

StarBurst Multimessenger Pioneer

Mission Overview

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NASA Marshall Space Flight Center

09/13/2024



Marshall Space Flight Center



THE UNIVERSITY OF ALABAMA IN HUNTSVILLE

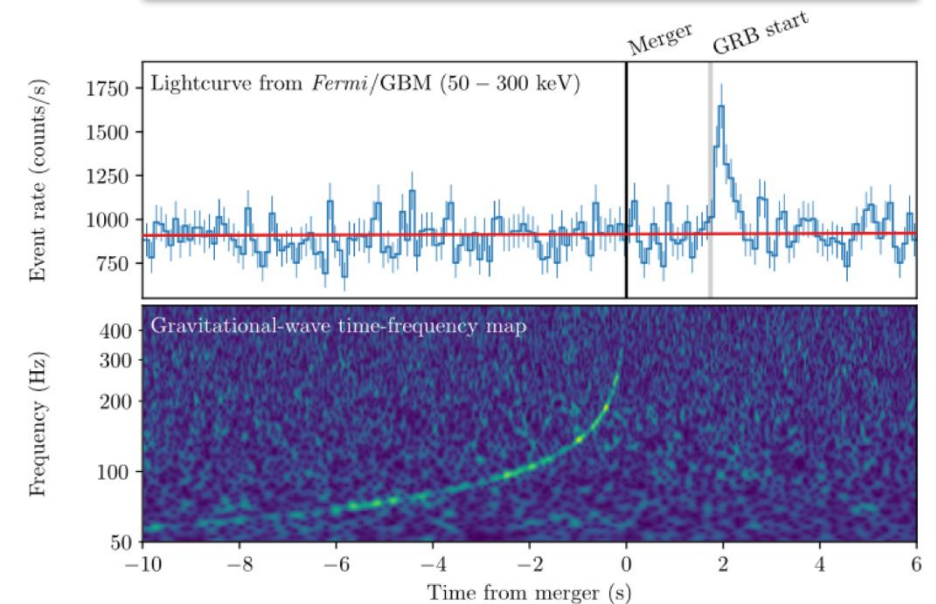
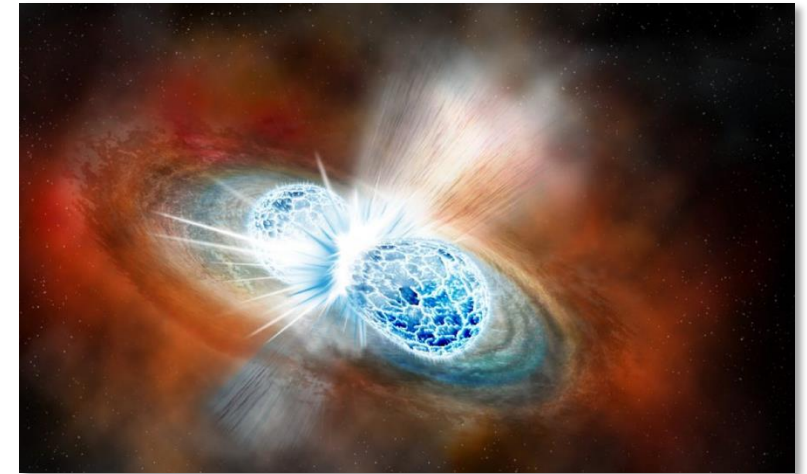


