



Challenging the high energy emission zone in **FSRQ**

PKS1222+216 (4C 21.35) VHE detection by **MAGIC**

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and

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on behalf of the MAGIC Collaboration

D. Wood, Y. Tanaka

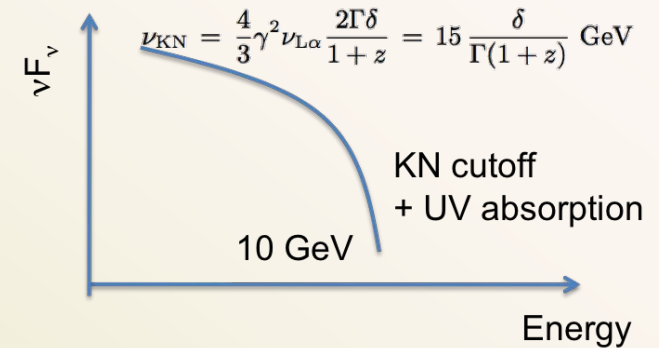
on behalf of the Fermi/LAT Collaboration



*Fermi Symposium
11 May 2011 - Roma*

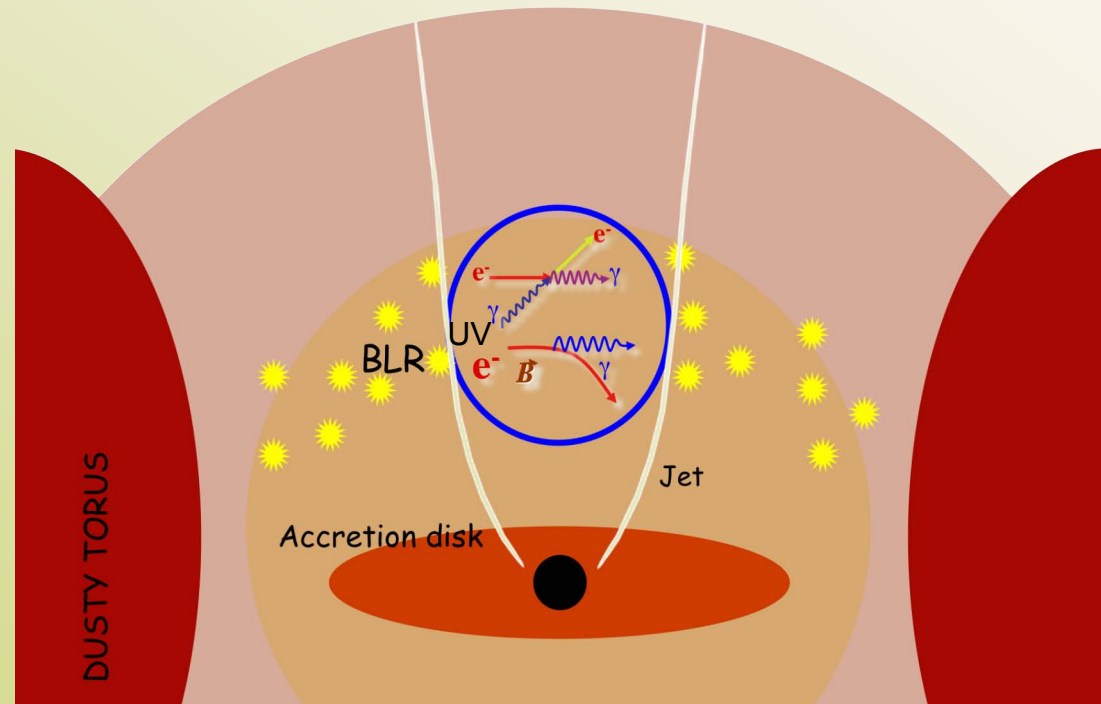
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The “Blazar Zone” in FSRQ



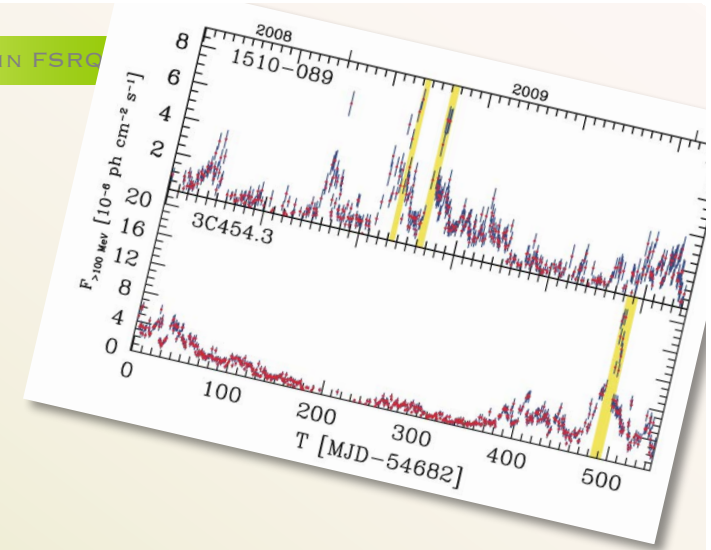
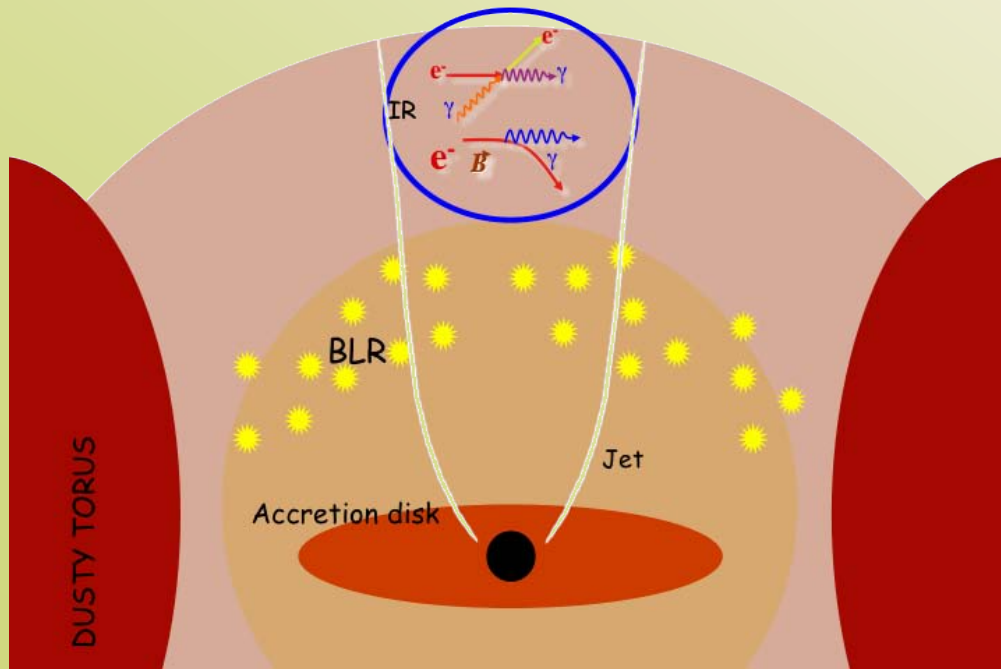
The “canonical” scenario

- ✧ Emission zone <0.1pc in conical jet within BLR
- ü Short time variability
- ☹ Internal absorption
e.g. *Liu&Bai 2006, Reimer 2007, Tavecchio&Mazin 2009*
- ☹ Reduced scattering efficiency Klein-Nishina (KN) e.g. *Albert et al. 2008, Tavecchio&Ghisellini 2008*



The “Blazar Zone” in FSRQ

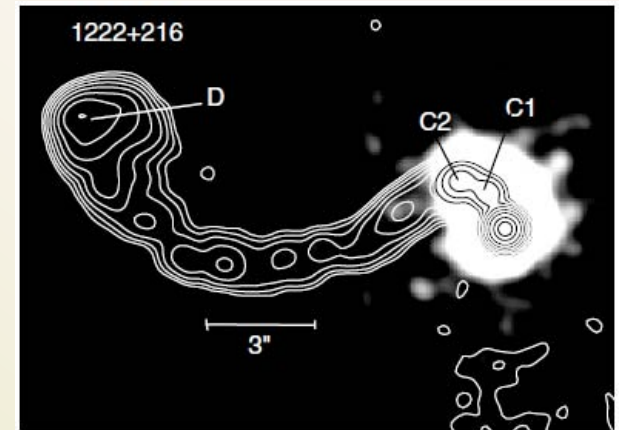
The far-dissipation scenario



- 2 Emission zone $\sim 1-10$ pc
dusty/IR torus
- ✧ Internal absorption in \sim TeV
Donea&Protheroe 2003, Ghisellini&Tavecchio 2009
- ü γ -ray \sim radio flares \rightarrow optical thin $\rightarrow > 1$ pc
- ☹ Fast variability $\sim < 1$ day e.g.
3C 454.3, PKS 1510-089 *Tavecchio 2010*

