

MARS ODYSSEY

ULYSSES

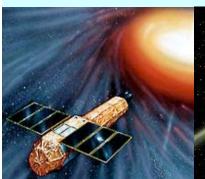
WIND

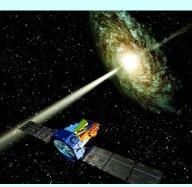
GLAST

MESSENGER

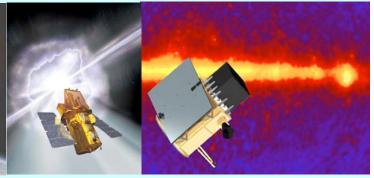
ADDING THE GLAST GAMMA-RAY BURST MONITOR TO THE 3rd INTERPLANETARY NETWORK

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SUZAKU

INTEGRAL

RHESSI

SWIFT

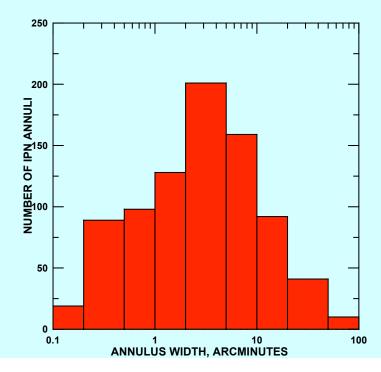
AGILE

THE GLAST GRBM

- 200 bursts/year
- Threshold sensitivity 0.35 photons cm⁻² s⁻¹ for untriggered events (cf. BATSE untriggered threshold ~0.1 photons cm⁻² s⁻¹)
- Location accuracies:
 - − ~8° onboard, real-time
 - − ~5°, near real-time
 - ~3° after manual intervention IPN, ~arcminutes

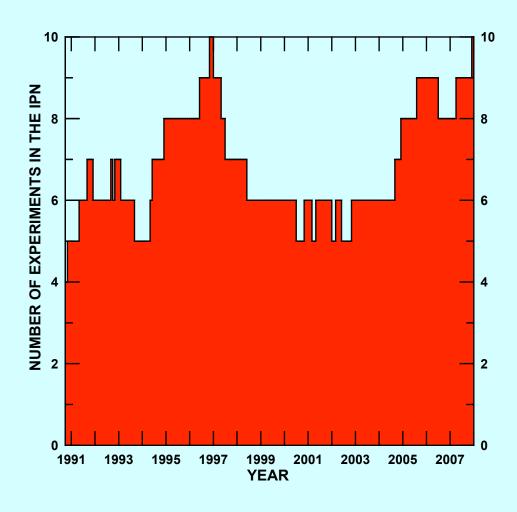
A QUICK REVIEW OF IPN – BATSE RESULTS

- •919 BATSE GRBs observed and localized by the IPN, + many SGR bursts
- •Data used to refine the BATSE localization error model (Pendleton et al. 1999)
- •IPN Results published in four catalogs



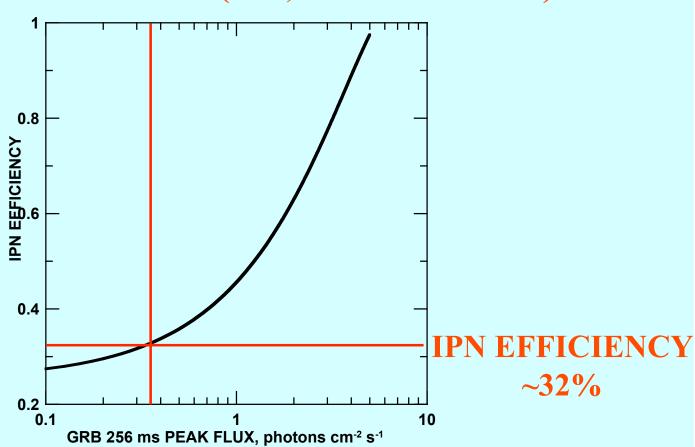
THE 3rd INTERPLANETARY NETWORK IN 2007

 Mars Odyssey, Ulysses, Wind, MESSENGER, Suzaku, INTEGRAL, RHESSI, Swift, + AGILE, GLAST, and a possible Space Station experiment

