

Welcome



Fermi Data Analysis Workshop

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Workshop Objectives



- Today we'll cover *Fermi* analysis basics:
 - Data content, selection cuts, caveats
 - Analysis methodologies, synopsis of tools
 - ML method→ point source analysis
 - Light curve & pulsar analysis
- Emphasis on hands on analysis
 - roving support staff
- Feedback & discussion
- GI Program: Guidelines for Proposer



Agenda, Staff



10:00 - 10:10	Overview, Workshop Objectives	Shrader
10:10 - 10:25	Data Selection & Exploration	Ferrara
10:25 - 10:30	Data Selection Caveats	McEnery
10:30 - 10:45	Point Source Analysis	Davis
10:45 - 11:00	Using the Catalog for Analysis	Ferrara
11:00 - 11:30	Hands on Session 1	All
11:30 - 11:40	Q&A, Review	All
11:40 - 12:30	Hands on Session 2	All
12:30 - 14:00	Working Lunch/Lunch Break	
14:30 - 14:50	Fermi Mission Status	Hays
14:50 - 15:10	Catalogs and MW Opportunities	Digel
15:10 - 15:30	Extragalactic Results	Thompson
15:30 - 15:50	Gamma-Ray Bursts	Gehrels
15:50 - 16:10	Results on Galactic Sources	Harding
16:10 - 16:30	GI Program Overview & Q&A	Shrader
13.10	or regiant overview a gart	<u> </u>



Prerequisites



- Science Tools Installation hopefully done prior, but we can help as needed
 - Workshop web page is useful resource
- Sample datasets on workshop web page
 - can substitute alternative data selections, but be cognizant of run-time, S/N issues
- Access to Fermi SSC web site
 - Data analysis documentation sets
 - Threads, Cicerone, Reference ('fhelp') docs



A Few Fermi Analysis Basics



What's different about Fermi data analysis?

- Structured sky backgrounds
- Energy-dependent point spread function
- Instrument response function(s) IRFs
 - Multiple dependencies: instrument design, event reconstruction, background & quality selections
- Wide field of view, continuously variable aspect



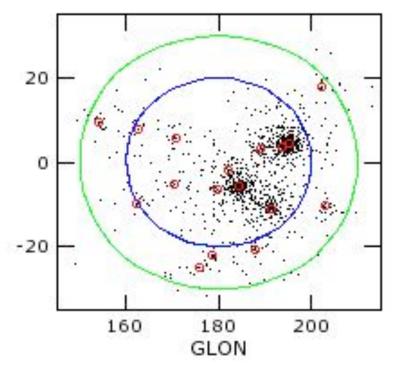
Energy Dependent PSF



• Sources must be fit simultaneously.

- Broad and energy-dependent PSFs: $\sigma_{68} < 3.5^{\circ}$ for 100 MeV (on axis) and $< 0.1^{\circ}$ for 10 GeV
- Emission from nearby point sources overlap.
- Intrinsic source spectrum affects the degree of source confusion.
- "Source region" must be significantly larger than the "region-of-interest" (ROI).

• Anticenter region:

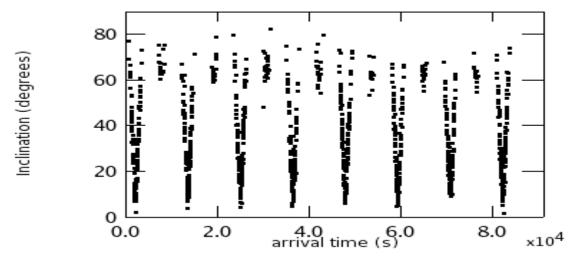




Continuously Varying Aspect



- Each event effectively has its own response function:
 - Large FOV, ~ 2.4 sr
 - Strong variation of response as a function of photon incident angle, $A_{eff} \propto \cos \theta$
 - Scanning mode of operation: 95 min orbit \Rightarrow continuous aspect changes of 4°/min.



