

# Fermi

Gamma-ray Space Telescope

**Users Group Meeting  
12 May 2010**

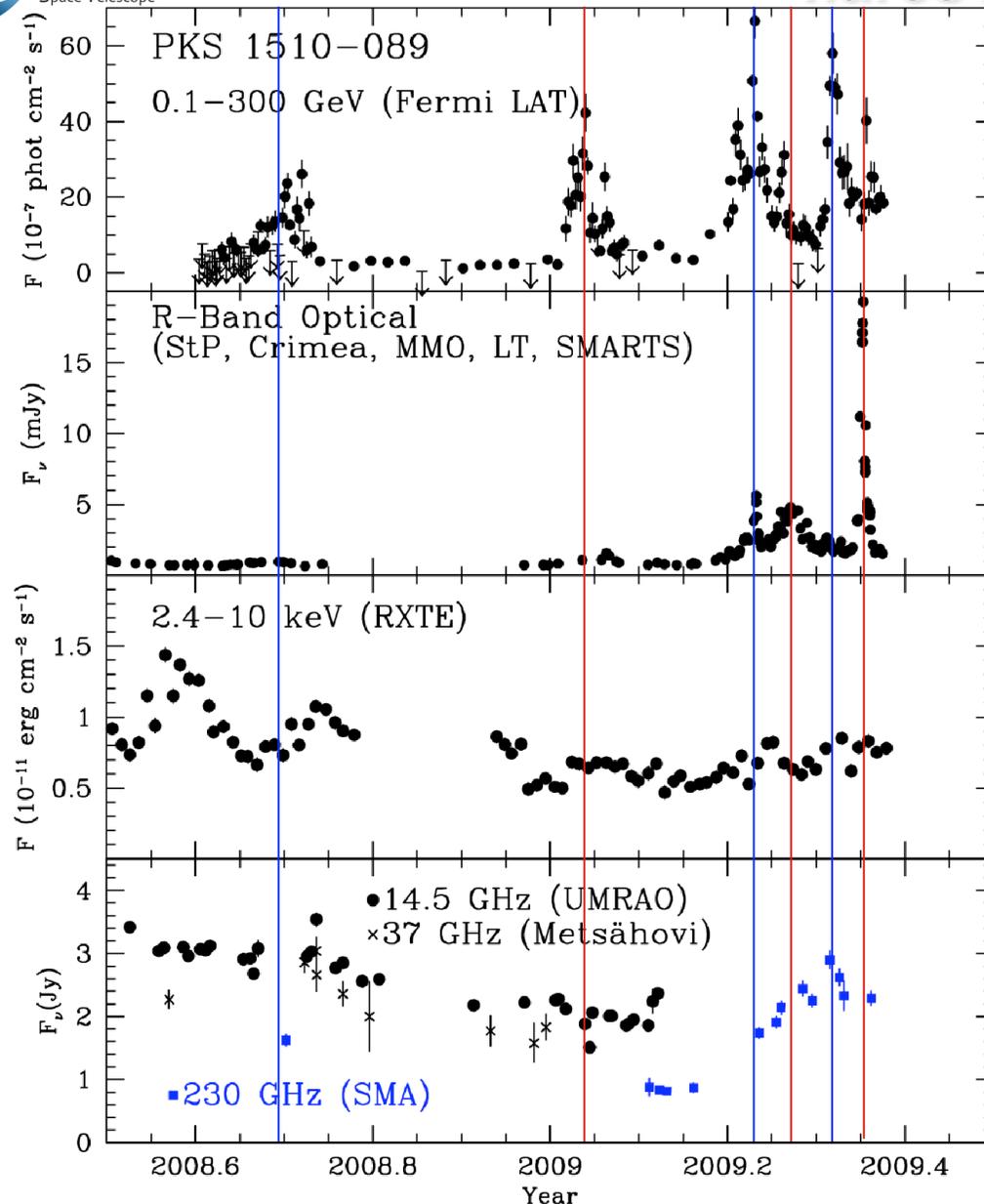
**Multiwavelength Studies**

**Dave Thompson**

## Outline

1. Examples of MW work involving Fermi
2. Project efforts to promote MW work
3. Suzaku/Fermi discussions

# Why not exact correspondence between $\gamma$ -ray & lower- $\nu$ flares?



Optical/mm-wave: synchrotron

→ Depends on  $B$  &  $N_0$

$\gamma$ -ray: inverse Compton

→ Depends on  $N_0$  & seed photon field

In 1 day, knot in 1510-089 moves by more than 1 pc!

