

# Borrowing Statistical Strength: Methods from the Great Observatories for the New Challenges of GLAST

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and

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(CHASC)

# What is the Purpose Of this Session?

- Understand statistical issues of GLAST data analysis.
- What solutions do exist?
- Start discussion and collaborations between astro- and statistics groups.



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# PAST: Chandra (AXAF) Challenges at SCMA II, 1996

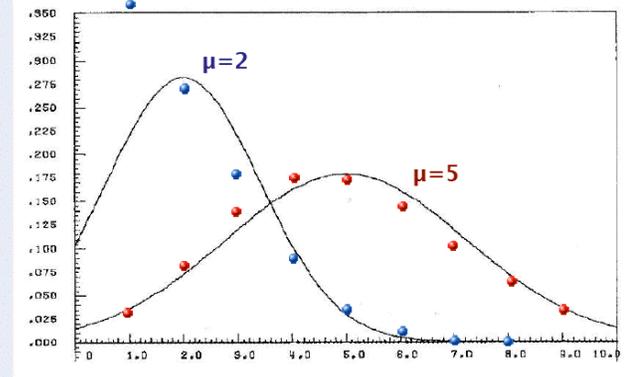
## AXAF Data Analysis Challenges

Aneta Siemiginowska<sup>1</sup>, Martin Elvis<sup>1</sup>, Alanna Connors<sup>2</sup>, Peter Freeman<sup>3</sup>, Vinay Kashyap<sup>3</sup>, and Eric Feigelson<sup>4</sup>

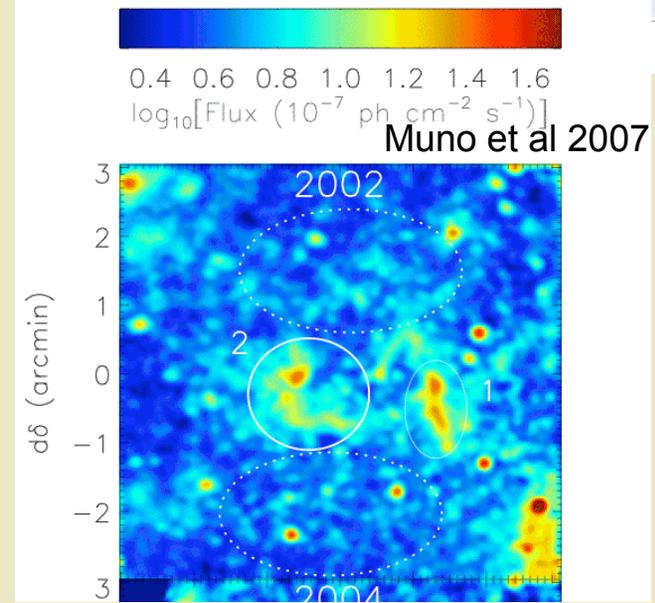
**ABSTRACT** The high quality of the AXAF X-ray data provides new challenges for the X-ray data analysis. It is clear that an “old” approach is not enough to fully exploit the capabilities of the AXAF instruments. We describe a few of the statistical and computational problems that we have so far identified. Some of them appear to be theoretically solvable

# PAST: Chandra (AXAF) Challenges at SCMA II, 1996

- 1) Modeling the data:
  - X-ray data => low counts Poisson
  - chi2 does not work - how good is my fit
  - Correlated residuals
- 2) Instrumental Issues:
  - Pileup
  - Calibration uncertainties - non-linear
  - Complex Instrumental response
- 3) Source Detection:
  - Feature significance
  - Boundaries/edges of a feature
  - Upper Bounds



Comparison of Poisson distributions (dotted) of mean  $\mu = 2$  and 5 with normal distributions of the same mean and variance (Eadie *et al.* 1971, p. 50).



# Detecting Light Echo from SgrA\*

CHANDRA X-ray Images



Aneta Siemiginowska

Credit: NASA/CXC/Caltech/M.Muno et al.

