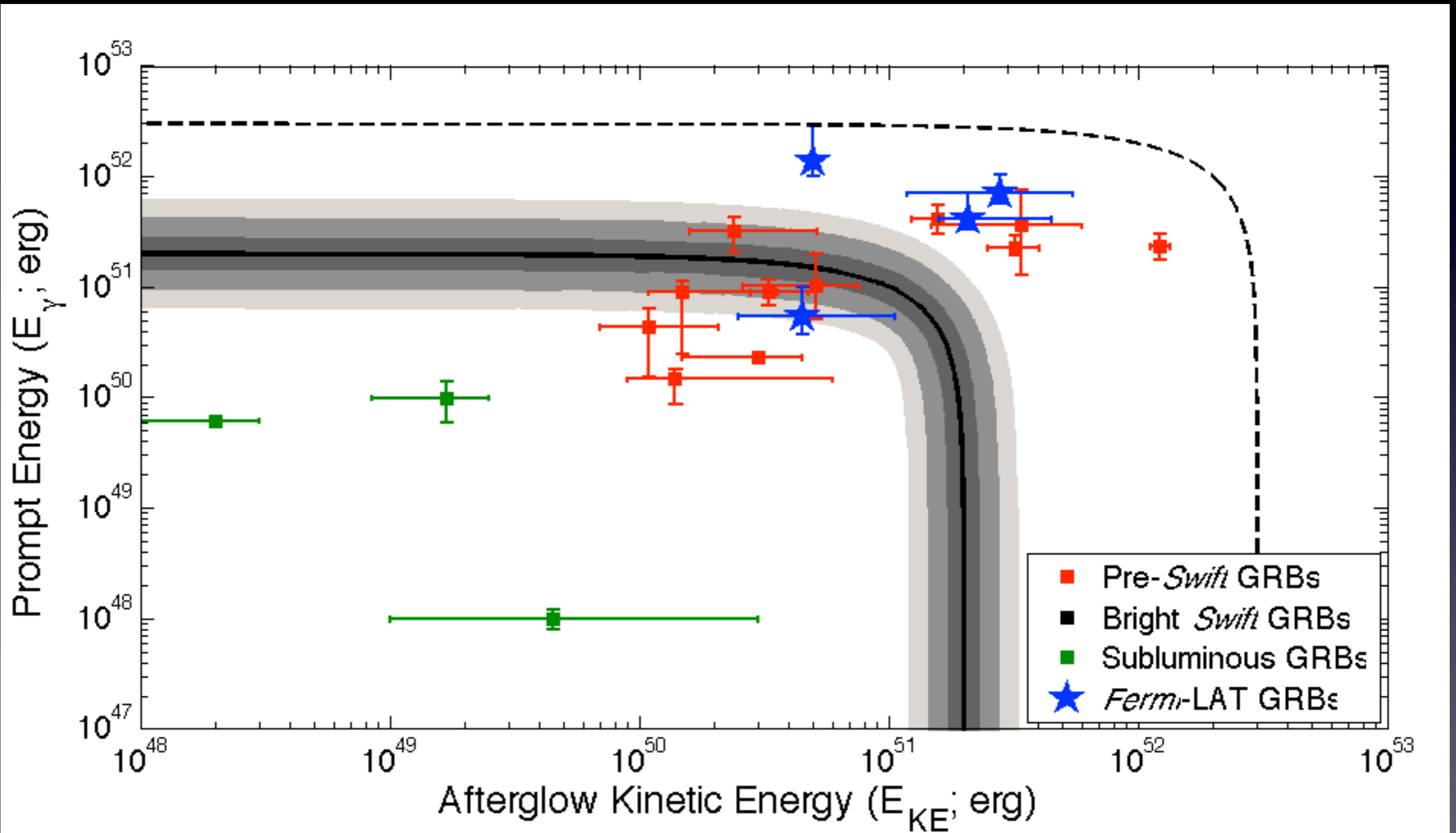
GRB 090926A:  $\theta \sim 7^\circ$ ;  $E_{\text{rel}} \sim 2 \times 10^{52}$  erg