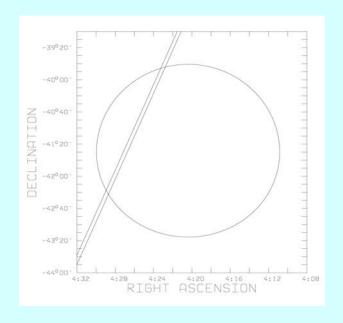


IPN EFFICIENCY (FLUX UNITS)

GLAST THRESHOLD (0.35, UNTRIGGERED)

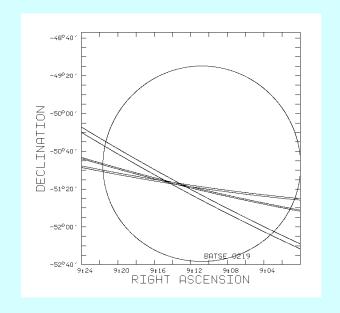


GRBM + 1 DISTANT IPN SPACECRAFT



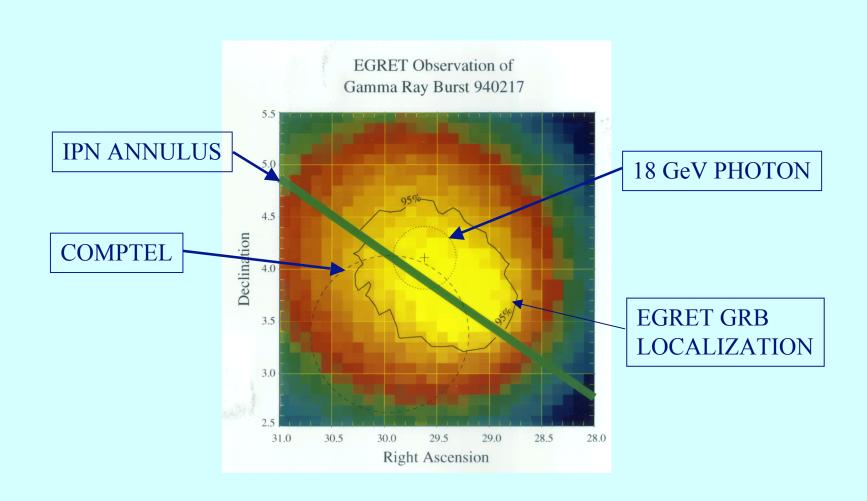
REDUCTION IN AREA X 30 - 100

GRBM + 2 OR MORE SPACECRAFT



REDUCTION IN AREA 3 – 4 ORDERS OF MAG.

EVEN A SINGLE IPN ANNULUS CAN MAKE A DIFFERENCE

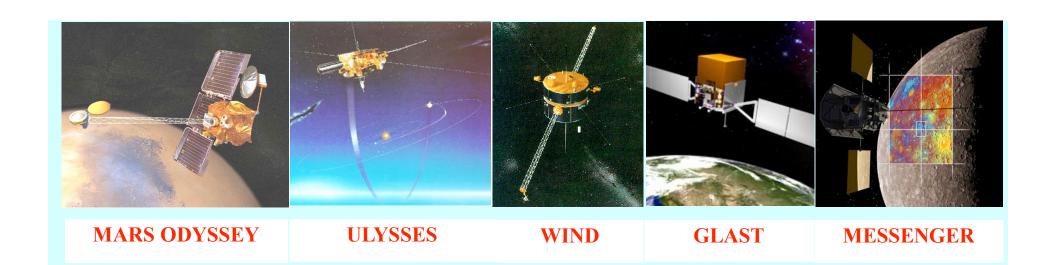


UTILISATION OF GLAST & IPN LOCALIZATIONS

- REFINING THE GRBM LOCALIZATION UNCERTAINTY
- LAT OBSERVATIONS OF GRBM ERROR BOXES
- AMANDA, ICE CUBE, RICE, ANITA v OBSERVATIONS
- MILAGRO SEARCHES FOR VHE EMISSION
- LIGO
- SWIFT ToO OBSERVATIONS

SUMMARY

- GRBM WILL BE THE 28th EXPERIMENT TO JOIN THE 3rd IPN SINCE 1990
- IPN/GRBM PERFORMANCE CAN BE PREDICTED BASED ON BATSE RESULTS
- GRBM ERROR BOXES CAN BE SUBSTANTIALLY IMPROVED FOR ~70 GRBs/YEAR
- IPN/GRBM LOCALIZATIONS CAN BE USED BY A WIDE COMMUNITY





IPN EFFICIENCY (FLUENCE UNITS)

