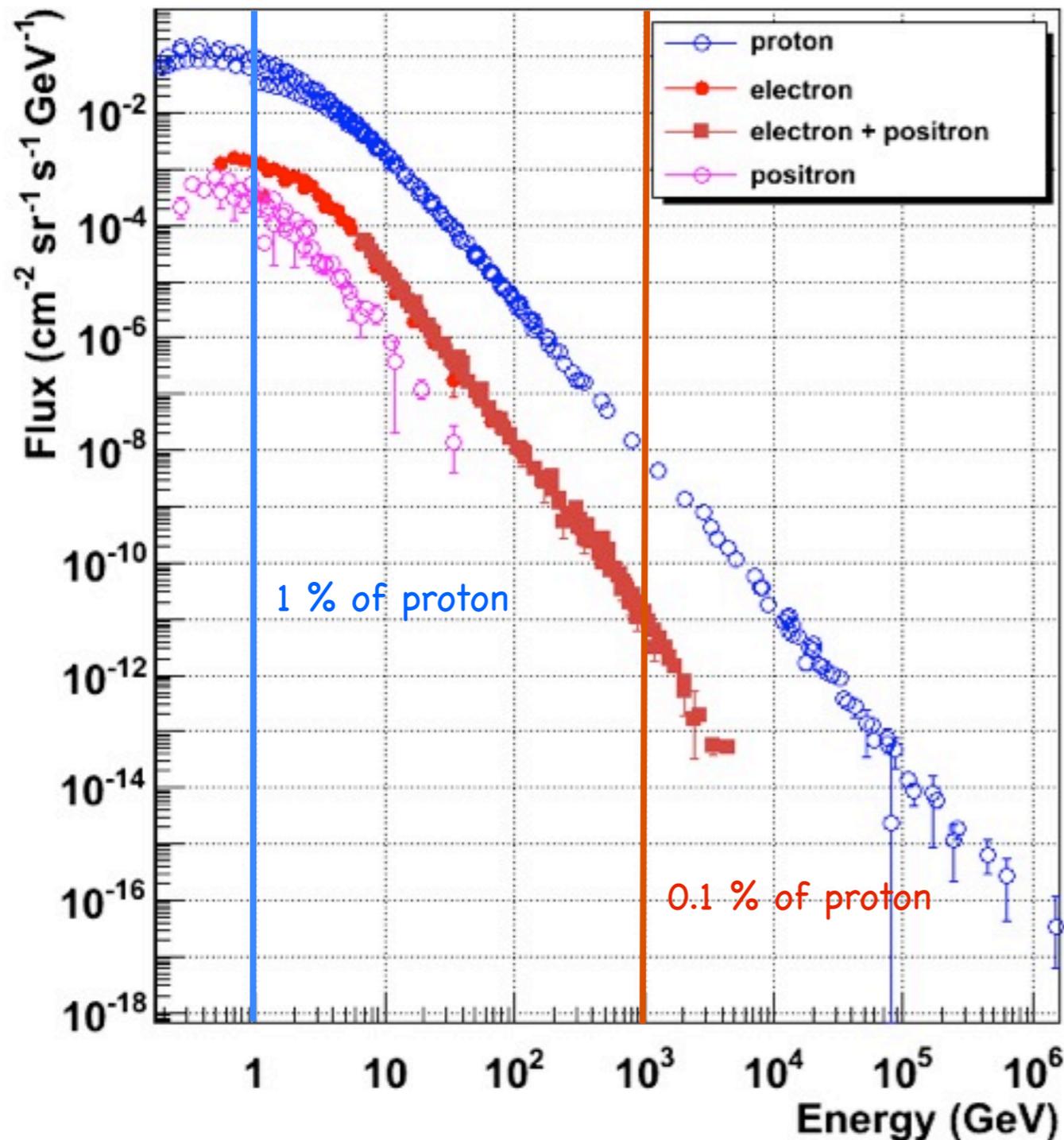


Multi-TeV measurement with CREST experiment

*Presented by Nahee Park
Enrico Fermi Institute, University of Chicago*

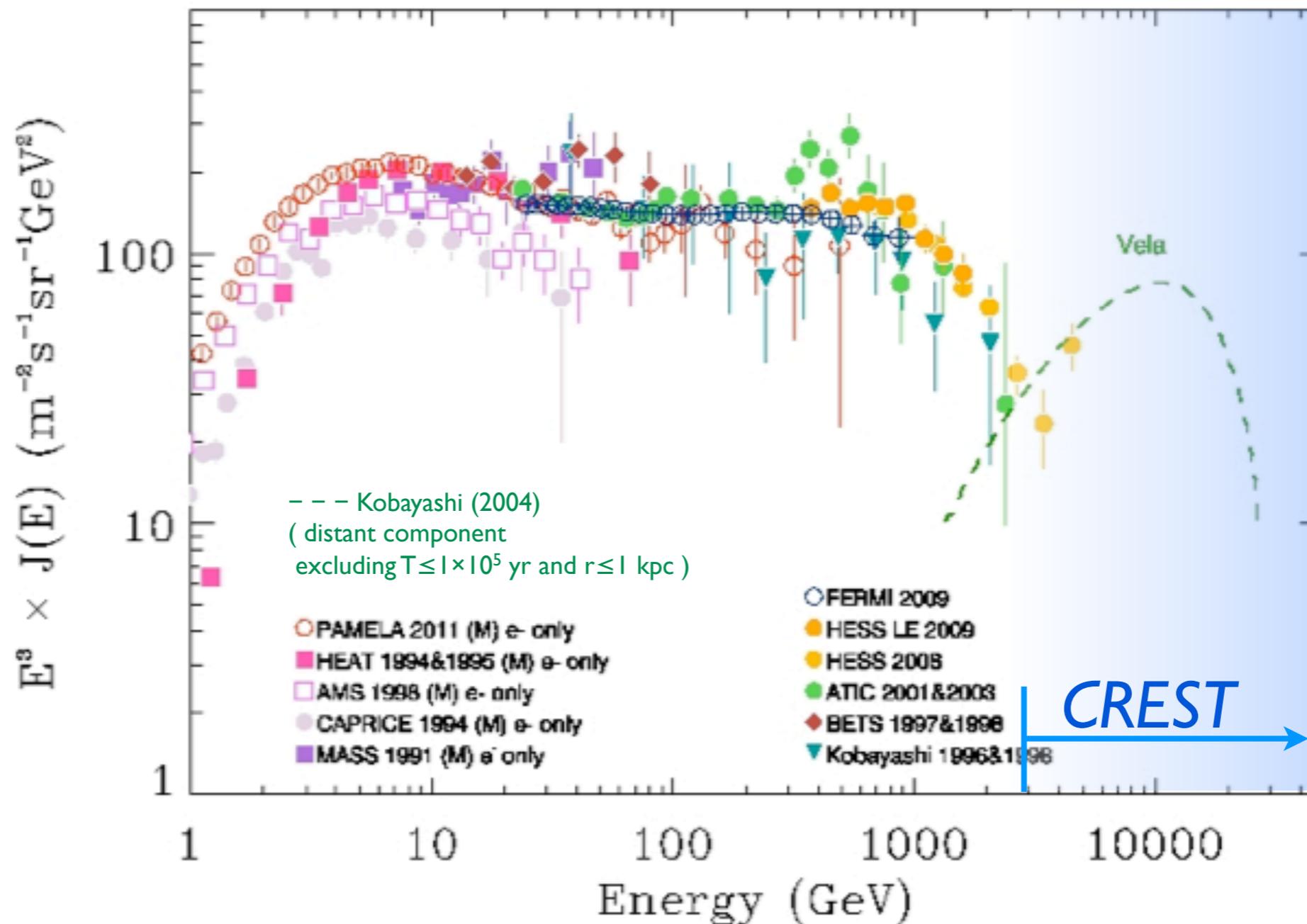
Cosmic Ray Electrons



- Primary + Secondary
 - Substantially primary (positron fraction ~ 10 %)
- ~ 1% of proton intensity at 1 GeV, rapidly decreased than proton
 - Energy loss of high energy electron is proportional to E²
 - TeV electron horizon : ~ 1 kpc (10⁵ yr propagation)
 - Possible local source : Vela, Cygnus loop, Monogen, SNRs

Compiled data up to Jan. 2010 from CR database (A.W.Strong et al, 2009 ICRC)