Universities Space Research Association

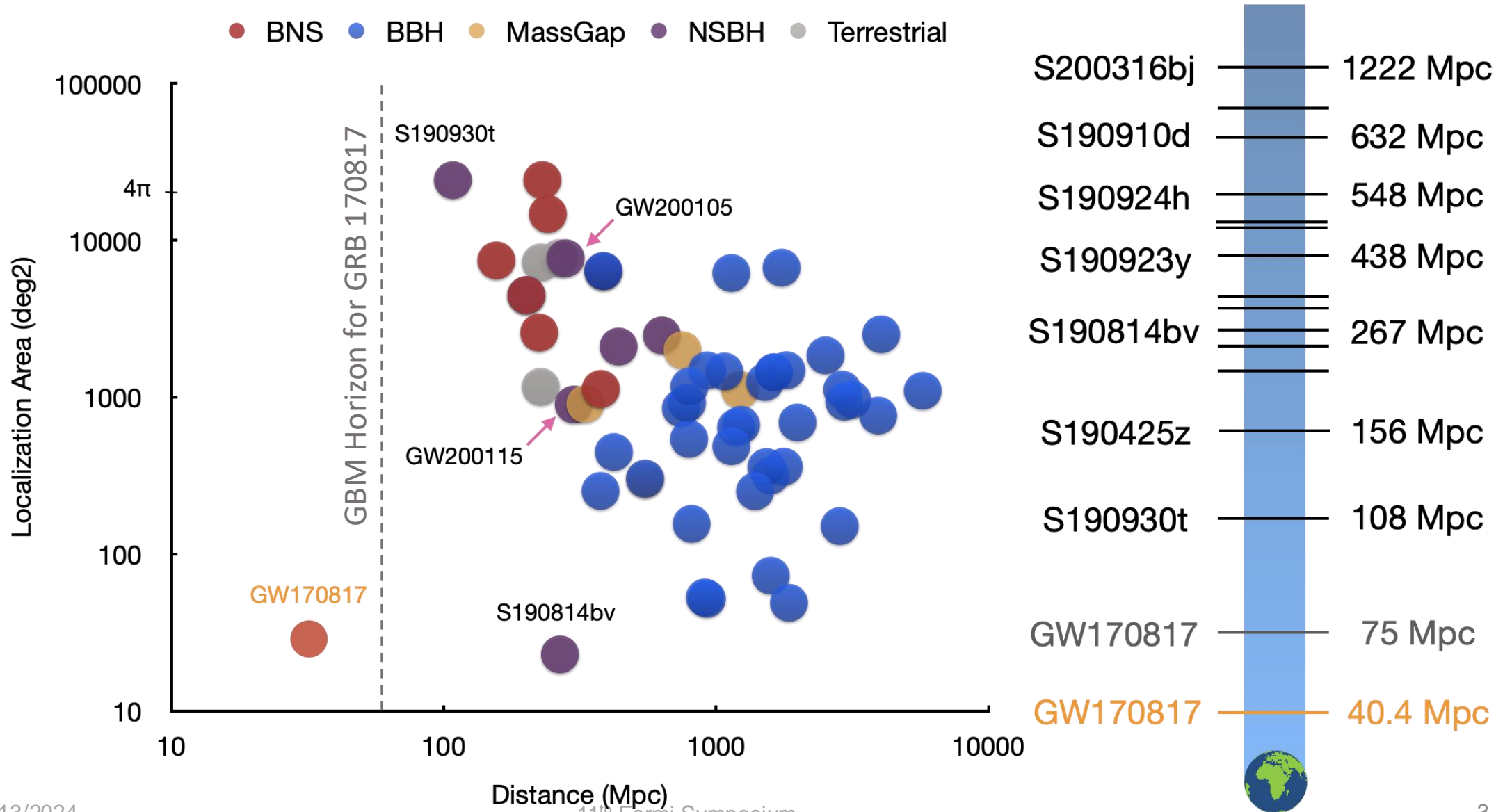


StarBurst Multimessenger Pioneer

- StarBurst is one of NASA's new Pioneer class missions designed to detect a large sample of SGRBs coincident with GW detections
- StarBurst achieves this through large area scintillation-based detectors read out by SiPMs and low-cost COTS electronics
- Science objectives focus on what can be learned from the temporal and spectral properties of the gamma-ray signal
- Science Objectives:
 - 1) Constrain the progenitors of SGRBs
 - 2) Probe the remnants of NS mergers
 - 3) Constrain the neutron star equation of state
 - 4) Probe the structure of relativistic outflows produced in neutron star mergers
- Launch readiness in mid-2026 for a 1-2 year mission



BNS/BHNS Events in O2 & O3



Mission Overview

Primary Institutions

Marshall Space Flight Center (MSFC)

- Principle Investigator (PI)
- Project Manager (PM)
- Chief (CE) & System Engineering (LSE, SE)
- Central Electronics Box (Instrument), Enclosure & Electronics
- Flight Software, development and testing
- Environmental testing (Instrument)
- Thermal Analysis

Naval Research Lab (NRL)

- Instrument design and fabrication
- Detector electronics

Space Flight Laboratory (SFL)

- Spacecraft bus
- Observatory: AI&T, Flight Analysis
- Mission Operations Center (MOC)
- Launch Site Support

University of Alabama (UAH)

- Instrument Flight Software lead
- Instrument calibration

University Space Research Association (USRA)

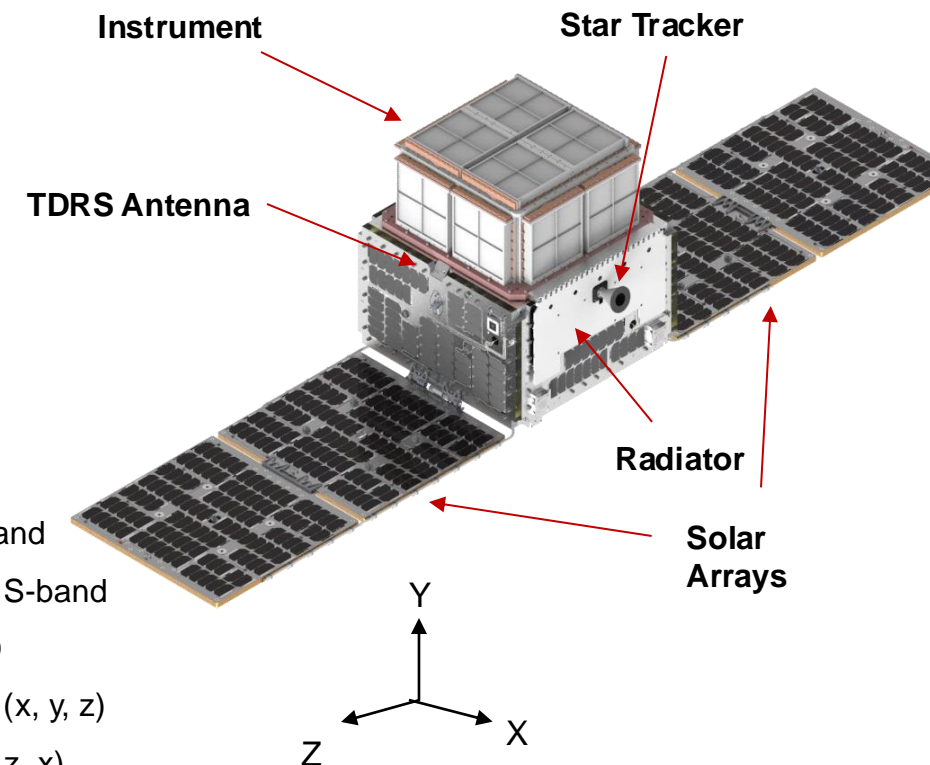
- Science Operation Center (SOC)

Quick Facts

- Energy range: 30-000 keV
- Energy resolution: 25% at 662 keV
- Field-of-view: $> 2\pi$ sr
- Pointing strategy: Fixed zenith
- Relative time resolution: 2 μ s
- Absolute time accuracy: 8 ms
- GRB localization: 3° for GRB 170817
- Instrument mass (CBE): 143.7 kg
- Observatory mass (CBE): 304.5 kg
- Nominal & peak power (CBE): 82 W
- Launch readiness date: Mid 2026
- Mission duration: 1 year
- Orbit: Low inclination (0° - 50°) LEO
- Low-latency communications: TDRS via S-band
- High-latency communications: KSAT-Lite via S-band
- Ground stations: KSAT-Lite @ TBD and TBD
- Size (stowed): 971 mm x 957 mm x 917 mm (x, y, z)
- Deployed: 971 mm x 3689 mm x 917 mm (y, z, x)