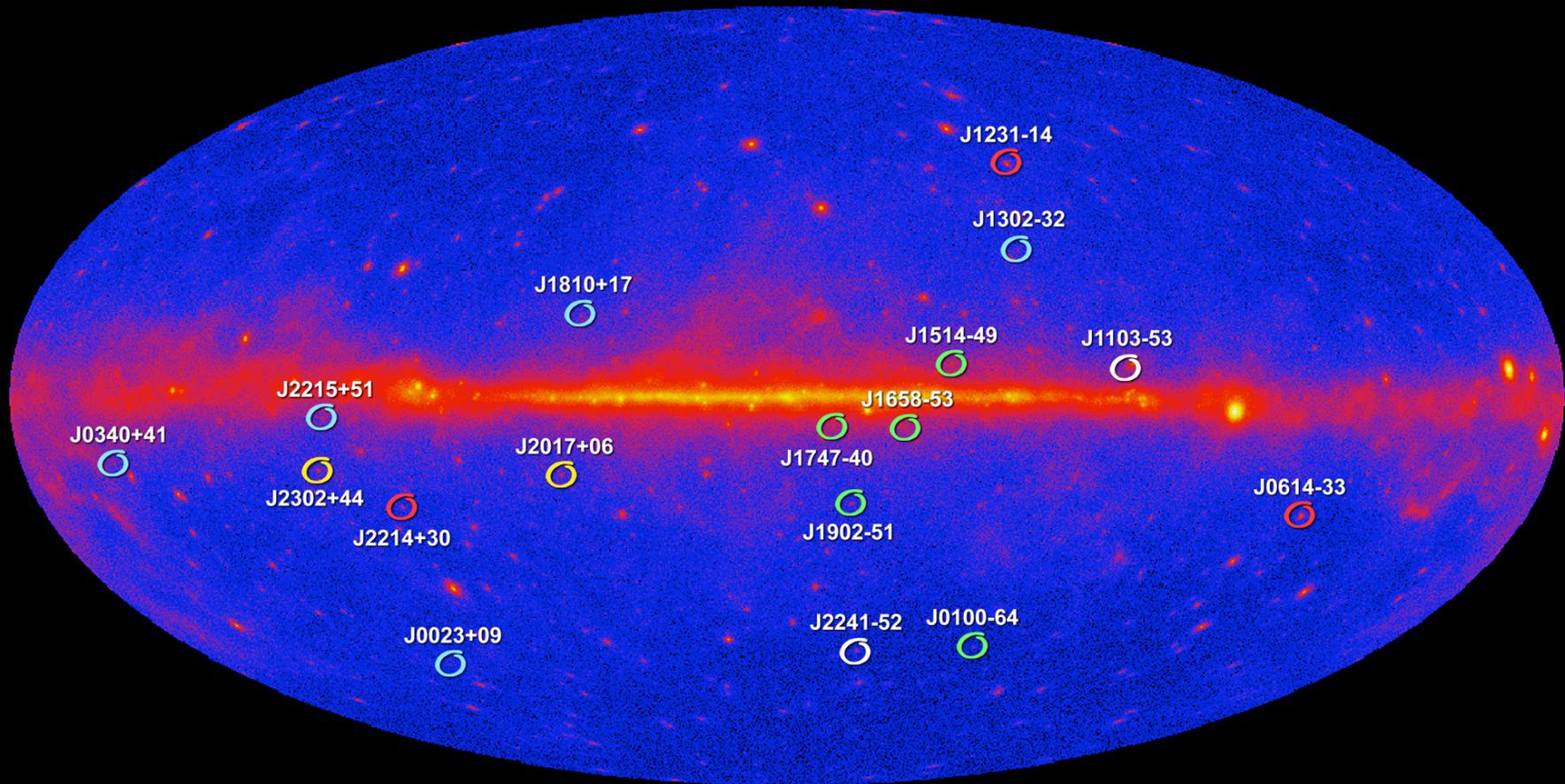
→ Number of external seed photons must change dramatically to get  $\gamma$ -ray flare without synchrotron flare