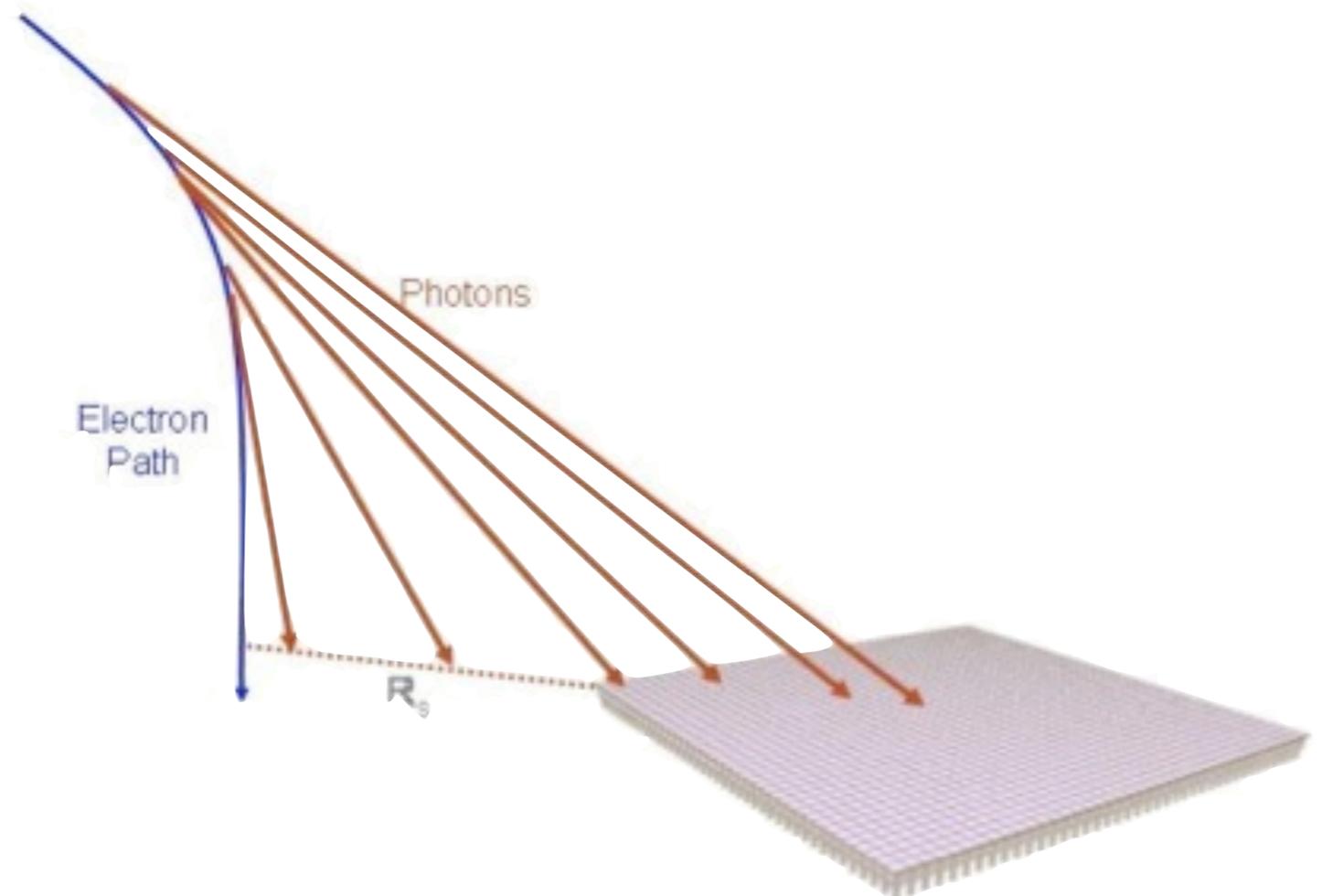
TeV measurement @ Earth



*Multi-TeV region largely unexplored,
where the potential is greatest for detecting nearby cosmic accelerators...*

Cosmic Ray Electron Synchrotron Telescope

- High energy electron ($> \text{TeV}$) measurement via synchrotron radiation
 - Detect x-ray synchrotron photons generated in the magnetic field of the Earth as primary electron passes through
 - Advantage
 - Increase of the effective area of instrument
 - Rejection of proton signal
- Designed for long duration balloon flight