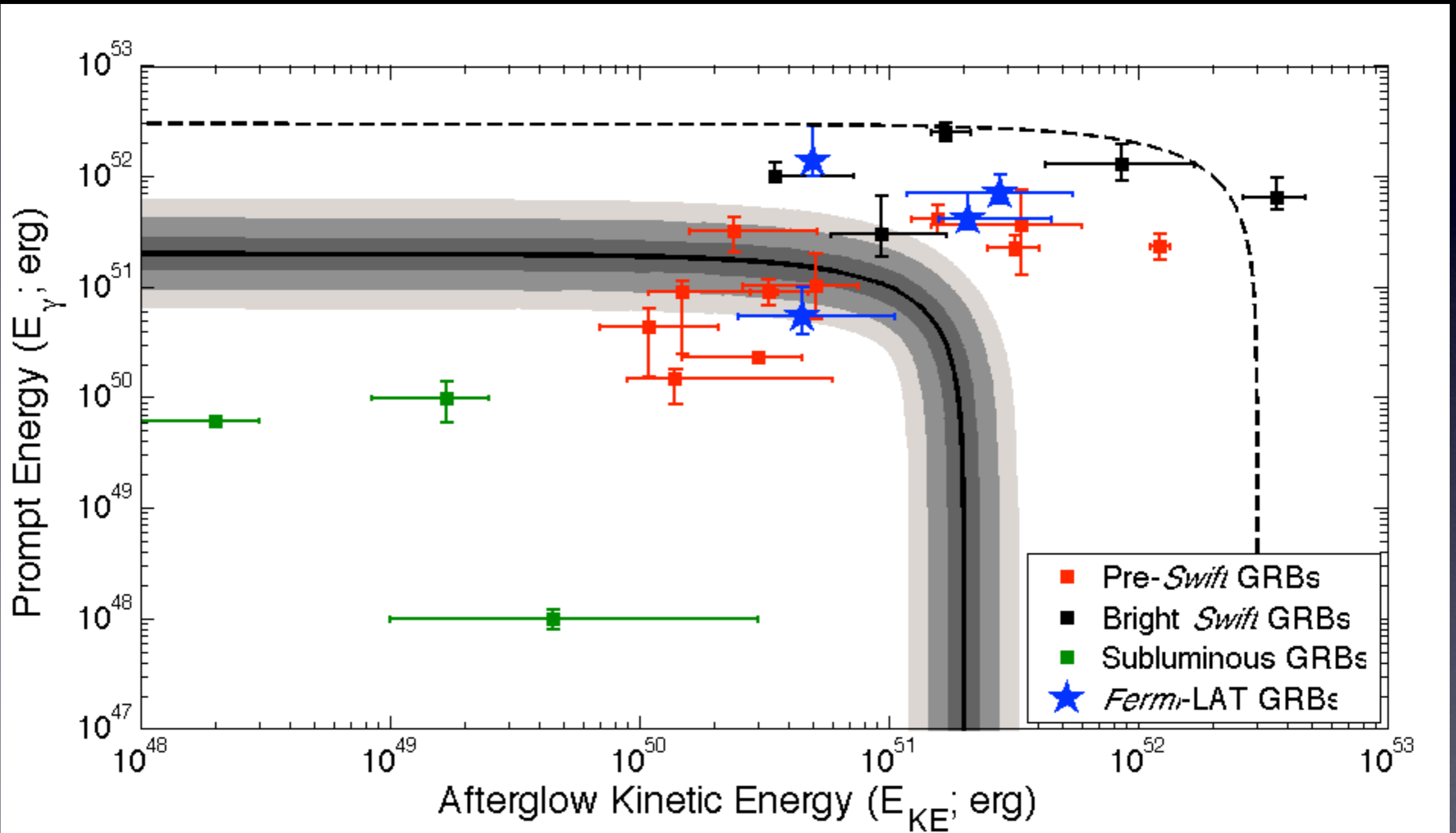
# Results: Energetics



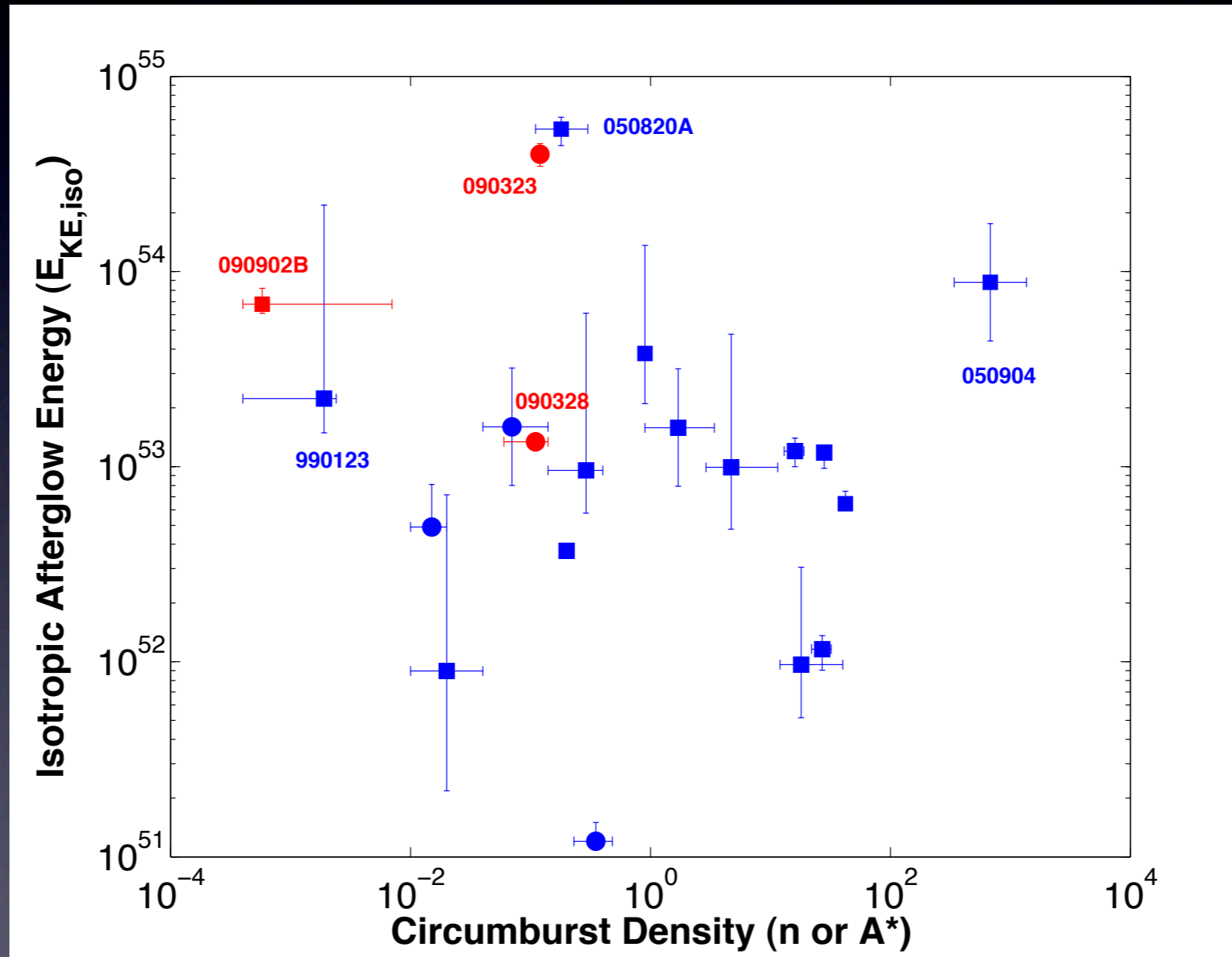
# Results: Energetics



# Results: Energetics



# Results: Circumburst Environment



# Conclusions

- Use broadband afterglow observations to constrain collimation and energetics from 4 *Fermi* LAT GRBs
- All 4 tightly collimated ( $\theta \sim 3\text{-}7^\circ$ )
- At least 1 event (090926A) has  $E_{\text{rel}}$  in excess of  $10^{52}$  erg  $\Rightarrow$  collapsar origin
- Low circumburst densities (consistent with rapidly rotating progenitors)
- Radio calorimetry with EVLA will soon provide independent confirmation of inferred  $E_{\text{rel}}$