Still mysterious: why are X-rays relatively inactive in spring 2009 while all other wavebands are in outburst?

From Alan Marscher

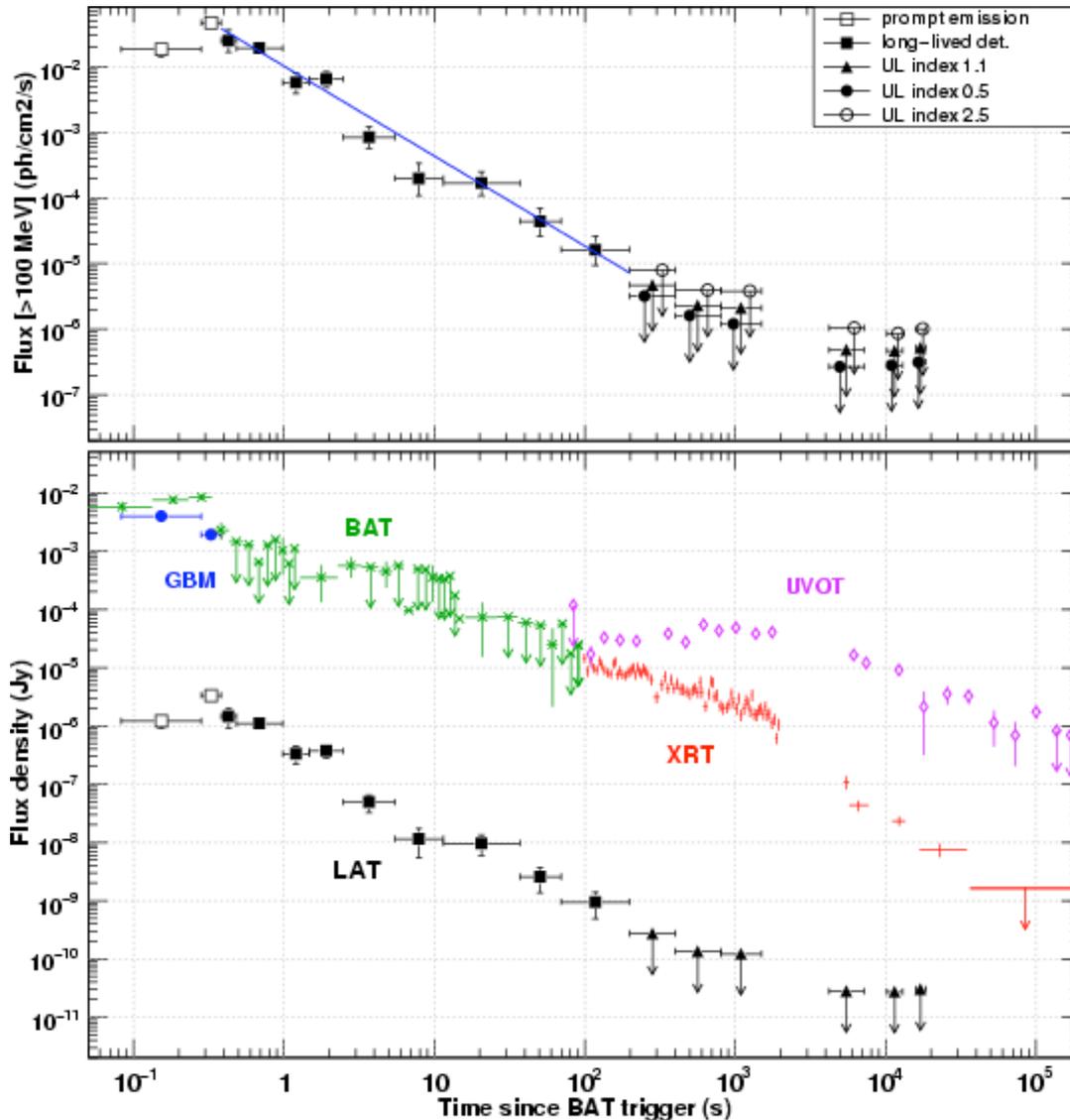
# Millisecond Pulsar Discoveries

## New Millisecond Radio Pulsars Found in Fermi LAT Unidentified Sources



-  Led by Fernando Camilo (Columbia Univ.) using Australia's CSIRO Parkes Observatory
-  Led by Mallory Roberts (Eureka Scientific/GMU/NRL) using the NRAO's Green Bank Telescope
-  Led by Scott Ransom (NRAO) using the Green Bank Telescope
-  Led by Ismael Cognard (CNRS) using France's Nançay Radio Telescope
-  Led by Mike Keith (ATNF) using Parkes Observatory

# GRB090510 : MW Afterglow



Significant emission (TS>25)  
up to T<sub>0</sub>+200s

No evidence of a spectral evolution

LAT lightcurve

be

s

fit by a power-law:  $a = -1.38 \pm 0.07$

Black : LAT

White : LAT (prompt)

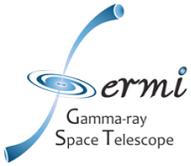
Blue : GBM (prompt)

Green : BAT (triggered on prompt)

Red : XRT (after T<sub>0</sub>+100s)

Violet : UVOT (after T<sub>0</sub>+100s)

De Pasquale et al. 2010



# Promoting Multiwavelength Research

## Bursts - GCN Notices, Circulars, and Reports

**For other results - public data, catalogs, and other published Fermi results are sometimes all that is needed.**

*Science* 2 April 2010:  
Vol. 328. no. 5974, pp. 73 – 75  
DOI: 10.1126/science.1184192

[< Prev](#) | [Table of Contents](#) | [Next >](#)