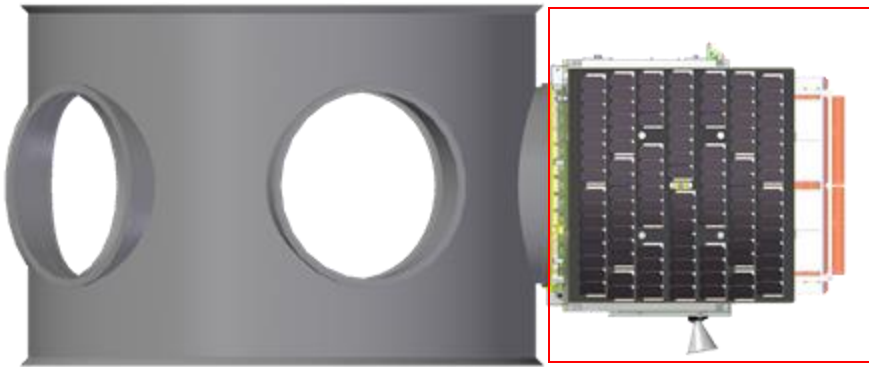
Observatory Description

StarBurst is a straightforward, low-risk, high-heritage, passive instrument that combines the strengths of MSFC, NRL, UAH, and USRA in developing, flying, operating, and managing gamma-ray missions

