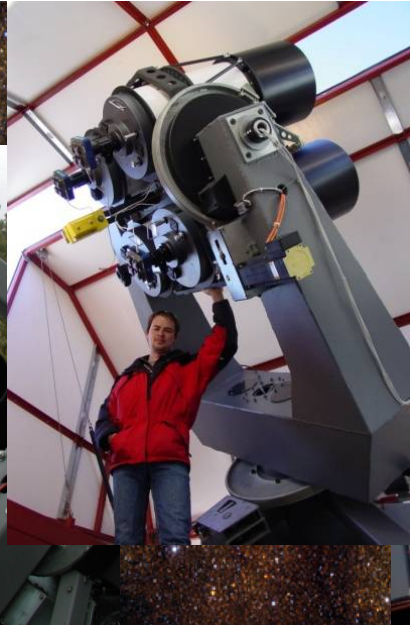


# The RAPTOR (*RAP*id Telescopes for *Optical* Response) telescopes and GeV/TeV gamma-ray astronomy





# RAPTOR, Thinking Telescopes



## Fast Response:

- Begin imaging anywhere in the sky in  $< 6s$
- Seven 0.4m class telescopes
- T: True Simultaneous Multi-Color: Clear, V, R, I
- Z: Fast cadence: 7Hz



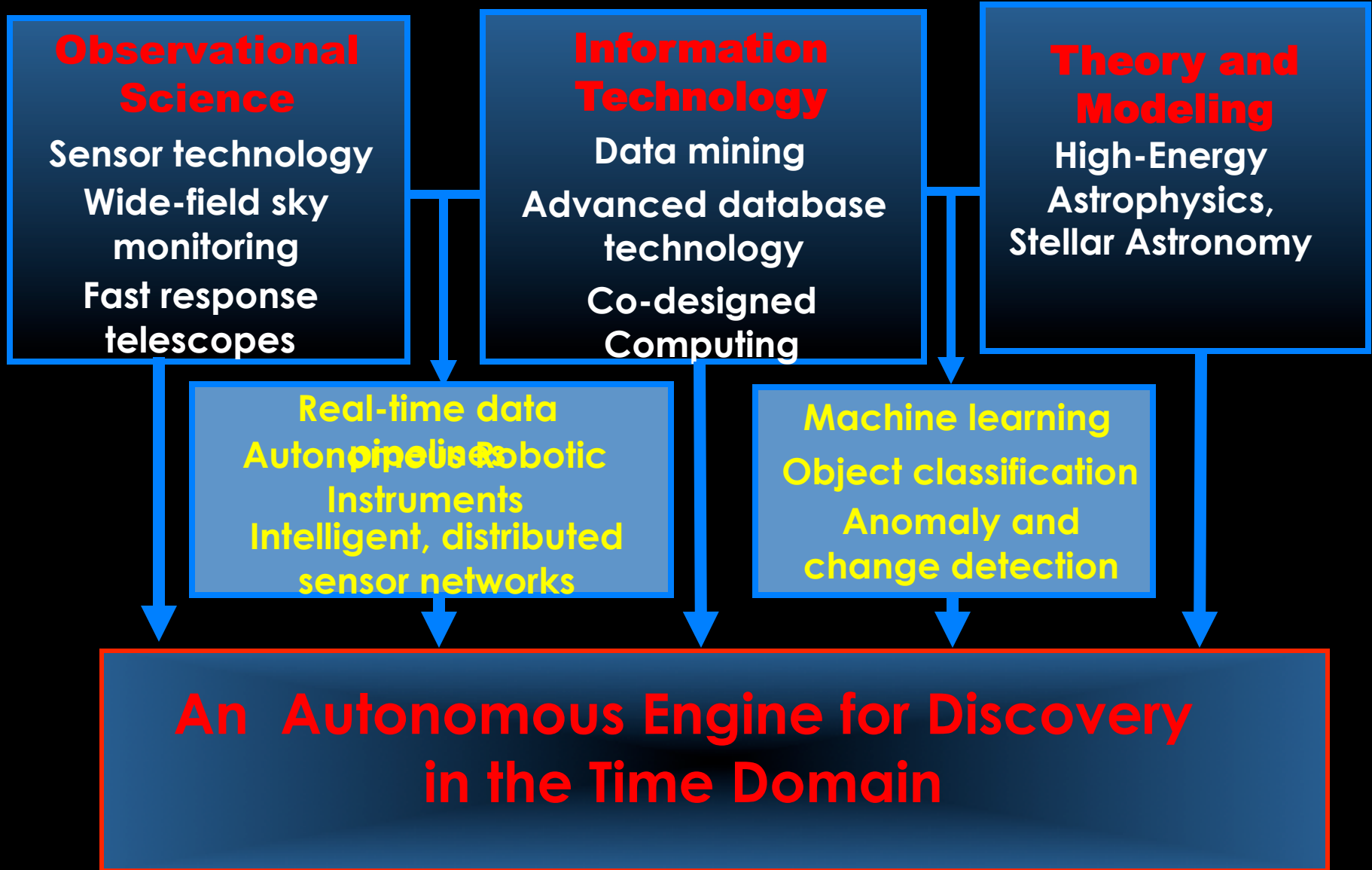
## Persistent Monitoring:

