

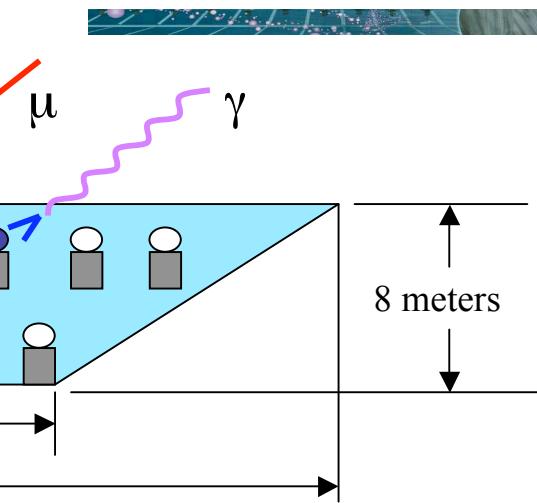
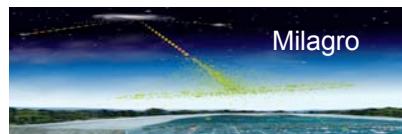
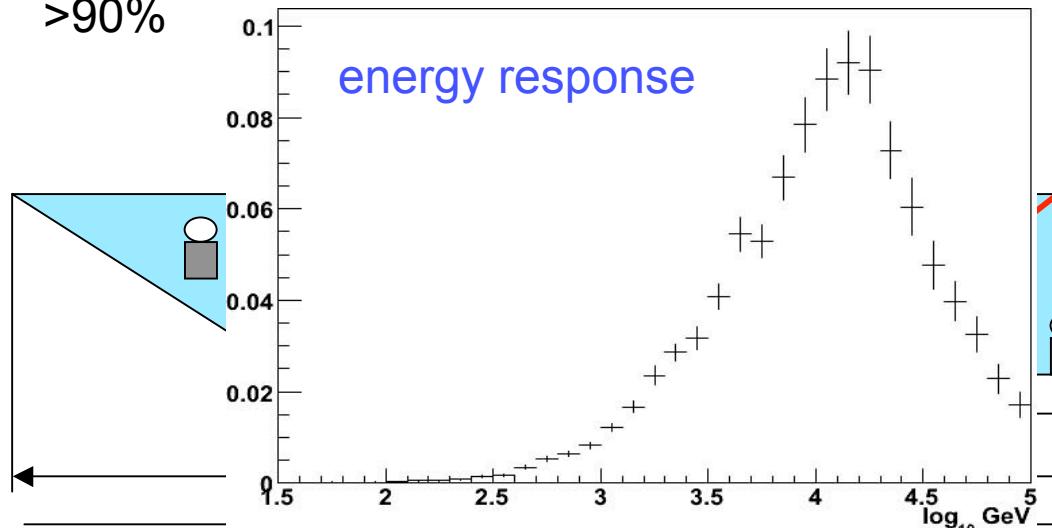
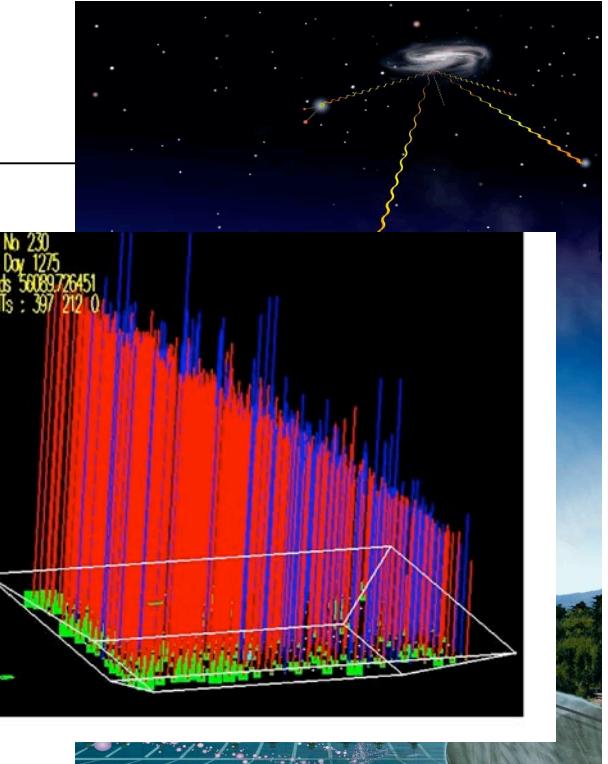
Survey of the Galactic Plane at 12 TeV with Milagro and the Discovery of MGRO J1909+06 and MGRO J2033+42

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for the Milagro Collaboration, Andrew Strong, Igor Moskalenko
GLAST Symposium
February 5-8, 2007

see Aous Abdo's plenary session talk and Gary Walker's poster for more details

How Does Milagro Work?

