



Fermi LAT Event Reconstruction Gallery

Fermi Summer School June 2011



- Subsystem Reconstruction
 - CAL
 - TKR
 - ACD
- Event Level Reconstruction
 - Energy
 - Direction
 - Event Classification
- Caveats:
 - This talk shows a mix of things we are doing, and things we will do in the future
 - Lots of pictures, not many details



CAL RECONSTRUCTION



CAL Roles (Slide from Instrument Talk)

- Primary Roles:
 - Energy reconstruction
 - Contributes to event trigger
- Other Roles:
 - "Energy Flow" axis at high energy
 - Seeds tracker pattern-recognition in complicated events
 - Background rejection
 - Shower topology e⁺e⁻ versus hadrons
 - Specific backgrounds
 - Up-going particles
 - Backsplash
 - Projection to ACD

CAL Reconstruction (Slide from Instrument talk)

- Apply per-crystal calibration
- Clustering: group hits into clusters (TBD)
 - Up to now treat whole CAL as single cluster
- Moments analysis

Gamma-ray

- Iterative procedure, minimize RMS w.r.t. shower axis
 - Cluster centroid (x,y,z)
 - Cluster axis (v_x,v_y,v_z)
 - Cluster moments and spread
 - Transverse, longitudinal RMS
- Energy Reconstruction (Multiple Methods)
 - Parametric correction for leakage out sides and gaps
 - Fit to cluster profile
 - Likelihood fit for event energy





Minimal Spanning Tree Clustering





MST: Link Creation





MST: Link Trimming





MST: Final Clusters





Cluster Moments Analysis: 1st Iteration





Cluster Moments Analysis: 2nd Iteration





Cluster Moments Analysis: 3rd Iteration





The Shower Profile from 100 MeV -> 100 GeV





Leakage Between Towers





TKR RECONSTRUCTION



TKR Roles (Slide from Instrument Talk)

- Primary Roles:
 - Direction reconstruction
 - Main event trigger
- Other roles:
 - Projection to CAL, ACD
 - Background rejection
 - pair-conversion
 - conversion vertex found?
 - (pre-)shower topology, e⁺e⁻ versus hadrons
 - specific backgrounds
 - backsplash from CAL
 - Up-going heavy ions stopping in TKR

TKR Reconstruction (Slide from Instrument Talk)

Hit clustering

Gamma-ray

- combine adjacent hit strips in clusters
- Start with CAL direction, if available
 - useful seed for high energy events, which are complicated
- Combinatoric search for straight(ish) lines
- Propagate lines to next plane, add hits as possible
- Kalman fit/filter technique
 - Combine information (hits) with loss of information (multiple scattering)
 - Requires energy estimate to handle multiple scattering
- Order tracks by "quality"
 - Favor longest, straightest track
 - Most likely to come from event origin
- Vertexing: try to combine 2 best tracks into single item





Starting the Combinatorial Search





Progressing Layer to Layer



The default region is 9σ (set very wide at this stage)



A typical 10 GeV Event in the TKR









Vector Links (Tree-Based)





Trimming the Tree





Vertexing (More Tomorrow)





ACD RECONSTRUCTION



- Primary Roles:
 - Offline background rejection
 - Hardware & onboard filter veto
- Other Roles
 - Identifying Heavy Ion (C,N,O + up) calibration events



- Apply tile calibrations
- Look for reason to veto event
 - Track extrapolation to ACD hit?
 - Compare ACD energy to CAL energy
 - Catches events where TKR direction is bad



Pattern Recognition I: Projecting the ACD

atan2(sqrt((x*x)+(y*y)),z+600.)*sin(atan2(y,x)):atan2(sqrt((x*x)+(y*y)),z+600)*cos(atan2(y,x))

Space Telescope



This figures shows a fisheye view of the ACD

The point is to project the ACD onto a plane

For Kicks: here is the transformation r = sqrt(x*x+y*y); zVal = z + 600.; rho = atan2(r, zVal); u = rho * (x / r);v = rho * (y / r);

Pattern Recognition II: building the hash set

atan2(sqrt((x*x)+(y*y)),z+600.)*sin(atan2(y,x)):atan2(sqrt((x*x)+(y*y)),z+600)*cos(atan2(y,x))

Space Telescope



During initialization we divide the projection into small squares, and calculate which ACD elements occupy which squares

This figure is heuristic, actual algorithm uses 20x20 grid and artificial expands the ACD elements by 50mm

Average square contains 1-3 tiles

Pattern Recognition III: getting the search set

 $atan2(sqrt((x^*x)+(y^*y)),z+600.)^*sin(atan2(y,x)):atan2(sqrt((x^*x)+(y^*y)),z+600)^*cos(atan2(y,x)))$

Gamma-ray Space Telescope



Blue dot shows point where track crosses nominal ACD volume

Purple circle shows search area around intersection point

Red shaded region shows searched squares

This is a very conservative algorithm. Most searches result in 5-12 ACD elements being found.

We calculate tile plane intersections and 3D POCAe for all of each of these elements.



EVENT LEVEL RECONSTRUCTION



- Extract useful information from tracks, clusters, and associations
 - We have used 200-500 different quantities for each event.
 Some examples:
 - Tkr1Chisq (χ^2 of best track in event)
 - Tkr1SSDVeto (# of hits in cone above best track)
 - AcdTrk1ActiveDist (distance from track to hit ACD tile)
 - CalTransRms (Transverse shower size in CAL)
 - CalTrackDoca (distance from track to CAL centroid)
 - etc...



Event Level Analysis



<u>**Complex**</u> multivariate analysis

Uses Classification Trees (CT) in conjunction with cuts

30+ individual cuts, in addition to CTs

Broken into many sub-sections



Outputs of the event level analysis





Data / MC comparisons





Tuning Cuts using Flight Data



Photon Rich Sample (Galactic Ridge)

Energy Dependent cut rejects 5% of event at all energies



Photon Poor Sample (High Latitude)

Cut rejections many more events



OTHER ISSUES



- Some algorithms require or benefit from information from other sub-systems
 - Determines execution order
 - Complicates validation
 - Independent samples are always nice for validation
- Huge number of events to reconstruct
 - Acquire events for downlink at 400Hz
 - => Need 200 cores running at 3 Hz to have 50% margin in processing time to keep up w/ data
 - Finite processing time for each event