### REPORTS

## **Evidence for Strong Extragalactic Magnetic Fields from Fermi Observations of TeV Blazars**

**Andrii Neronov<sup>\*</sup> and Ievgen Vovk**

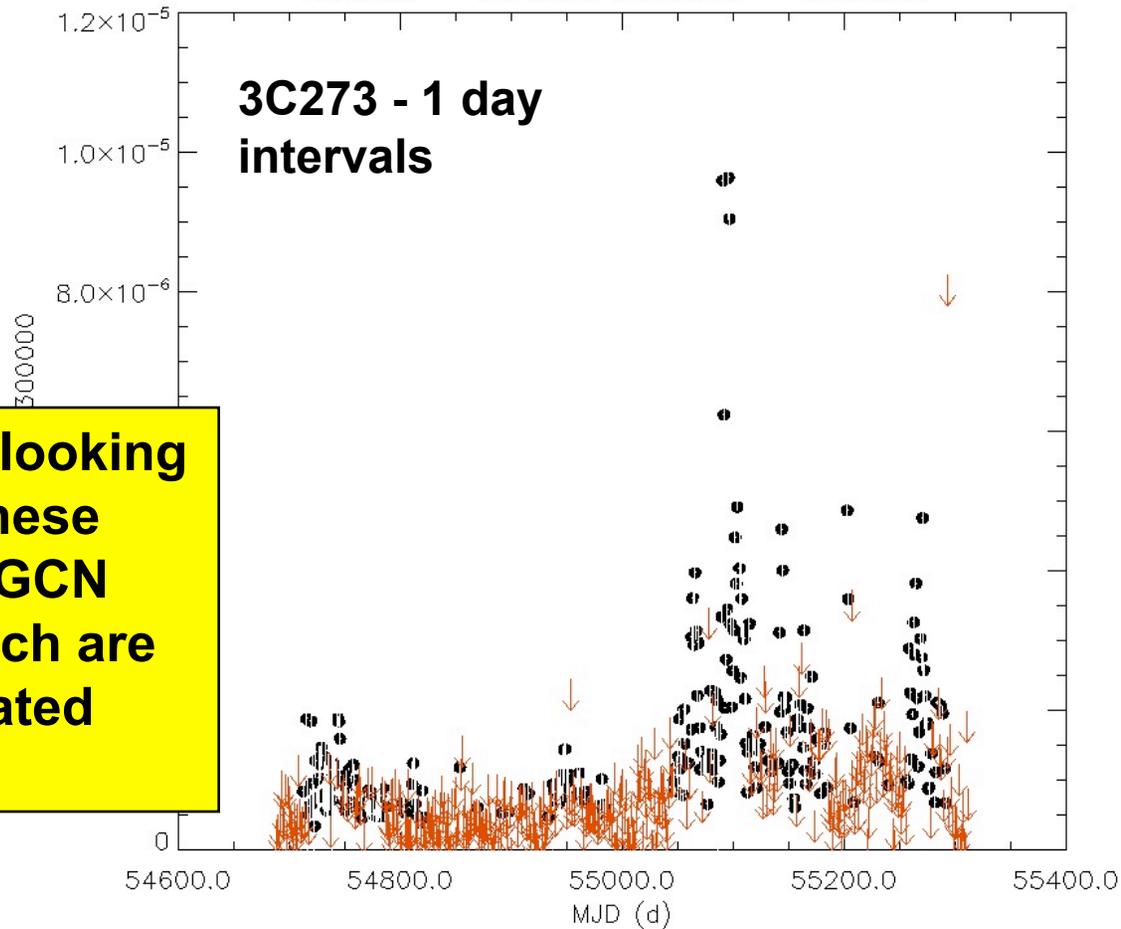
# Promoting Multiwavelength Research

For many observers, the question is, “What is interesting in the gamma-ray sky right now?”

Source = 3C 273 Duration = 86400.0

Several ap  
Monitored  
values are

We are currently looking at a way to use these files to generate GCN flare notices, which are useful for automated telescopes.



**Outside**

[GCN](#)  
[IAUCs](#)

**Other**

MacOS: [Dashboard Widget](#)  
Follow ATel on [Twitter](#)

## The Astronomer's Telegram

for reporting and commenting on new astronomical observations

[Post a New Telegram](#) | [Search](#) | [Information](#) | [Mirror Software](#)

[Telegram Index](#)

[Register To Post](#) | [Email and RSS Subscriptions](#) | [Forget your password?](#)