- Real-time transient detection
- Q: Full Sky to 10<sup>th</sup> Rmag in 20s
- K: 16<sup>th</sup> Rmag in 30s, Full Sky 7m

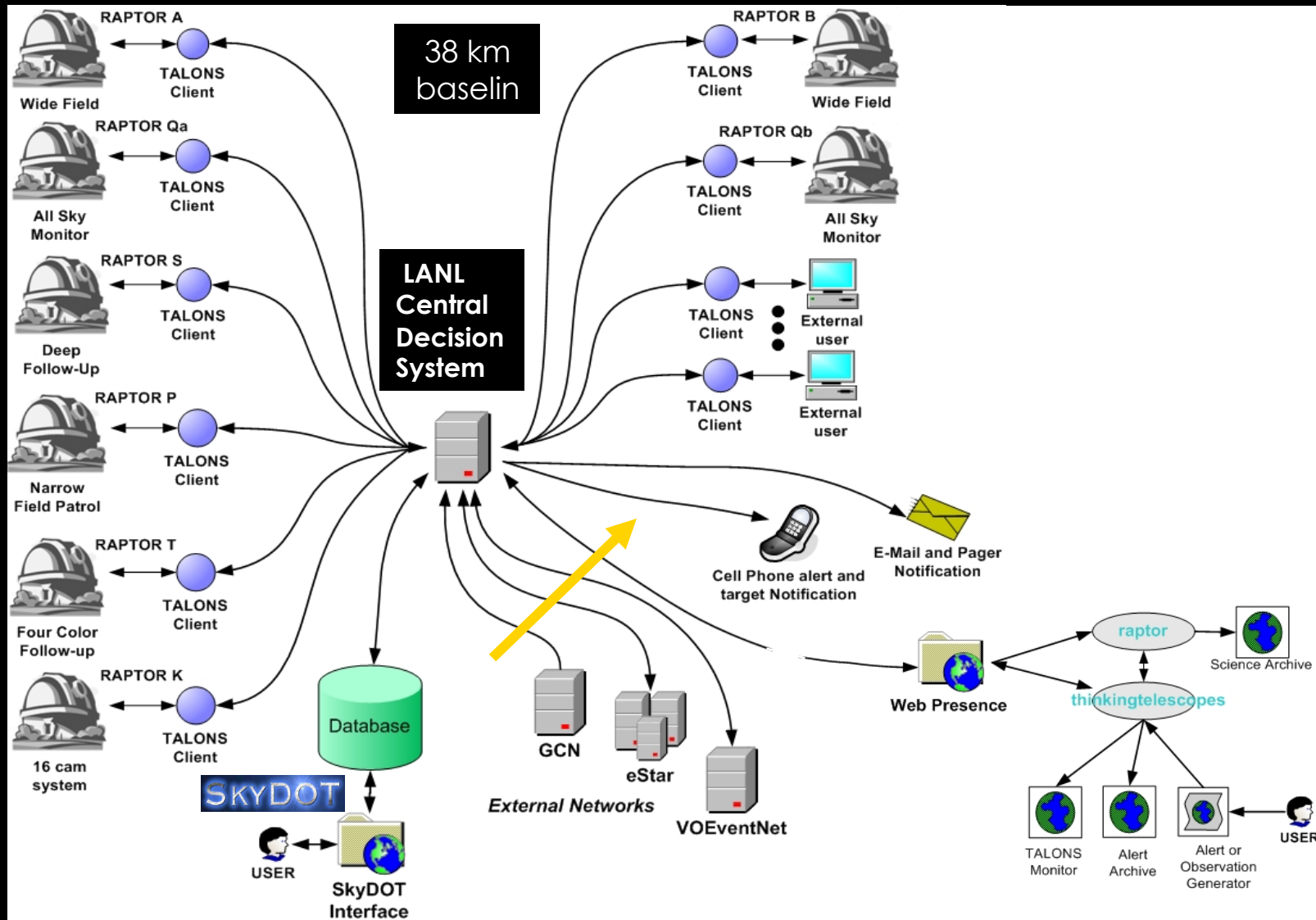




# The RAPTOR/Thinking Telescopes Project---Autonomous Robotic Astronomy



# Distributed "Thinking" Telescopes Network

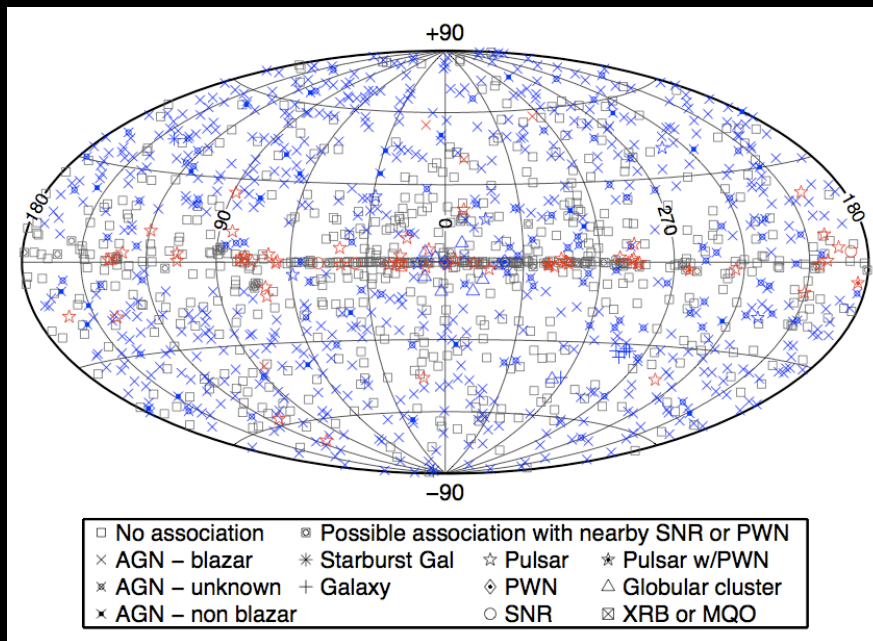