GLAST Symposium, Feb. 7, 2007

# TODAY:

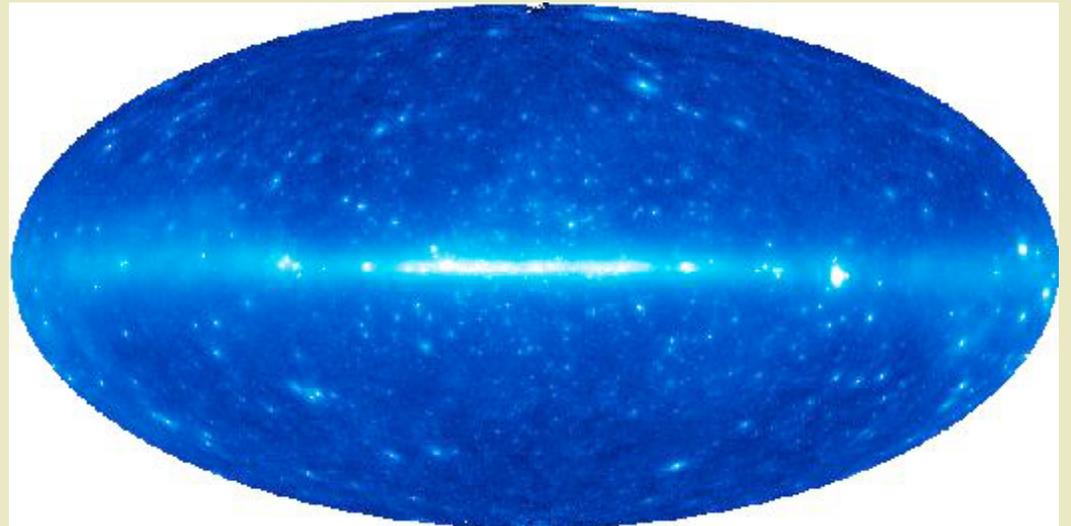
## Have we solved any Challenges?

- Data Analysis in Poisson Regime - modeling, testing for additional known component (line or point source)
- Goodness of Fit
- Correlated Residual
- Calibration uncertainties
- Feature Detection
- References:
  - Connors & van Dyk at SCMA 2006 "How to win with non-gaussian data"
  - Park et al 2006, ApJ: "Bayesian Estimation of Hardness Ratios"
  - Drake et al 2006 SPIE: Monte Carlo Processes for Including Chandra Response uncertainties.
  - Protassov et al 2002 ApJ, "Statistics Handle with Care"
  - Van Dyk et al 2001 ApJ, "Analysis of Energy Spectra with Low Counts"
- CHASC Web page:
  - <http://hea-www.harvard.edu/AstroStat/>

## FUTURE:

# Do we have solutions for GLAST?

- GLAST Challenges
  - Gamma-Rays => Poisson data
  - High background
  - Complex Instrumental Response
- Detection issues
- Confidence Bounds
- Hypothesis Testing



# Web Resources:

- CHASC Web page:
  - <http://hea-www.harvard.edu/AstroStat/>
- Center for Astrostatistics, Penn State:  
<http://astrostatistics.psu.edu/statcodes/>
- Other  
<http://www.ee.duke.edu/~willett/software.html>

# Astrostatistic Session

- \* *Statistical Challenges for GLAST LAT Data - James Chiang*
- \* *Multi-Scale Image Reconstruction with Low-Count Poisson Data - David van Dyk*
- \* *Low-Count Poisson Goodness-of-Fit and Feature Detection: An EGRET Example with EMC2 - Alanna Connors*
- \* *Computing Upper Bounds for Contaminated Weak Sources: The Banff Challenge - Xiao-Li Meng*
- \* *Modern Statistical Methods for GLAST Event Analysis - Robin Morris*
- \* *Algorithms for Detection and Modeling of Sources, GRBs, Quantum Gravity and Dark Matter Signatures - Jeff Scargle*