CREST Collaboration



University of Chicago : S.Wakely, N. H. Park, D. Müller

Indiana University : C.R. Bower, J. Musser

Northern Kentucky University : S. Nutter

Penn State University : T.Anderson, S. Coutu, M. Geske

University of Michigan : M. Schubnell, G. Tarlé, A. Yagi, J. Gennaro

Signal and Background

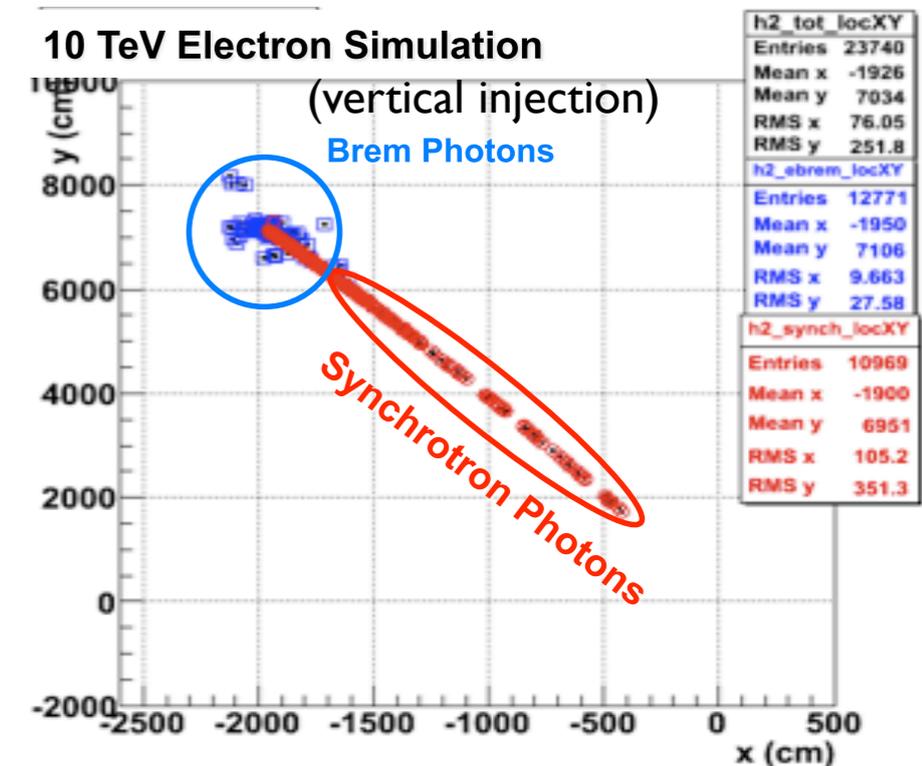
Signal

- Synchrotron radiation generated from electron
 - Line of photons arriving nearly simultaneously
 - Mean photon energy related to primary electron energy

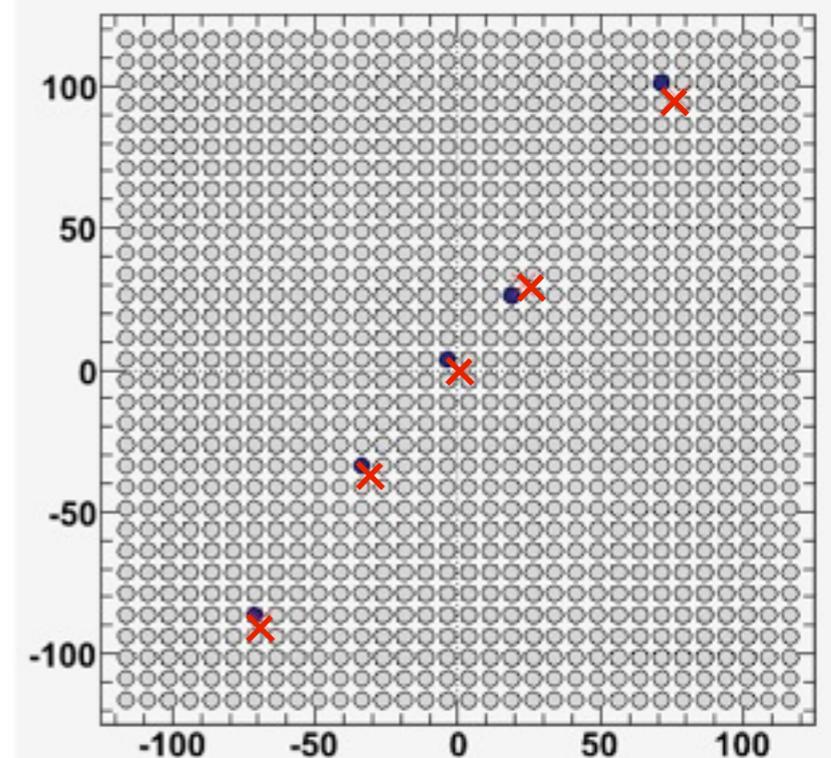
Background

- Cosmic and shower-produced x-rays and large charged particle flux
 - Random single x-ray coincidences
 - Interactions in the detector and frame
- Bremsstrahlung photons from low energy electrons