Aspen Meeting---Galactic GeV/TeV sources

# Potential *Fermi* Science Areas

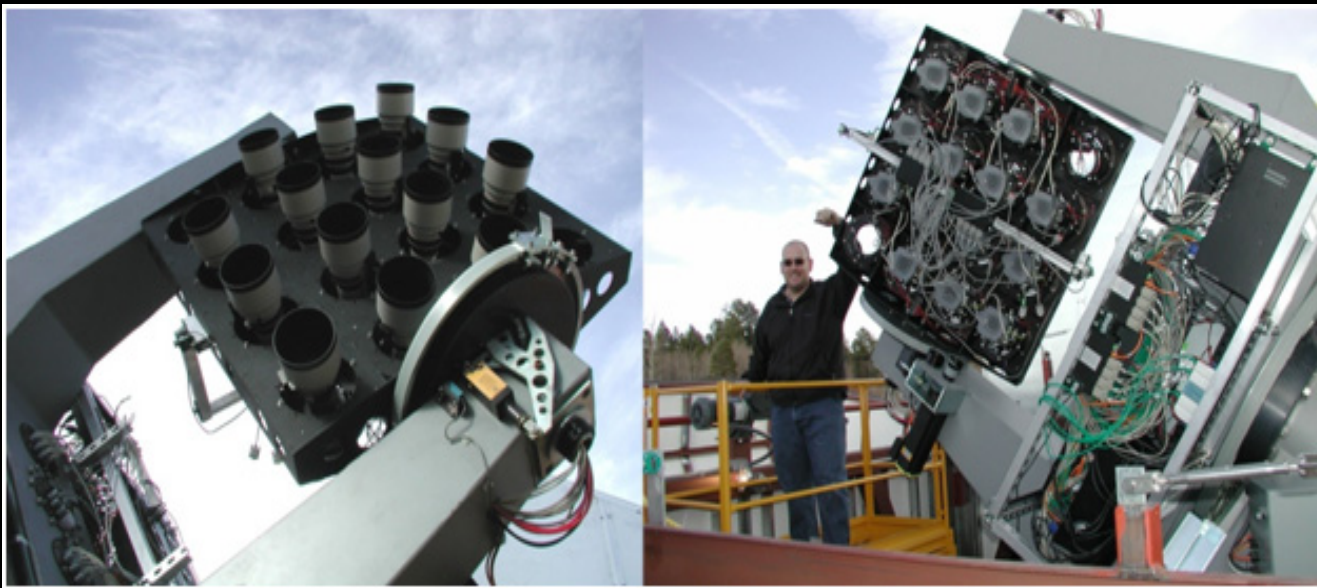
- Revealing the identities of unidentified Fermi GeV gamma-ray sources
- GeV emitting Blazars
- GeV emitting explosions in Galactic Sources
- GeV emitting Gamma-ray Bursts (GRBs)