**Present Time: 30 Apr 2010; 12:46 UT**

[ [Previous](#) | [Next](#) ]

### **Fermi and Swift observation of a GeV/X-ray/UV flare from PKS 2233-148**

ATel #2589; [K. V. Sokolovsky \(MPIfR/ASC Lebedev\)](#), [D. Donato \(NASA/GSFC\)](#), [F. K. Schinzel \(MPIfR\)](#) on behalf of the [Fermi Large Area Telescope Collaboration](#); and [Y. Y. Kovalev \(ASC Lebedev/MPIfR\)](#)

*on 29 Apr 2010; 19:11 UT*

*Distributed as an Instant Email Notice (Request for Observations)*

*Password Certification: Frank Schinzel ([schinzel@mpifr-bonn.mpg.de](mailto:schinzel@mpifr-bonn.mpg.de))*

**Subjects: Radio, Ultra-Violet, X-ray, Gamma Ray, >GeV, Request for Observations, AGN**

The Large Area Telescope (LAT), one of the two instruments on the Fermi Gamma-ray Space Telescope, has observed an increasing gamma-ray flux from a source positionally consistent with the BL Lac type object PKS 2233-148 (R.A.= 22:36:34.08716, Dec.= -14:33:22.1897, J2000; Petrov et al. 2008, AJ, 136, 580).