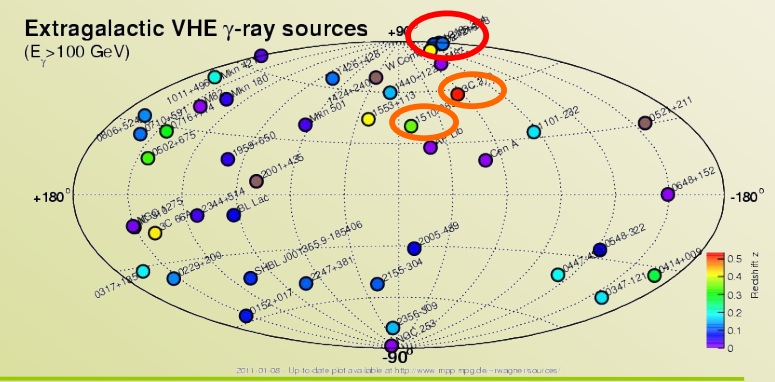
PKS 1222+21

$Z=0.432$ Lum. dist. 2.3 Gpc



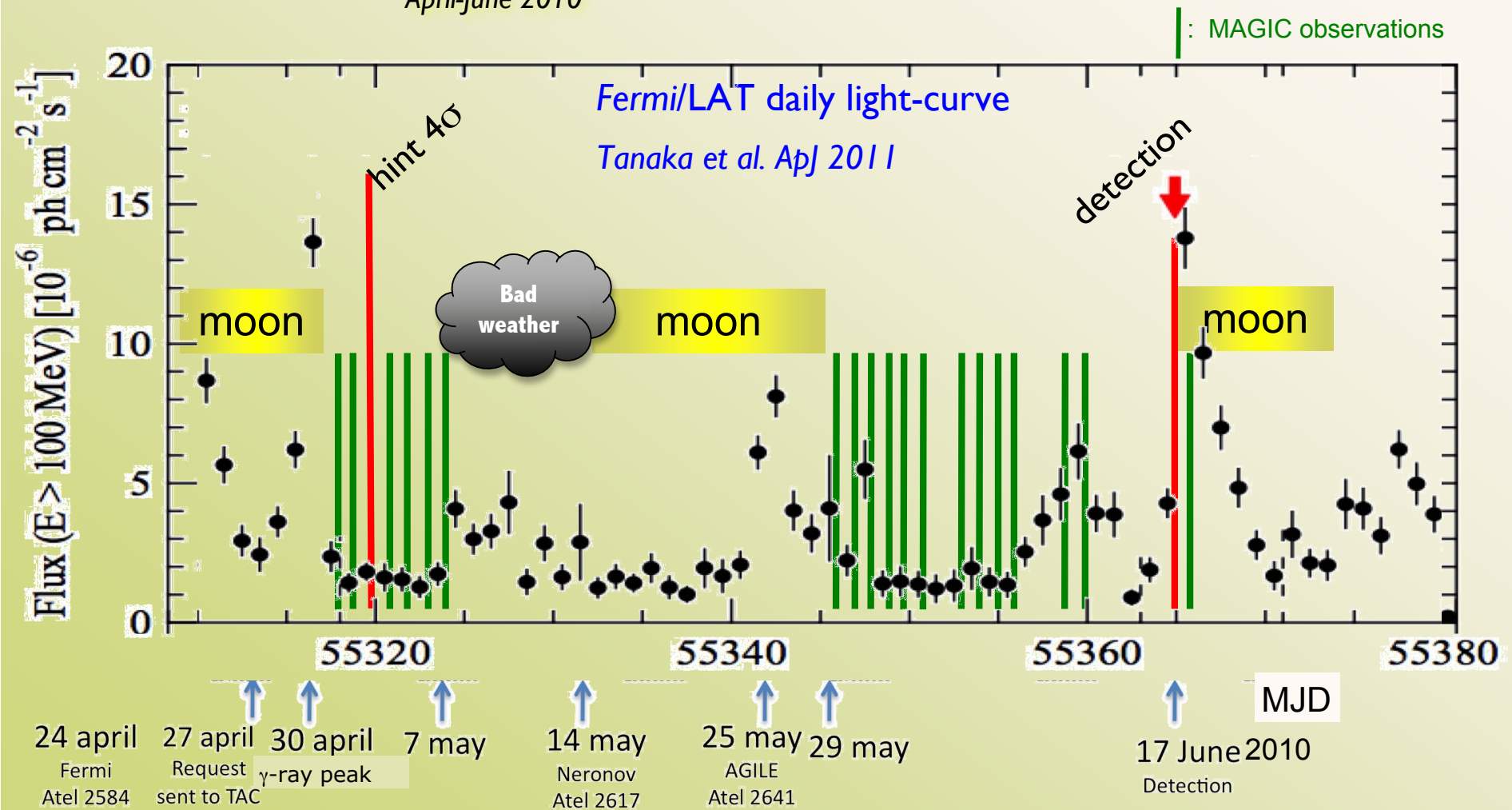
- ✧ Distorted quasar (bended jet) *Saikia 1993*
- ✧ Jet aligned to our line of sight $\sim 5-10^\circ$
- ✧ Superluminal motion 17-21c *Homan 2001, Lister 2009*
- ✧ X-ray emission *Jorstad&Marscher 2006*
- ✧ Gamma-ray blazar
 - ✧ 3EG J1224+2118, IFGL J1224.7+2121,
 - ✧ 3 FSRQ in VHE; **MAGIC** VHE detection: Atel #2684

$M_{\text{BH}} \sim 1.5 \times 10^8 M_\odot$
 $L_{\text{BLR}} \sim 10^{45} \text{ erg/s}$
 $L_{\text{disk}} \sim 10^{46} \text{ erg/s}$
 $R_{\text{BLR}} \sim 10^{17} \text{ cm}$



