Variable galactic y-ray sources

Guillaume Dubus

with many thanks to Fermi/LAT collaboration

Fermi Symposium 2011, Rome Institut de Planétologie et d'Astrophysique de Grenoble



Variable galactic y-ray sources

[pulsars]

gamma-ray binaries: spectrum of LS sources, lightcurve of PSR B1259-63 **unidentified transients in the Galactic Plane microquasars:** Cygnus X-3 flares **Novae:** Symbiotic V407 Cyg **colliding wind binaries:** eta Carina ?

- interacting massive star + compact object
- dominant gamma-ray emission
- likely pulsar winds in binaries



LS I +61°303 spectral energy distribution (Zdziarski et al. 2009)





HESS J0632+057



detection at periastron

 \rightarrow A. Abdo

13h00m

HESS J1303-631

PSR B1259-63

Dec [deg]

-63

-64

-65

13h20m

13h10m

Februar 2004

GeV modulation, spectrum, variability

 \rightarrow D. Hadasch

X-ray, TeV modulation found no GeV reported

Bongiorno et al. 2011 Acciari et al. 2011



new y-ray binary

 \rightarrow R. Corbet





LS I+61 303

HESS J0632+057



detection at periastron

 \rightarrow A. Abdo

Dec [deg

GeV modulation, spectrum, variability

16

l (deg)

16.5

X-ray, TeV modulation found no GeV reported

Bongiorno et al. 2011 Acciari et al. 2011



new γ-ray binary \rightarrow R. Corbet

\rightarrow D. Hadasch



XSS J12270-4859?



Tam et al. 2010

low mass y-ray binaries ?

Hill et al. 2011



HESS J0632+057



detection at periastron

 \rightarrow A. Abdo

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key: orbital modulations

Orbital modulations

- Anisotropic inverse Compton on star photons
- Pair production on star photons
- Doppler boost [if bulk relativistic motion]



- \rightarrow identify emission mechanism
- \rightarrow distinguish variability in phys. conditions from var. due to observer geometry

modulations suggest inverse Compton



modulations suggest inverse Compton



- modulations suggest inverse Compton
- spectrum suggests two populations of HE particles



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pulsar magnetospheric pulsar wind
emissionemissionnebula(but modulation, variability ??)

 \rightarrow D. Hadasch

- modulations suggest inverse Compton
- spectrum suggests two populations of HE particles



→A. Abdo



Fermi/LAT lightcurve (Abdo et al. 2011)

- 10³⁶ erg/s spindown pulsar in 3.5 year orbit around Be star
- Fermi/LAT detection at periastron mid-December 2010

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Fermi/LAT lightcurve (Abdo et al. 2011)

- 10³⁶ erg/s spindown pulsar in 3.5 year orbit around Be star
- Fermi/LAT detection at periastron mid-December 2010
- brighter in February (near inf. conjunction)

→A. Abdo

Fermi/LAT lightcurve (Abdo et al. 2011)





inv. Compton on Be star photons works only near periastron

→A. Abdo

Fermi/LAT lightcurve (Abdo et al. 2011)





• inv. Compton on Be star photons works only near periastron

– Be disk (π_0 , free-free...) ? HE sync ? Doppler boost ? IC on other fields ?...

• Nearly all spindown power radiated away in γ -rays at peak

Transients in the Galactic Plane

Fermi/LAT all-sky survey mode. Automated search runs every 6 hours, 1 day, 1 week. weekly report http://fermisky.blogspot.com list of binaries specifically monitored \rightarrow poster Glanzman & Dubois