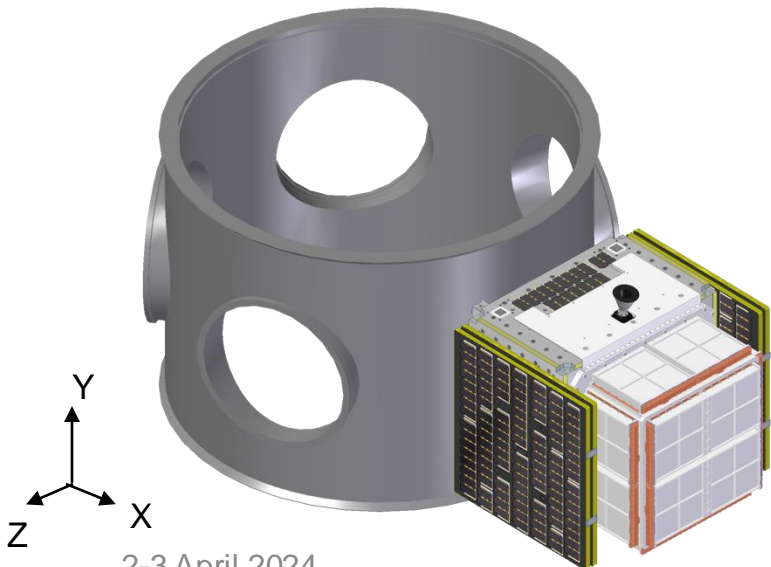


System Architecture & Design Concept

ESPA Grande Ring



Stowed StarBurst Observatory

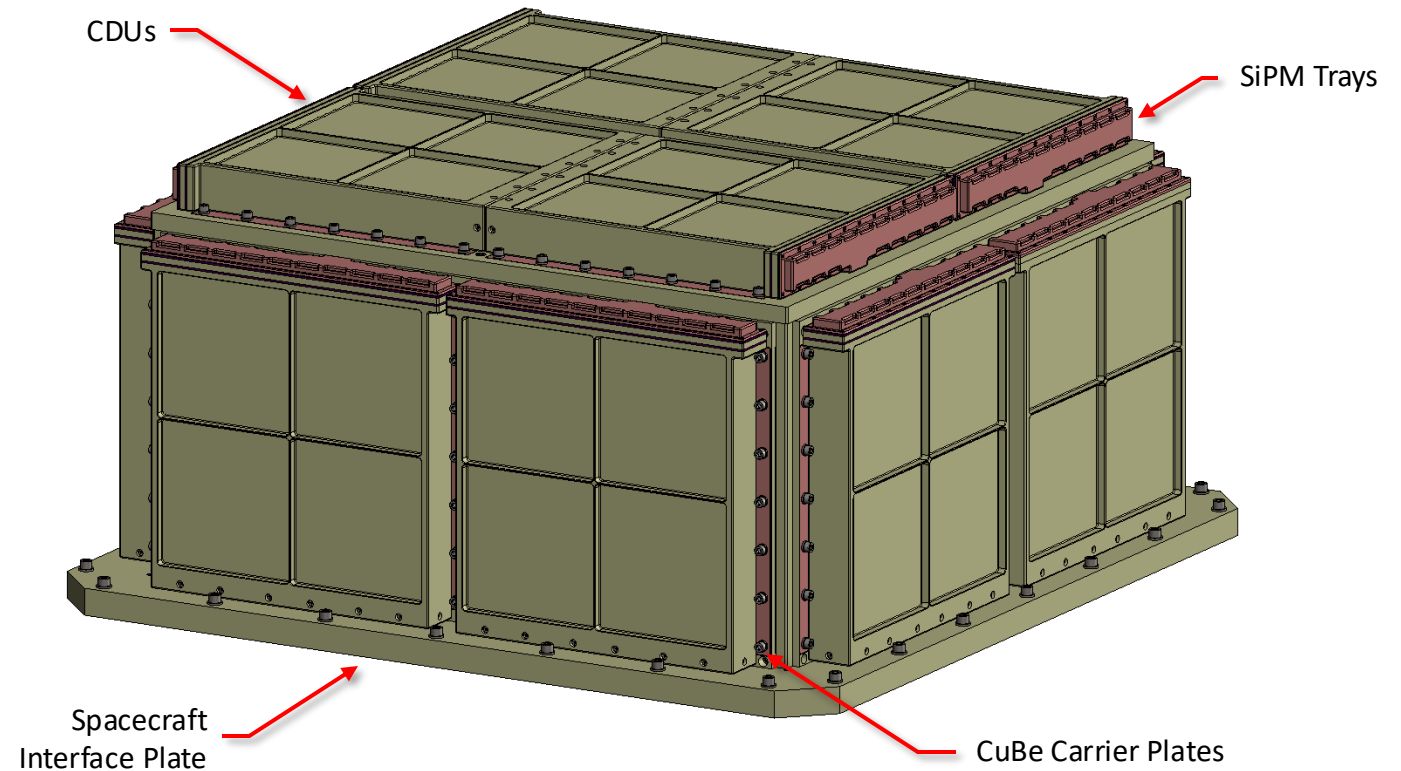


2-3 April 2024
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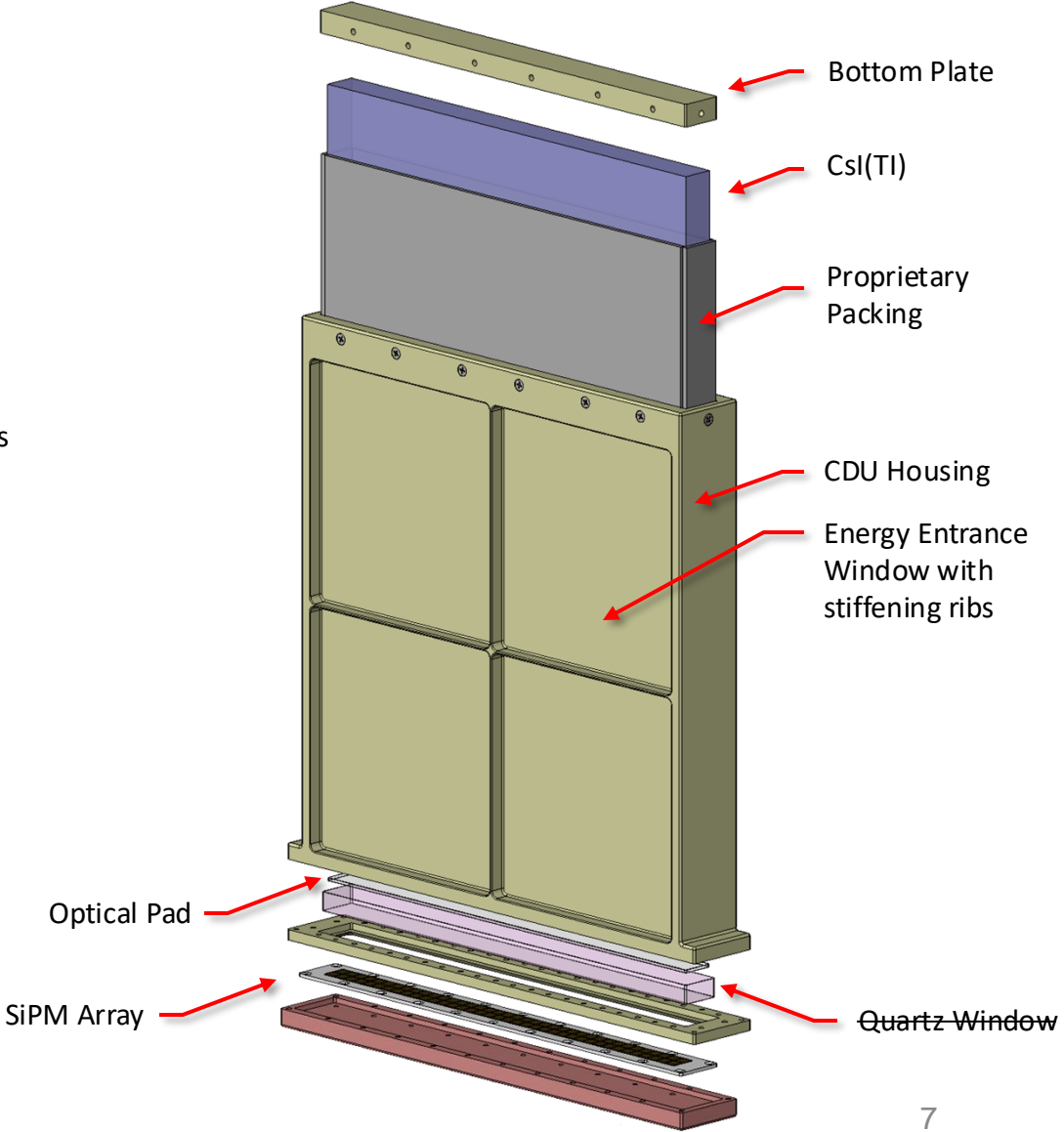
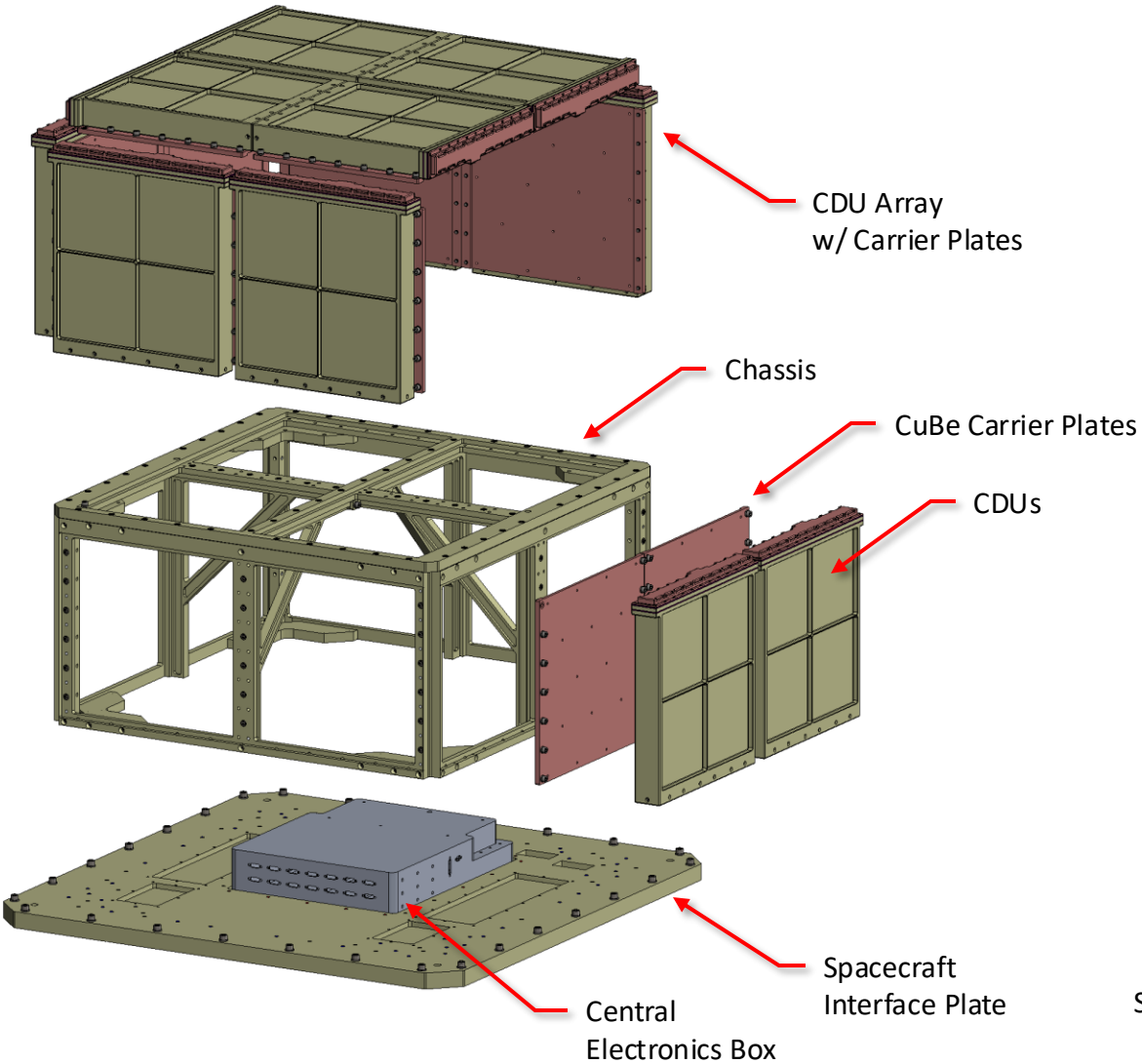
- StarBurst Observatory designed to attach to the ESPA Grande Ring
 - Port mass capacity 465 kg
 - Volume available 42" x 46" x 38" (y, z, x)
 - Spacecraft interface 24" circular
- Mass is 304.5 kg (CBE), 335 kg (with contingency)
 - Instrument 143.7 (CBE), 160 kg (with contingency)
 - Spacecraft bus 160.8 (CBE), 160 kg (with contingency)
- Size
 - Stowed: 971 mm x 957 mm x 917 mm (y, z, x)
 - Deployed: 971 mm x 3689 mm x 917 mm (y, z, x)
- Orbit: LEO
 - Altitude: 365 km to 545 km
 - Inclination: 0 - 60°