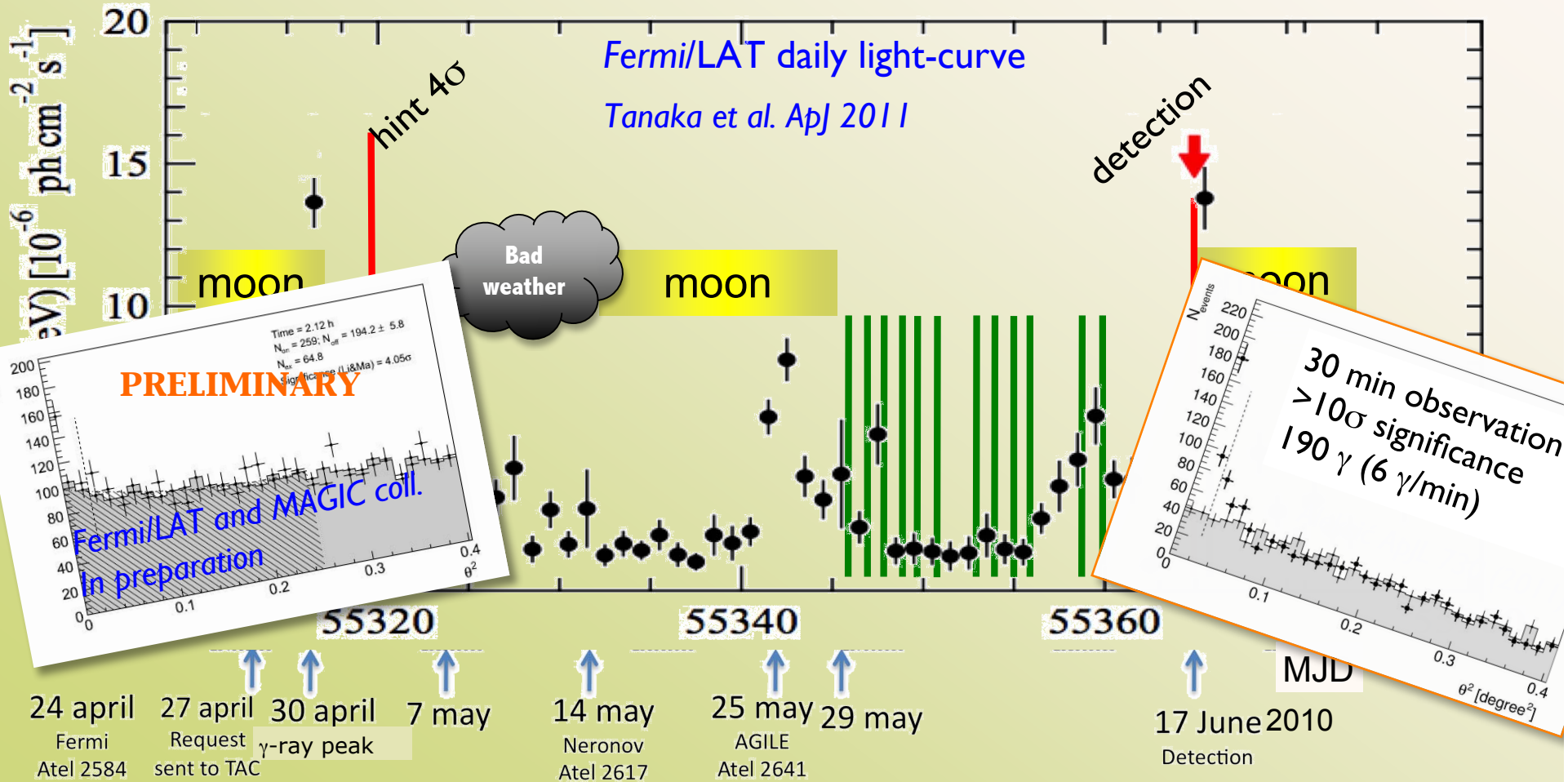
MAGIC campaign

April-June 2010

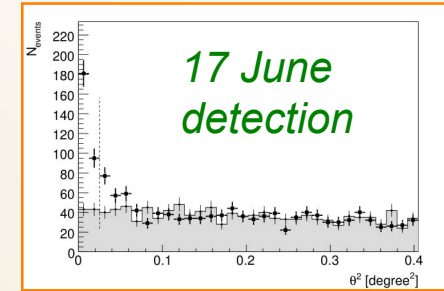


MAGIC campaign

April-June 2010



VHE spectrum



Aleksić et al. *ApJL* 730 (2011)

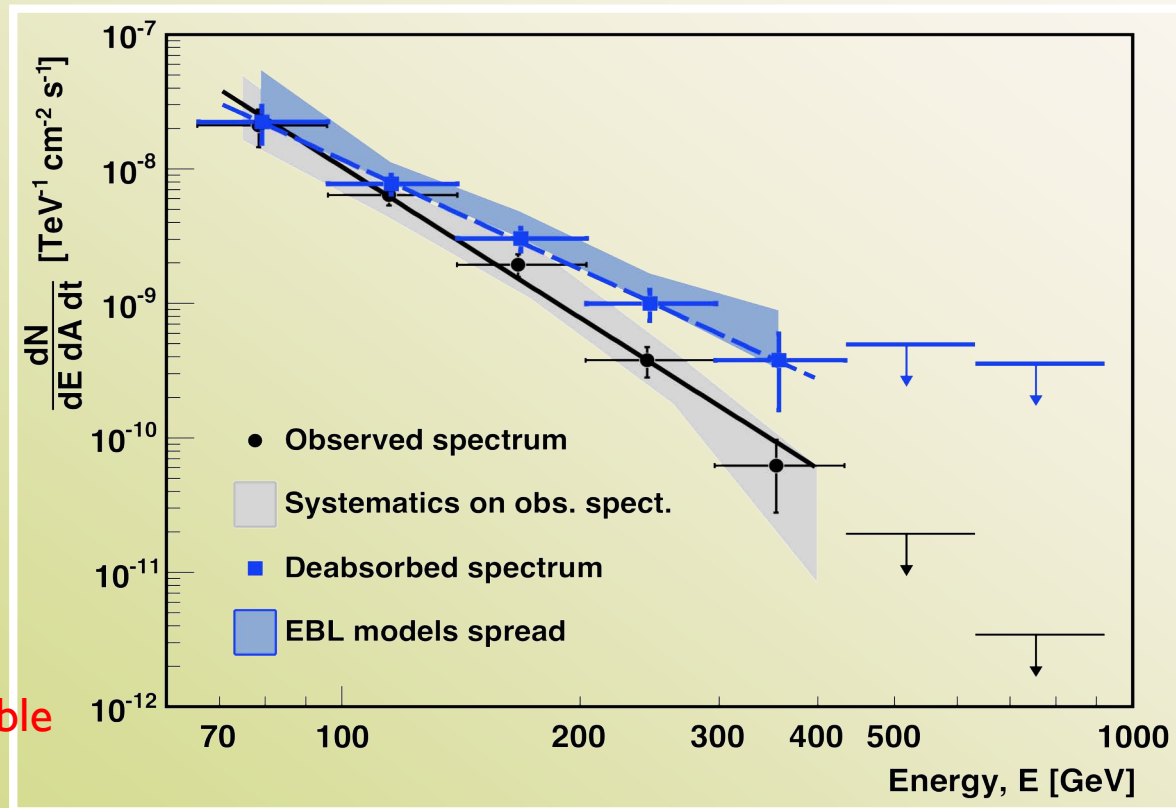
Measured spectrum:

$$\alpha = 3.75 \pm 0.29$$

De-absorbed Spectrum

(Dominguez+10):

$$\alpha = 2.72 \pm 0.34$$

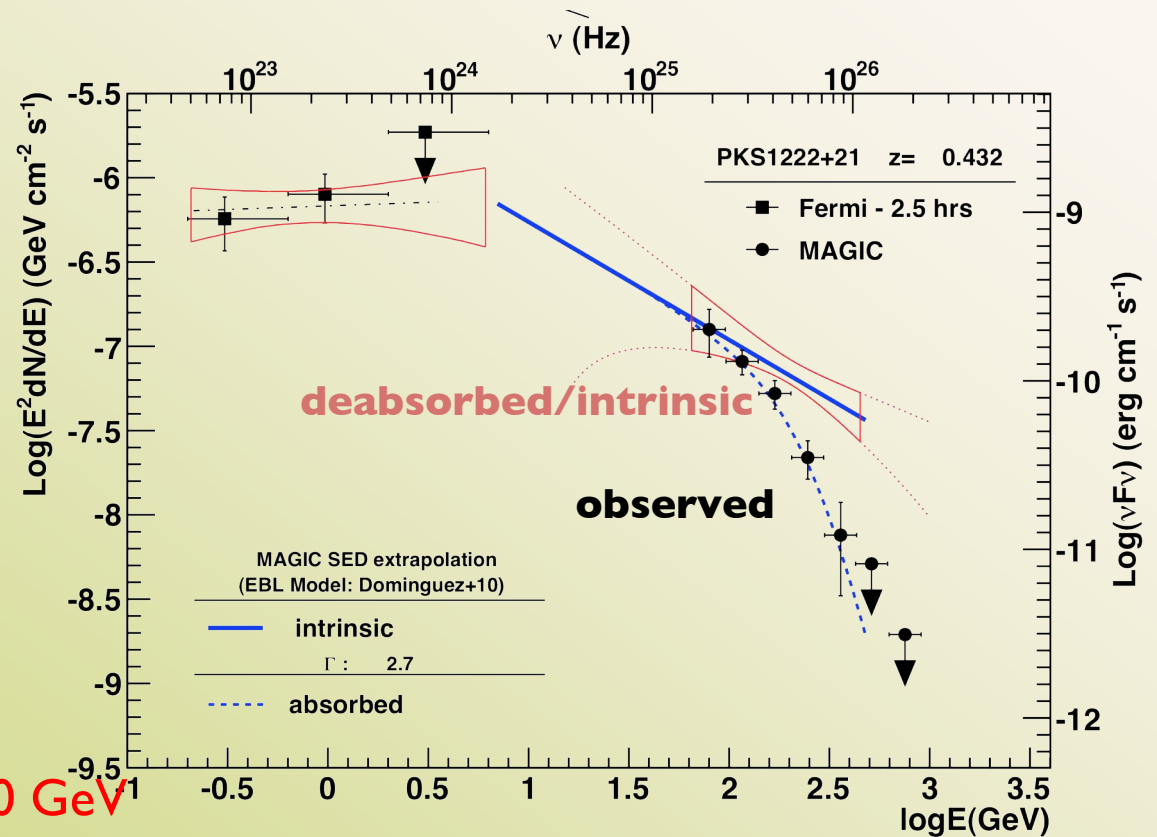


De-absorbed spectrum compatible with a simple power law

High energy SED

Aleksić et al. ApJL 730 (2011)

- Simultaneous Fermi/LAT 2.5 hrs encompassing MAGIC obs.



