

# Physics of gamma-ray loud AGN jets

through high cadence, multi-frequency polarization monitoring

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Ioannis Myserlis

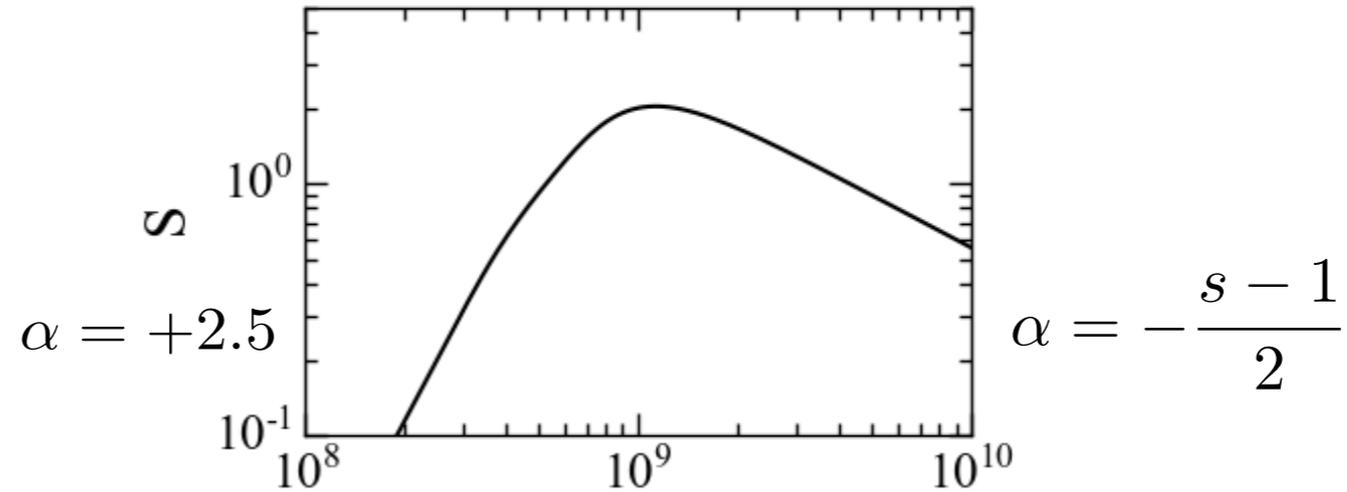
E. Angelakis, L. Fuhrmann, A. Kraus, V. Pavlidou, D. Blinov, V. Karamanavis, J. A. Zensus

# Synchrotron polarization

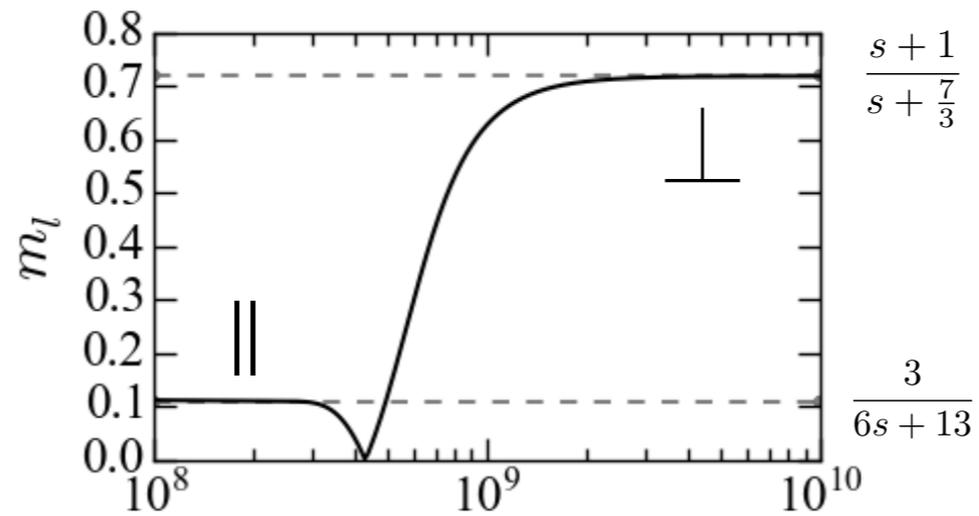
Energy distribution

$$N(E)dE = \kappa E^{-s}dE$$

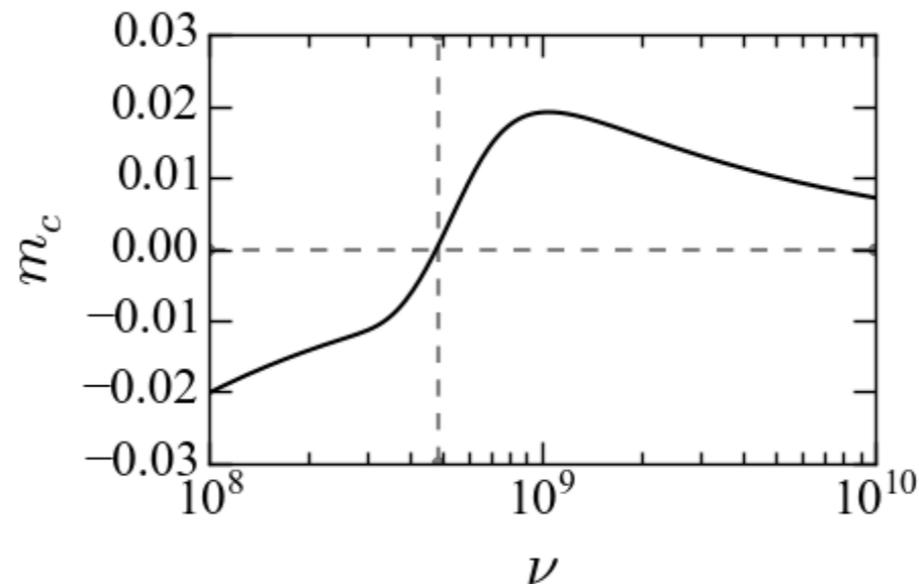
$$S(\nu) \propto \nu^\alpha$$



$\chi$  parallel to  
projected **B**-field



$\chi$  perpendicular to  
projected **B**-field



# High precision linear and circular radio polarimetry

Realistic parametrization of the telescope response

Instrumental polarization removal

- LP: Instrument model
- CP: Zero-level using unpolarized and polarized sources

Post-measurement corrections

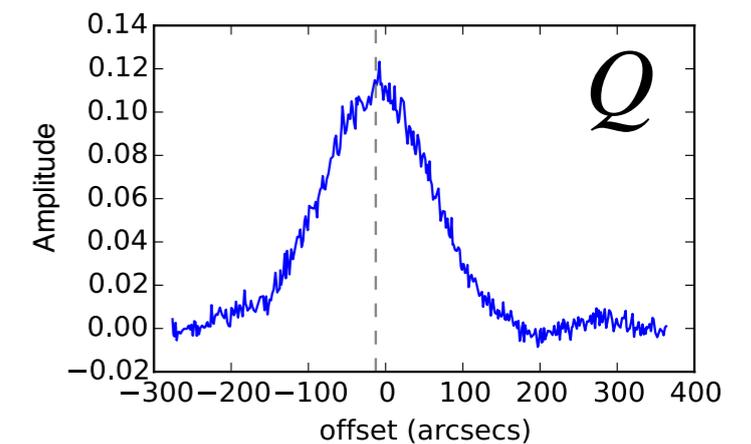
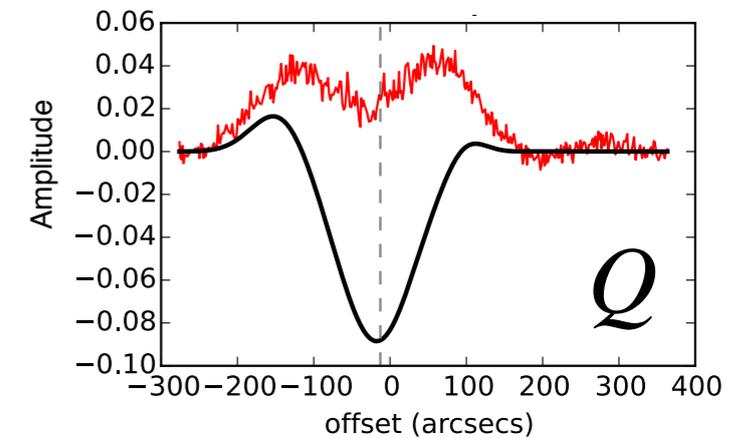
- Atmospheric opacity
- Gravitational deformations of the telescope's surface