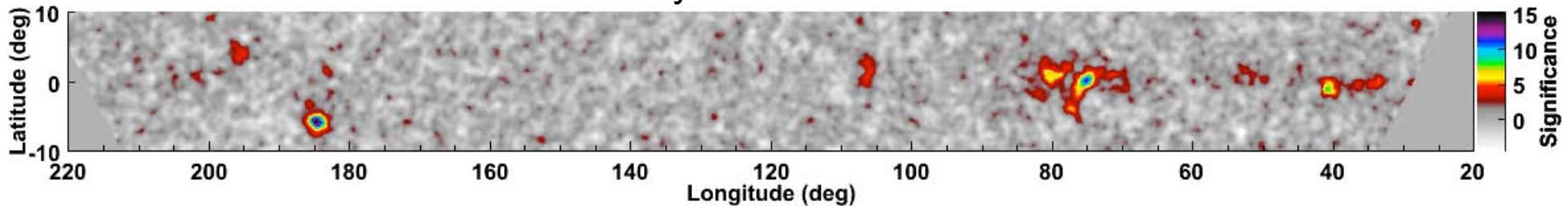
- Detect Particles in Extensive Air Showers from Cherenkov light created in 60m x 80 m x 8m pond containing filtered water
- Reconstruct shower direction to $\sim 0.5^\circ$ from the time different PMTs are hit
- 1700 Hz trigger rate mostly due to Extensive Air Showers created by cosmic rays
- Field of view is $\sim 2 \text{ sr}$ and the average duty factor is $>90\%$



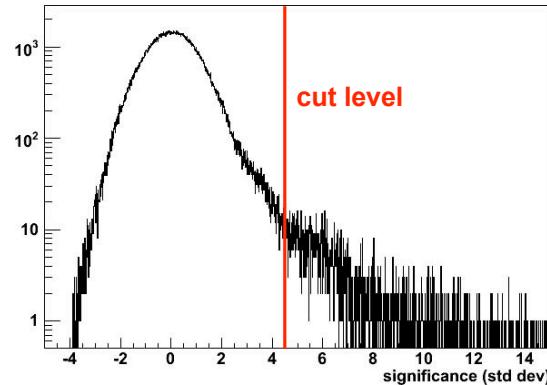
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>4.5 σ Regions in the Galaxy

6 years of data

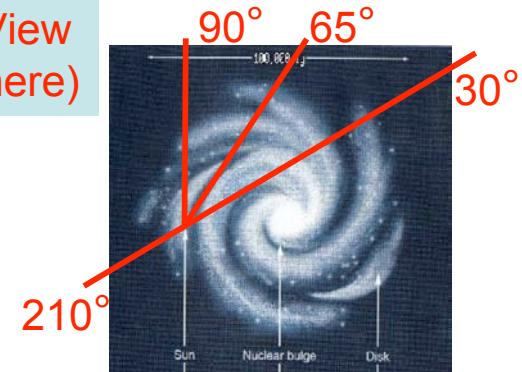


Distribution of Excesses in the Galactic Plane

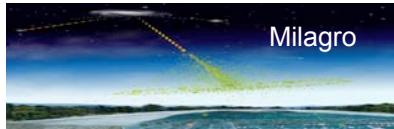


Milagro's Field of View
(Northern Hemisphere)

Milagro sees the Galactic plane
from longitude $\sim 30^\circ$ to $\sim 220^\circ$