→ *Requires good timing resolution*



A simulated event with synchrotron photons



Detector Design

● Crystal Array

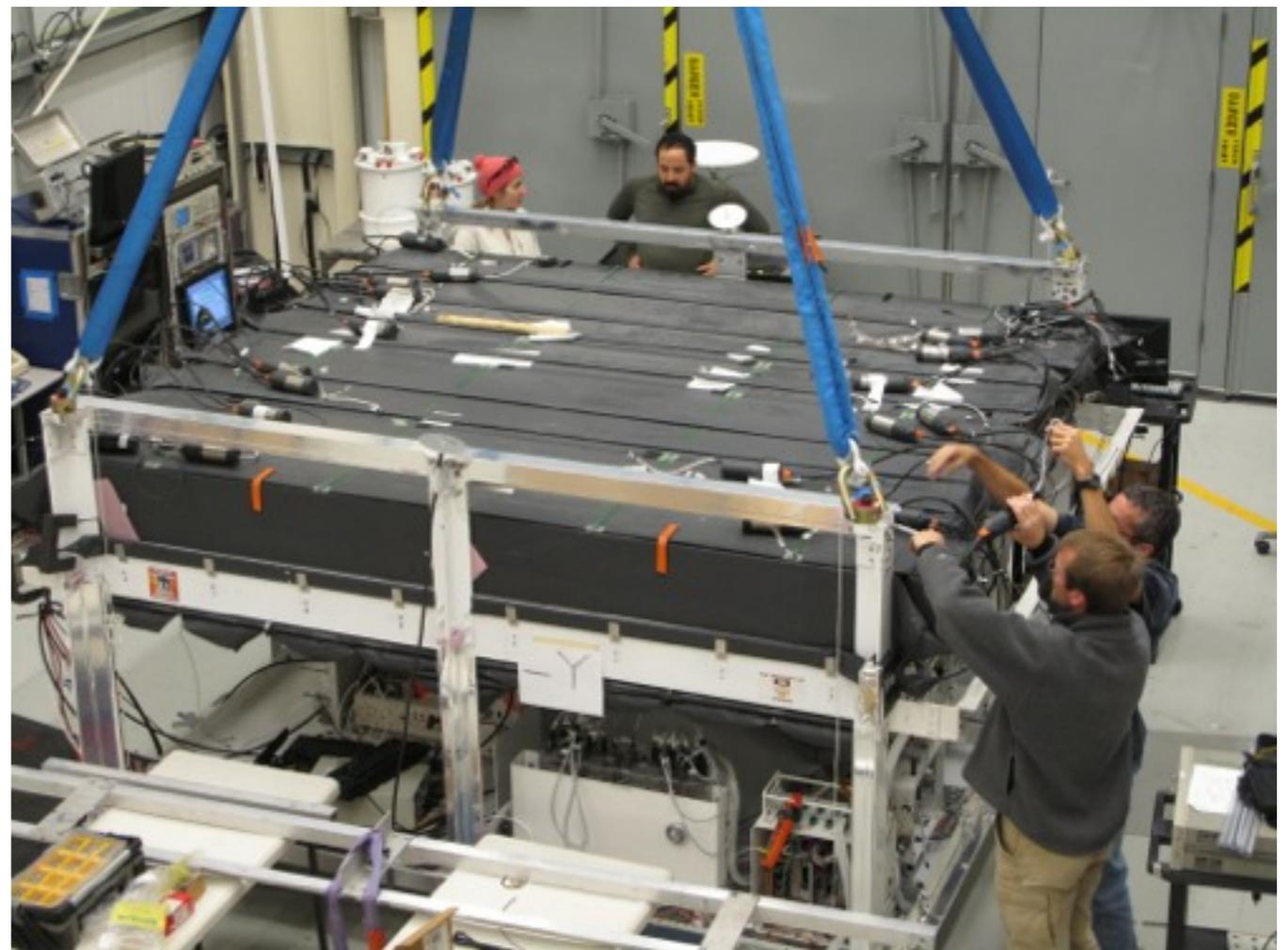
- 1024 BaF₂ crystals w/ 2" PMT readout, embedded in foam matrix