✓ Standards Stokes parameter stability: 1-2%

✓  $m_l$  and  $m_c$  uncertainties: 0.1-0.2%

✓  $\chi$  uncertainty: 1-2°

✓ Tables with linear and circular polarization parameters of standard sources



# The radio dataset

Part of the *F-GAMMA* dataset

*Fuhrmann et al. (2007), Angelakis et al. (2008)*

87 sources

July 2010 - January 2015

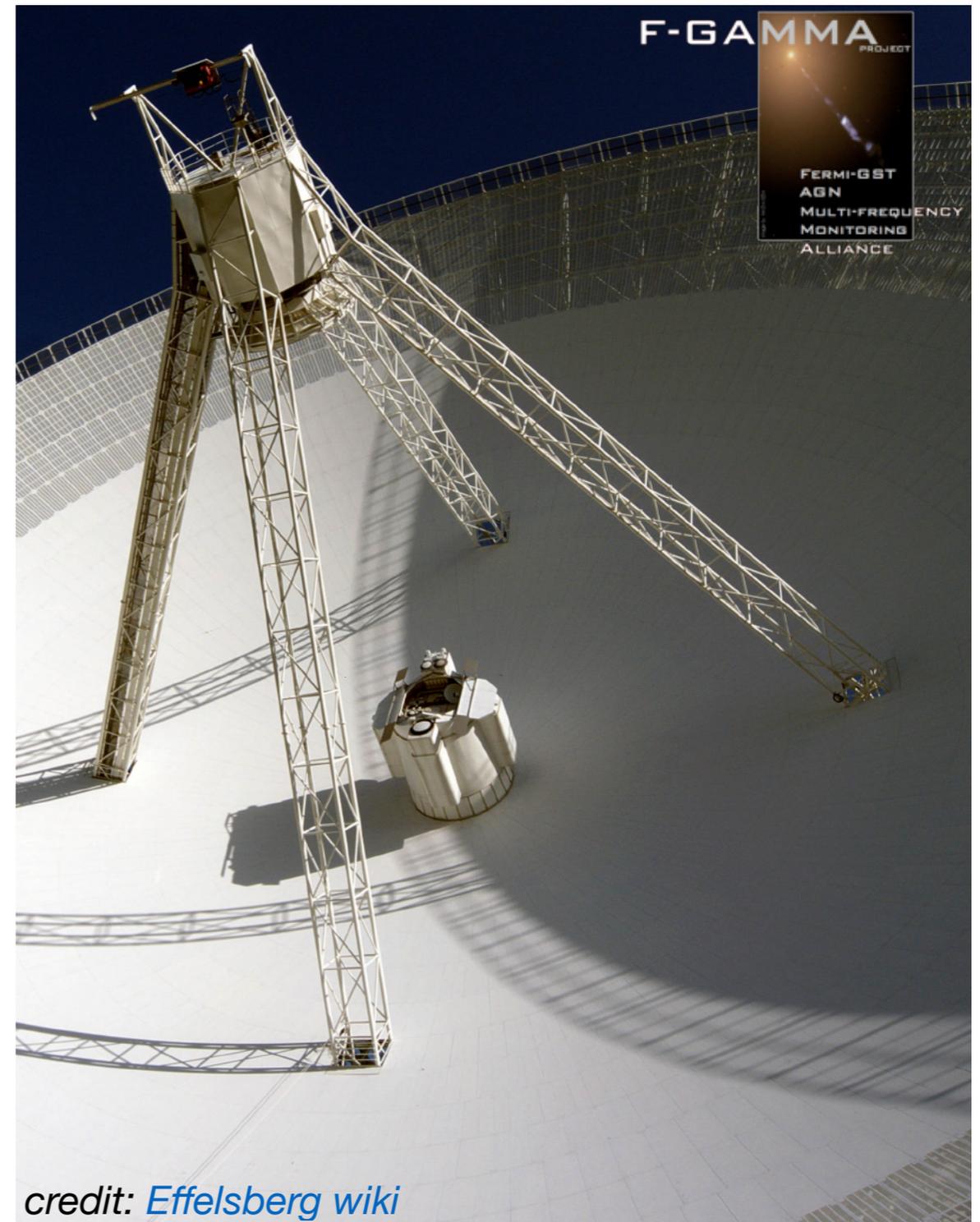
Cadence: 1.6 months

Linear polarization

2.64, 4.85, 8.35, 10.45 GHz

Circular polarization

4.85, 8.35 GHz



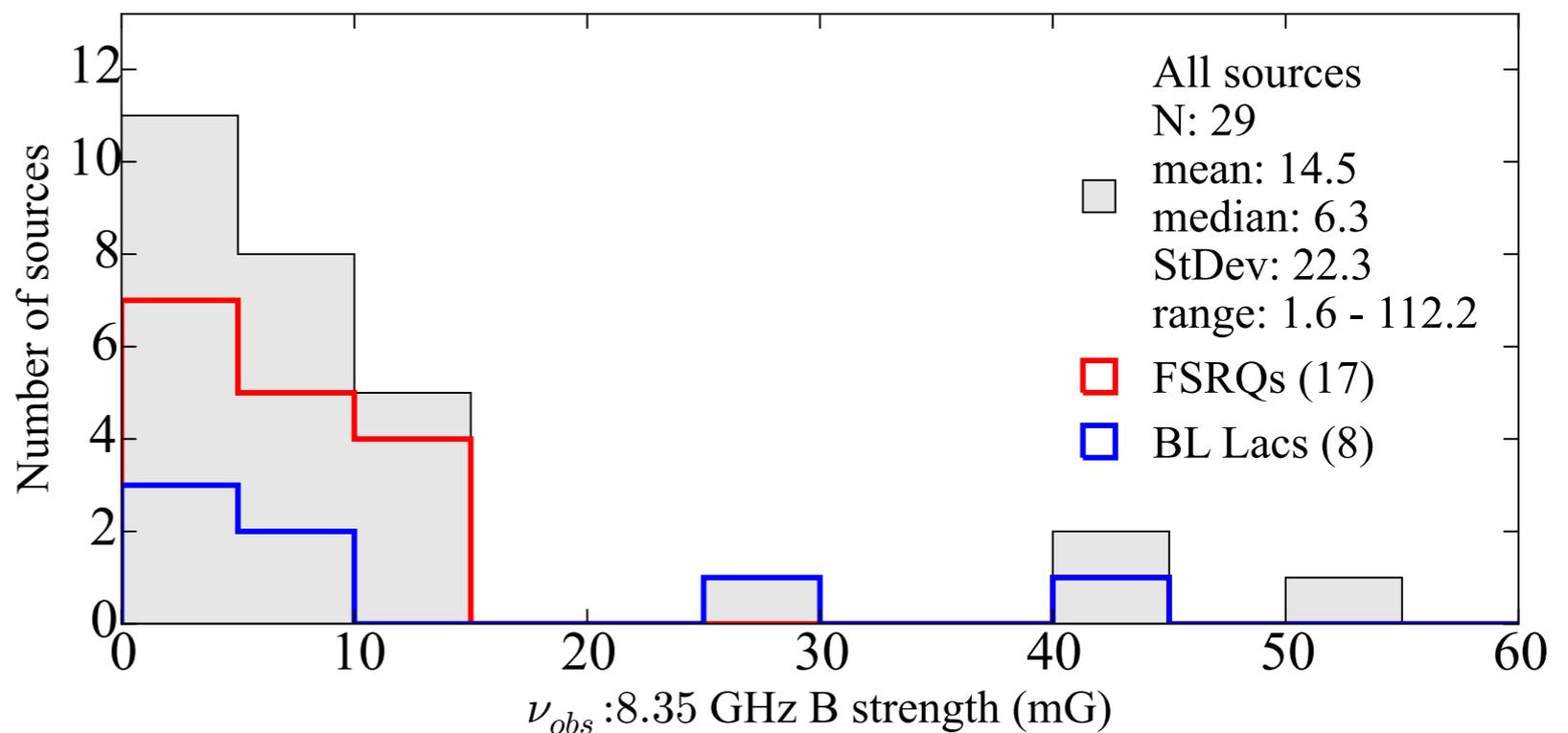
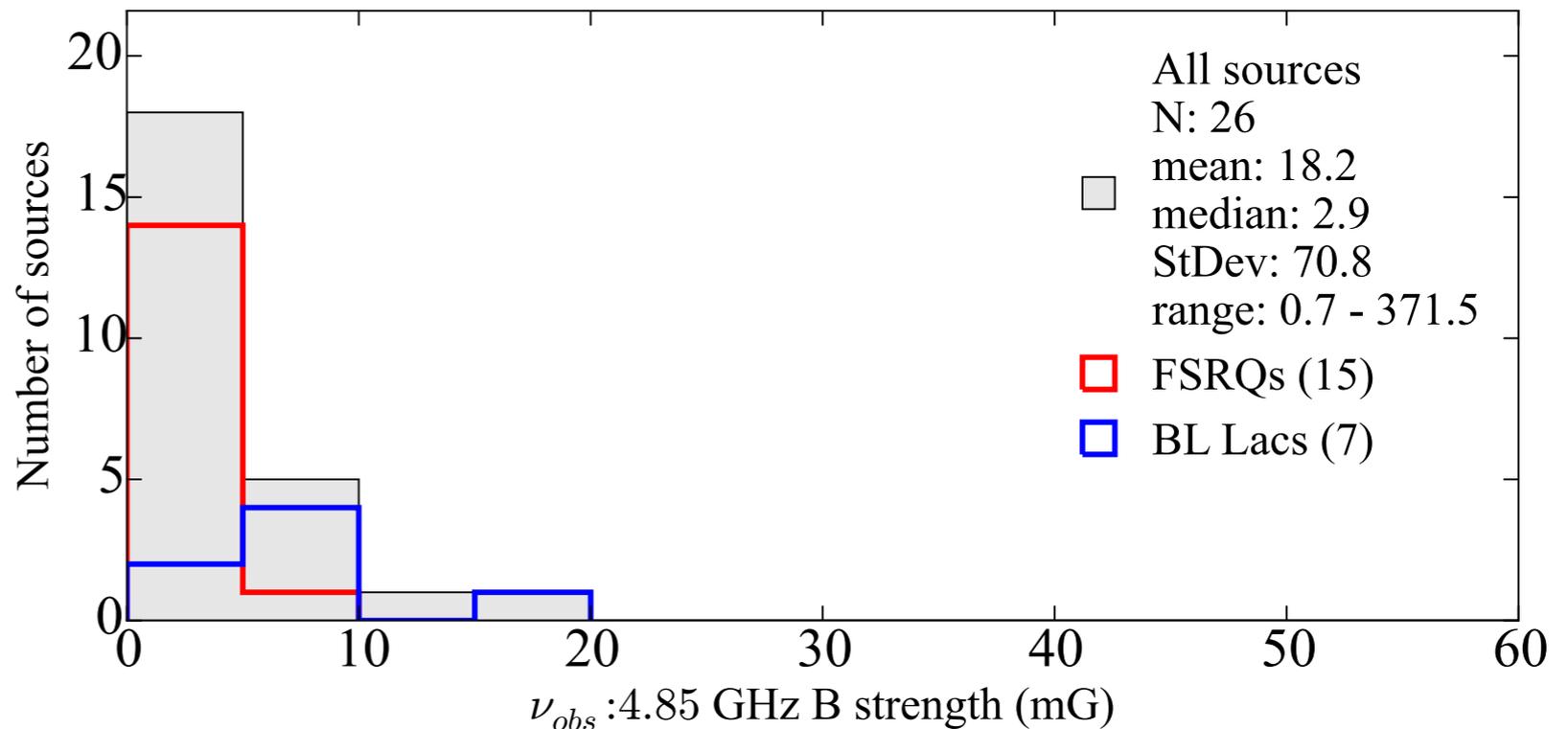
# Magnetic field strength

## Assumptions

- Synchrotron circular polarization
- One particle population, e.g.  $e^-$  or  $e^+$
- Uniform magnetic field

$$B \sim \frac{m_c^2 \nu_{\text{obs}} (1 + z)}{2.8D}$$

*Homan et al. (2009)*



# Magnetic field strength

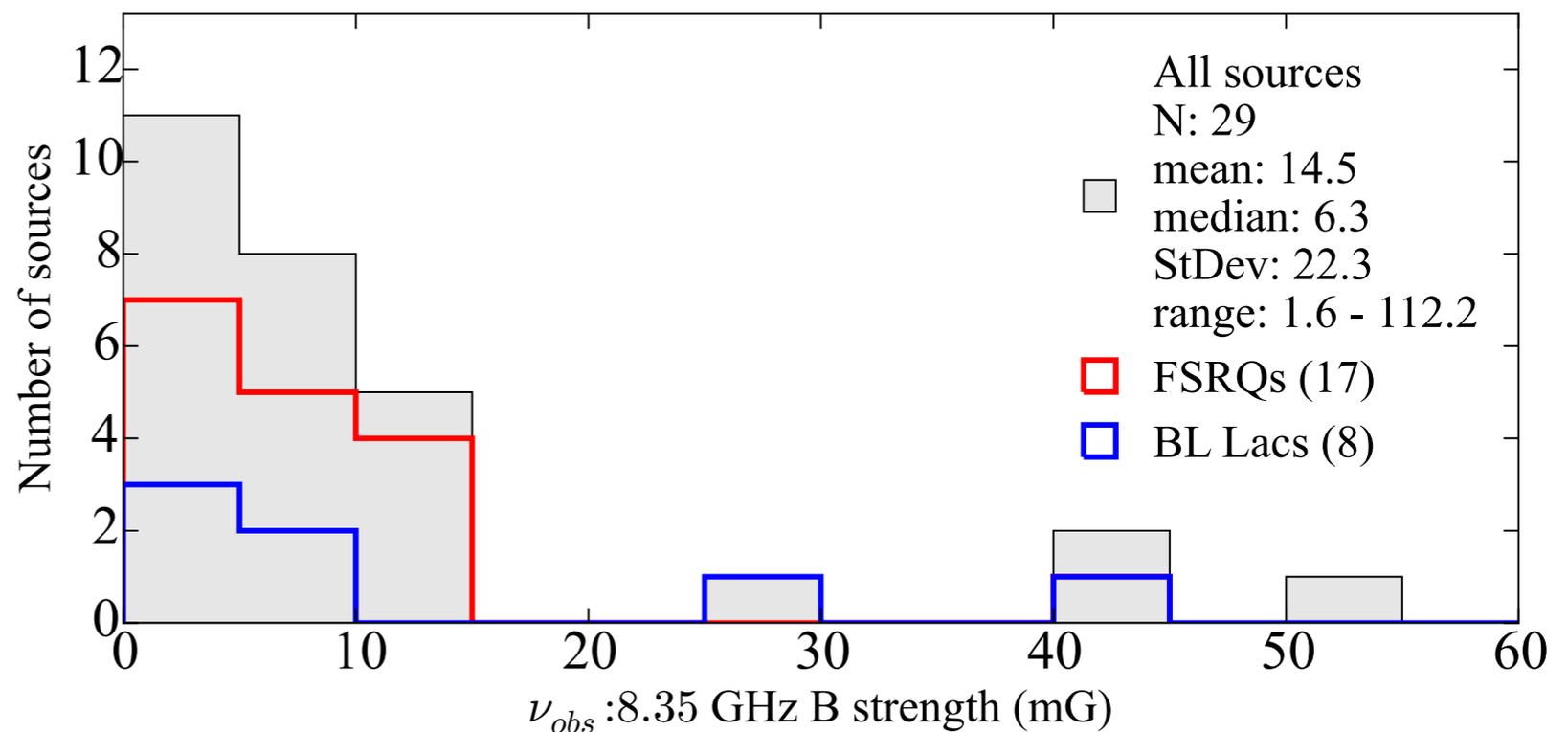
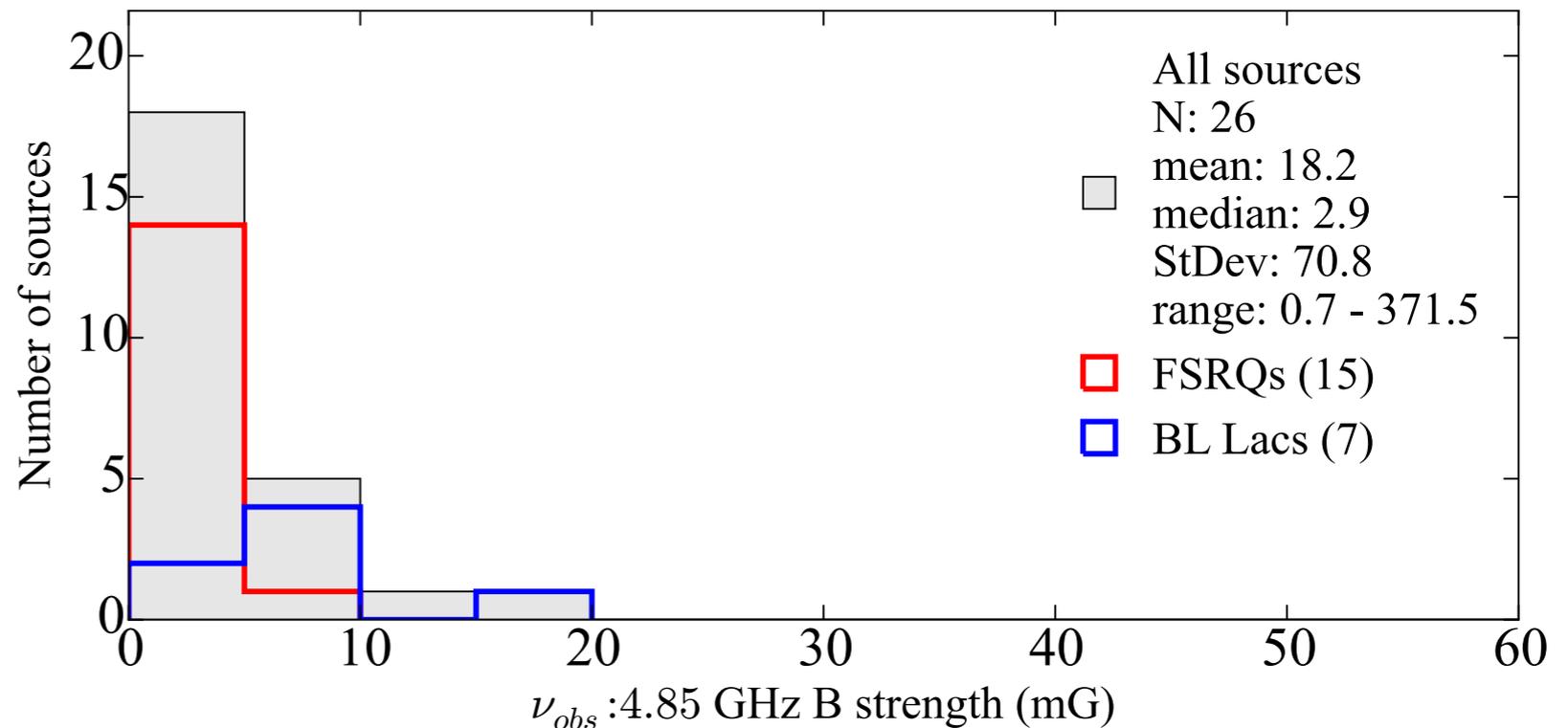
## Assumptions