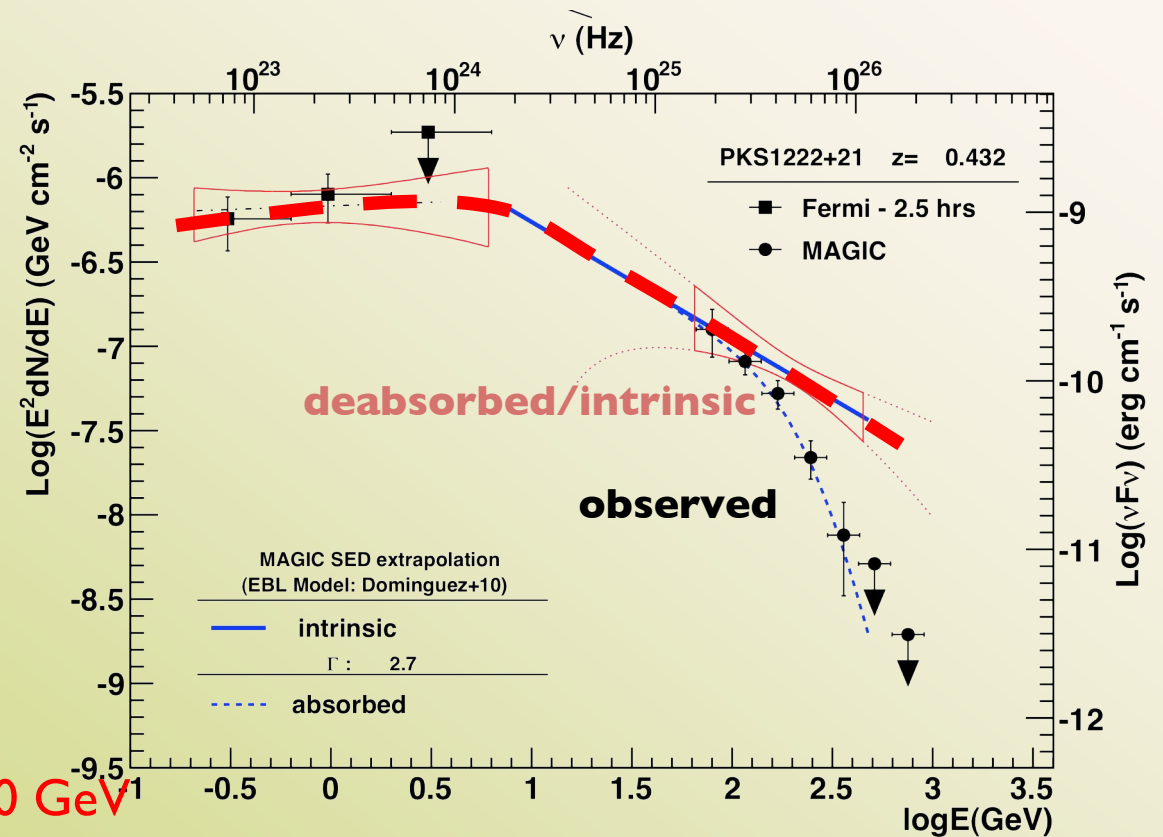
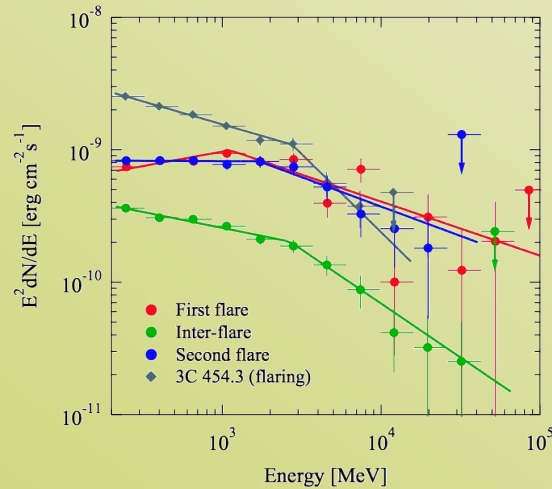
Single component from 2 to 400 GeV
Cutoff excluded at $E < \sim 130$ GeV (95% CL)

High energy SED

Aleksić et al. ApJL 730 (2011)

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Tanaka et al. ApJ 2011



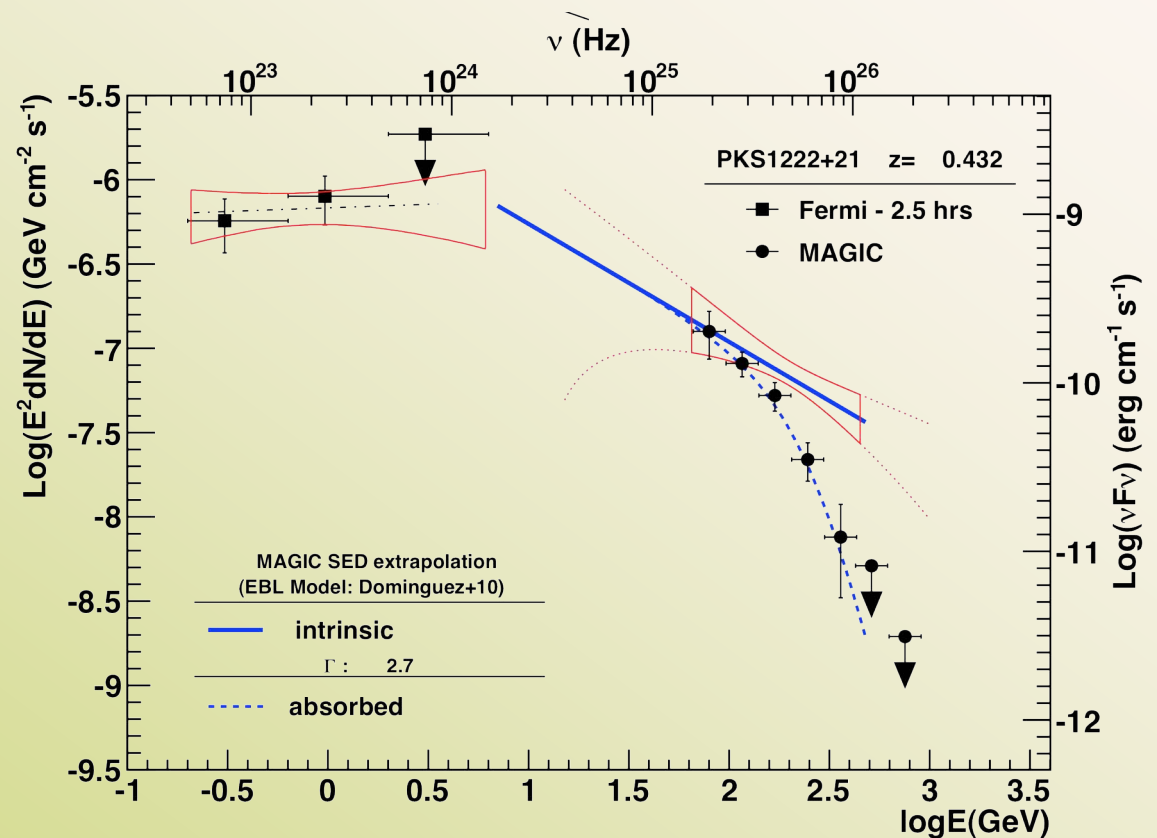
Single component from 2 to 400 GeV
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High energy SED

Aleksić et al. ApJL 730 (2011)

- ✧ If 1-10 GeV breaks due to photon-photon interaction in the BLR
- ⊗ γ -ray production should be in the BLR region!