- Photon energies from ~30 keV to 30 MeV
- Designed to have 1 nsec timing resolution

● Veto paddles

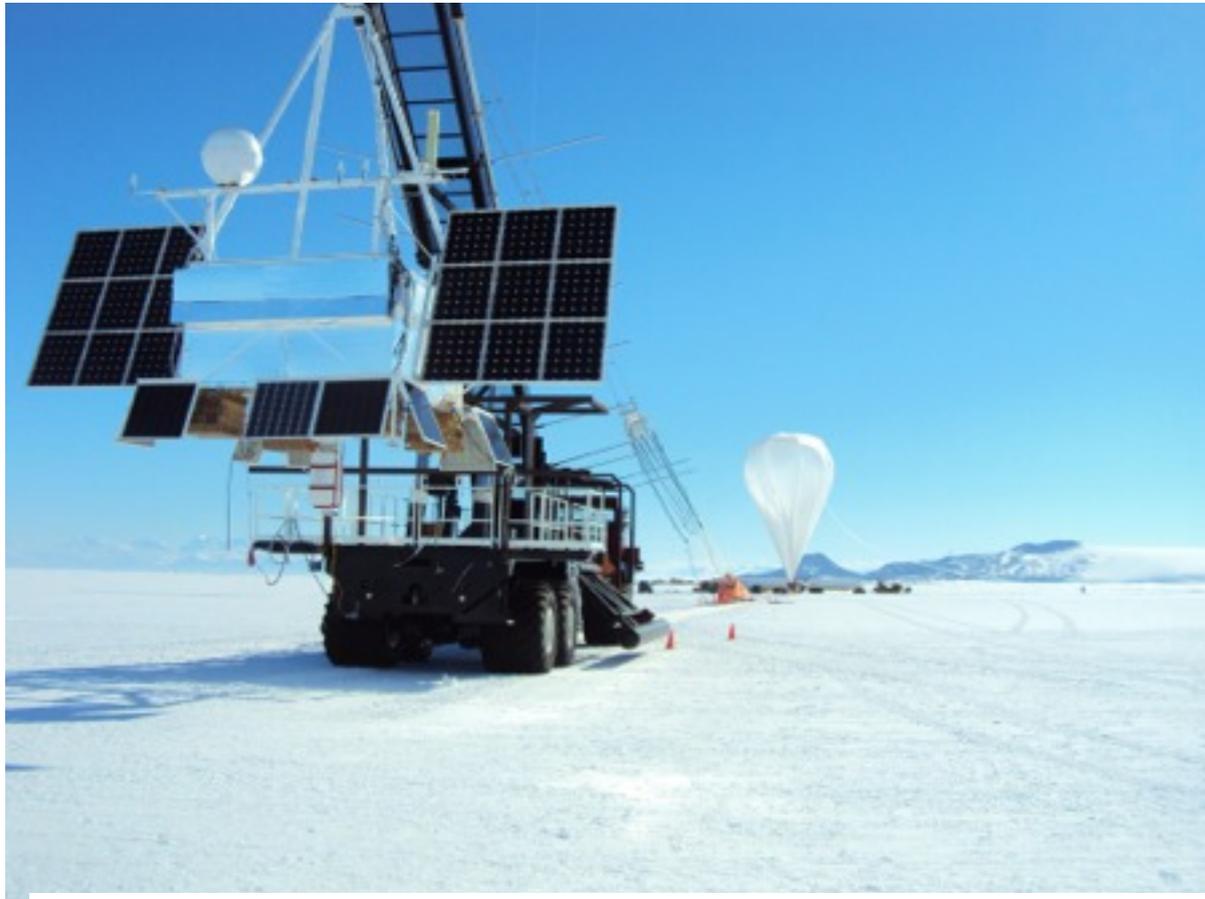
- > 99% hermetic
- Thin plastic scintillator with waveshifting fiber readout into 2" PMTs