StarBurst Instrument Design Overview

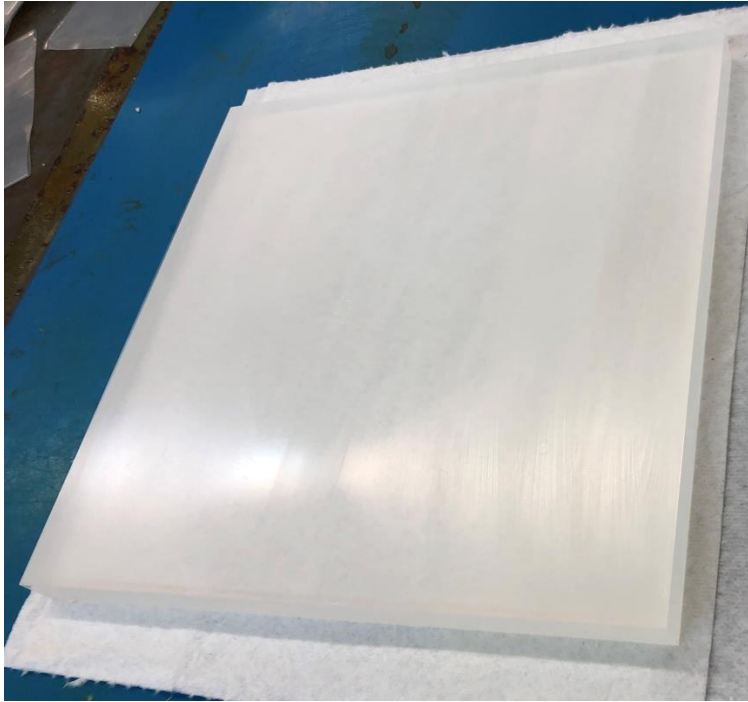
- Crystal Detector Units (CDUs)
 - 12 large CsI(Tl) panels
 - 240 mm x 240 mm x 13 mm
 - Mounted to CuBe carrier plate
- Silicon Photomultiplier (SiPM) readout
 - Edge readout
 - Active summing circuit
- Copper-Beryllium (CuBe) SiPM trays
- Copper-Beryllium (CuBe) carrier plate
- Central Electronics Box (CEB)
 - Housed within the detector array
 - Bridgeport Instruments SlimMorpho multichannel analyzers
- Spacecraft Interface Plate (SIP)