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$$B \sim \frac{m_c^2 \nu_{\text{obs}} (1+z)}{2.8D}$$

*Homan et al. (2009)*

➔ Median  $B$ : 3-6 mG



# Jet plasma composition

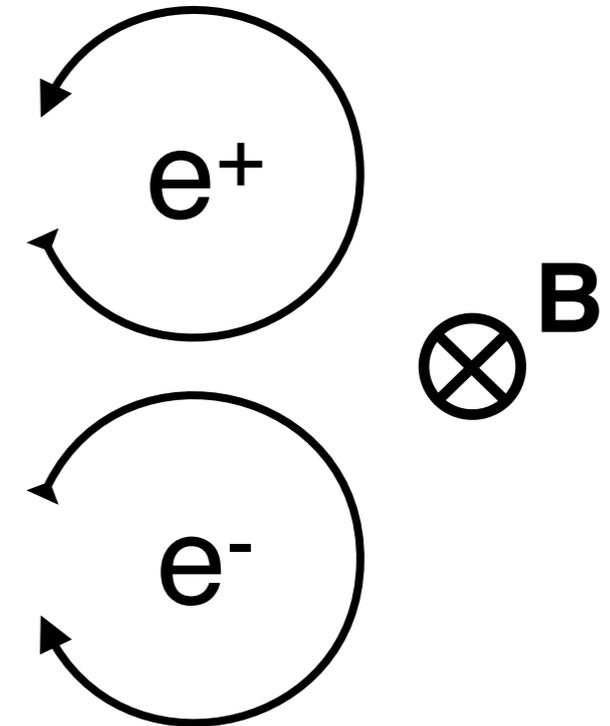
## Magnetic field strength

- Predicted: 10-100 mG  
*O'Sullivan & Gabuzda (2009a)*
- Observed: 3-6 mG

## $e^- - e^+$ admixture

Lepton number  $\ell = \frac{n_- - n_+}{n_- + n_+}$

*Homan et al. (2009)*



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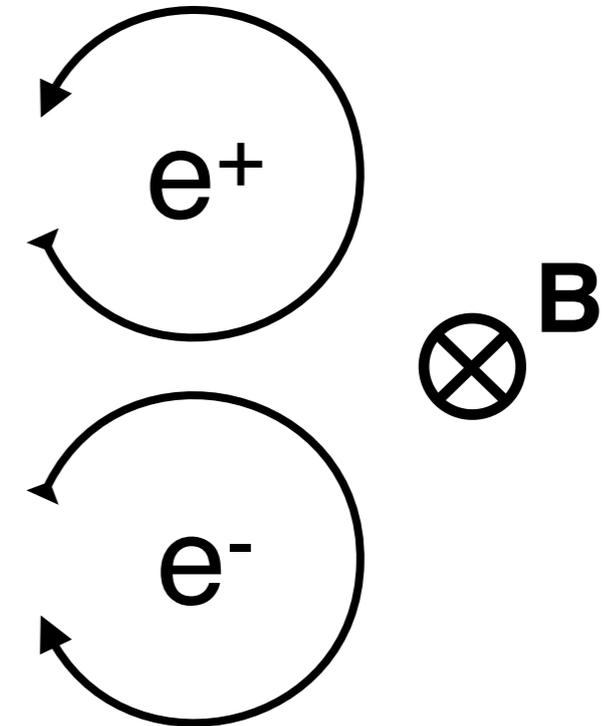
## $e^- - e^+$ admixture

Lepton number  $l = \frac{n_- - n_+}{n_- + n_+}$

*Homan et al. (2009)*

➔  $l \sim 0.2-0.5$

➔ Admixture 1:1.5-2

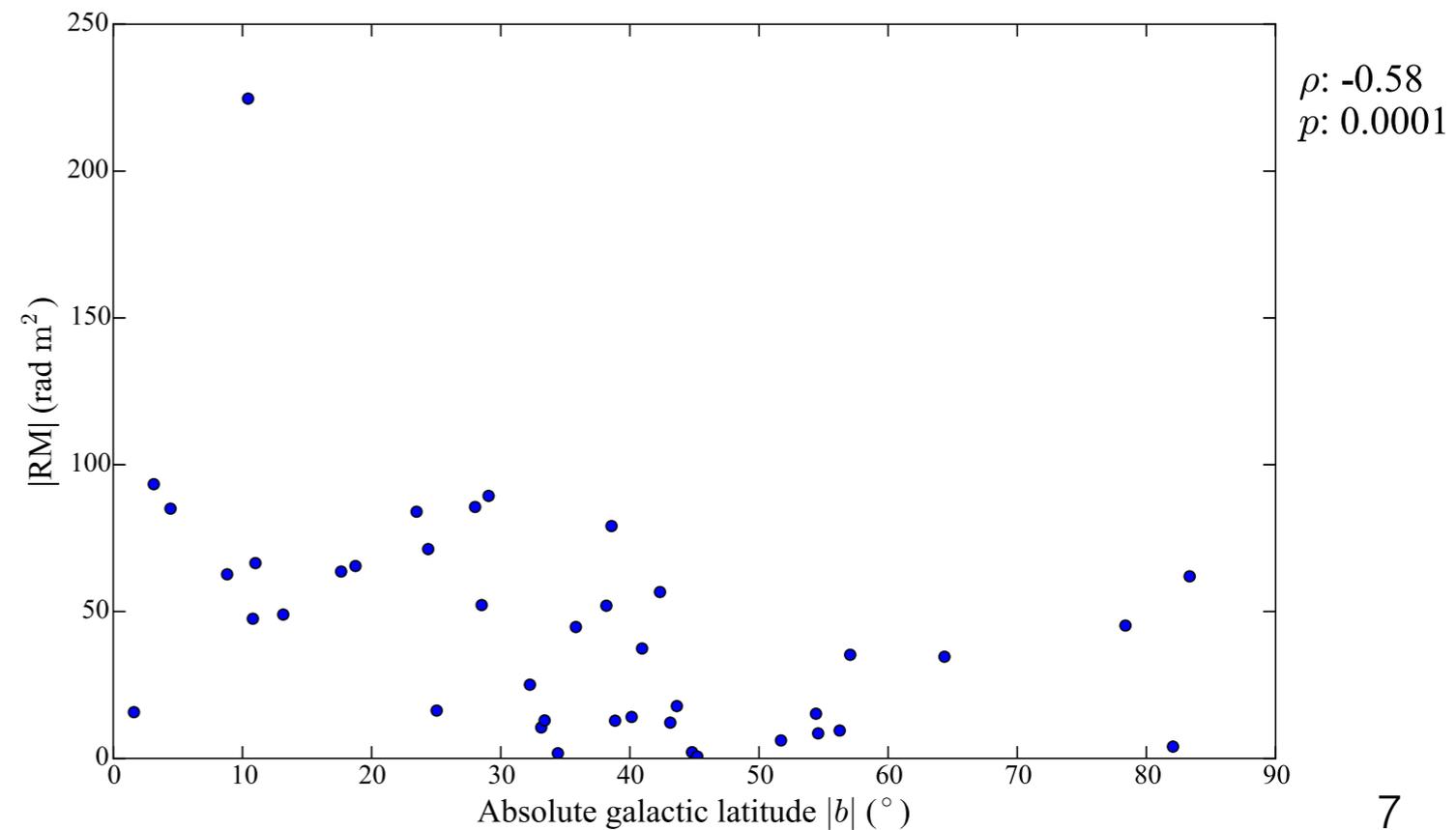
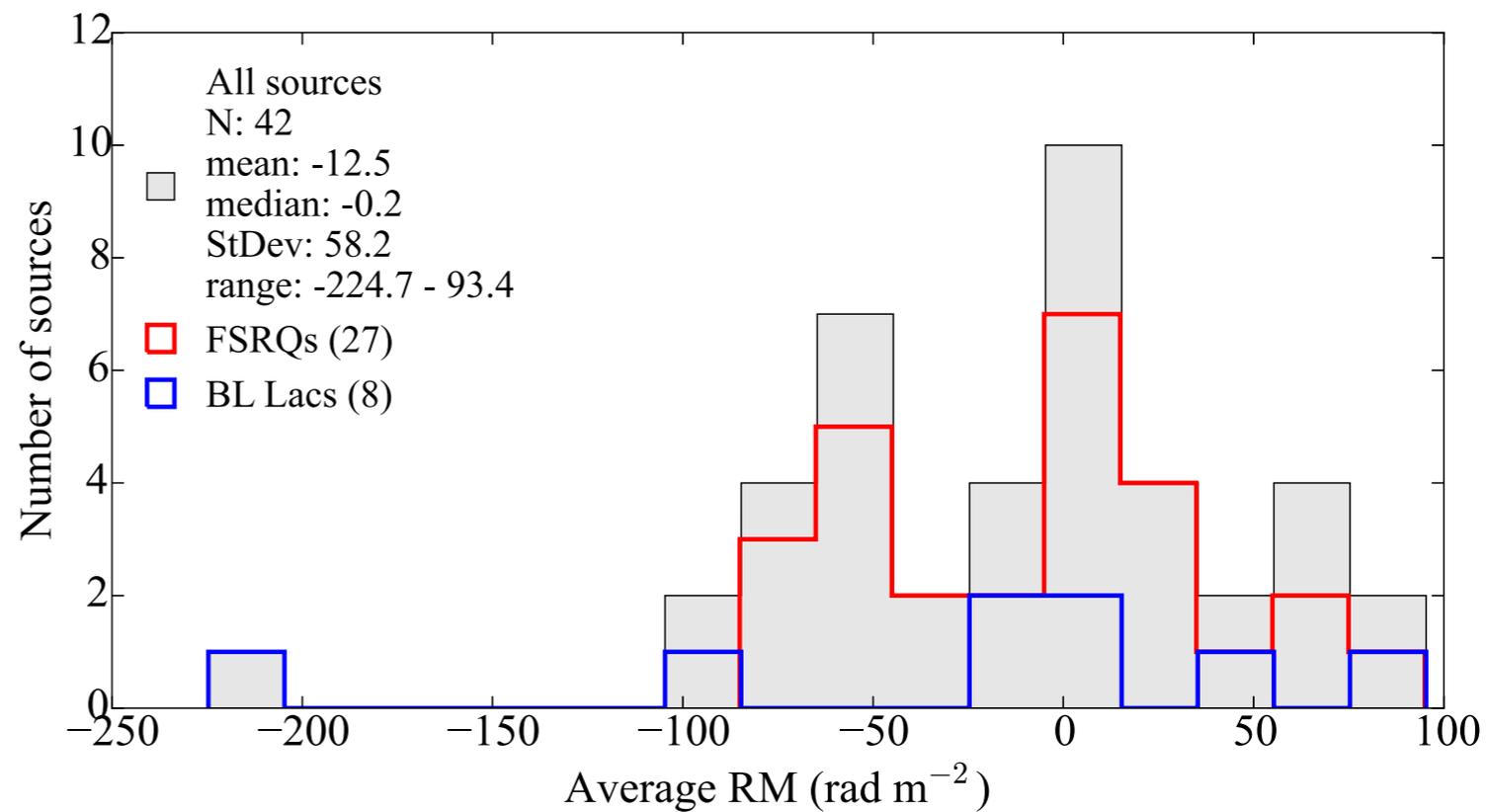
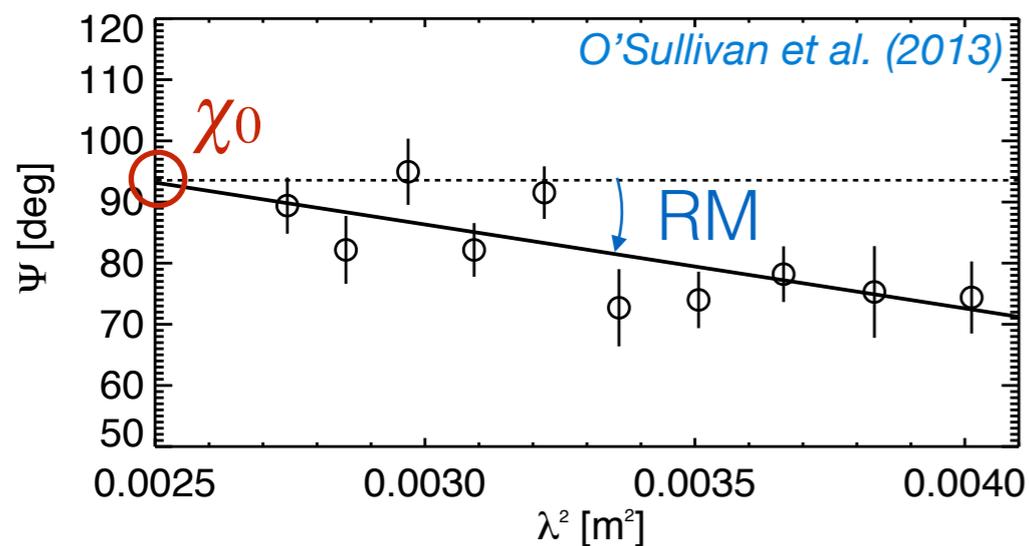