● Expected Performance

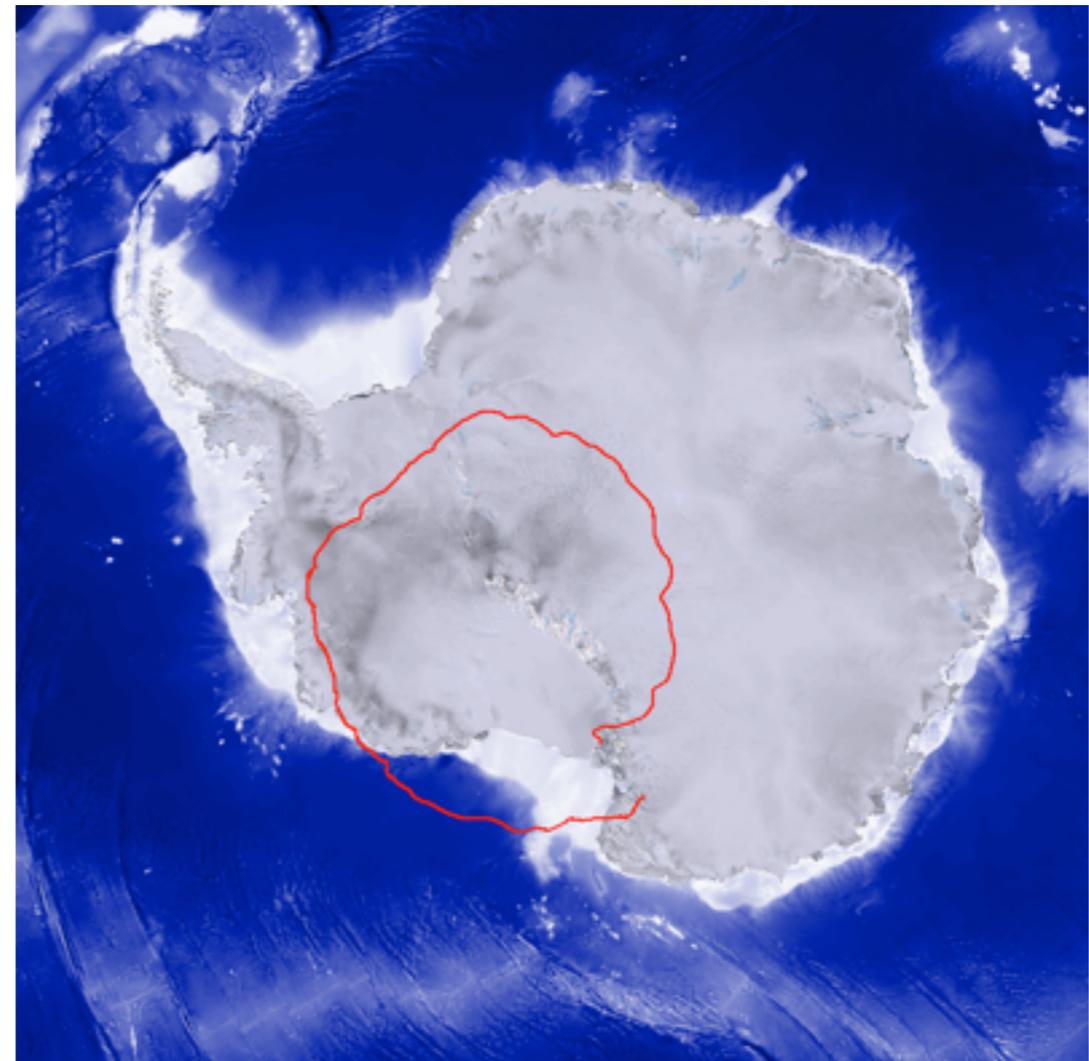
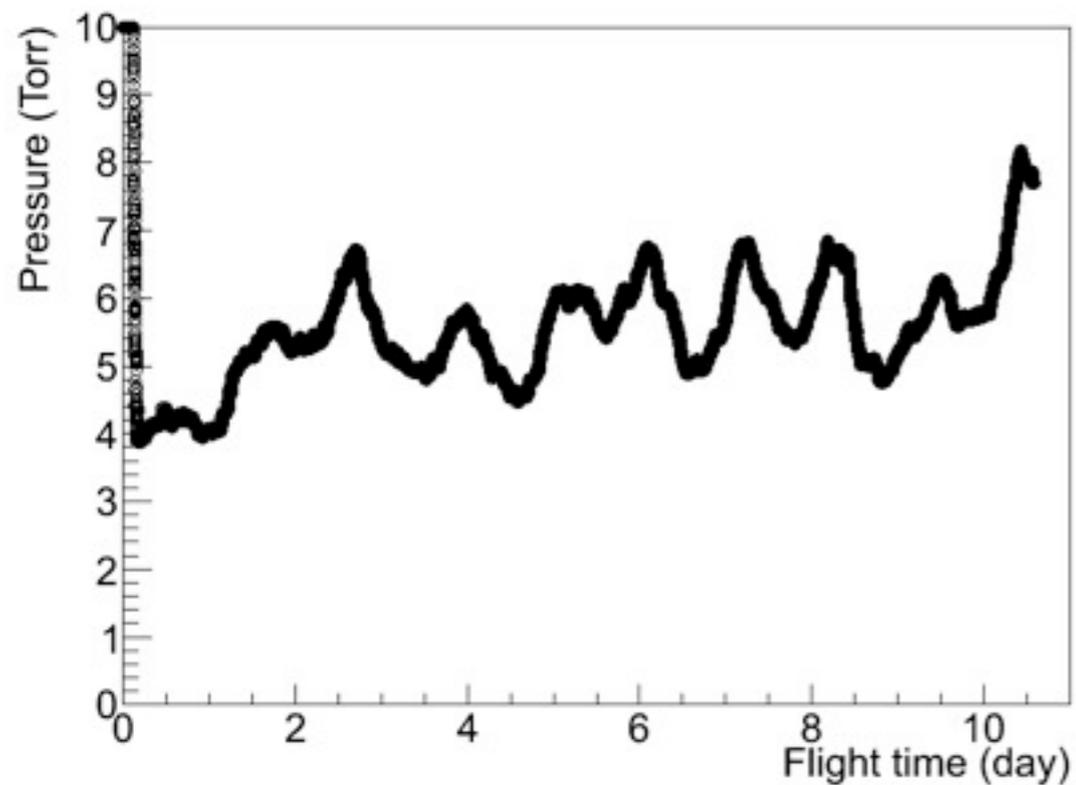
- Sensitivity on synchrotron coming from electron up to ~ 50 TeV or so