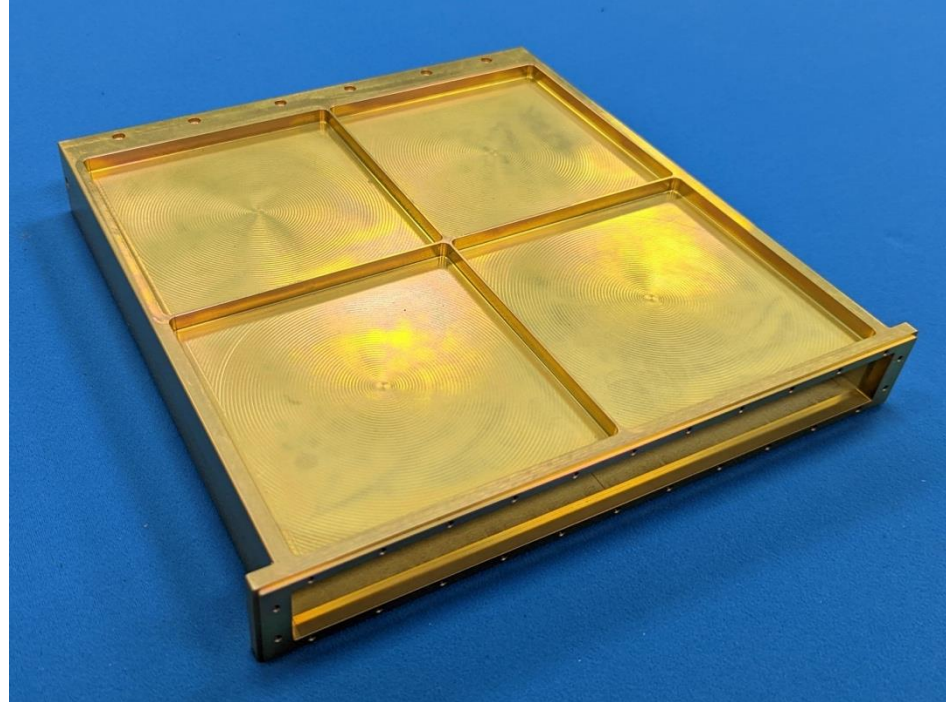
StarBurst Instrument Design Overview



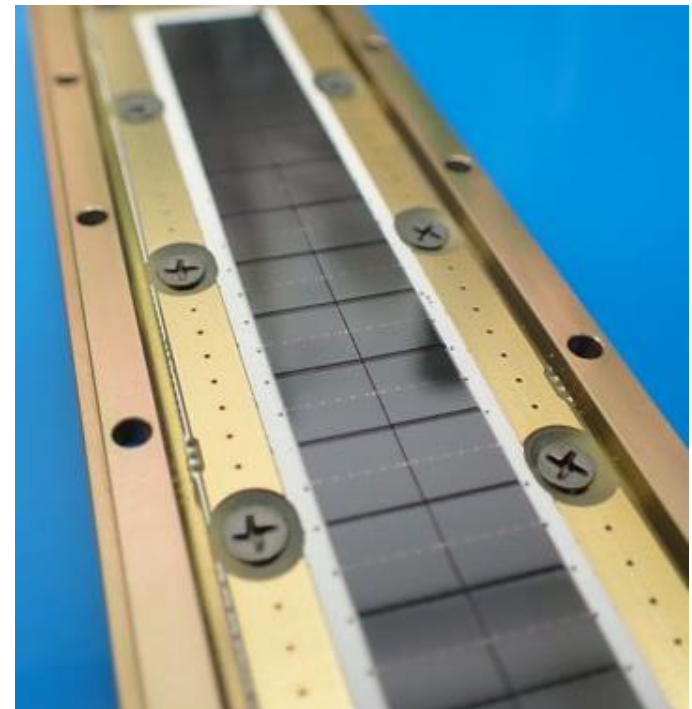
CDU Hardware



Bare CsI(Tl) crystal



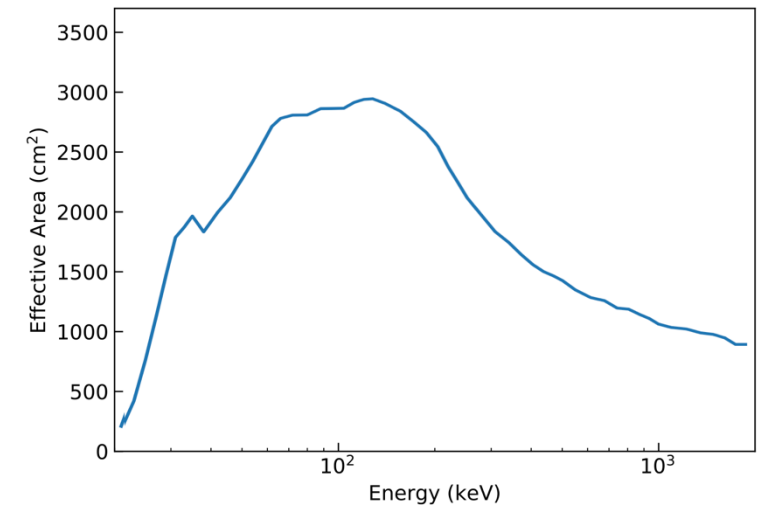
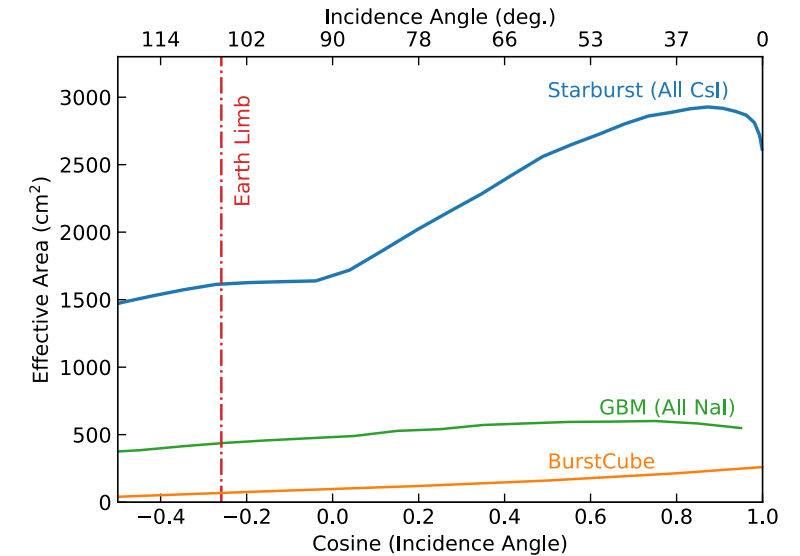
CDU housing