Poutanen & Stern 2010

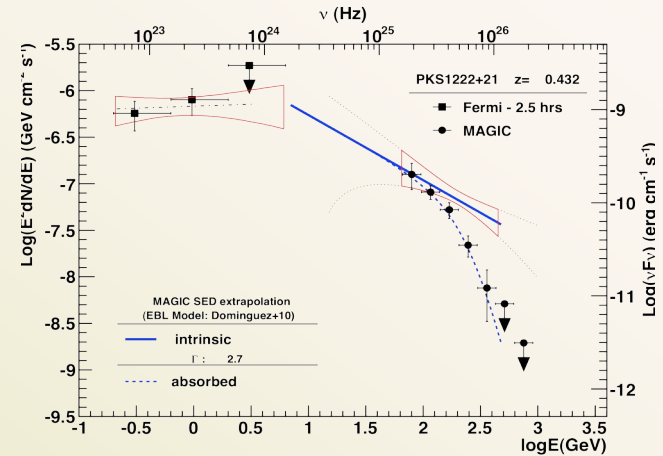


High energy SED

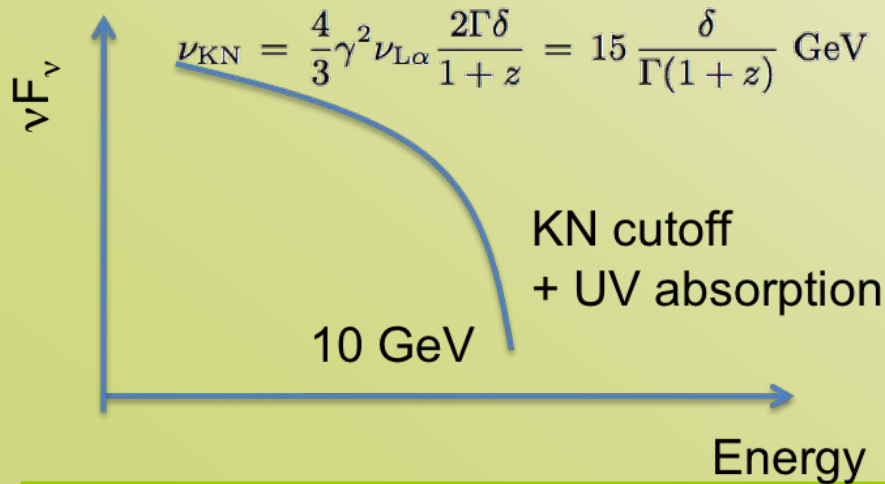
- Internal absorption + Klein-Nishina break

Ghisellini & Tavecchio 2009

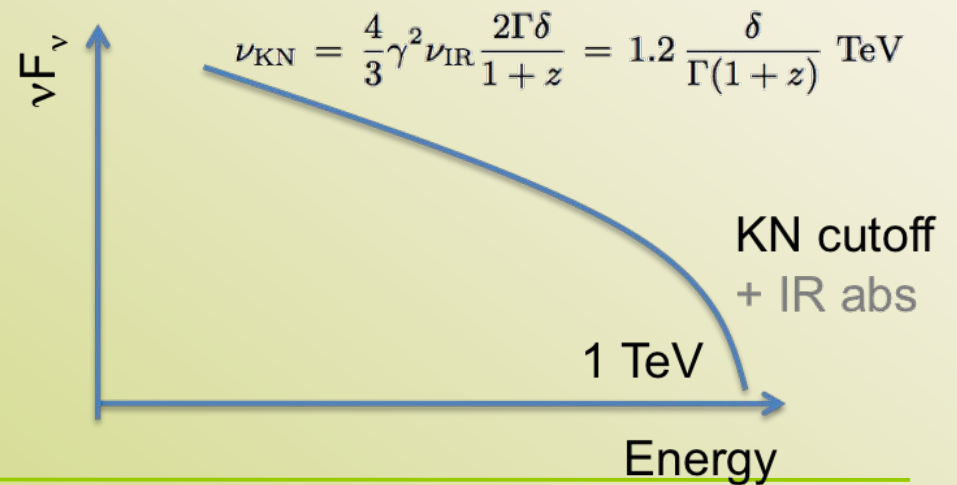
Liu & Bai 2006



Inside BLR



Outside BLR



High energy SED

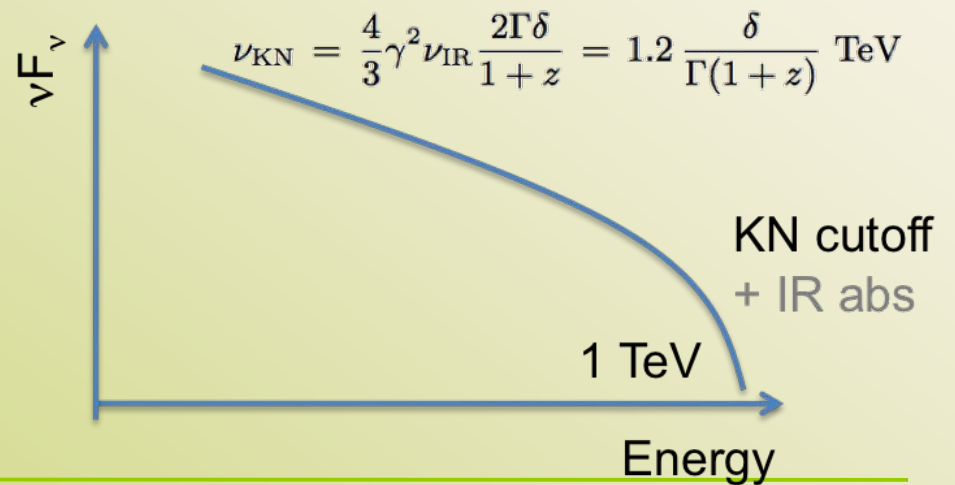
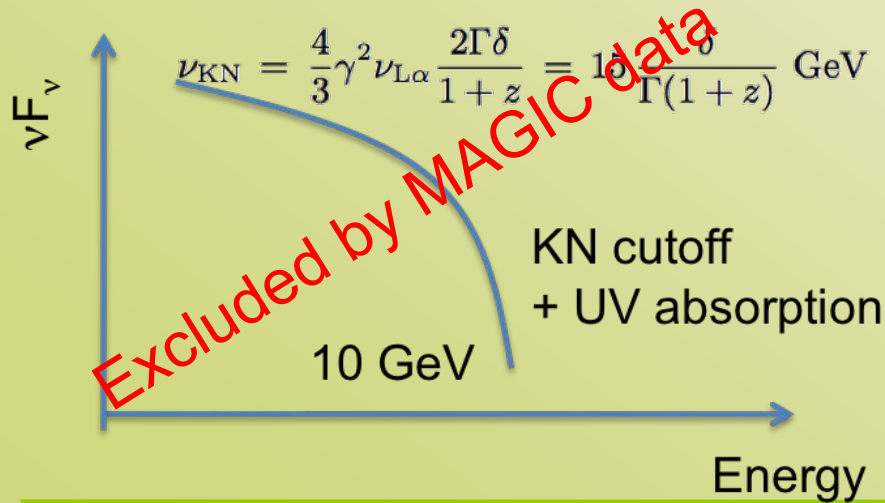
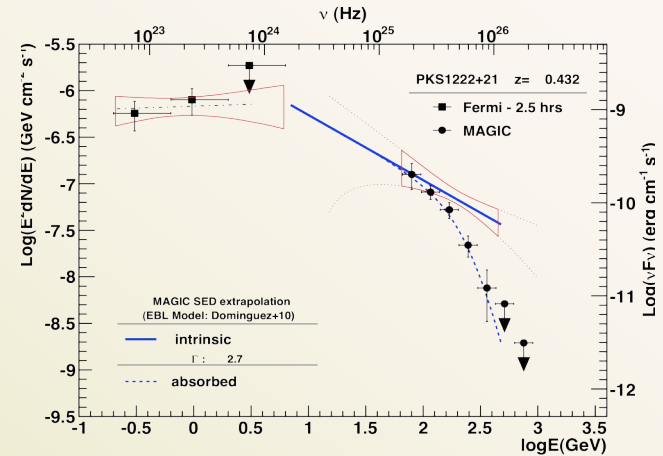
- Internal absorption + Klein-Nishina break