Antarctica Flight



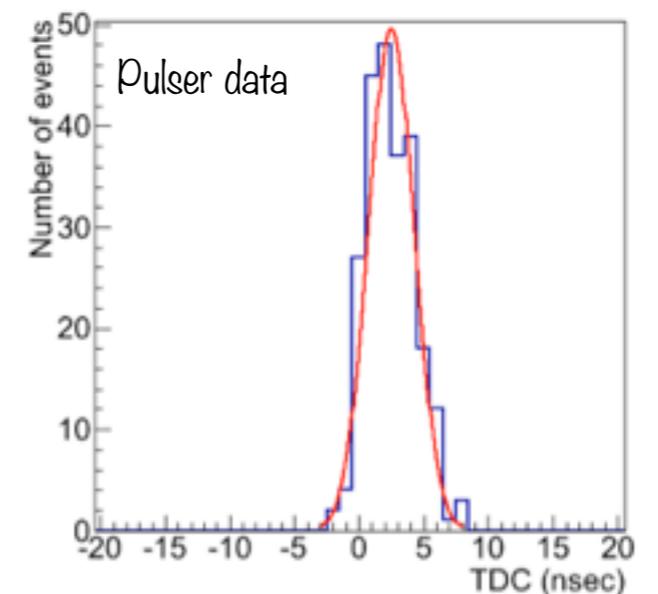
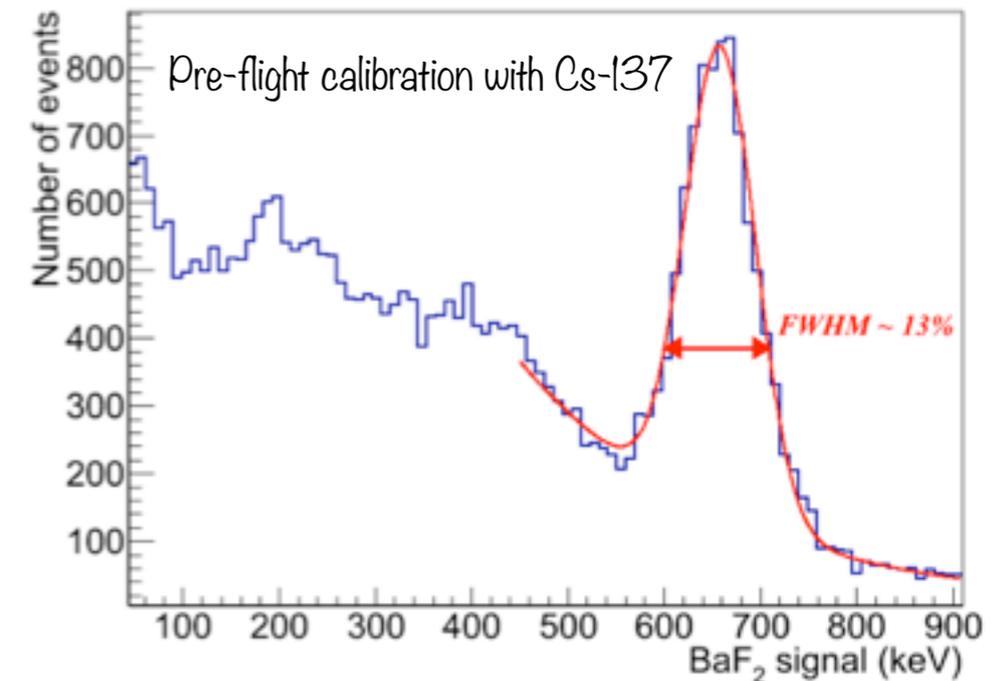
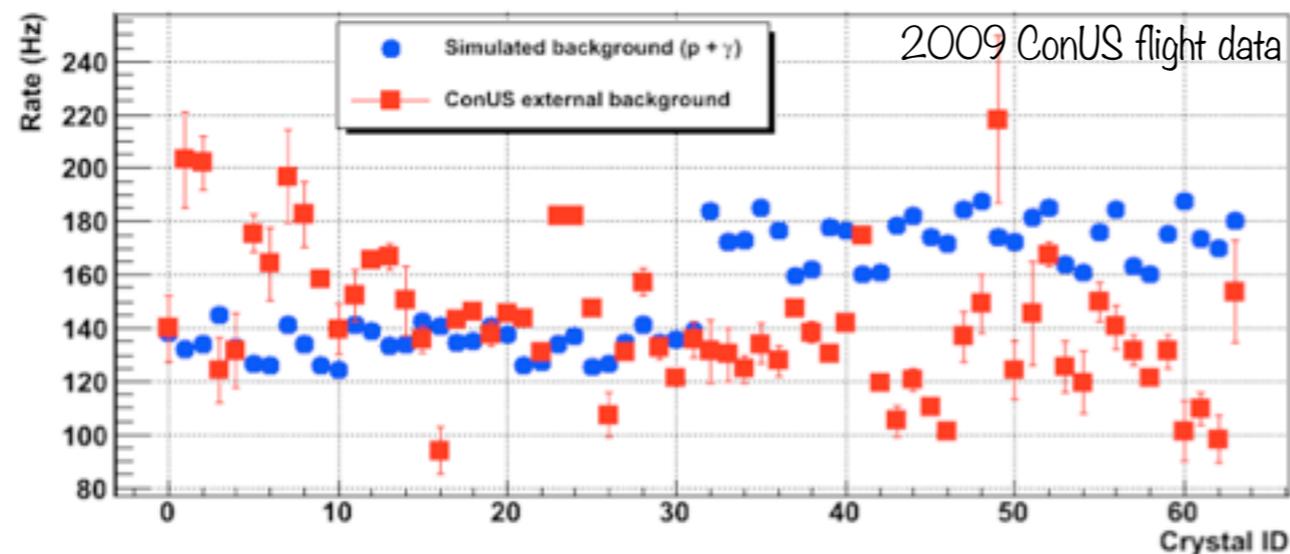
- Antarctica flight in 2011/12 season
 - Launch on Christmas day on 2011
 - Flight time : ~ 10 days
 - Recovery done on Feb. 2012



Current Status

Analysis

- Flight calibration
 - Timing calibration : by using LED pulser run & adjacent hits in crystal (calibration trigger)
 - Energy calibration : by using Radium impurities in crystal and 511 keV line
- Comparison between flight data and simulation



Investigation on hardware improvements

- Lighter detector
- Better Compton scattering shield