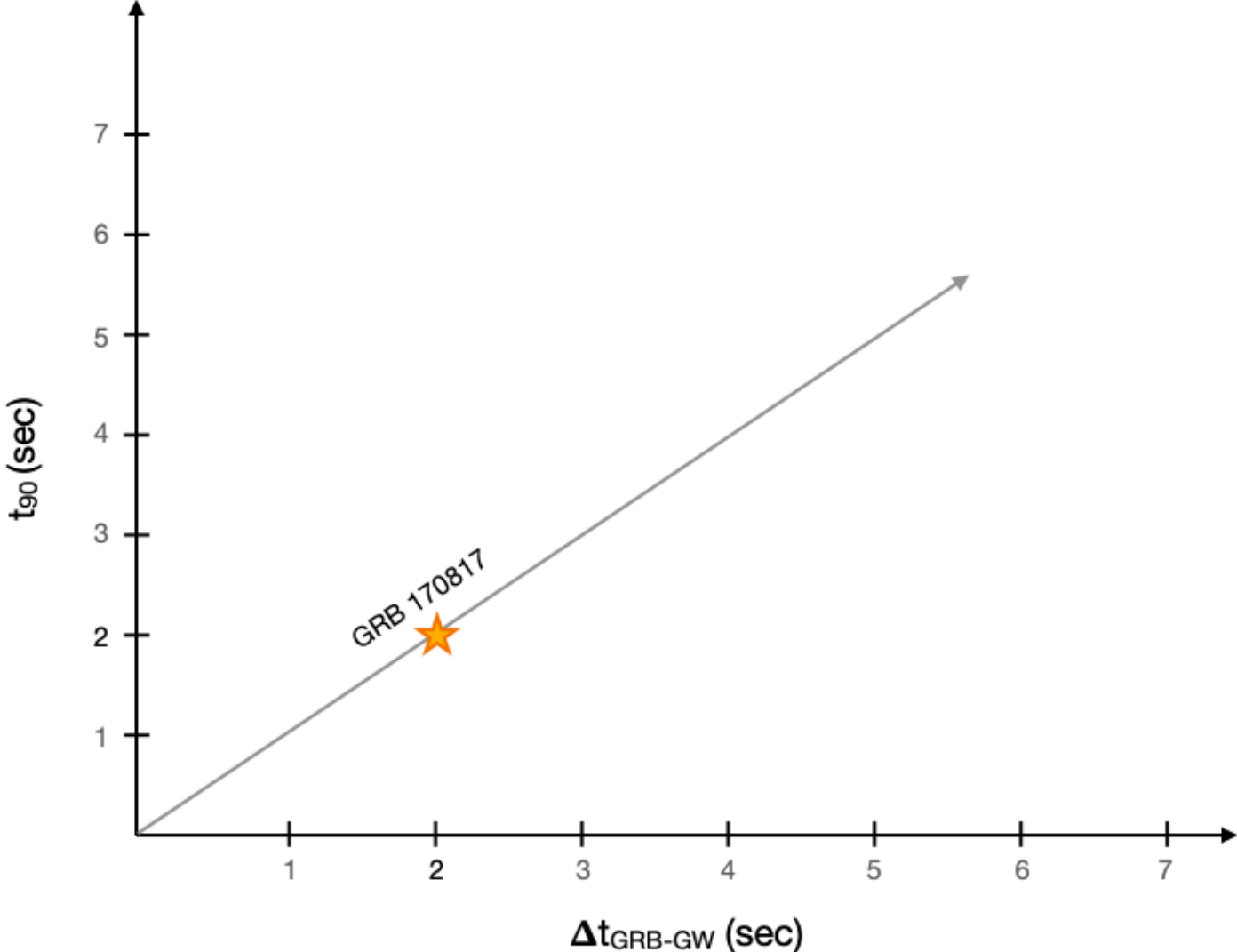
SiPM array in closeout tray

Instrument Performance

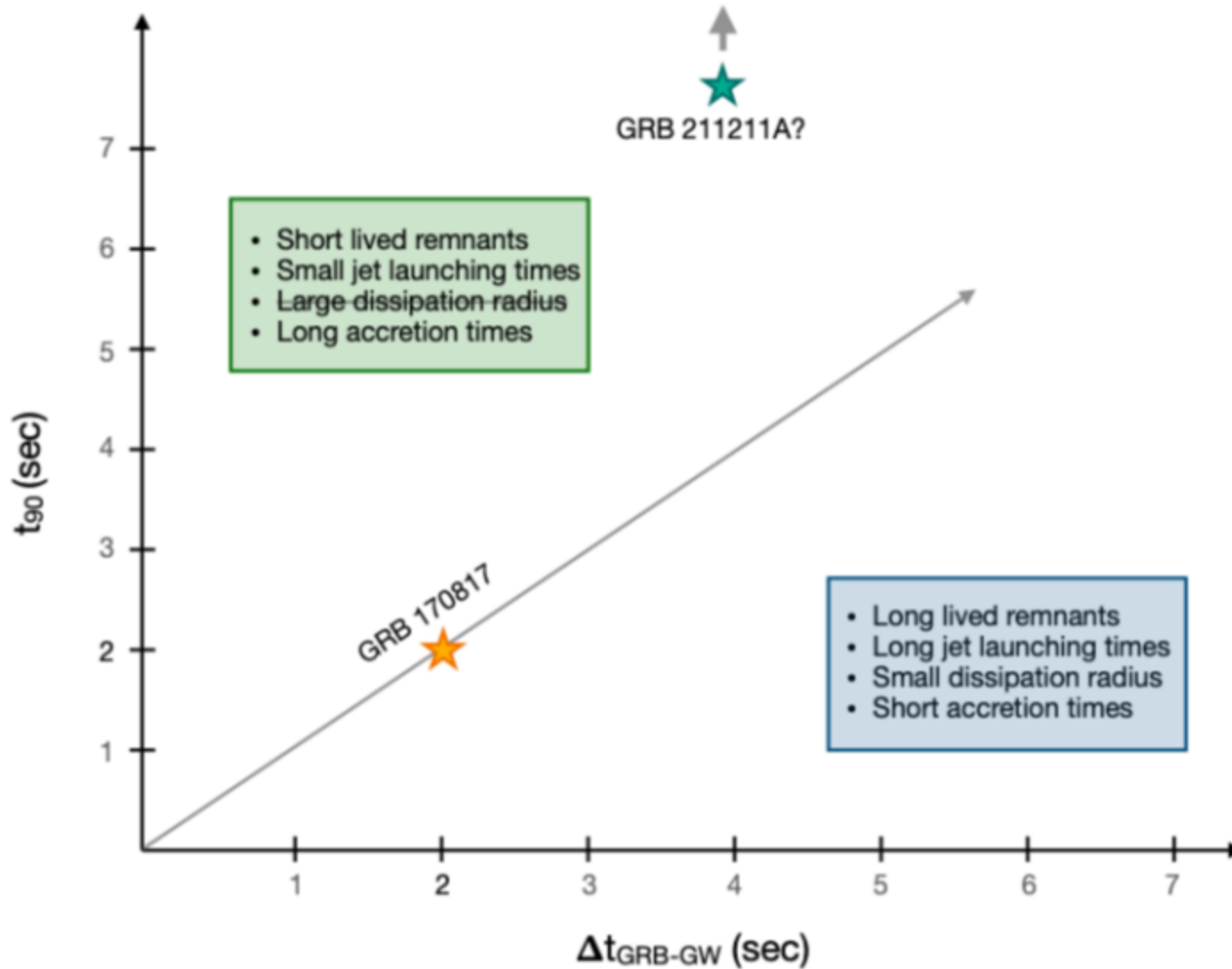
- Effective Area
 - Azimuth average effective area peak of $\sim 3000 \text{ cm}^2$
 - StarBurst effective area $\sim 500\%$ of GBM
- GRB Detection Rate
 - StarBurst: 158 SGRBs/yr (790 LGRBs)
 - Swift: 8.6 SGRBs/yr, GBM: 40 SGRBs/yr
- Joint detection rates
 - Based on E. Howell et al (2018) and Abbott et al. (2020)
 - Roughly 3.7% of the median A+ BNS detection rate
 - Median rate of 7 GW-SGRBs/yr (2.6—25.2 @ 90% CL)