# Unidentified Fermi Sources



- More than 1000 new sources have been discovered, most (~600) are unidentified.
- Gold standard for identification is finding of optical counterpart through correlated variability
- Most localized to ~1 degree. Each has many possible counterparts even at depth of 16<sup>th</sup> magnitude
- In survey mode, Fermi scans the sky in 180 minutes.
- Challenge for optical systems --- scan the sky at a similar cadence.

# Full Sky Persistent Monitoring RAPTOR-K

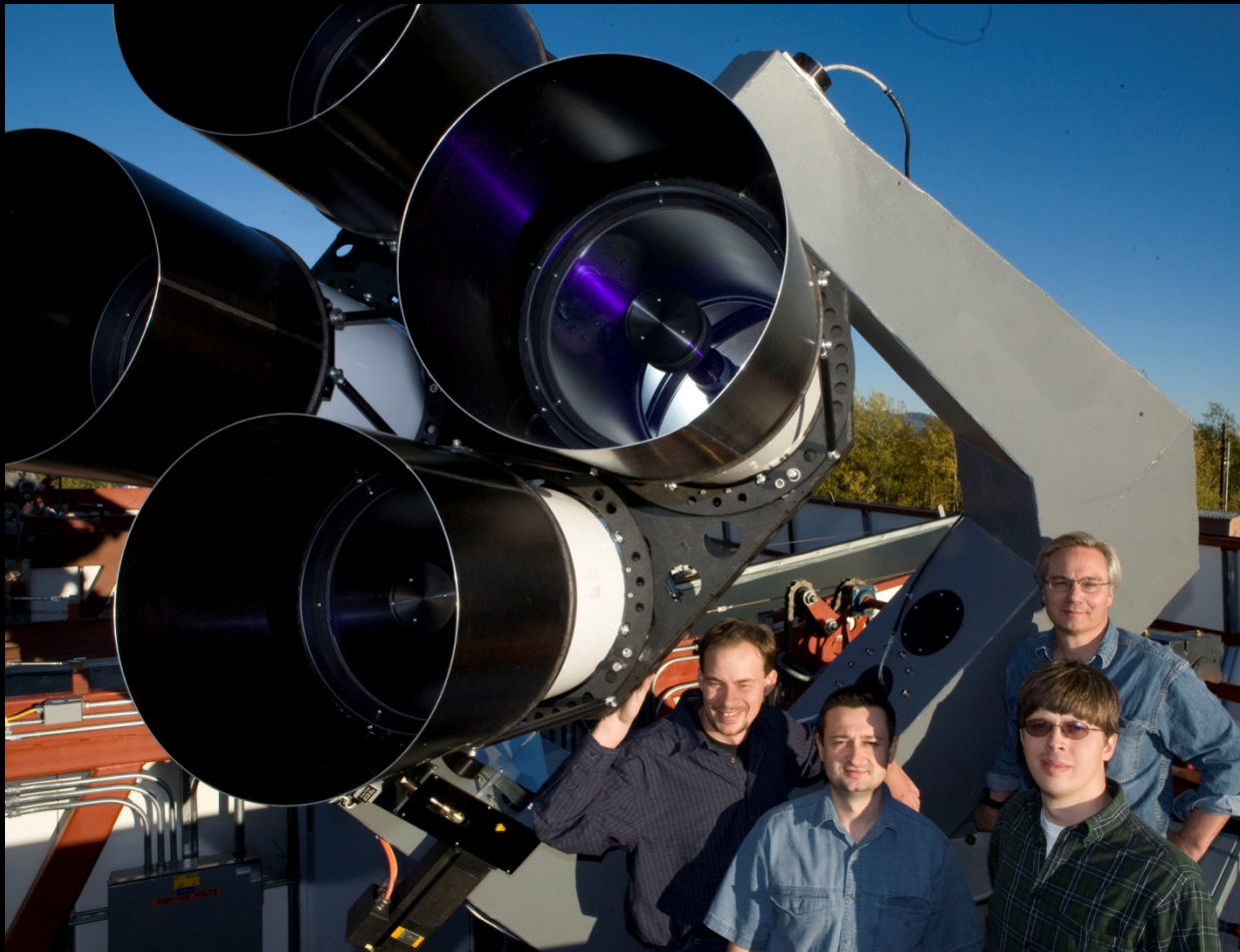


- Scans the full sky at site in 10 minutes
- Sensitivity of  $R \sim 16^{\text{th}}$  magnitude in 30s
- 16 telescope array
- $10^3$  deg-sq FOV



# Persistent monitoring of localized ( $<0.5$ deg) sources

## RAPTOR-T



- Each site monitors about 100 sources of interest twice per night .

- Simultaneous cVRI Imaging,  $C_R=19.9$  in 90-sec.

- RAPTOR-T North (built and operating since 2008)

- RAPTOR-T South (built)

- (Also monochrome monitoring with RAPTOR-S and two RAPTOR-W telescopes)



# Potential *Fermi* Science Areas

- Revealing the identities of unidentified Fermi GeV gamma-ray sources
- GeV emitting Blazars
- **GeV emitting explosions in Galactic Sources**
- GeV emitting Gamma-ray Bursts (GRBs)

# How do you establish the existence of new classes of transients?

- Seen simultaneously by multiple instruments in the same wavelength/energy band.
- Or, in a single instrument, see enough of them to exclude instrumental or man-made effects.
- Seen simultaneously in different energy band with a both having a distinct (and uncommon) temporal signature.

Are there Explosive, >100 MeV gamma-ray emitting, galactic sources? (not counting the Sun)

## EGRET Detection of Cen X-3 outburst

Vestrand, Sreekumar and Mori (ApJ, 483, L49 (1997))

- Found as part of an EGRET GI program to survey suspected TeV emitting x-ray binary systems
- Detected an outburst of >100 MeV with a flux of  $9 \times 10^{-7}$  photons  $\text{cm}^{-2} \text{s}^{-1}$  ( $\sim 5\sigma$  significance) during a two week interval in October 1994.
- Cen X-3 is a disk-fed pulsar in orbit around an O-type supergiant companion ---2.09 day orbit period, 4.8 second spin period