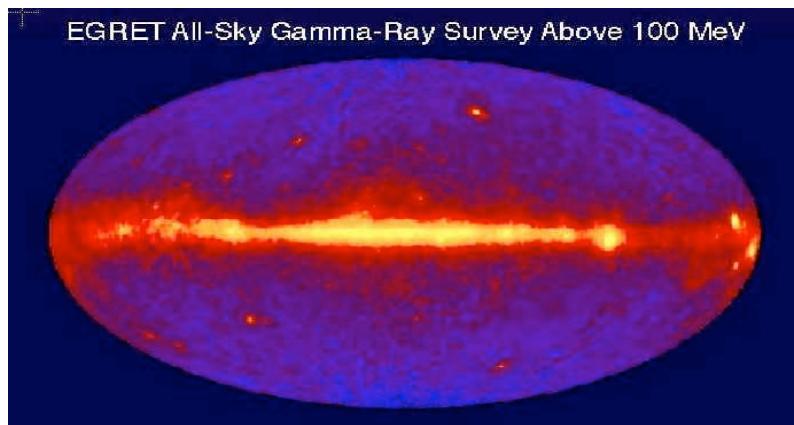


- 7 regions >4.5 σ
 - Expect 0.2 spots >4.5 σ in $l \in [30^\circ, 216^\circ]$, $b \in [-5^\circ, 5^\circ]$

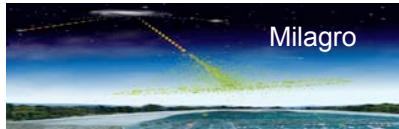


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Diffuse Emission from the Galactic Plane