# Low energy magnetized plasma content

Faraday rotation  $\Delta\chi$

$$\Delta\chi = \text{RM}\lambda^2$$

$$\text{RM} = \frac{e^3}{2\pi m^2 c^4} \int_0^d n_e(s) B_{\parallel}(s) ds$$

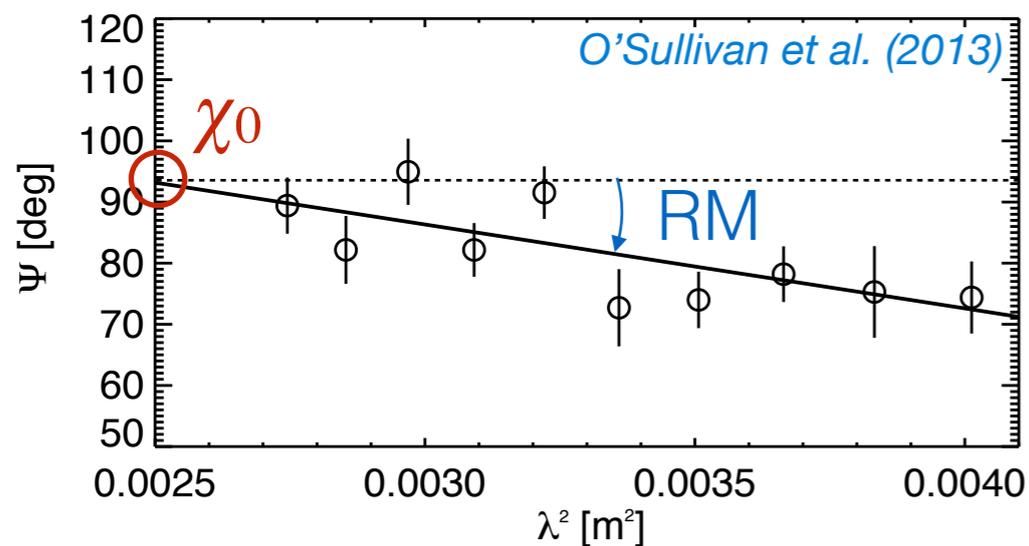


# Low energy magnetized plasma content

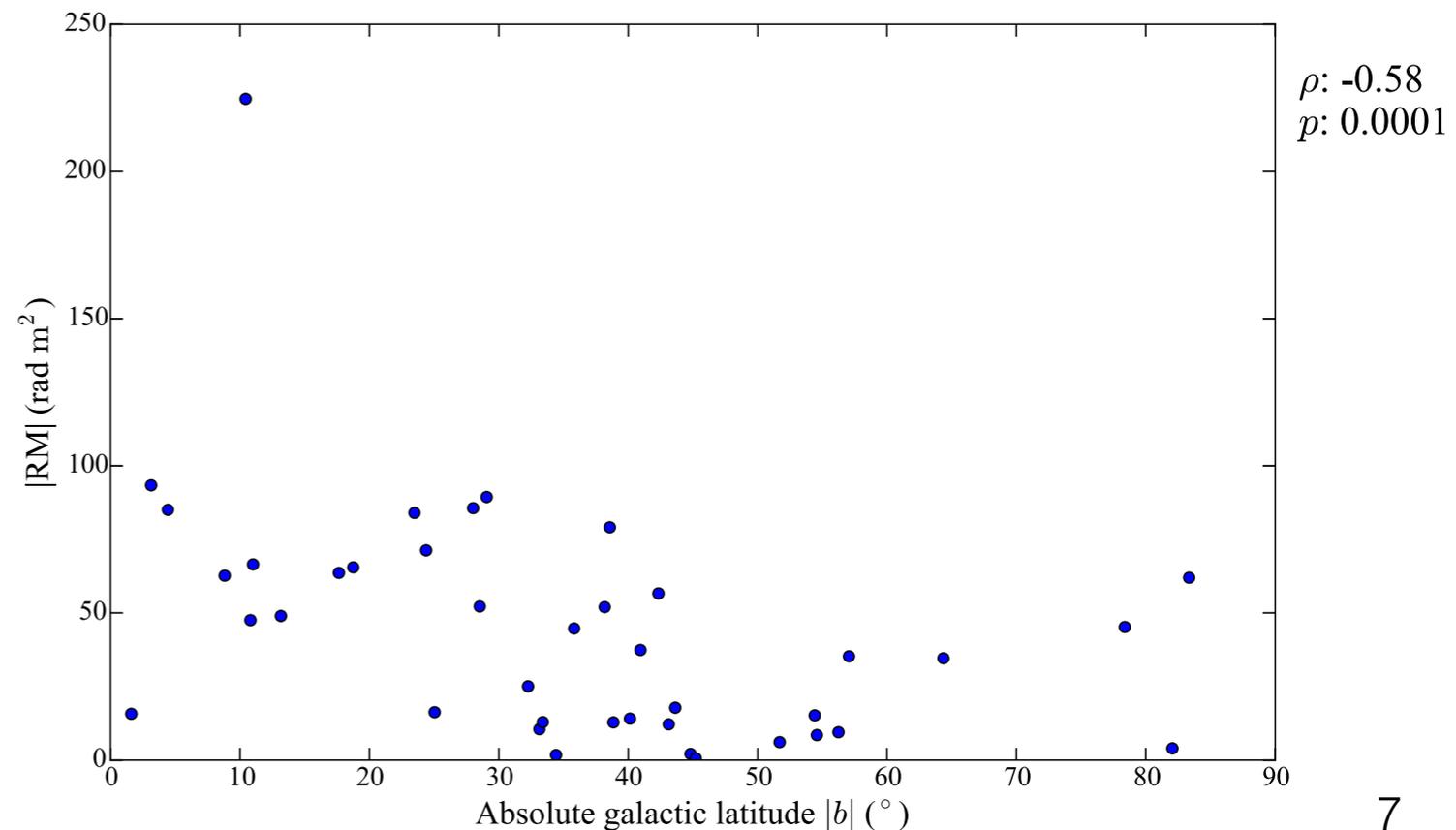
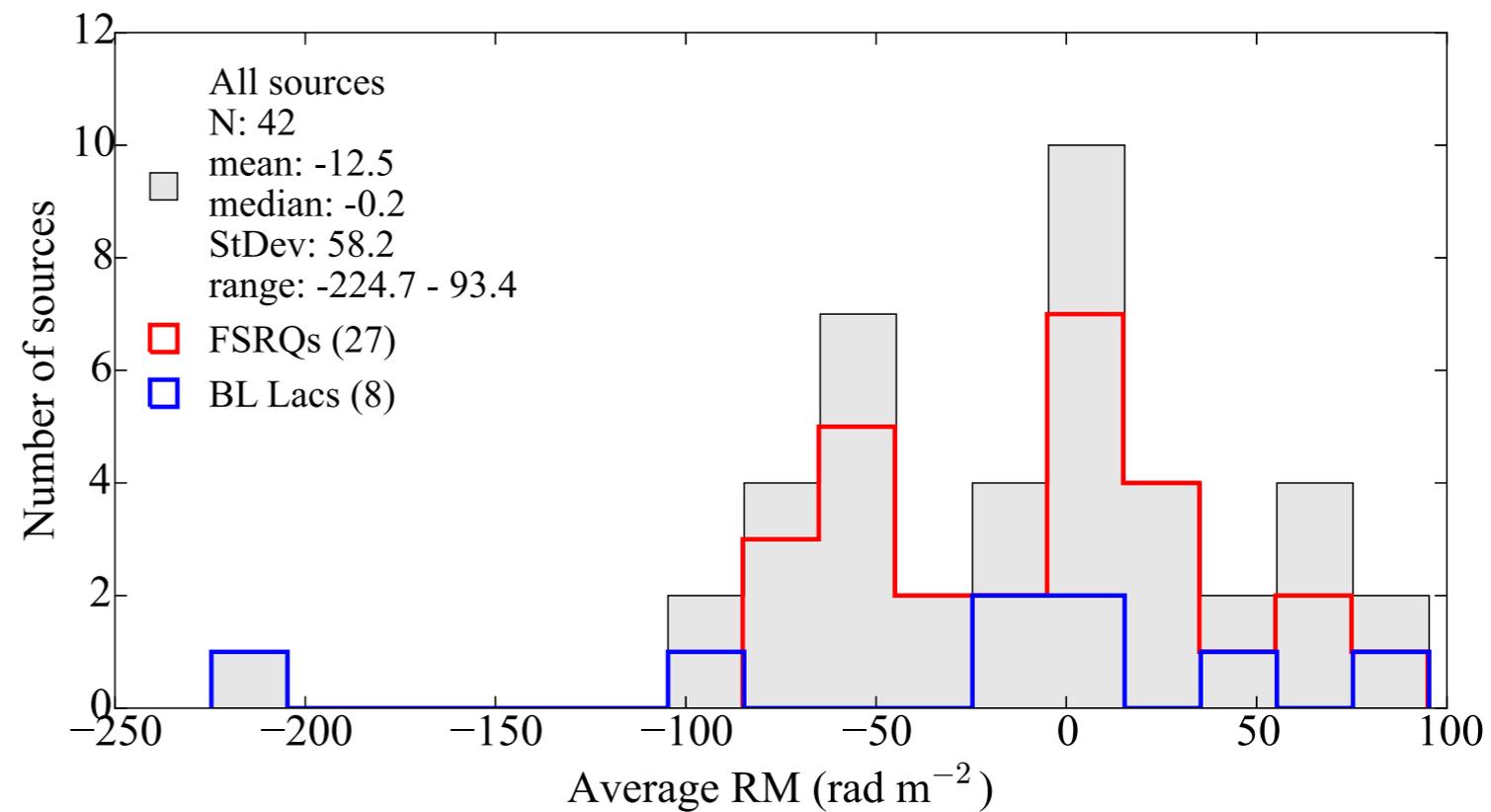
Faraday rotation  $\Delta\chi$