Ghisellini & Tavecchio 2009

Liu & Bai 2006

~~Inside BLR~~

Outside BLR

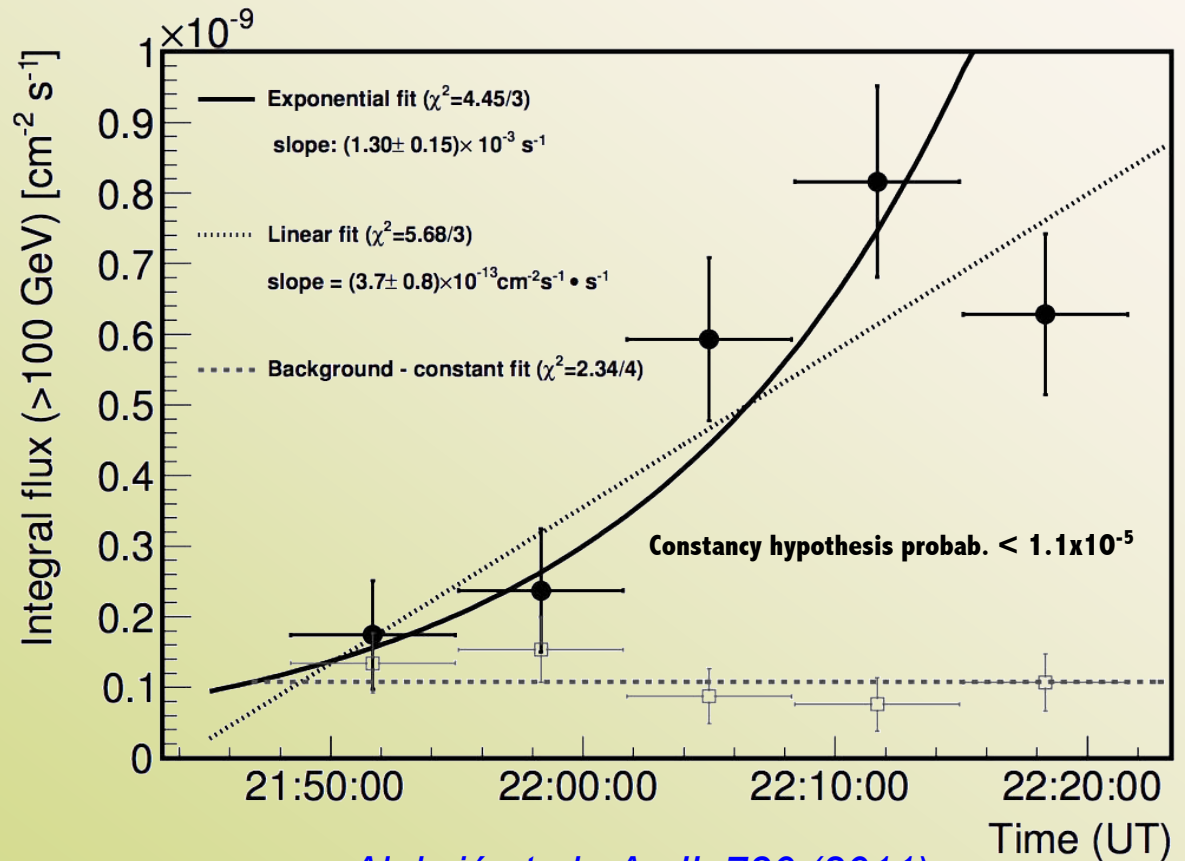


Light curve

- ✧ Very fast variability
- ✧ Doubling time:
~ 10 minutes
- ✧ Size emission region:
- ✧ $R \sim 10^{14}$ cm

$$R < ct_{\text{var}} \frac{\delta}{(1+z)} \approx$$

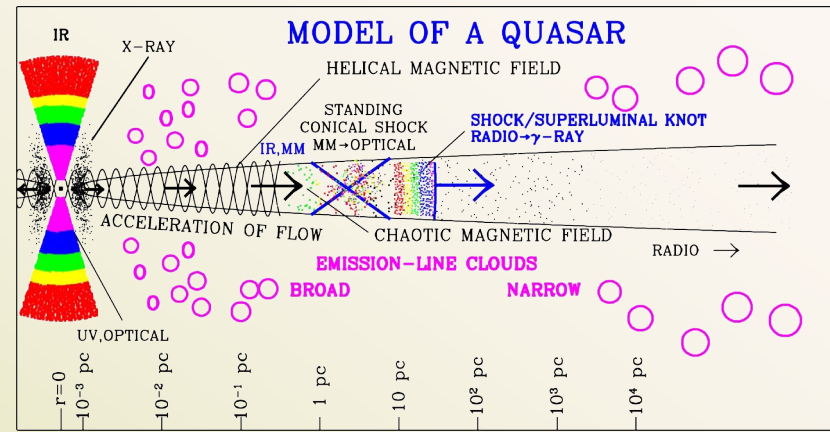
$$1.3 \times 10^{14} \left(\frac{\delta}{10} \right) \left(\frac{t_{\text{var}}}{10 \text{ min}} \right) \text{ cm}$$



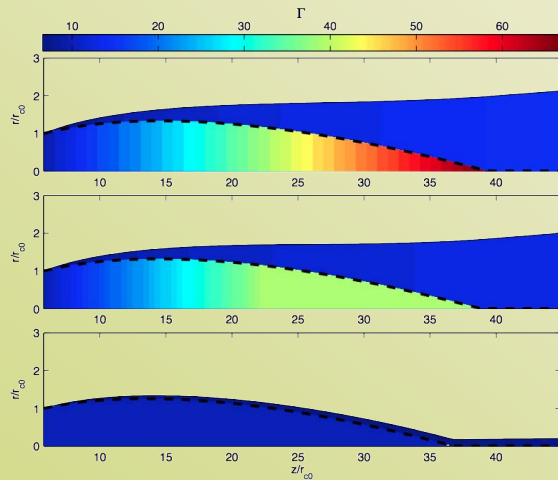
Aleksić et al. ApJL 730 (2011)