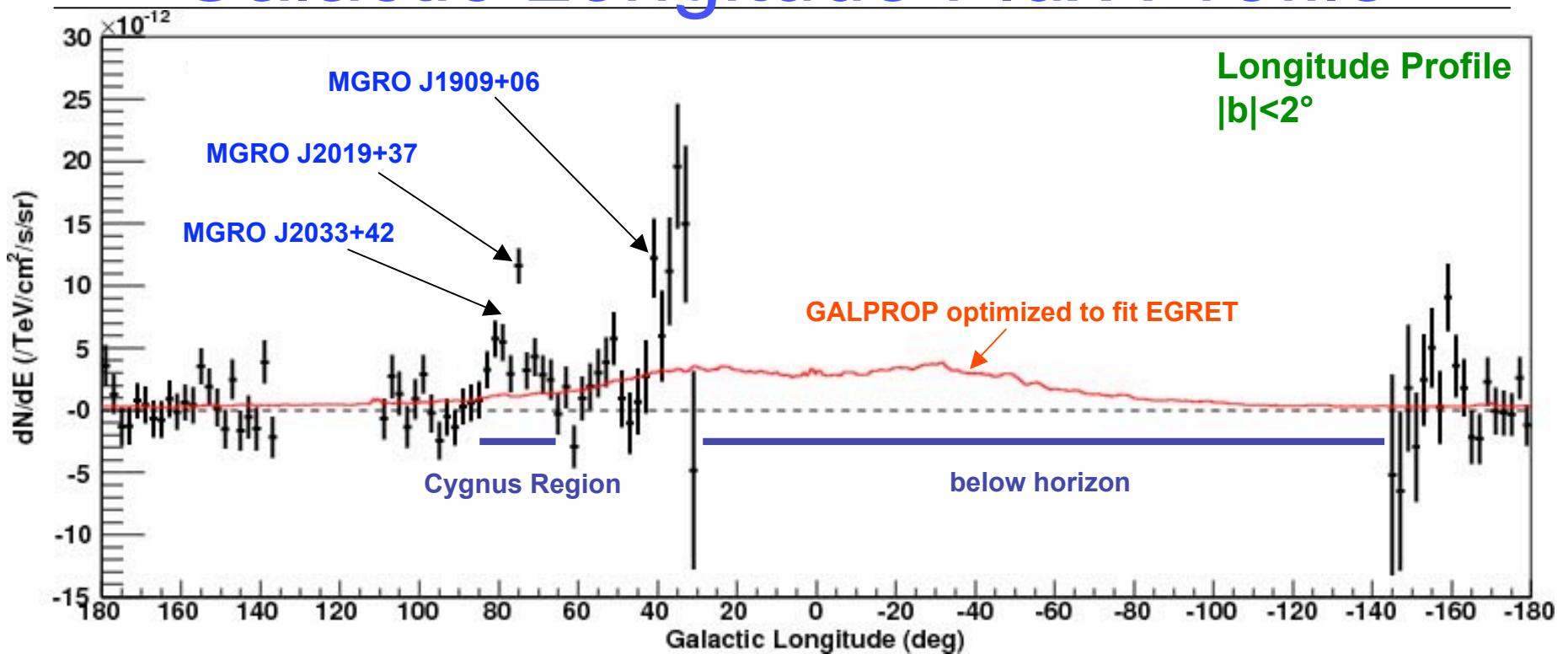


- EGRET observations to 20 GeV indicate a GeV excess
 - Is there an excess at TeV energies?
- Gamma rays can measure cosmic ray flux elsewhere by observing cosmic ray interactions that produce γ rays
 - Proton + matter interactions
 - Inverse Compton scattering
- See Aous Abdo's talk about the Cygnus Region



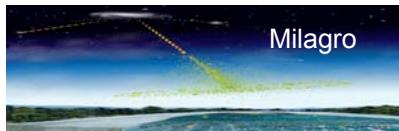
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Galactic Longitude Flux Profile

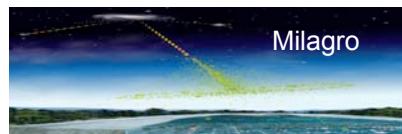
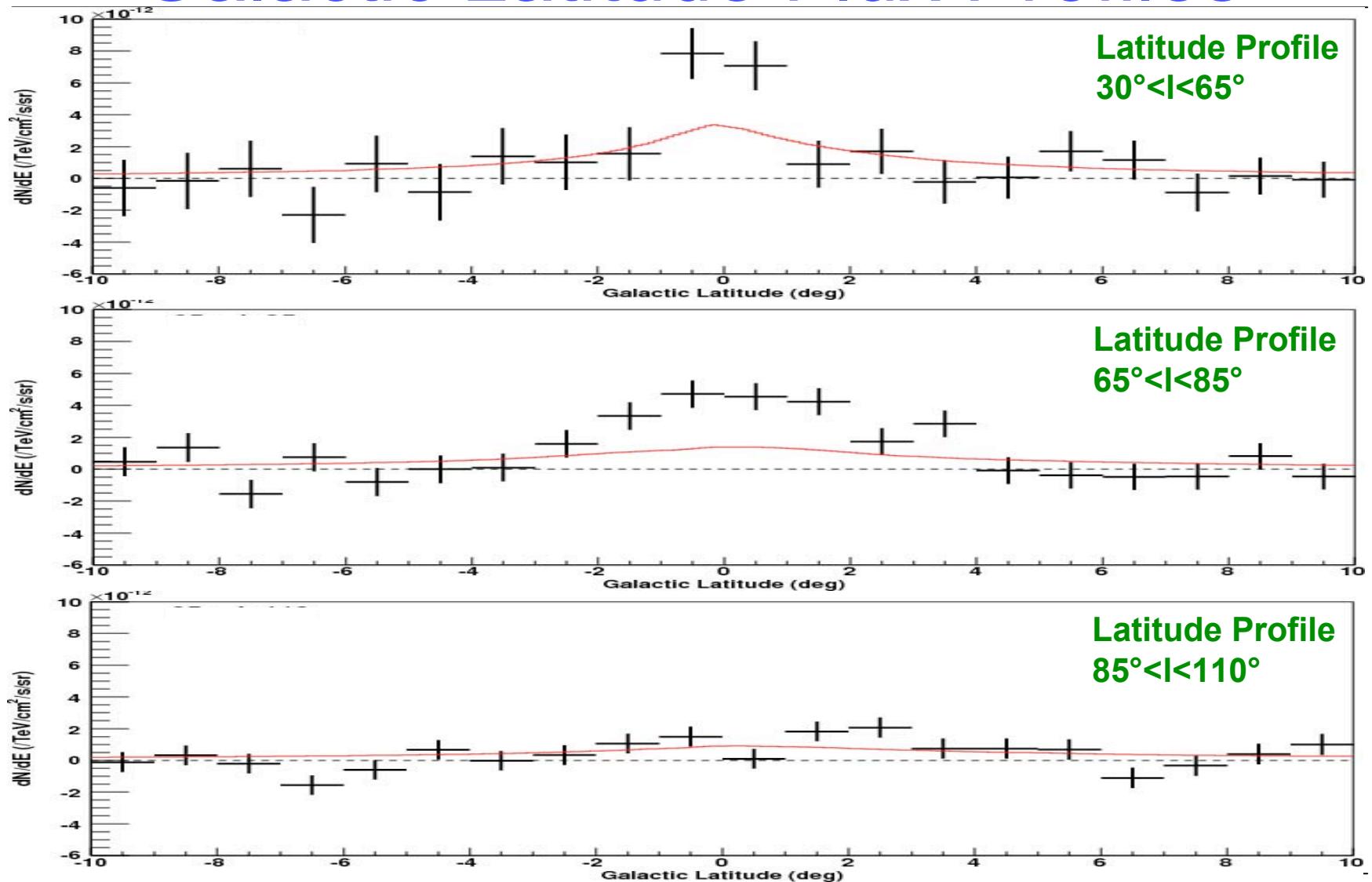