$$\Delta\chi = \text{RM}\lambda^2$$

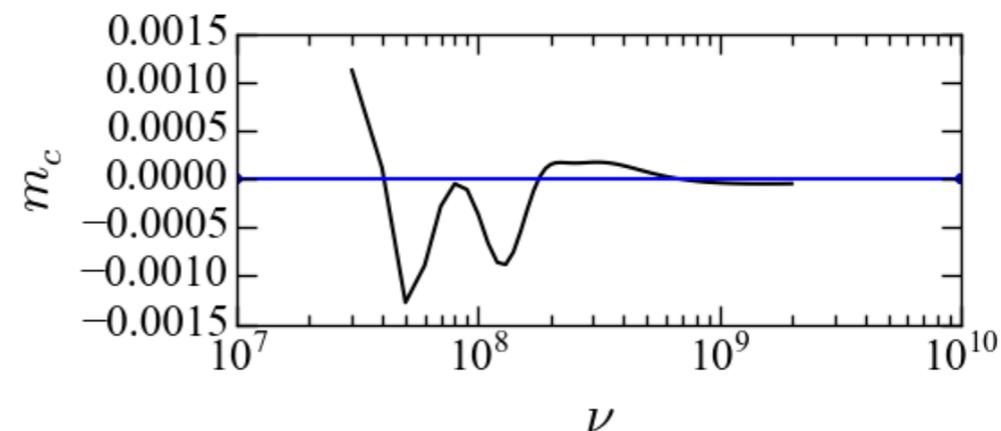
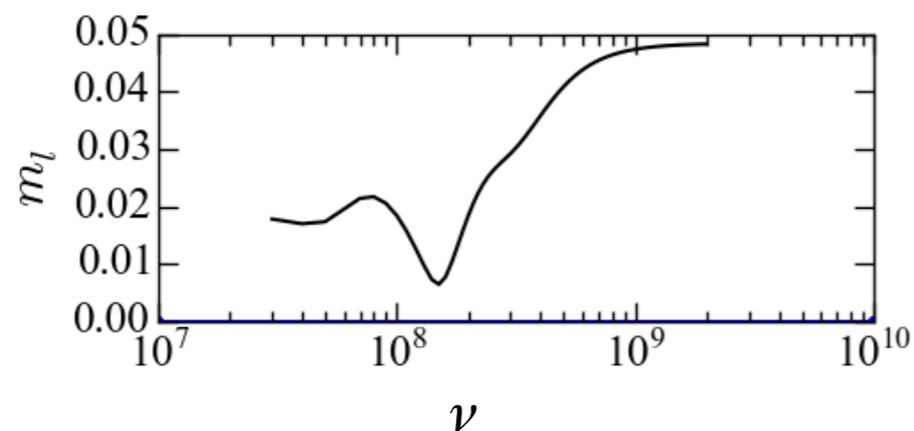
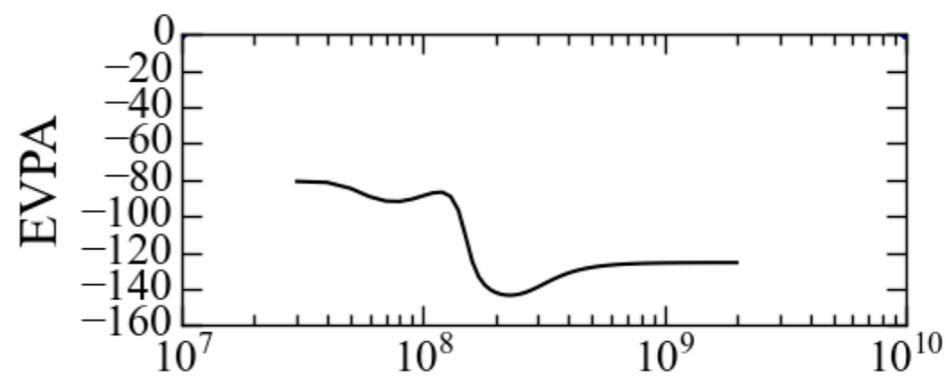
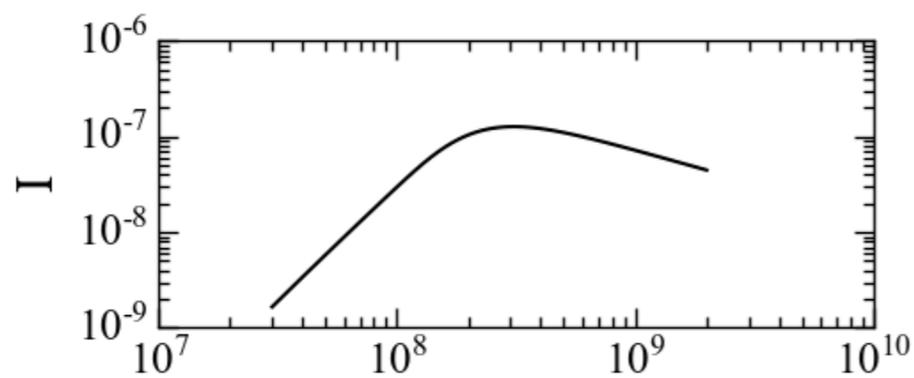
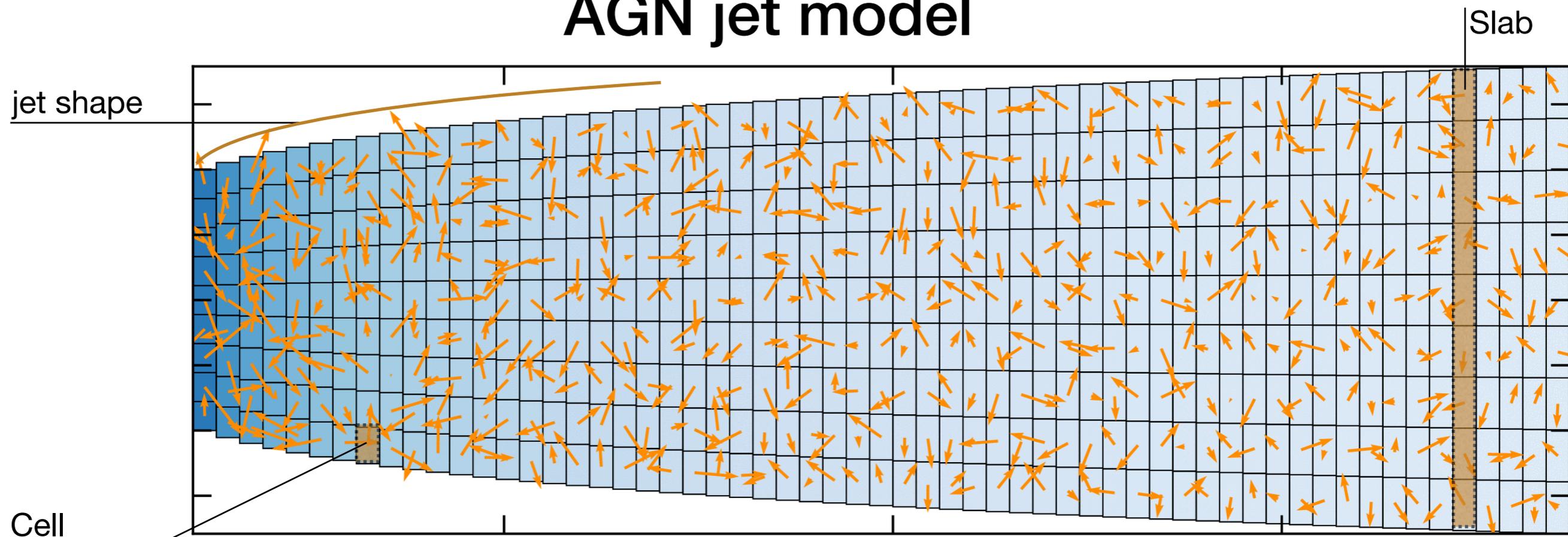
$$\text{RM} = \frac{e^3}{2\pi m^2 c^4} \int_0^d n_e(s) B_{\parallel}(s) ds$$



