Fermi in the New Era of Radio Astronomy

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Current suite of radio arrays

- Very Large Array
- Westerbork Synthesis Radio Telescope
- Australian Telescope Compact Array
- Giant Metrewave Radio Telescope
- Ryle Telescope
- European VLBI Network
- Very Long Baseline Array

Larger – deeper – faster

- Very Large Array → Jansky Very Large Array
- Westerbork Synthesis Radio Telescope → Apertif
- Australian Telescope Compact Array
- Giant Metrewave Radio Telescope
- European VLBI Network → e-EVN
- Very Long Baseline Array

New kids on the block

- Low Frequency Array
- Long Wavelength Array
- Murchison Widefield Array
- MeerKAT
- Australian Square Kilometer Array Pathfinder

Square Kilometer Array

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The (near) future Square Kilometer Array looks radio bright!

Exploring the low frequency radio sky

- Epoch of Reionisation (redshifted HI & CO lines):
 - first structure formation during dark ages
- Deep extragalactic surveys (continuum & lines):
 - high-z galaxies, clusters, cosmic star formation history
 - AGN physics & evolution
- Cosmic magnetism (polarization surveys):
 - magnetic field evolution in galaxies over cosmic time
- Ultra high energy cosmic rays
- Solar science & space weather
- Transient sources

The transient low frequency radio sky

Incoherent emission

- Relatively slow variability
- Found mostly in images
- Explosive events & jet sources
 - Gamma-ray bursts
 - Supernovae
 - Magnetars
 - X-ray binaries
 - Active Galactic Nuclei
 - Tidal disruption events

Coherent emission

- Relatively fast variability
- Found mostly in time series
- Largely unexplored, exciting new science
 - Theoretical predictions, e.g. GRBs
 - Possible Lorimer bursts

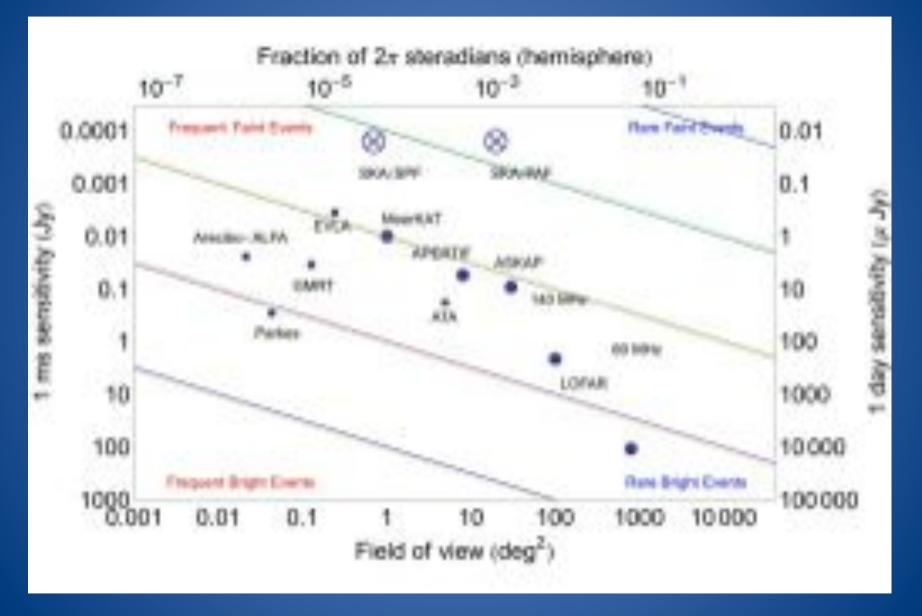
Low Frequency Array (LOFAR)