- Flux calculations assume a Crab spectrum (-2.62)
- 3 sources detected, MGRO J1909+06, ¹MGRO J2019+37, and MGRO J2033+42

¹ Abdo *et al.*, arXiv:astro-ph/0611691, submitted to ApJ Letters



Galactic Latitude Flux Profiles



Milagro

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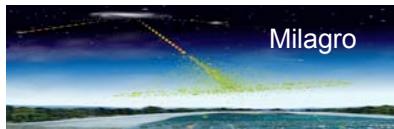
Galactic Plane Fluxes

| object | position box (deg) $ b <2$ | ¹ significance (std deviations) | ² flux ($\times 10^{-12}$) (/TeV/cm ² /s/sr) | GALPROP ($\times 10^{-12}$) (/TeV/cm ² /s/sr) |
|----------------|-------------------------------|-----------------------------------------------|-------------------------------------------------------------------------|---------------------------------------------------------------|
| Galactic Plane | $30 < l < 110$ | 7.5 | $3.0 \pm 0.4_{\text{stat}}$ | 1.6 |
| | $30 < l < 65$ | 4.8 | $4.3 \pm 0.9_{\text{stat}}$ | 2.5 |
| | $65 < l < 85$ | 8.4 | $4.2 \pm 0.5_{\text{stat}}$ | 1.2 |
| | $85 < l < 110$ | -- | <1.1 (95% CL) | 0.7 |
| | $136 < l < 216$ | -- | <1.1 (95% CL) | 0.3 |