Beyond the BLR



Bromberg&Levinson 2009



✧ Recollimation

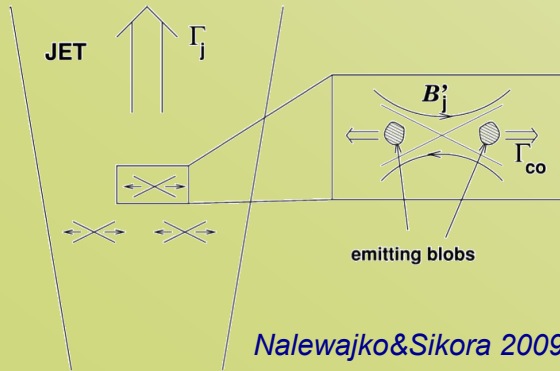
Marscher 1980, Bromberg&Levinson 2009

✧ Reconfinement shocks

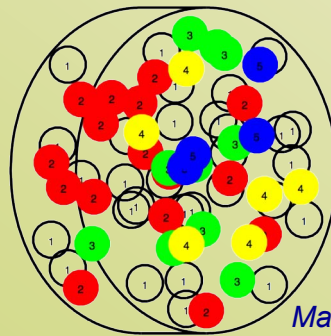
Nalewajko&Sikora 2009, Stawarz 2006

✧ Compact region embedded

Giannios 2009, Marscher&Jorstad2010, Ghisellini&Tavecchio 2008

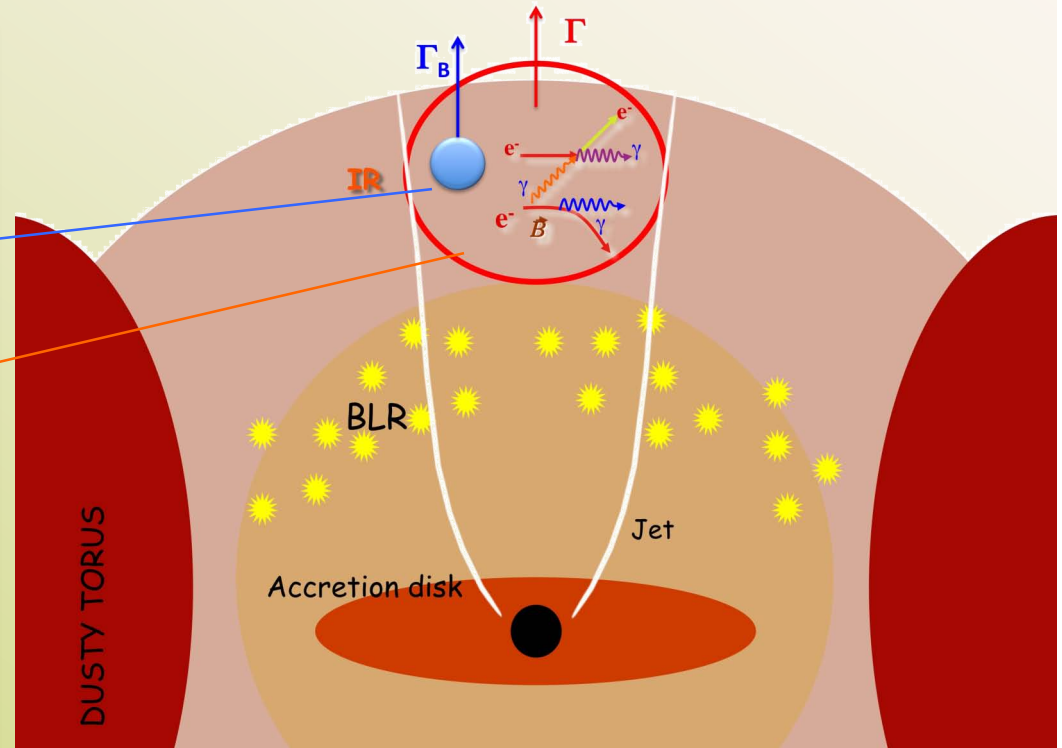
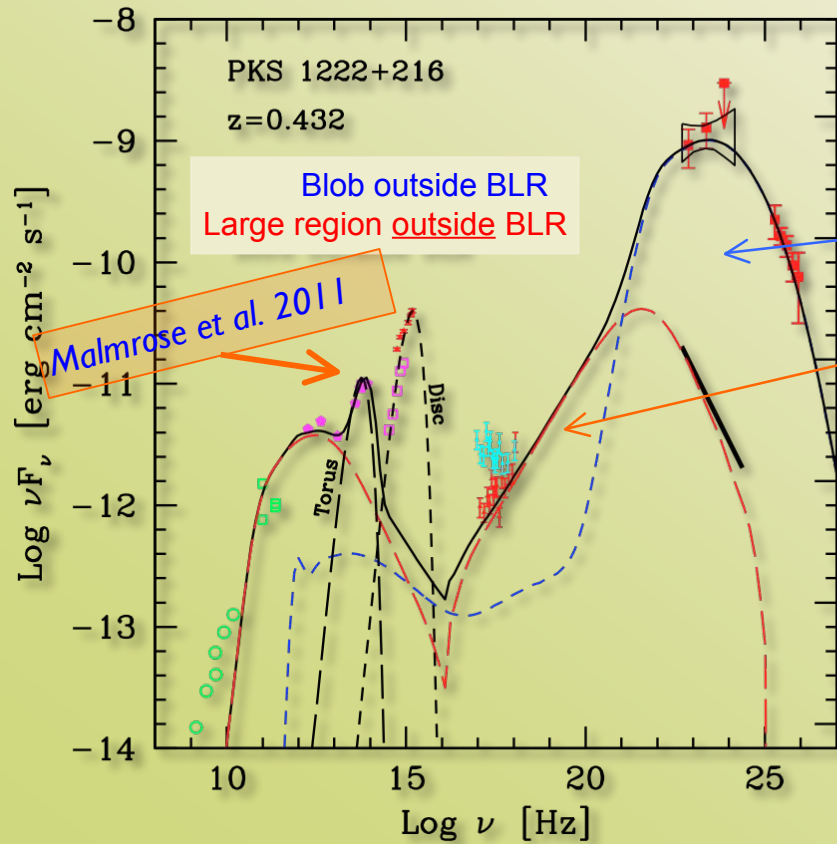


Nalewajko&Sikora 2009



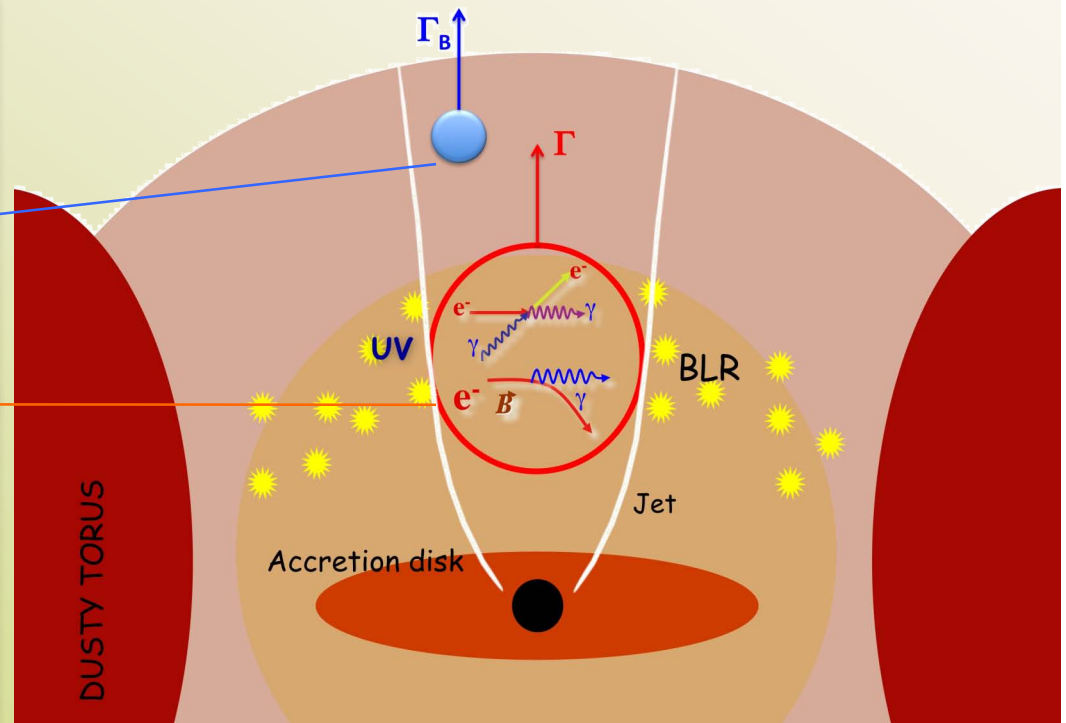
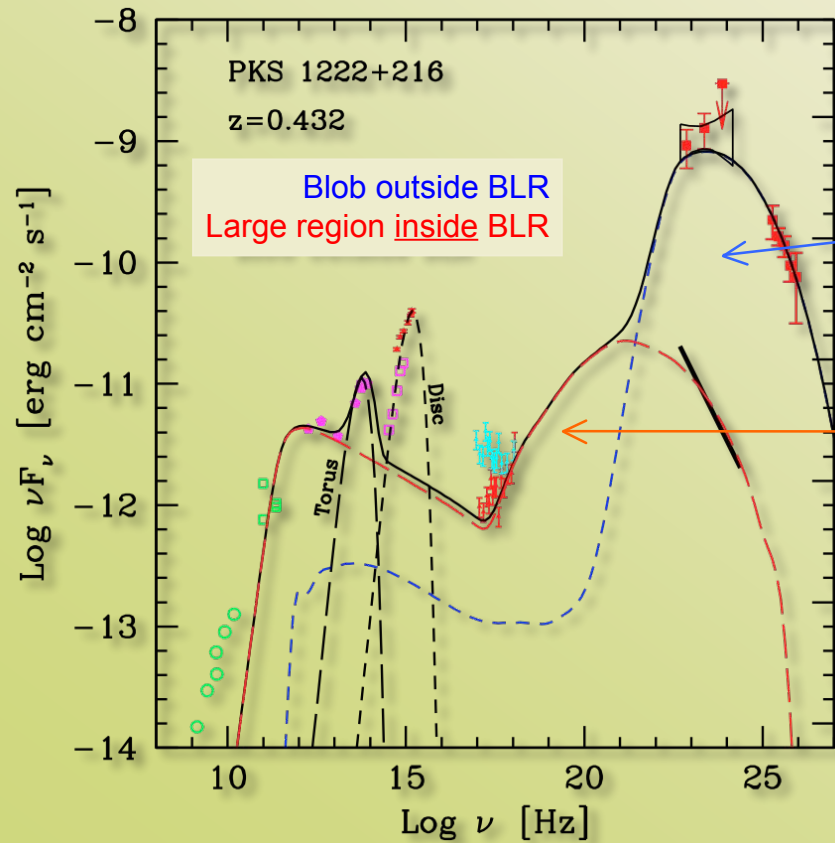
Marscher&Jorstad 2010

