

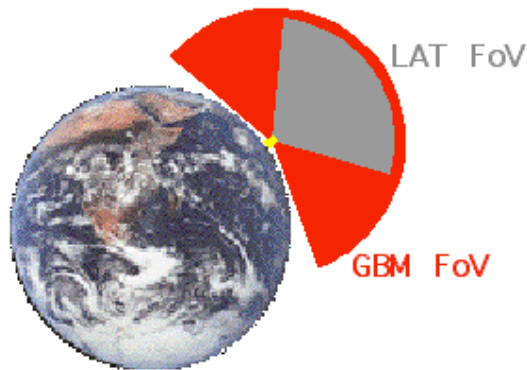
GRBs with GLAST and Swift



Jan. 17, 2007

GLAST Workshop / Hans Krimm

Numerology



GBM expects ~ 215 bursts/year
LAT ~ 50 bursts/year

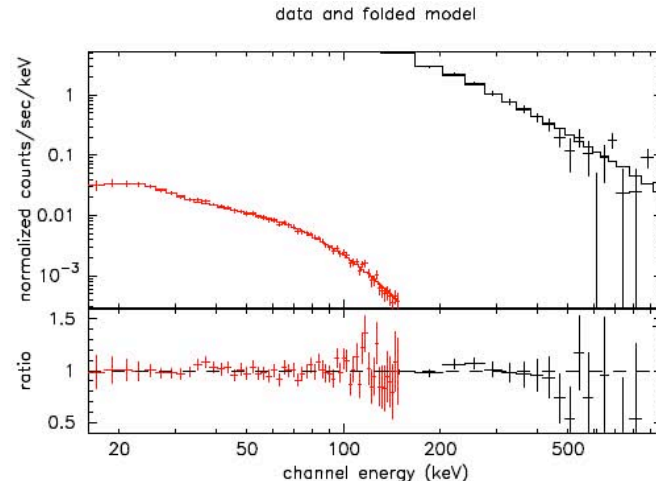
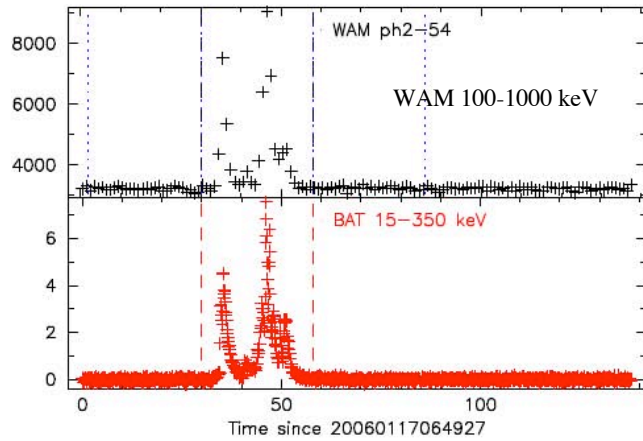
GBM is ~ 9 sr (LAT 2.4 sr)
BAT is 2 sr (partially coded)

So for GBM-BAT coincidences we expect
 $215 \text{ yr}^{-1} \times (2 \text{ sr} / 9 \text{ sr}) \approx 48 \text{ yr}^{-1} \approx 4/\text{month}$ ($\sim 1/\text{month}$ in LAT FoV)
Expect BAT to see all the GBM bursts that are in its FOV.

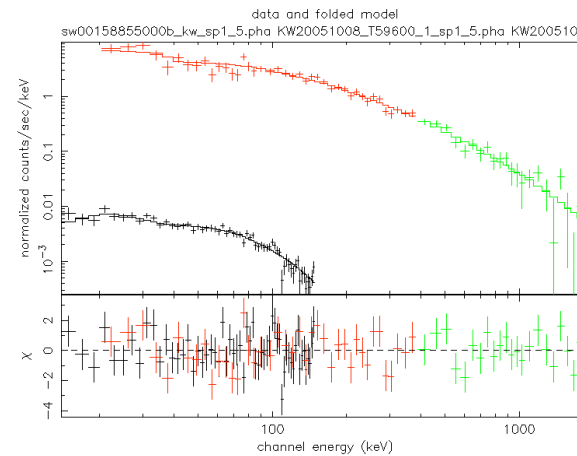
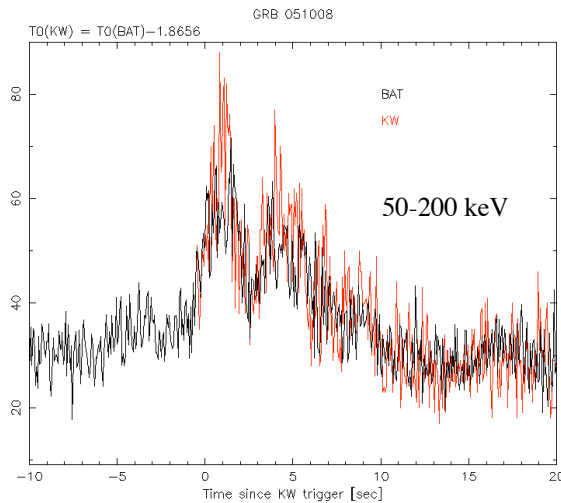
We currently see $\sim 1/\text{month}$ coincident with WIND/Konus and
 $\sim 1\text{-}2/\text{month}$ coincident with Suzaku/WAM.