¹ pre-trials significance

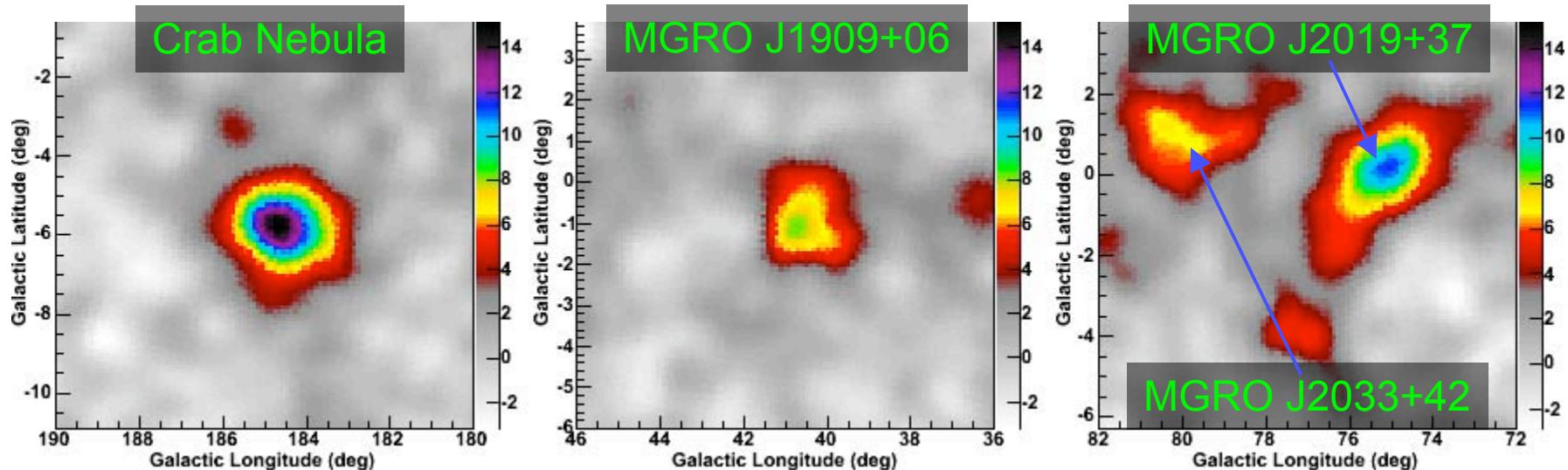
² add 30% systematic error to fluxes

- 30% systematic error is applied
 - Crab flux is consistent with HEGRA measured flux
 - But the trigger rate is higher than expected from the cosmic-ray flux
- A significant excess over GALPROP is seen for $30^\circ < |l| < 85^\circ$ regions



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Galactic TeV Sources



- Milagro detects 3 TeV sources in the Galactic plane
 - MGRO J1909+06 at 8.2σ ($>6.5\sigma$ post-trials)
 - MGRO J2019+37 at 11.3σ ($>10.2\sigma$ post-trials)
 - MGRO J2033+42 at 7.1σ ($>5.2\sigma$ post-trials)



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“Hot Spot” Fluxes

| object | ¹ position (l, b) | ² significance (std deviations) | ³ flux ($\times 10^{-14}$) (/TeV/cm ² /s) |
|---------------|---------------------------------------------------------------|-----------------------------------------------|----------------------------------------------------------------------|
| Crab | $-175.4 \pm 0.1_{\text{stat}}$, $-5.7 \pm 0.1_{\text{stat}}$ | 15.2 | $4.8 \pm 0.5_{\text{stat}}$ |
| MGRO J2019+37 | $75.1 \pm 0.1_{\text{stat}}$, $0.3 \pm 0.1_{\text{stat}}$ | 11.3 | $2.4 \pm 0.4_{\text{stat}}$ |
| MGRO J1909+06 | $40.5 \pm 0.1_{\text{stat}}$, $-1.0 \pm 0.1_{\text{stat}}$ | 8.2 | $4.1 \pm 0.9_{\text{stat}}$ |
| MGRO J2033+42 | $80.4 \pm 0.4_{\text{stat}}$, $1.0 \pm 0.3_{\text{stat}}$ | 7.1 | $1.7 \pm 0.4_{\text{stat}}$ |
| | $76.3 \pm 0.1_{\text{stat}}$, $-1.9 \pm 0.2_{\text{stat}}$ | 5.8 | $0.9 \pm 0.2_{\text{stat}}$ |
| | $77.2 \pm 0.2_{\text{stat}}$, $-4.0 \pm 0.2_{\text{stat}}$ | 5.6 | $1.2 \pm 0.2_{\text{stat}}$ |
| | $34.1 \pm 0.3_{\text{stat}}$, $0.0 \pm 0.2_{\text{stat}}$ | 5.1 | $5.5 \pm 1.4_{\text{stat}}$ |
| | $106.4 \pm 0.5_{\text{stat}}$, $1.7 \pm 0.8_{\text{stat}}$ | 4.5 | $1.0 \pm 0.4_{\text{stat}}$ |

¹ add 0.3° systematic error to positions

² pre-trials significance

³ add 30% systematic error to fluxes

- Sources in the Cygnus Region
 - MGRO J2019+37 is coincident with EGRET 3EG J2016+3657 and 3EG J2021+3716
 - MGRO J2033+42 is coincident with HEGRA TeV 2032+4130 ($dN/dE = (0.5 \pm 0.2) \times 10^{-14} / \text{TeV}/\text{cm}^2/\text{s}$) and EGRET 3EG J2033+4118