- ➔ Galactic origin of RM
- ➔ No low energy plasma at the source

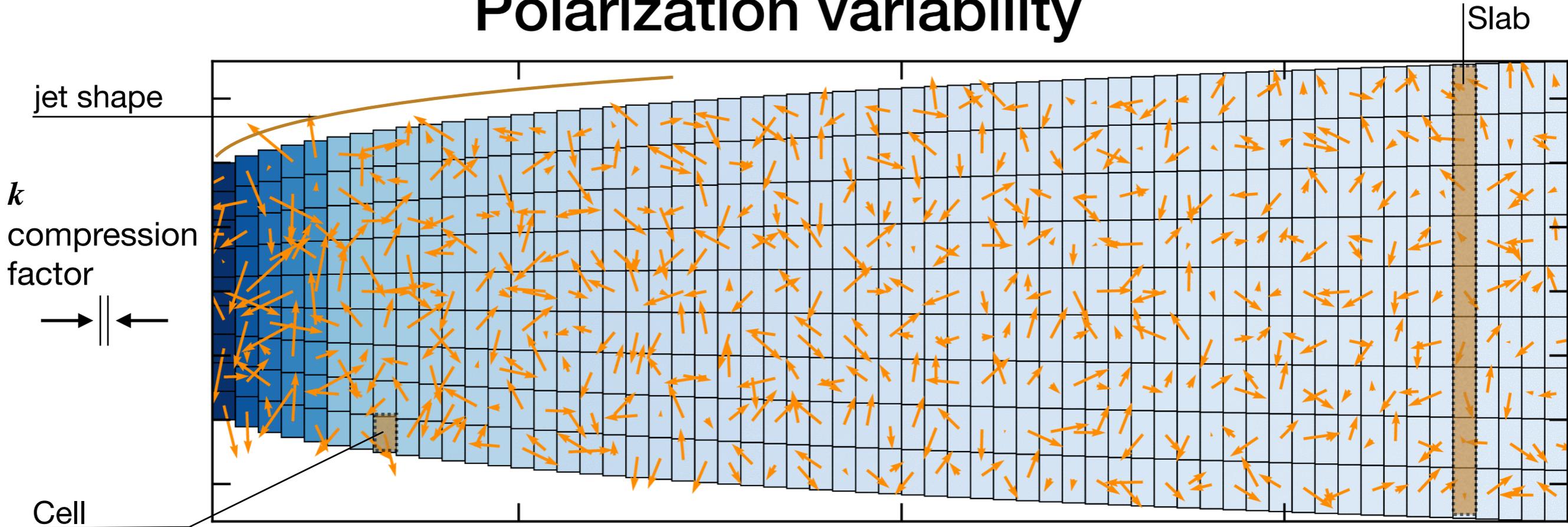


# AGN jet model



Line of sight

# Polarization variability



Density

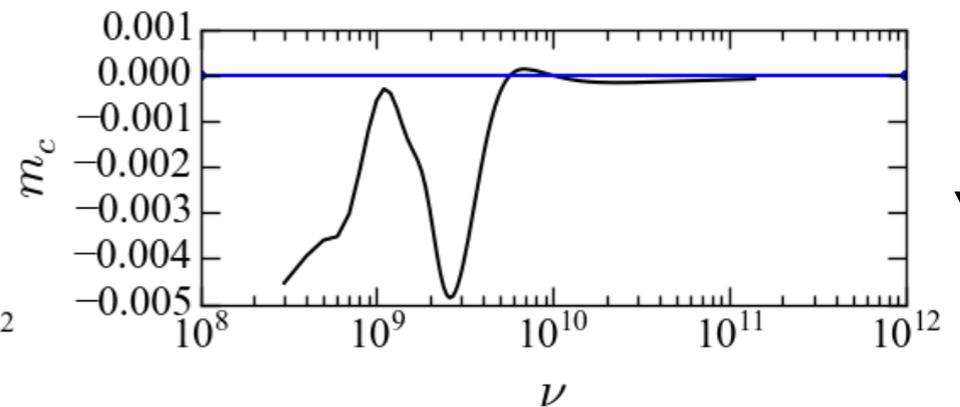
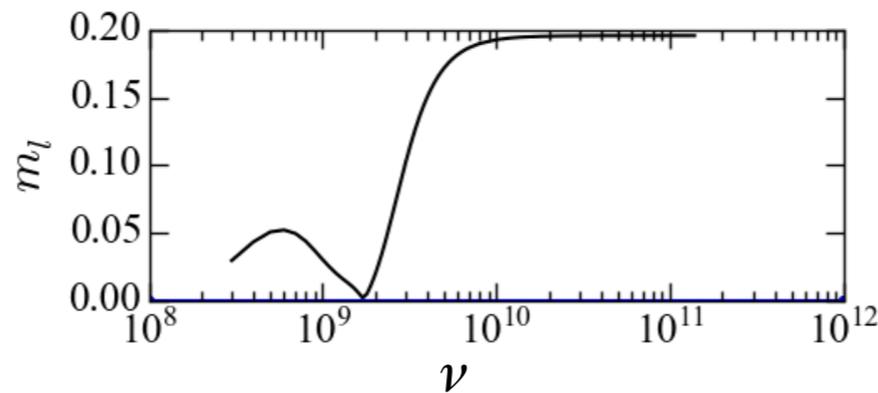
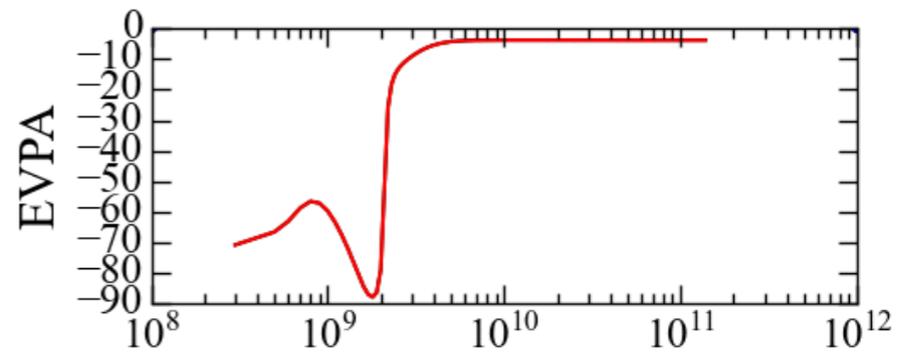
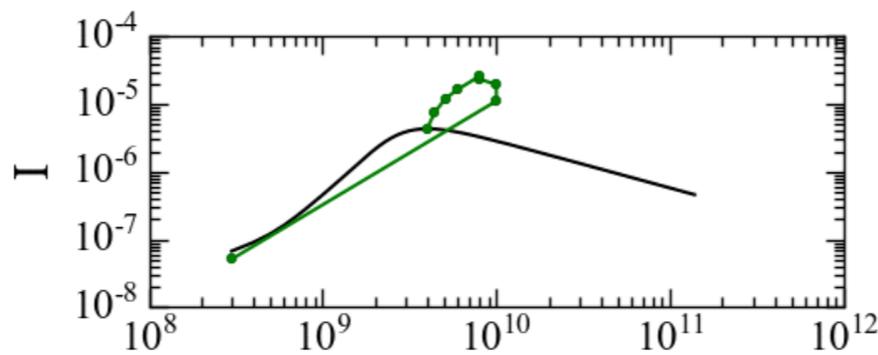
$$n'_0 = n_0 k^{-\frac{s+3}{6}}$$

Lower energy cutoff

$$E'_{\min} = E_{\min} k^{-\frac{1}{3}}$$

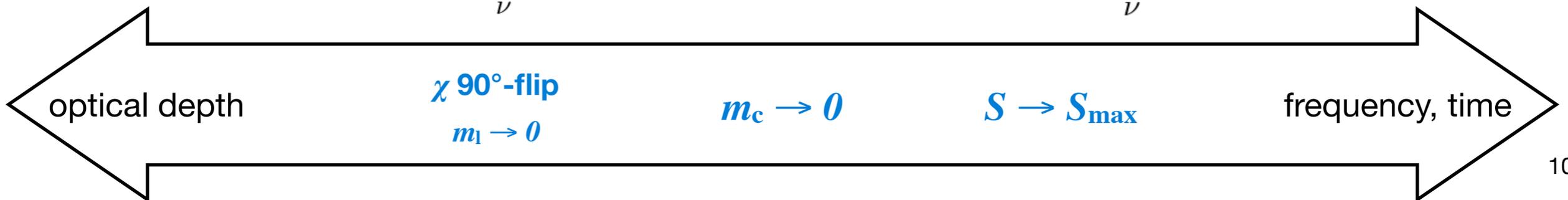
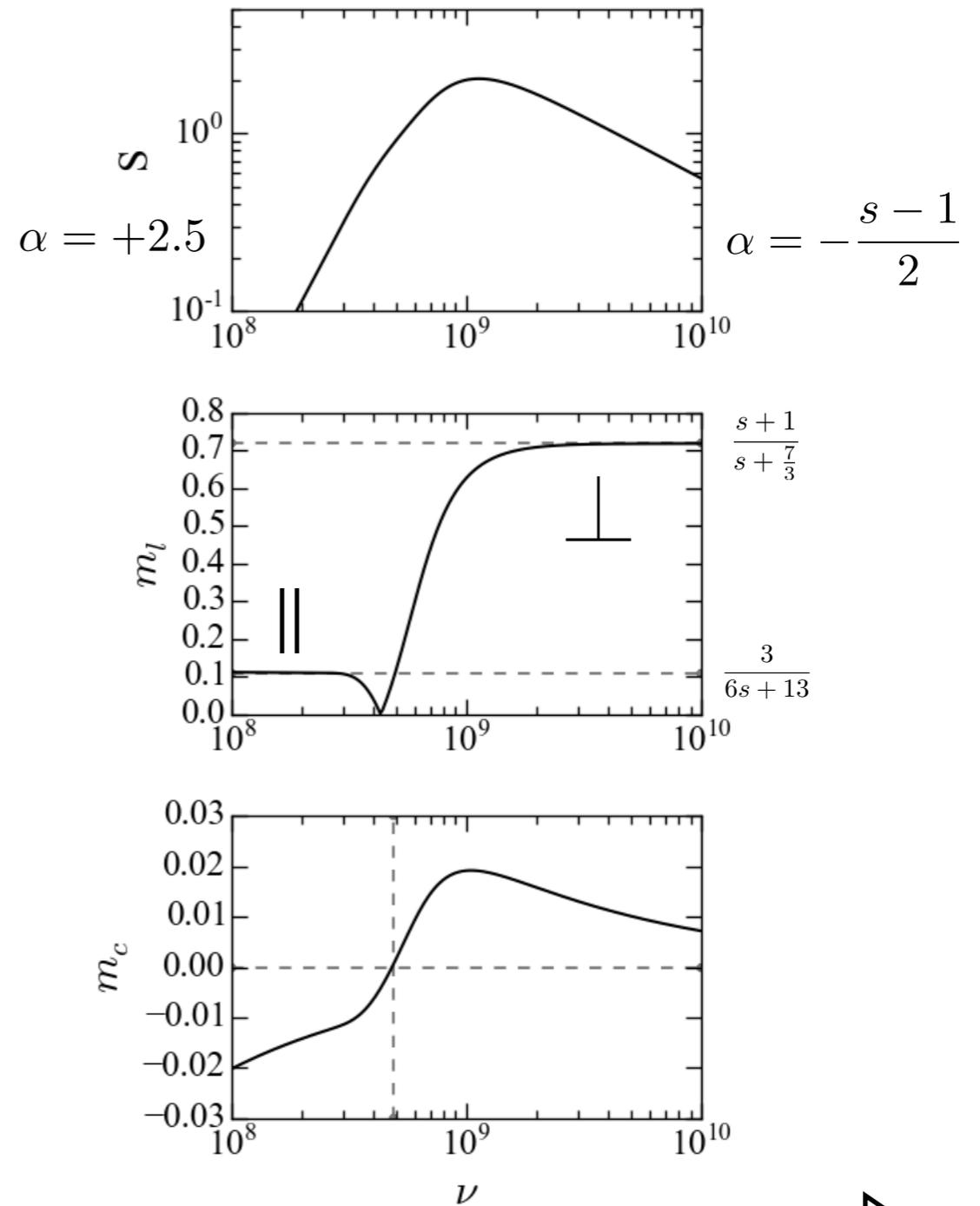
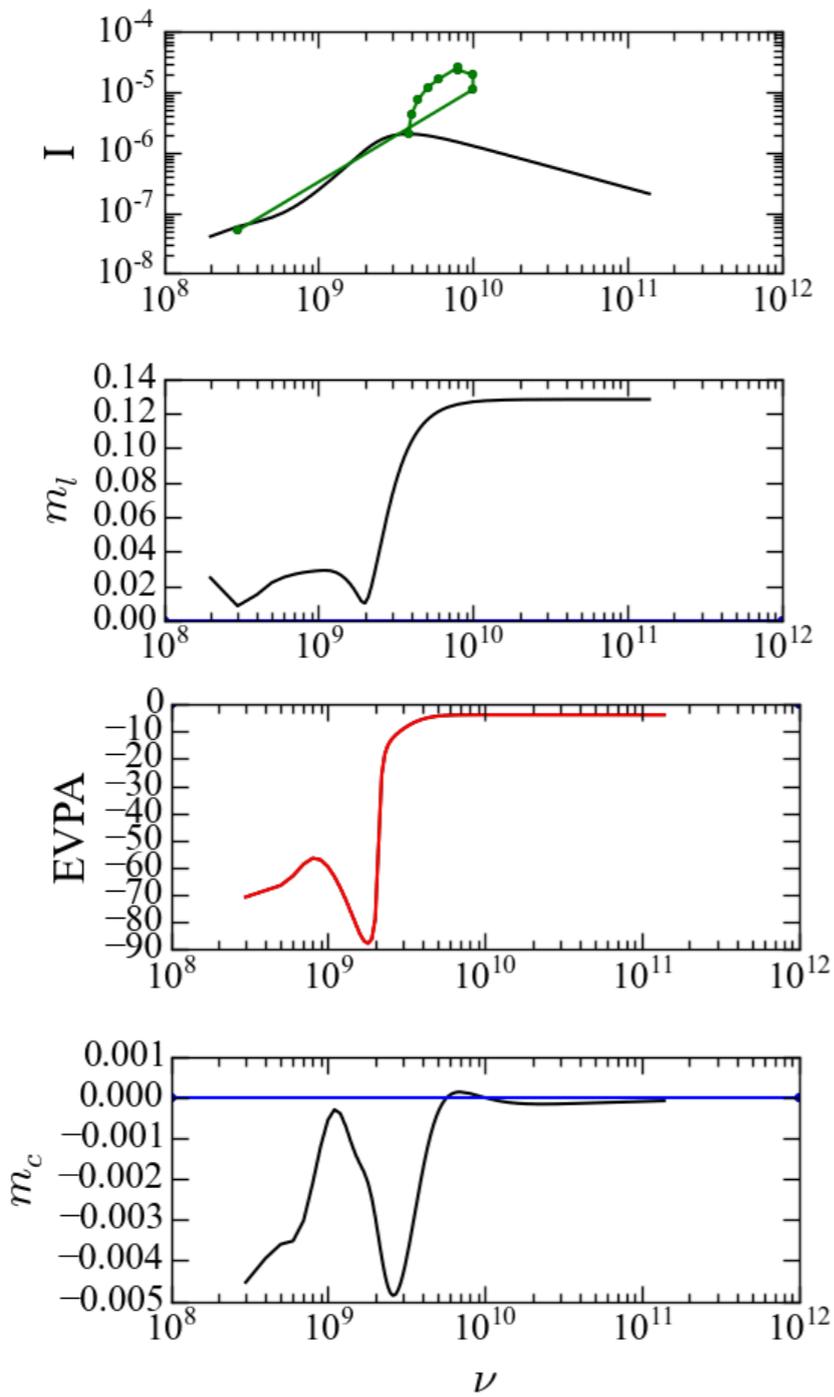
B-field strength