Examples of joint fits with Konus and WAM

WAM: Shield for Suzaku HXD -- 100 keV -- 5 MeV



GRB 06117
Jointly
observed with
Swift/BAT
and
Suzaku/WAM



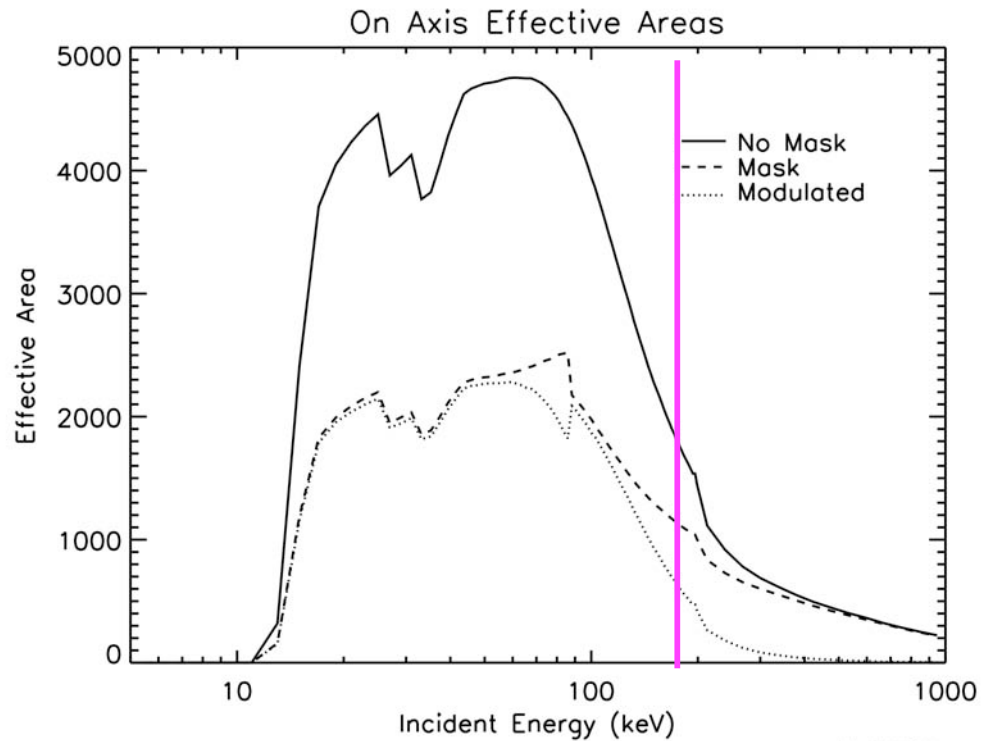
GRB 051008
Jointly
observed with
Swift/BAT
and
WIND/Konus

Konus: Scintillator detectors -- 10 keV -- 10 MeV

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Importance of extended energy range