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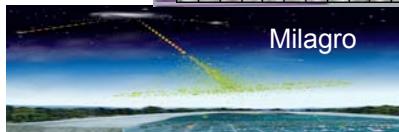
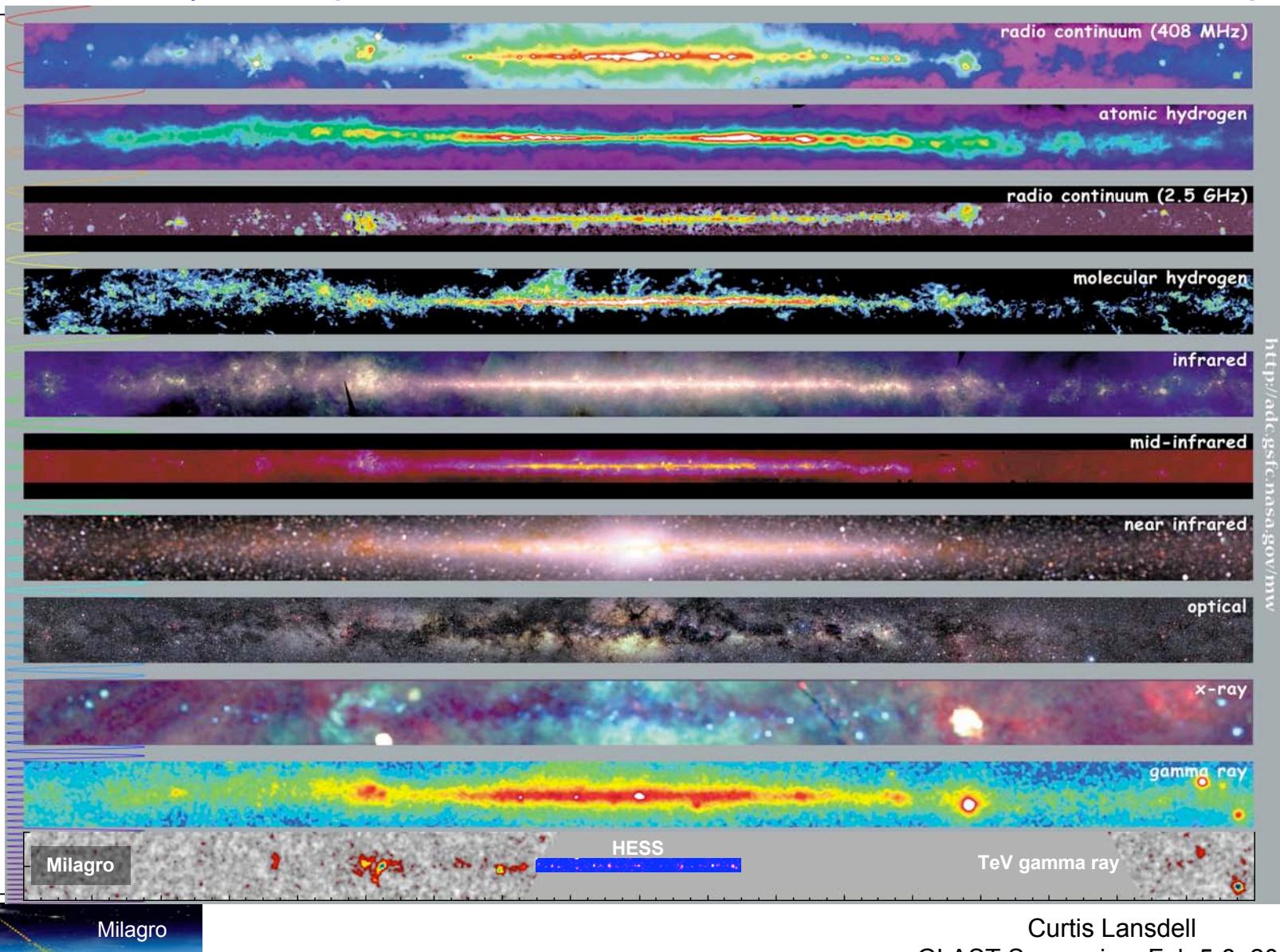
Summary

- TeV sources discovered in the Galactic plane
 - MGRO J1909+06 with $>6.5\sigma$ post-trials
 - MGRO J2019+37 with $>10.2\sigma$ post-trials
 - MGRO J2033+42 with $>5.2\sigma$ post-trials
- 7 interesting spots in the Galactic plane at TeV energies
 - Expect 0.2 spots above 4.5σ by chance
 - Observe 7, all within longitude = $[30^\circ, 110^\circ]$
 - 4 spots in the Cygnus Region (includes MGRO J2019+37 and MGRO J2033+42) – see Aous Abdo's plenary session talk tomorrow
 - $(44 \pm 9_{\text{stat}})\%$ of the total flux in the Galactic plane as seen with Milagro is contained in these spots
 - GALPROP fit to EGRET gives $\sim 50\text{-}70\%$ of the total Milagro flux in the plane in the same field of view



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TeV γ Rays: New Window for the Sky



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