C. Shrader, NASA/GSFC



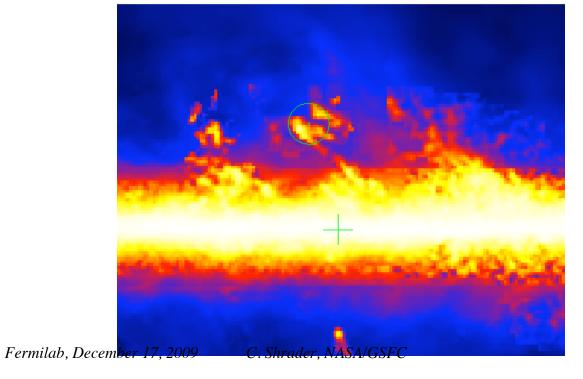
Diffuse Emission



- Emission results from cosmic ray interactions with interstellar gas.
- Models rely on HI & CO observations for the gas distribution

• These observations reveal structures on angular scales

similar to the PSF:





Useful Information



• Web URLS:

- http://fermi.gsfc.nasa.gov/workshops/data_analysis_dec09/
 Workshop web site
- <u>http://fermi.gsfc.nasa.gov/ssc/</u> FSSC home
- http://fermi.gsfc.nasa.gov/ssc/data/access/ Data access
- http://fermi.gsfc.nasa.gov/ssc/data/analysis/ Data analysis page
- http://fermi.gsfc.nasa.gov/cgi-bin/ssc/faq/glastfaq.cgiFAQs



Cycle-3 Timeline



Announcement (as part of ROSES 2008) Septem

Release online proposal aids &

documentation

Notices of Intent (optional)

Proposals Due

Proposal Peer Review

Stage-II (budget proposal) solicitation

Budget deadline, processing & grants

administration

Fermi Cycle 3 Begins

September, 2009

November 5, 2009

November 16, 2009

February 5, 2010

April, 2010

May 2010

June-July 2010

Mid August, 2010





Let's get started ...





Extra Slides





- + NASA Homepage
- + GSFC Homepage
- + Fermi Homepage

SEARCH Fermi:

Search

+ GO

Important resource:

FSSC data analysis web page



+ FSSC Home

Data

Data Policy

Data Access

Data Analysis

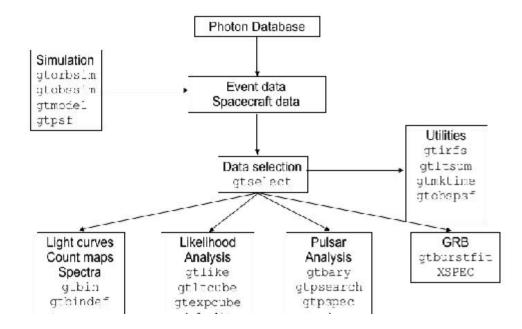
- + System Overview
- + Caveats
- + Software Download
- + Documentation
- + Analysis Threads
- + User Contributions

Newsletter

FAQ

Overview: LAT Data Analysis Tools

Through a collaborative effort between the Fermi Science Support Center and the LAT instrument team a suite of instrument specific science analysis tools has been developed for public release. This software will be distributed and maintained by the FSSC. It has been designed within the framework of the HEADAS FTOOLS methodology, to ensure cross-mission compatibilities wherever possible and to minimize the learning curve for users of other high-energy astrophysics mission data sets. The general analysis flow is illustrated in this graphic:



Fermi





Sample data sets posted online:

Vela Pulsar, 3C 454.3



Workshop Data

- 3c453
- Vela





Science Tools Download Page:

Hopefully this has already been done(?) but if not, refer to this page for supported platforms, installation instructions



We have binary distributions for:

- Scientific Linux 4.4 32 bit libc 2.3.4
- · Scientific Linux 5 32 bit libc 2.5
- . Scientific Linux 4 64 bit libc 2.3.4
- Scientific Linux 5 64 bit libc 2.5
- MAC OS X 10.4 nowerno