$$B' \sim kB$$



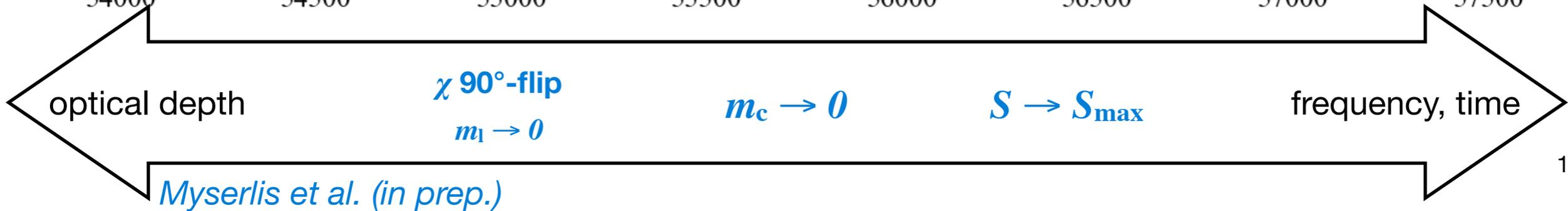
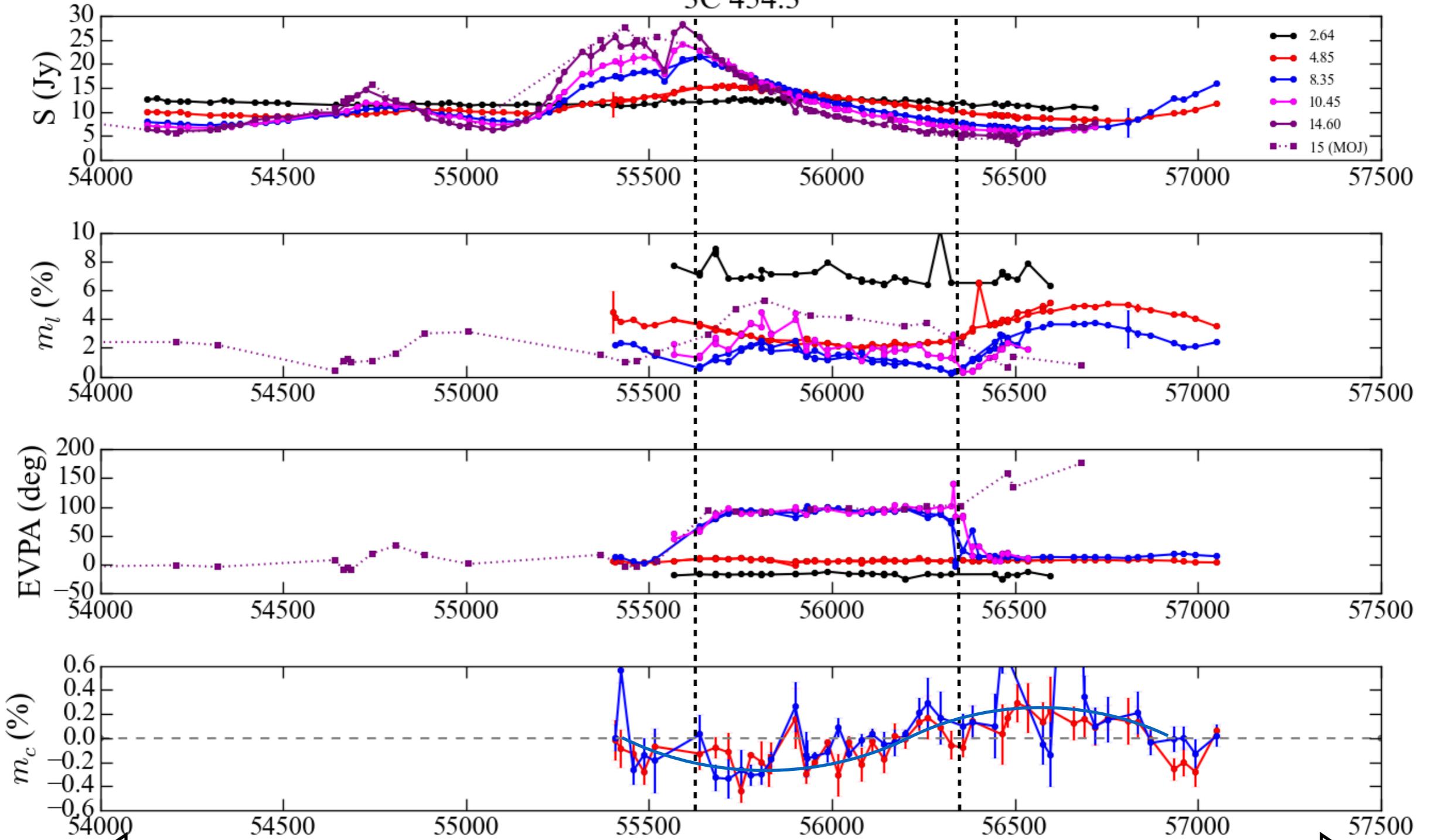
Line of sight

# Polarization variability



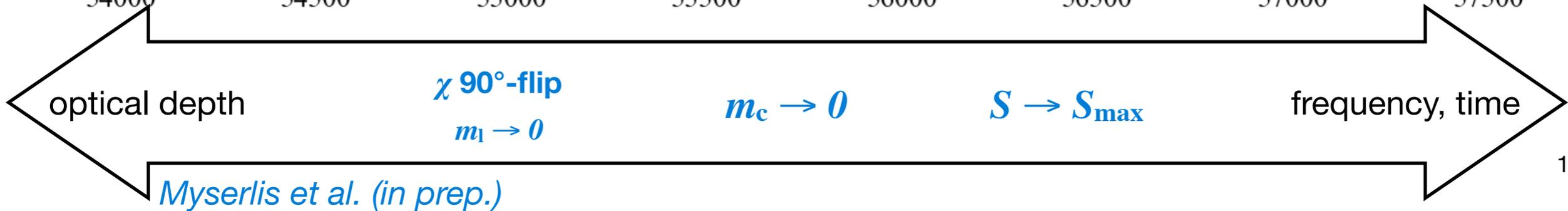
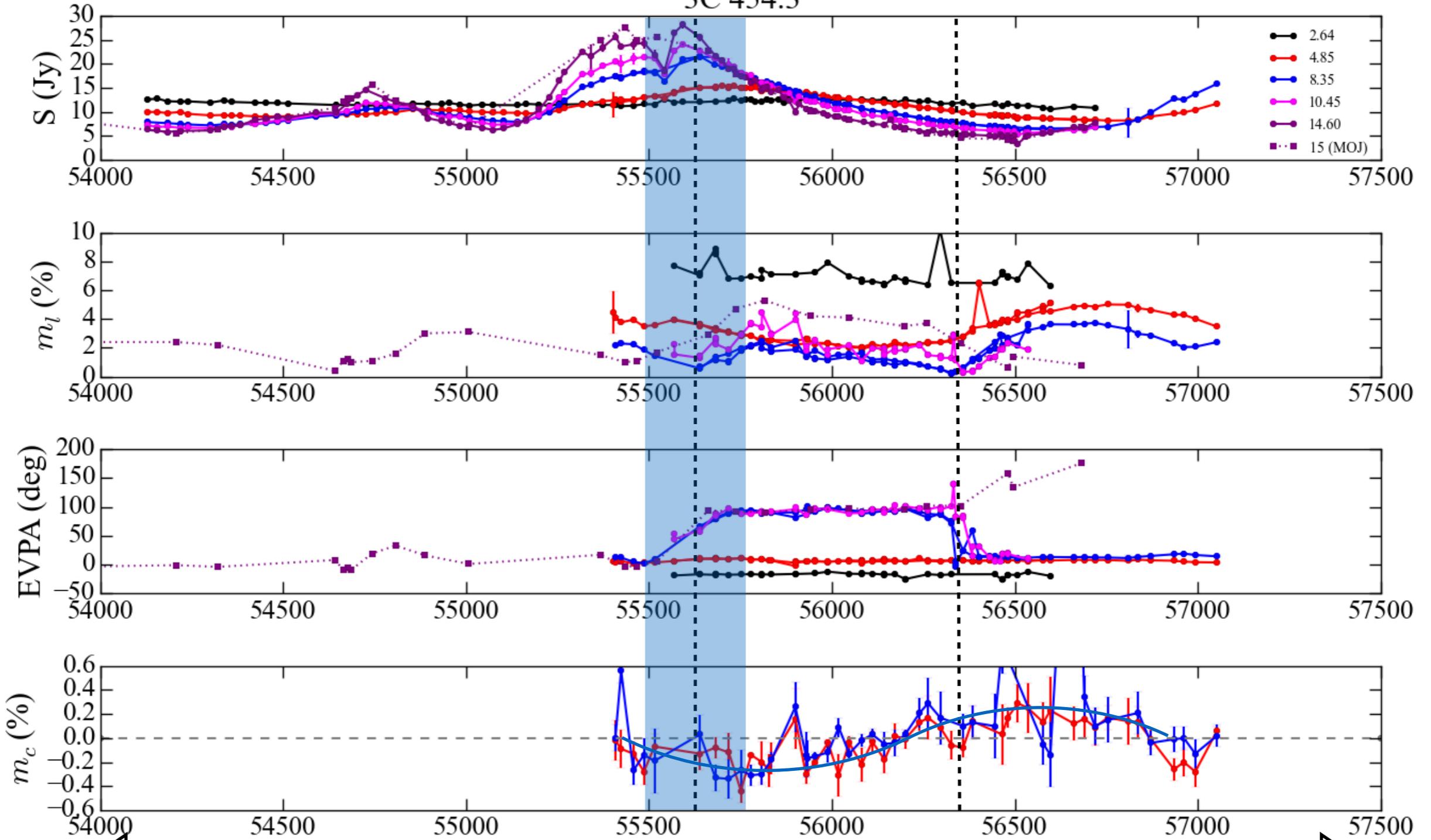
# Polarization variability

3C 454.3

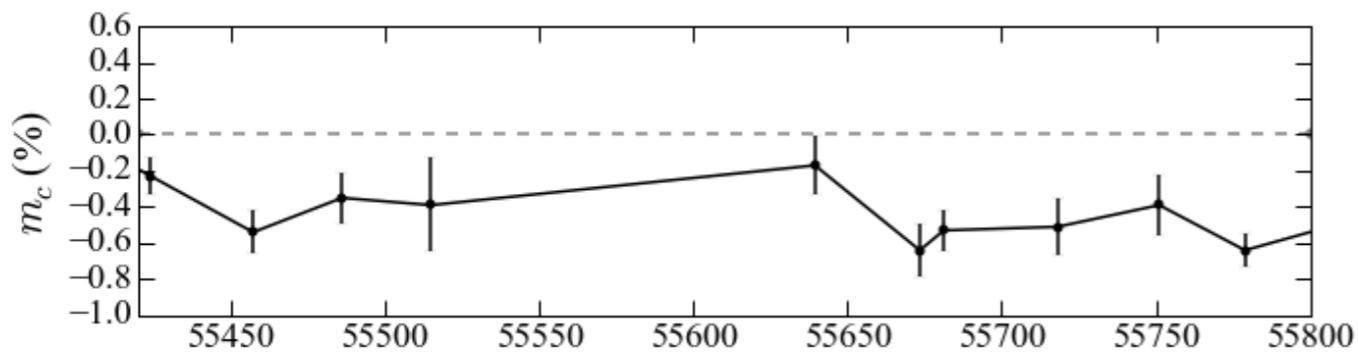
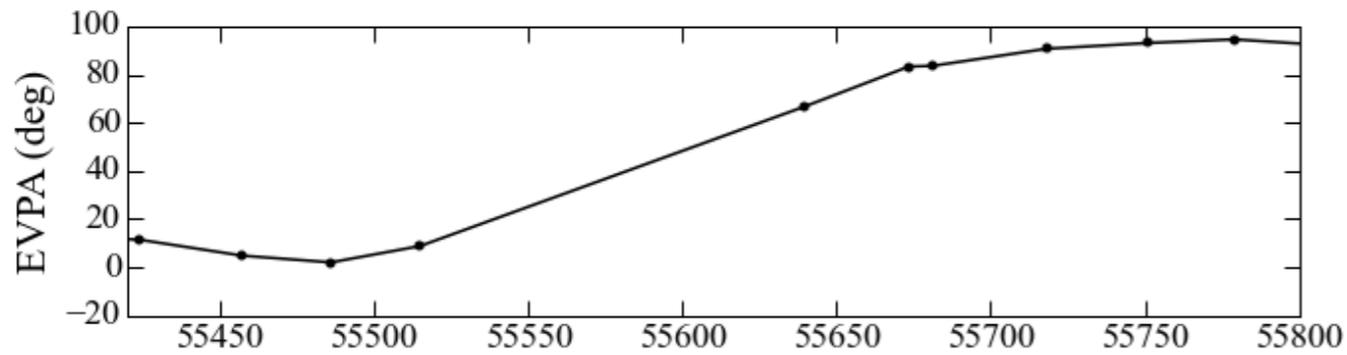
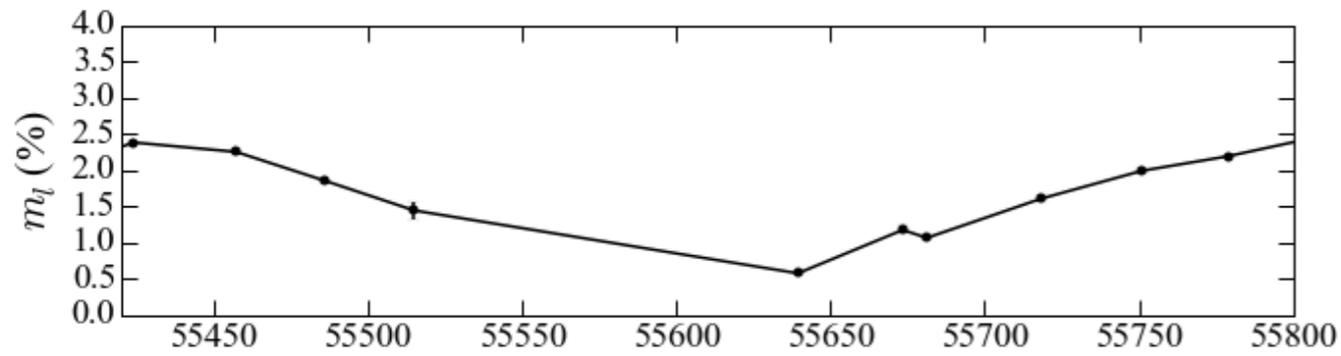
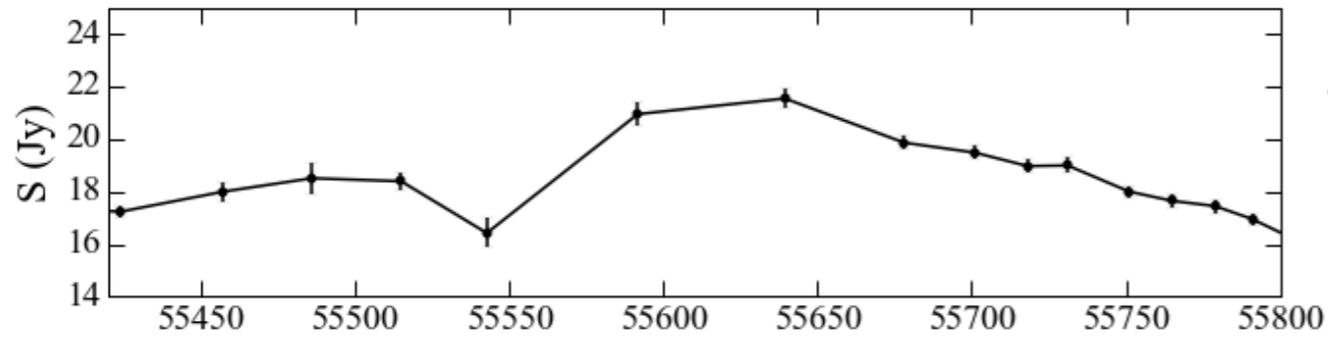