BATSE E_{peak} distribution

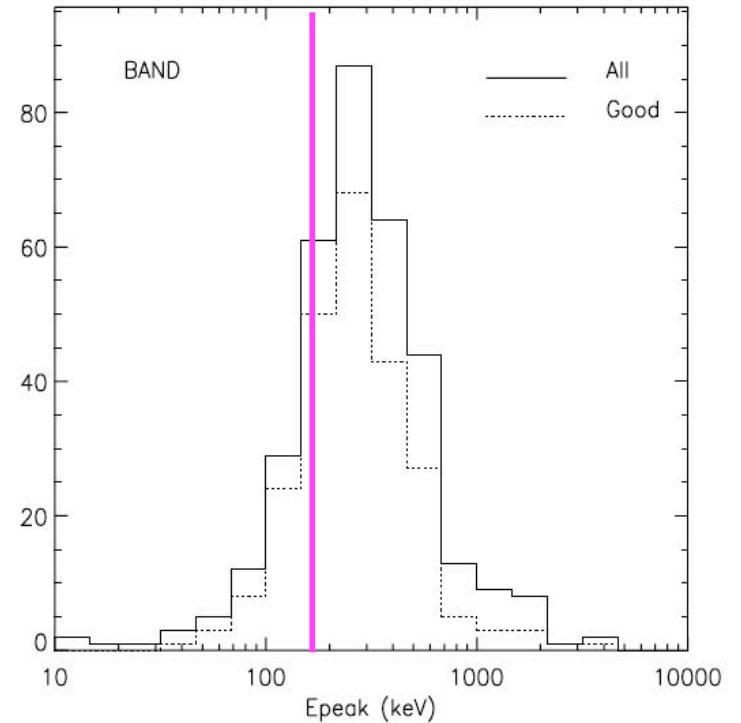
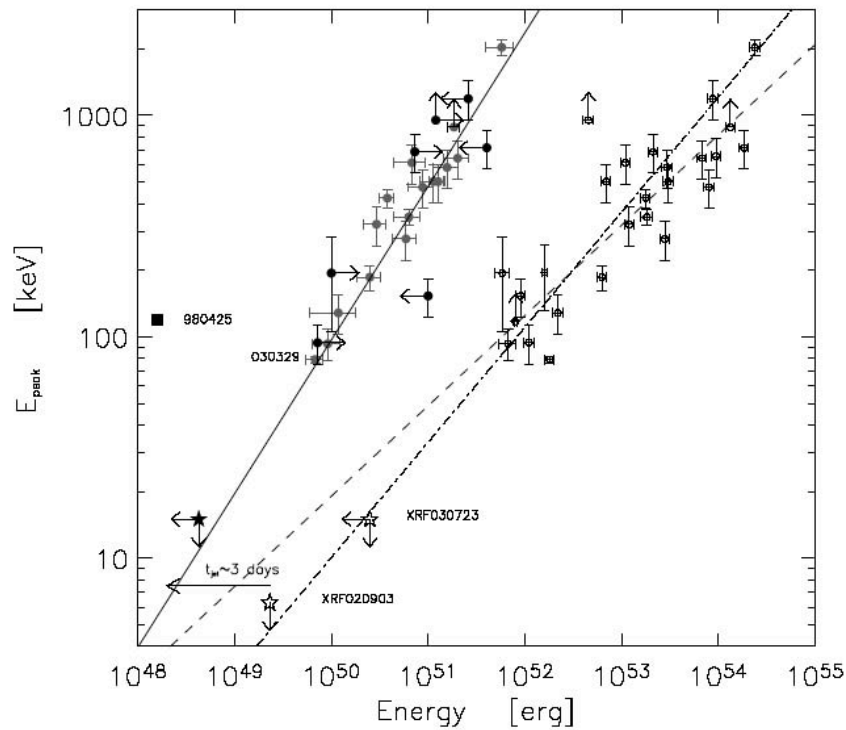