- Detected during an interval of rapid x-ray pulsar spin-down



# Cen X-3 Outburst

- $L_\gamma \approx 5 \times 10^{36}$  ergs/sec----- $L_x \approx 10^{38}$  ergs/sec
- Spectrum— power law with index  $1.8 \pm 0.4$

## Was it a background Blazar? Maybe but...

- Six Parkes sources within the 95% contour, using the standard gamma/radio flux ratios for EGRET blazars one predicts a 5 Ghz flux of  $\sim 4500$  mJy---- all potential candidates have fluxes less than 70 mJy. (But maybe it is an outlier...)

# Evidence for gamma-ray modulation at drifting x-ray spin period

- No sign of 2.09 day orbital modulation
- No sign of nominal (steady) 4.8 second orbital modulation
- But pulsar was undergoing an interval of rapid spin-down during the EGRET observations. When the gamma ray arrival times were folded with the simultaneous BATSE measurements of the drifting x-ray spin period, we found modulation at the 3.5-sigma significance level for the EGRET detected gamma-rays.
- We did the timing analysis at four other positions in the field as controls, those locations showed no modulation.

# Are there new classes of Galactic GeV /TeV transients?

## GeV emission from Recurrent Nova?

- Nova, detected in the optical at 8<sup>th</sup> mag on March 10, 2010--V407 Cygni
- White dwarf accreting from Red giant star wind, material accretes and triggers thermonuclear explosion
- In ATEL #2487 Cheung et al. report the emergence of Fermi J2102+4542 on March 12 and associate it with the V407 Cygni
- Did a strong shock in the wind generate GeV emitting particles?
- Suggests that bright optical outbursts could signal the pending onset of gamma-ray outbursts. (as it sometimes does for x-ray outbursts)





# Conclusion

- Time domain astronomy employing advances in robotic instrumentation, real-time knowledge extraction, and global networking of heterogeneous observing platforms is fundamentally changing the way we search and interrogate astrophysical transients.
- Autonomous, Robotic, Optical instruments like RAPTOR have the potential to make important contributions to the study of explosive GeV/TeV transients. I would be very interested in talking about potential collaborations.