# Polarization variability

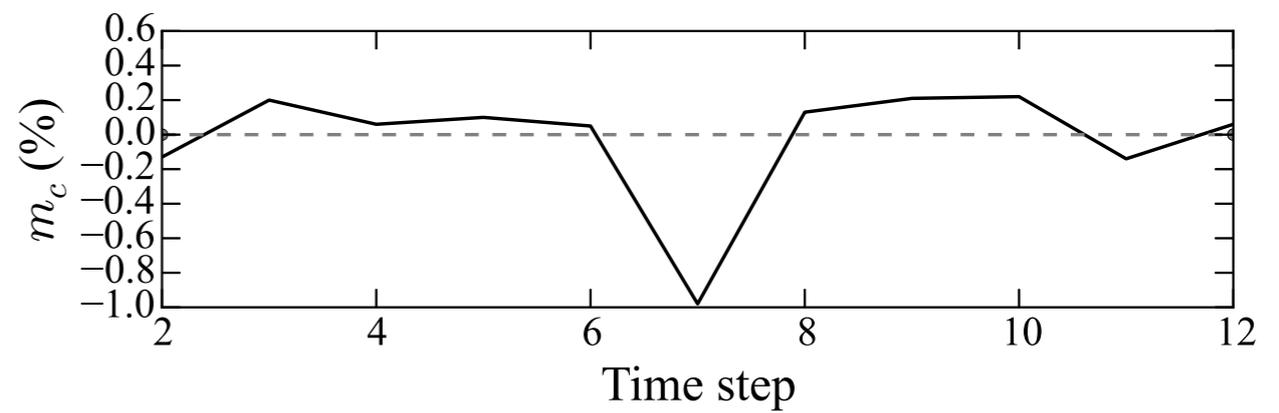
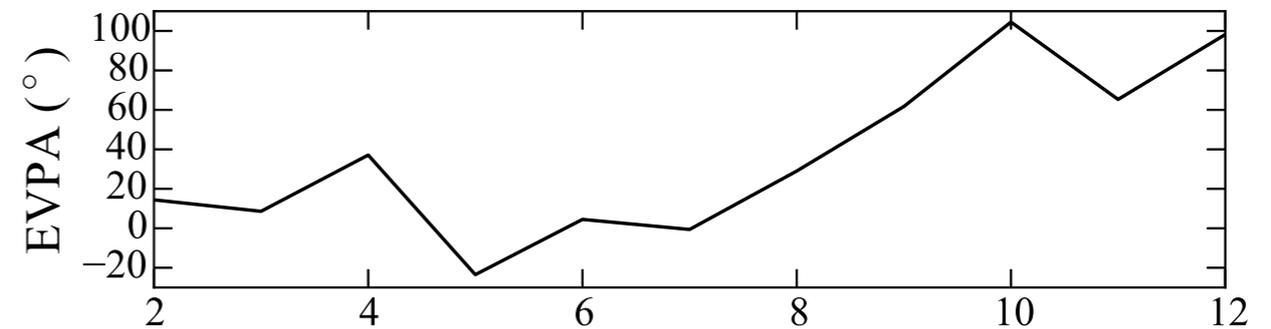
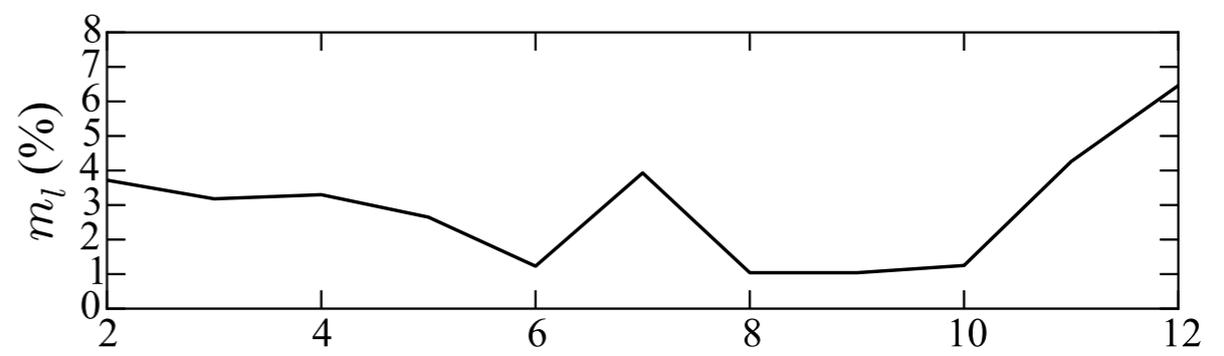
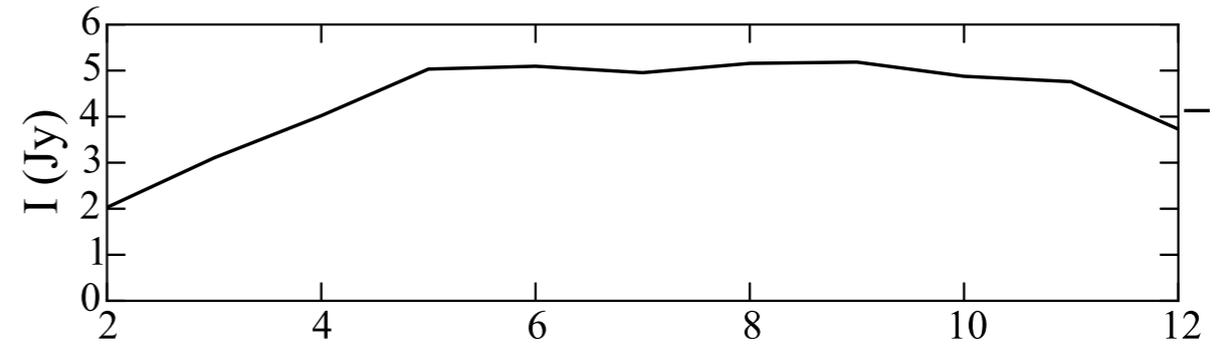
3C 454.3



# Data



# Simulation



# The study case of 3C 454.3

## Shock parameters

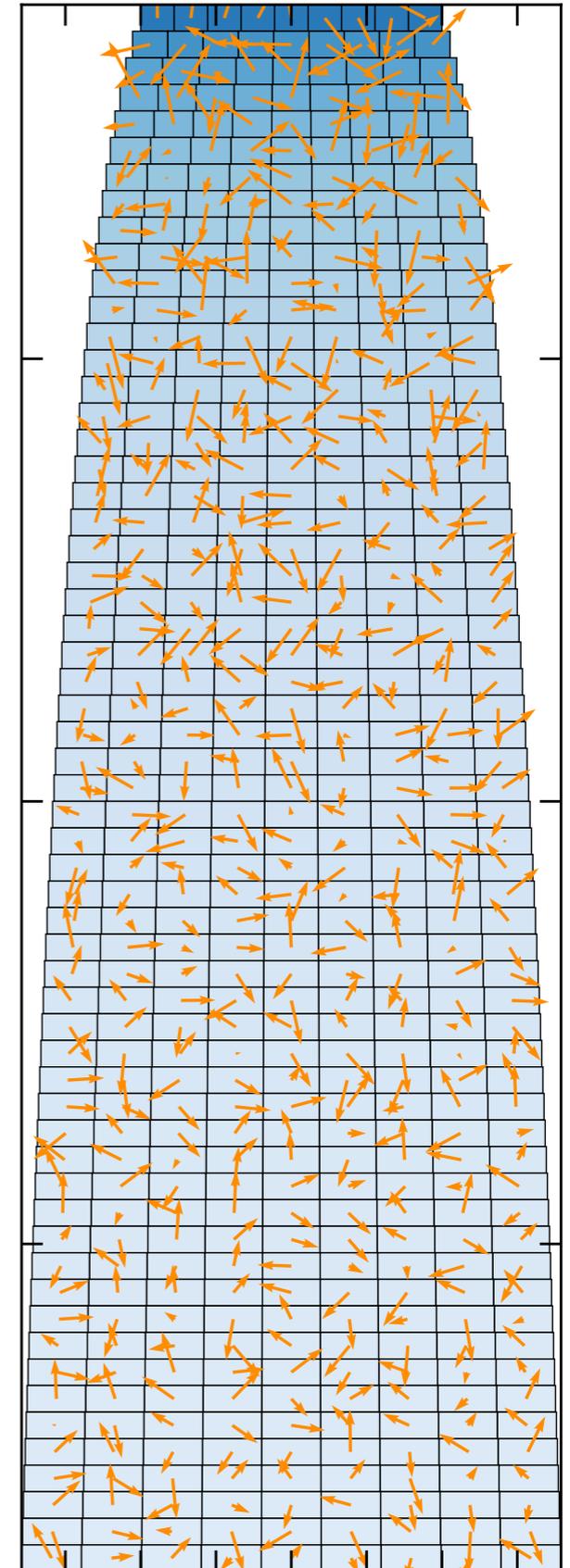
- Compression factor:  $k = 0.8$
- Doppler factor:  $D \sim 30$

Consistent with  $D_{\text{var}}$  at 37 GHz

*Hovatta et al. (2009)*

## Jet plasma parameters

- Density:  $n_0 = 10^1 - 10^2 \text{ cm}^{-3}$
- Magnetic field coherence length: 9 pc



# Conclusions

## Novel data analysis pipeline

- ➔ realistic parametrization of the telescope response
- ➔ instrument model for spurious LP correction
- ➔ polarization standards stability: 2%
- ➔  $m_l$  and  $m_c$  uncertainty: 0.2%
- ➔ LP and CP standards' catalogues of  $m_l$ ,  $\chi$  and  $m_c$

## Physical conditions

- ➔ **B**-field strength: 3-6 mG
- ➔ plasma composition: for e<sup>-</sup>-e<sup>+</sup> plasma: 1:1.5-2
- ➔ galactic origin of RM: no low energy plasma

# Conclusions

Modeling the variability

- ➔ full-Stokes radiative transfer code accounting for all propagation effects
- ➔ reproduce all observables and their variability
- ➔ study case: 3C 454.3
  - ➔ reproduce full-Stokes variability
  - ➔ physical conditions:
    - ➔ shock compression factor:  $k = 0.8$
    - ➔ Doppler factor:  $D \sim 30$
    - ➔ plasma density:  $10 - 100 \text{ cm}^{-3}$
    - ➔ Magnetic field coherence length: 9 pc