Two zones EC



Tavecchio et al. submitted

Two zones EC



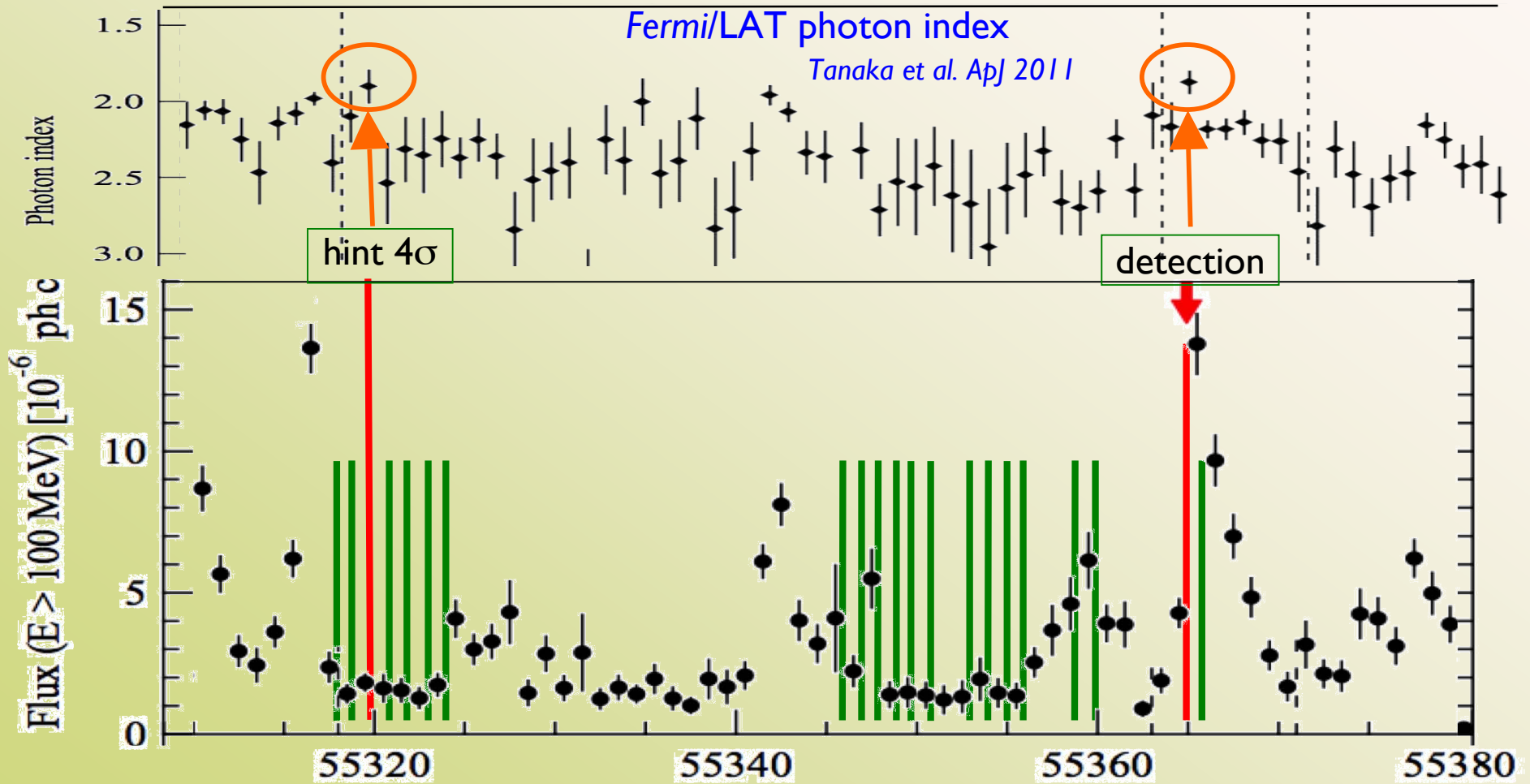
Tavecchio et al. submitted

Conclusions



More details on:
[Aleksić et al. ApJL 730 \(2011\)](#)
 MAGIC Discovery of Very-High Energy Emission
 from the FSRQ PKS 1222+21

- ✧ MAGIC detection of FSRQ PKS1222+21 during high gamma-ray activity observed by *Fermi*/LAT
- ü No VHE cutoff detectable
- ü Fast variability ~10 min
- Ø Severe constraints to FSRQ emission models
- Ø Challenge to “canonical” and “far-dissipation” scenarios
- ☺ Promise of substantial progress from simultaneous MAGIC-Fermi observations

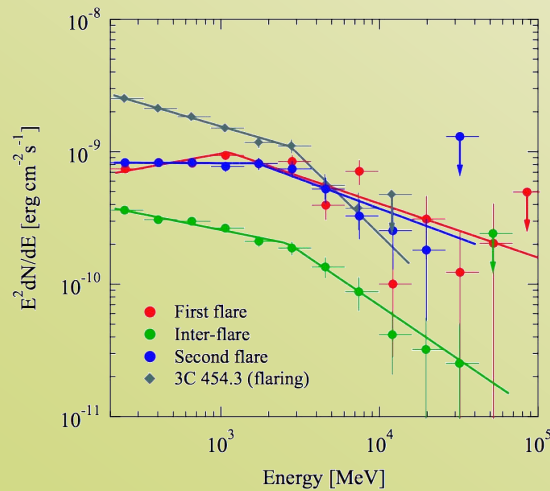


high VHE flux → flat GeV photon index

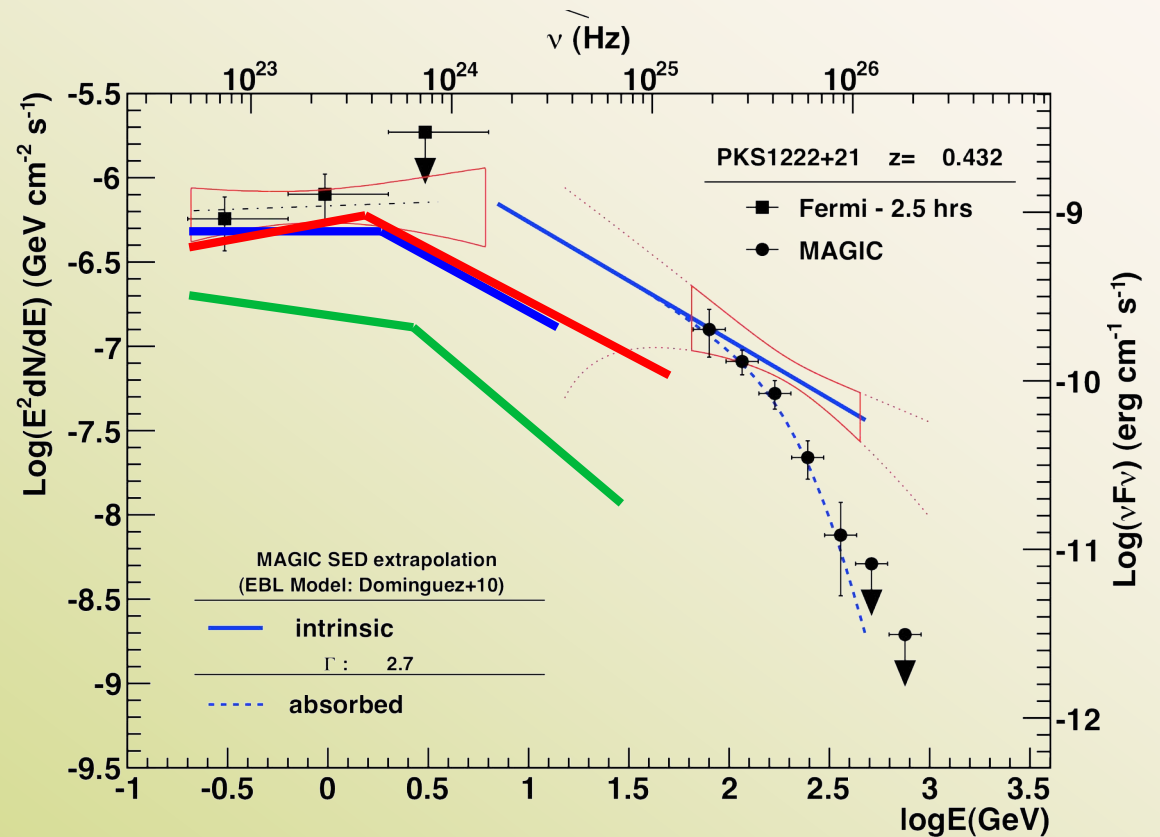
High energy SED

MAGIC Coll. *ApJL* 730 (2011)

Simultaneous *Fermi*/LAT 2.5 hrs encompassing MAGIC obs.



State	Γ_{LE}	Γ_{HE}	E_{br} [GeV]	ΔL
Intermediate+Active	2.18 ± 0.02	2.64 ± 0.06	$2.4^{+0.2}_{-0.2}$	-27.2
Intermediate	2.30 ± 0.05	2.69 ± 0.40	$2.5^{+1.1}_{-0.4}$	-6.6
First flare	1.80 ± 0.06	2.40 ± 0.07	$1.1^{+0.3}_{-0.2}$	-17.9
Inter-flare	2.24 ± 0.03	2.81 ± 0.14	$2.7^{+0.6}_{-0.6}$	-8.4
Second flare	2.00 ± 0.05	2.44 ± 0.10	$1.7^{+1.1}_{-0.4}$	-6.3



Tanaka et al. *ApJ* 2011