$$
\begin{aligned}
& \text { Fermi in the New Era } \\
& \text { of Radio Astronomy }
\end{aligned}
$$

## Alexander van der Horst

Astronomical Institute Anton Pannekoek
University of Amsterdam

』\{ $\begin{aligned} & \text { Astronomical institute } \\ & \text { ANTON PANNEKOEK }\end{aligned}$

## Current suite of radio arrays

- Very Large Array
- Westerbork Synthesis Radio Telescope
- Australian Telescope Compact Array
- Giant Metrewave Radio Telescope
- Ryle Telescope
- European VLBI Network
- Very Long Baseline Array
Larger - deeper - faster
- Very Large Array $\rightarrow$ Jansky Very Large Array
- Westerbork Synthesis Radio Telescope $\rightarrow$ Apertif
- Australian Telescope Compact Array
- Giant Metrewave Radio Telescope
- Ryle Telescope $\rightarrow$ Arcminute Microkelvin Imager
- European VLBI Network $\rightarrow$ e-EVN
- Very Long Baseline Array


## New kids on the block

- Low Frequency Array
- Long Wavelength Array
- Murchison Widefield Array
- MeerKAT
- Australian Square Kilometer Array Pathfinder

Square Kilometer Array

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> The (near) future looks radio bright!

## Exploring the low frequency radio sky

- Epoch of Reionisation (redshifted HI \& CO lines):
- first structure formation during dark ages
- Deep extragalactic surveys (continuum \& lines):
- high-z galaxies, clusters, cosmic star formation history
- AGN physics \& evolution
- Cosmic magnetism (polarization surveys):
- magnetic field evolution in galaxies over cosmic time
- Ultra high energy cosmic rays
- Solar science \& space weather
- Transient sources

The transient low frequency radio sky

## Incoherent emission

- Relatively slow variability
- Found mostly in images
- Explosive events \& jet sources
- Gamma-ray bursts
- Supernovae
- Magnetars
- X-ray binaries
- Active Galactic Nuclei
- Tidal disruption events


## Coherent emission

- Relatively fast variability
- Found mostly in time series
- Largely unexplored, exciting new science
- Theoretical predictions, e.g. GRBs
- Possible Lorimer bursts


## Low Frequency Array (LOFAR)



## International LOFAR Telescope



## Imaging survey speed



## LOFAR with -0.7 spectral correction



## LOFAR with -2.0 spectral correction



## LOFAR with -2.0 spectral correction

## Transient hunting with LOFAR

- Radio Sky Monitor / Zenith Monitoring Program

- Eight 7-beam LBA tiles (4500 deg²)
- Fourteen HBA tiles (1400 deg²)
- Phases with daily monitoring $\Rightarrow$ mJy sensitivity
- AARTFAAC
- 24/7 all-sky monitor with 6 central stations
- Piggy-back mode in all LOFAR observations
- LBA: whole sky, HBA: 1000 deg² $^{2}$
- Transient Buffer Boards
- 5 second storage
- Dispersion delay $\rightarrow$ subband approach


## First LOFAR transient



- ~100 mJy, varying by factor 10
- No counterpart or pulsations
- Nature unkown


Broderick et al. 2012

## LOFAR transient searches

- Multifrequency Snapshot Sky Survey
- Transient search in all fields, ind. 10 minute snapshots
- Simultaneous observations of North Celestial Pole: 280 images so far $\rightarrow$ no transients found at Jy level
- LOFAR Cycle 0 starting in December
- Proposed coordinated observations with PanSTARRS and Palomar Transient Factory
- Very recently: LOFAR UK-Chibolton responding to Fermi \& Swift GRB triggers (1 hour follow-up)
$\rightarrow$ first data taken after 20 seconds!


## LOFAR-GBM correlative studies

- Fermi Guest Investigator program
- AJvdH, Kouveliotou, Younes, Wijers, Fender, Stappers
- Large fields of view \& transient search capabilities
- Gamma-ray bursts:
- Searching for radio coherent emission $\rightarrow$ GBM triggers
- GRB energetics: prompt gamma-rays vs late-time radio
- Magnetars:
- Bursts and pulsed emission
- Giant flares
- Serendipity


## Conclusions

- Dawn of a new radio era:
- Upgrades of new facilities
- Square Kilometer Array pathfinders
- Large fields of view
- Unprecedented sensitivity in broad radio bands
- Extensions of the frequency \& time domains
- Synergy with Fermi:
- Extragalactic surveys
- Transients at various timescales