Figure from Y. Kaneko et al 2006, ApJS **166**, 298

Most GRBs have E_{peak} above BAT energy range

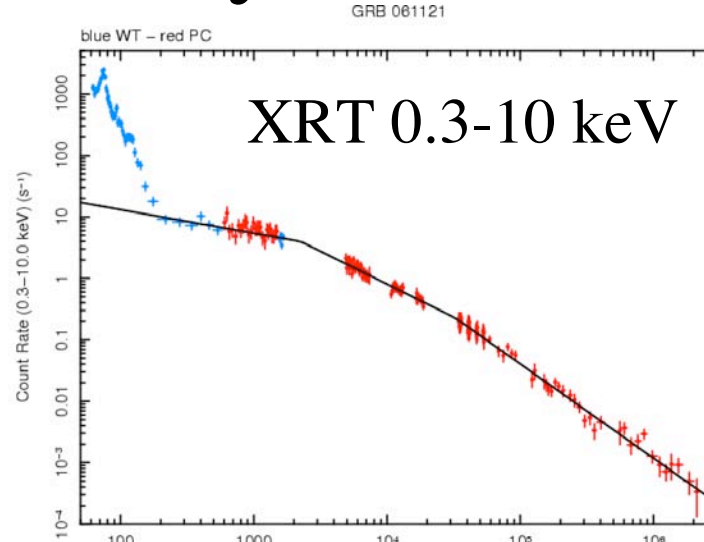
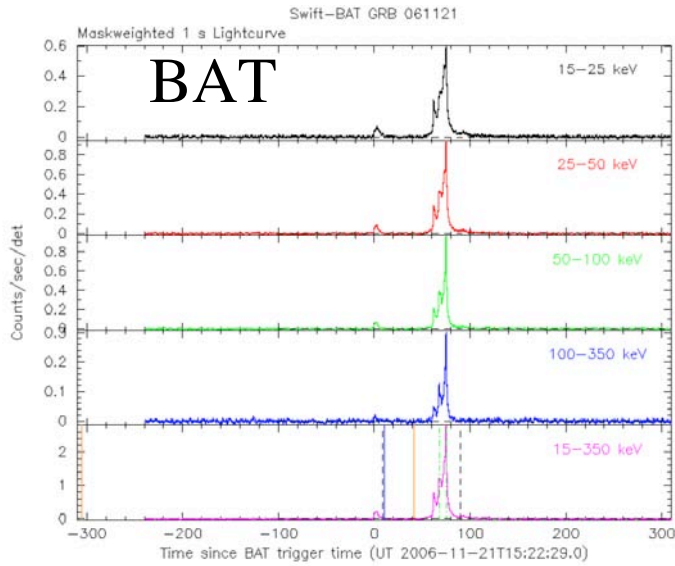
Need to determine validity of E_{peak} relations for Swift/GLAST bursts



Swift doesn't see jet breaks, but we do have redshifts for $\sim 40\%$ of bursts, so we can determine E_{iso} .

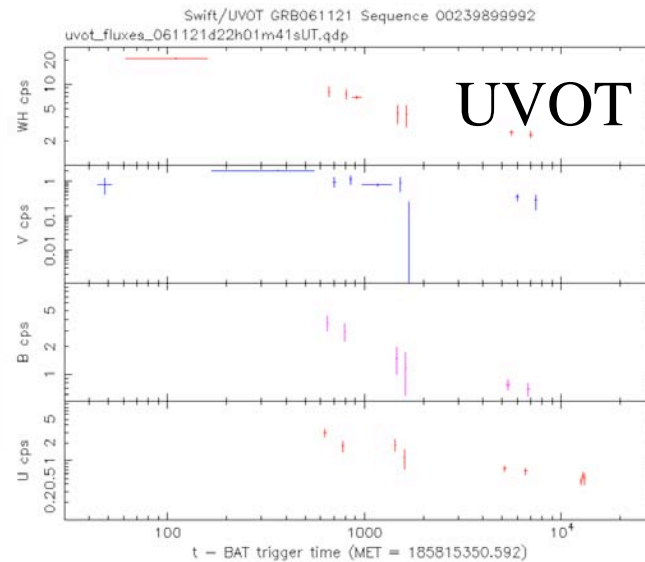
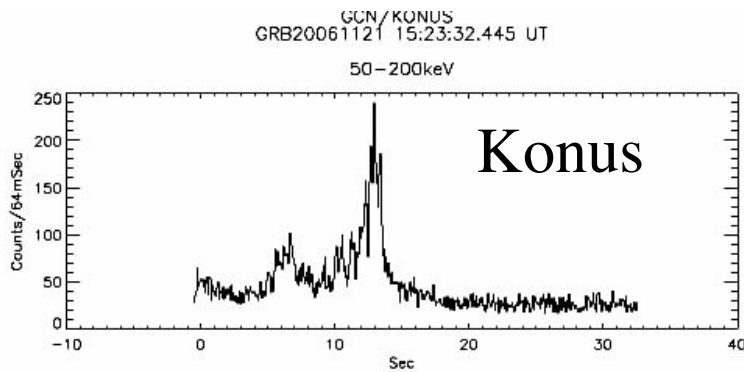
Figure from G. Ghirlanda et al 2004, ApJ **616**, 331

Possibility for very broad coverage



VLA (8.46 GHz) radio detection

With LAT high energy detection could cover 16 orders of magnitude



$z = 1.314$

GRB 061121

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Conclusions

- ⚡ Swift and GLAST to both operate for ≥ 3 years
- ⚡ Expect ~ 3 coincident bursts per month with GBM.
- ⚡ Swift/BAT team has developed experience with joints fits with Konus and Suzaku/WAM.
- ⚡ Can constrain E_{peak} for all coincident bursts and use redshift to determine burst luminosity, etc.
- ⚡ Some bursts will have simultaneous detections from R band to > 1 GeV.