# Flare Advocates - Follow-up Work

Othe  
1. Th

<http://fermisky.blogspot.com/>

FERMI GAMMA-RAY SKY

7

Othe  
2. Ju  
Daily R  
Highlight  
blog

TUESDAY, APRIL 20, 2010

## Fermi LAT weekly report N. 97

Covered period: 2010.Apr.12 2010.Apr.18

LAT Mission week: 97.57 - 98.57

Extragal  
• PKS 042  
• PKS 053  
• CGRaBS  
The source  
1.4+/-0.5  
• CGRaBS  
• 4C +21.3  
• 3C 454.3

- **3C 454.3** detected in high state again this week, although seems fading. At the beginning of the week, on April 13, it reached a daily maximum flux of  $(13.4 \pm 0.9) \times 10^{-6}$  while it has been detected with a flux of  $(7.3 \pm 0.7) \times 10^{-6}$  in the last day.
- **PKS 0537-441** active during the first days of the week (daily flux above  $10^{-6}$ ), then fading but maintaining a daily flux less above  $0.6 \times 10^{-6}$ .

LAT DATA

[LAT First Catalog](#)

[LAT Monitored Source List Light Curves](#)

[LAT Bright Source List](#)

[Browse interface to monitored source data](#)

[Contact Information by Individual Sources](#)

BLOG ARCHIVE

▼ 2010 (22)

▼ April (3)

[Fermi LAT weekly report N. 97](#)

[Fermi LAT weekly report N. 96](#)

[Extra note - Apr. 15, 2010](#)

▶ March (8)

▶ February (5)

▶ January (6)

aily  
ing up  
LAT  
is is  
ternal  
e some  
in

ublicly  
activity  
sky  
in sky

## A Surprise - A Gamma-ray Nova?

- In early March, the LAT Flare Advocates found a new, flaring gamma-ray source in the Cygnus region.

Electronic Telegram No. 2199

- Central Bureau for Astronomical Telegrams  
INTERNATIONAL ASTRONOMICAL UNION  
CBAT Director: Daniel W. E. Green; Room 209; Dept. of Earth and Planetary  
Sciences; Harvard University; 20 Oxford St.; Cambridge, MA 02138; U.S.A.  
e-mail: [cbat@iau.org](mailto:cbat@iau.org); [cbatiau@eps.harvard.edu](mailto:cbatiau@eps.harvard.edu)  
URL <http://www.cfa.harvard.edu/iau/cbat.html>

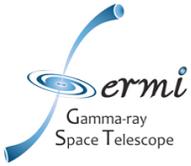
V407 CYGNI

Hiroyuki Maehara, Kwasan Observatory  
discovery by Koichi Nishiyama (Fukuoka, J  
Japan) of an apparent unusually bright ou  
star V407 Cyg on an unfiltered CCD image

Also seen in X-rays and radio!  
3E camera  
filtered  
mag 6.8) and 10.814 (mag 6.9) using a 0.4



The amateur astronomers who  
discovered the optical flare.



# Fermi - Suzaku Planning

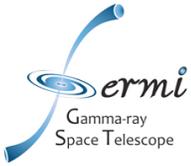
Dear Dr. Thompson, Dr. McEnery,

We have discussed Fermi-Suzaku joint proposals among the Suzaku-observation steering committee members in Japan. They have no objection and support it.

Although, we have plenty of time to discuss, if we start from Suzaku Cycle-6 GO and Fermi Cycle-4 GI, here I would like to show some our idea as the starting point.

(0) Definition of the Fermi Suzaku joint program

Proposers of Fermi GI program can ask Suzaku observation time. The proposal will be reviewed by the Fermi review board. The Suzaku team will support the review process mainly in view of technical feasibility (and science, TBD). The Suzaku team may not accept some of the proposals in case of technical difficulty or schedule conflicts.



## Summary

**The Fermi Project and the instrument teams are enthusiastic supporters of multiwavelength studies.**

**We are always looking for new ways to promote such research, and we welcome any suggestions from the Users Group.**