Variable galactic y-ray sources ATels

| Atel source | date | I | b | err | note |
|----------------|------------------------------|--------|-------|------|------------------------------------|
| J0903-3531 | 10/2008 | 259.59 | 7.7 | 0.08 | AGN ? |
| J0910-5041 | 10/2008 | 271.62 | -1.8 | 0.07 | AGN ? |
| J0109+6134 | 02/2010 | 125.12 | -1.23 | 0.09 | AGN |
| J1018-5856 | 03/2011 | 284.32 | -1.70 | 0.02 | gamma-ray binary |
| Cyg X-3 | 05/2010, 03/2011 | 79.85 | 0.79 | - | microquasar |
| V407 Cyg | 03/2010 | 86.96 | -0.55 | 0.12 | nova |
| Crab | 09/2010, 04/2011 | 184.43 | -5.79 | 0.06 | Crab flare |
| PSR B1259-63 | 12/2010, 01/2011 | 304.18 | -0.99 | | gamma-ray binary |
| J1057-6027 | 06/2009 | 289.3 | -0.64 | 0.07 | |
| GC region | 02/2011 | 0.09 | -0.23 | 0.36 | 20s |
| J0109+6134 | 02/2010 | 125.18 | -1.21 | 0.6 | AGN |
| J1037-5708 | 10/2010 | 285.5 | 1.1 | 0.5 | |
| J2206+6203 | 01/2010 | 104.9 | 5.1 | 0.4 | |
| J1410-6147 | 02/2008 | 312.2 | -0.3 | 0.5 | |
| J2206+6203 | 01/2010 | 104.9 | 5.1 | 0.4 | |
| PSR B1259-63 ? | 08/2010 | 304.7 | -0.8 | 0.6 | no Fermi at that time |
| Cygnus region | 11/2007 | 75.0 | -0.4 | 1 | |
| J2022+4032 | 04-05-06-10/2008 | 78.2 | 2.1 | 0.5 | ? PSR J2021+4026. Chen et al. 2011 |
| Cyg X-1 ? | 10/2009, 03./2010 | 70.3 | 2.5 | 1.2 | no Fermi. Sabatini et al. 2010 |
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Cygnus X-3 flares



→ poster Piano for AGILE

Cygnus X-3 flares



HE y-rays from a microquasar

- Relations between X-rays, radio, γ-ray
- Link between non-thermal proc. and jet formation ?
- Only confirmed detection from a microquasar





Y-ray and X-ray modulation

X-ray modulation from Thomson scattering in Wolf-Rayet wind



• γ-ray modulation due to inv. Compton on Wolf-Rayet photons ?

Y-ray and X-ray modulation

γ-ray modulation due to inv. Compton on Wolf-Rayet photons



- not too close, not too far : recollimation shock ?
- matter, radiation density : is Cyg X-3 unique ?

Gamma-rays from a nova → poster T. Cheung



Fermi Detects Gamma Rays from Nova Cygni 2010





Feb. 19 to March 9, 2010



Nova Cygni 2010 in Visible Light

March 7, 20:36 UT

March 10, 19:08 UT



optical nova outburst discovered by amateur astronomers Nishiyama & Kabashima



Gamma-rays from a nova

 \rightarrow poster T. Cheung

- thermonuclear runaway WD ejects 10^{-6} M $_{\odot}$ at ~ 3000 km/s
- symbiotic system: lots of matter & radiation



- Mini-supernova, 10⁴⁴ erg, π_0 from high energy p⁺ or IC on e⁻
- One every few years ? (Yungelson et al. 1995)



Identified variable galactic γ -ray sources tend to be binaries

gamma-ray binaries

spectrum of LS sources, lightcurve of PSR B1259-63 challenge models variability gives new insights into pulsar winds

microquasars

Cyg X-3 could be unique object to link jet formation with non-thermal processes symbiotic novae

particle acceleration in a mini supernova

colliding wind binaries tbc

unidentified transients in the Galactic Plane more fun to come !

Colliding wind binaries

- Gamma-ray emission from Eta Carinae ?
- Large kinetic energy in winds ~ 10³⁷ erg/s, 5.5 yr orbit
- AGILE flare near periastron but no variability in Fermi/LAT



Abdo et al. 2010

Fermi/LAT spectrum : pulsar wind nebula in Tr 16 cluster ?

• No conclusive evidence for colliding wind binaries yet.