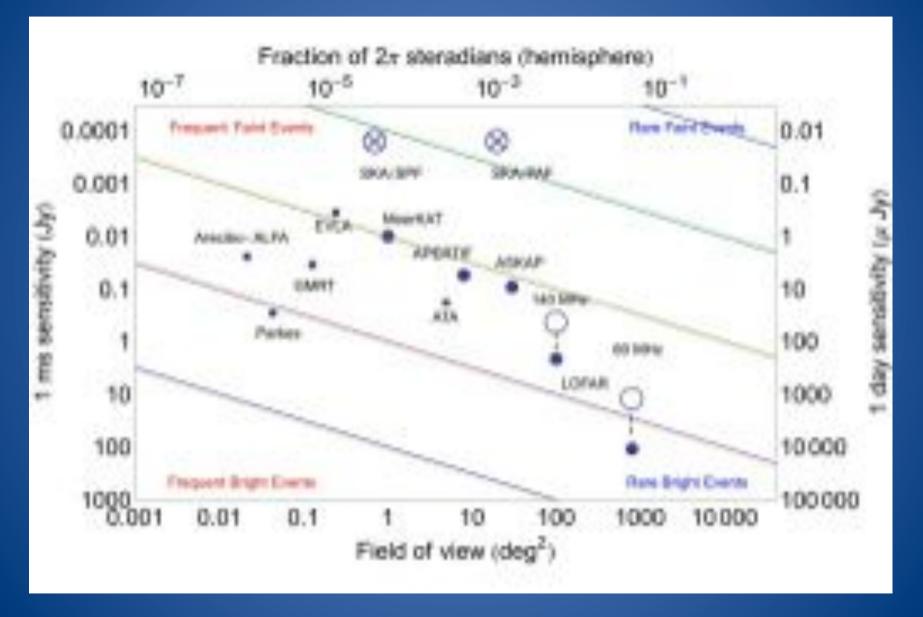
International LOFAR Telescope



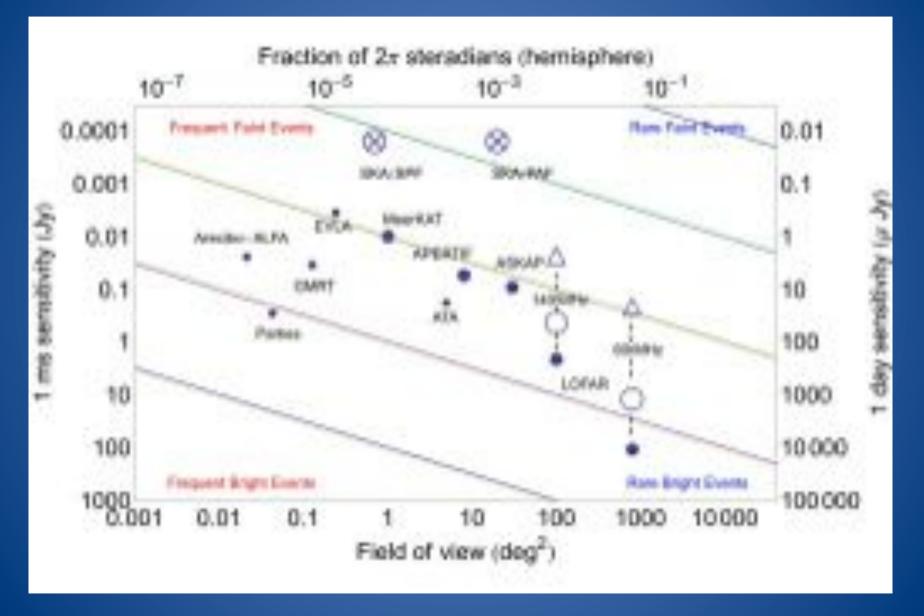
Imaging survey speed



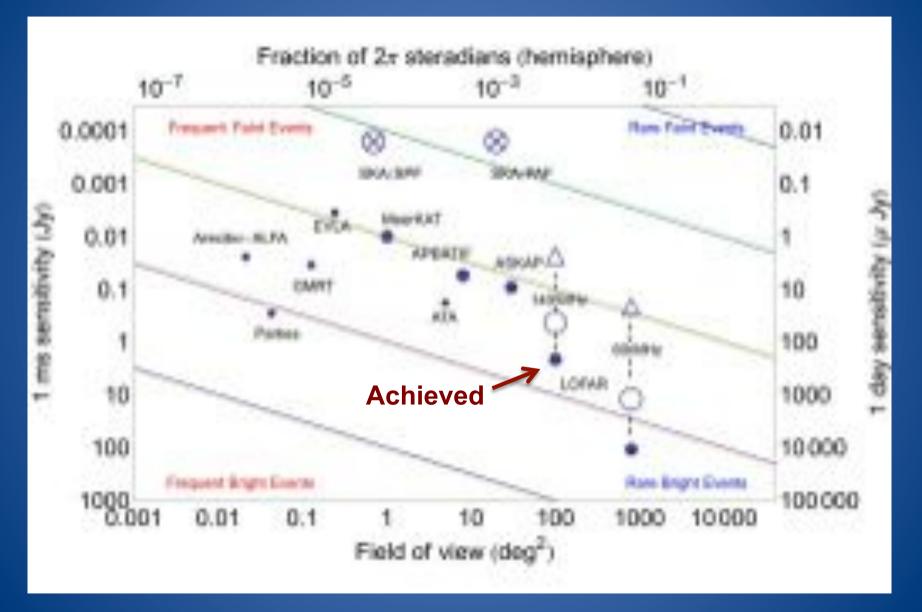
LOFAR with -0.7 spectral correction



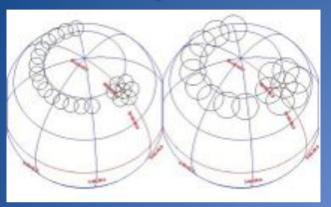
LOFAR with -2.0 spectral correction



LOFAR with -2.0 spectral correction



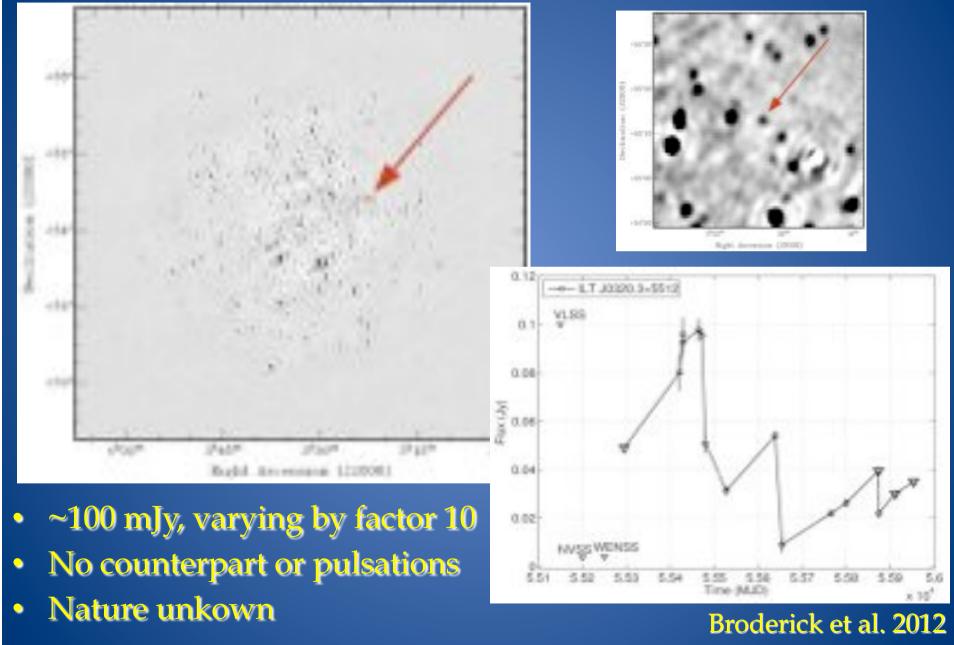
Transient hunting with LOFAR Radio Sky Monitor / Zenith Monitoring Program



- Eight 7-beam LBA tiles (4500 deg²)
- Fourteen HBA tiles (1400 deg²)
- Phases with daily monitoring
 → mJy sensitivity

- AARTFAAC
 - 24/7 all-sky monitor with 6 central stations
 - Piggy-back mode in all LOFAR observations
 - LBA: whole sky, HBA: 1000 deg²
- Transient Buffer Boards
 - 5 second storage
 - Dispersion delay \rightarrow subband approach

First LOFAR transient



LOFAR transient searches

- Multifrequency Snapshot Sky Survey
 - Transient search in all fields, incl. 10 minute snapshots
 - Simultaneous observations of North Celestial Pole:
 280 images so far → no transients found at Jy level
- LOFAR Cycle 0 starting in December
- Proposed coordinated observations with PanSTARRS and Palomar Transient Factory
- Very recently: LOFAR UK-Chibolton responding to Fermi & Swift GRB triggers (1 hour follow-up)

 first data taken after 20 seconds!

LOFAR-GBM correlative studies

- Fermi Guest Investigator program
 - AJvdH, Kouveliotou, Younes, Wijers, Fender, Stappers
- Large fields of view & transient search capabilities
- Gamma-ray bursts:
 - Searching for radio coherent emission → GBM triggers
 - GRB energetics: prompt gamma-rays vs late-time radio
- Magnetars:
 - Bursts and pulsed emission
 - Giant flares
- Serendipity

Conclusions

- Dawn of a new radio era:
 - Upgrades of new facilities
 - Square Kilometer Array pathfinders
 - Large fields of view
 - Unprecedented sensitivity in broad radio bands
 - Extensions of the frequency & time domains
- Synergy with Fermi:
 - Extragalactic surveys
 - Transients at various timescales