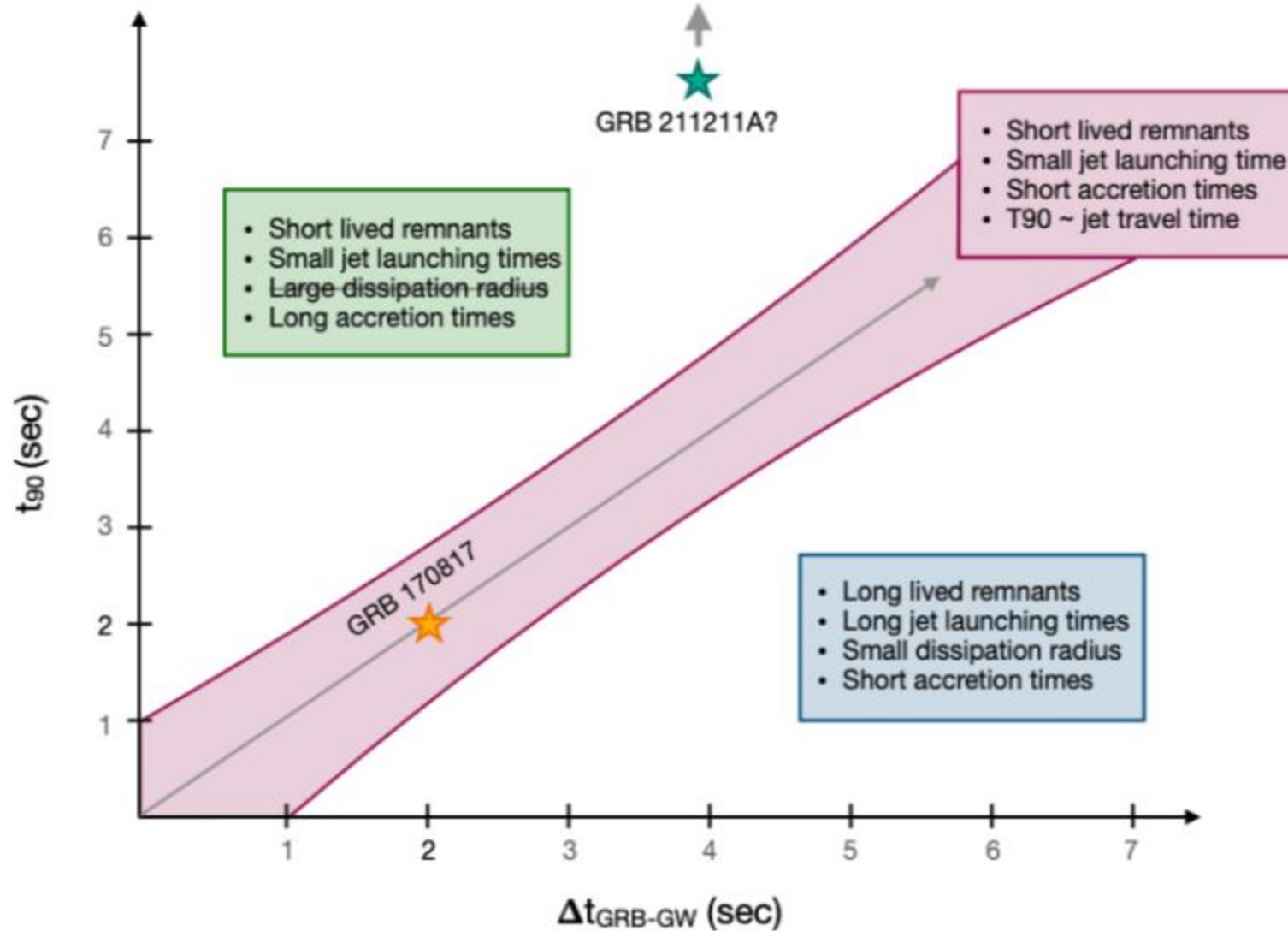
Future Science – GRB-GW Time Delays



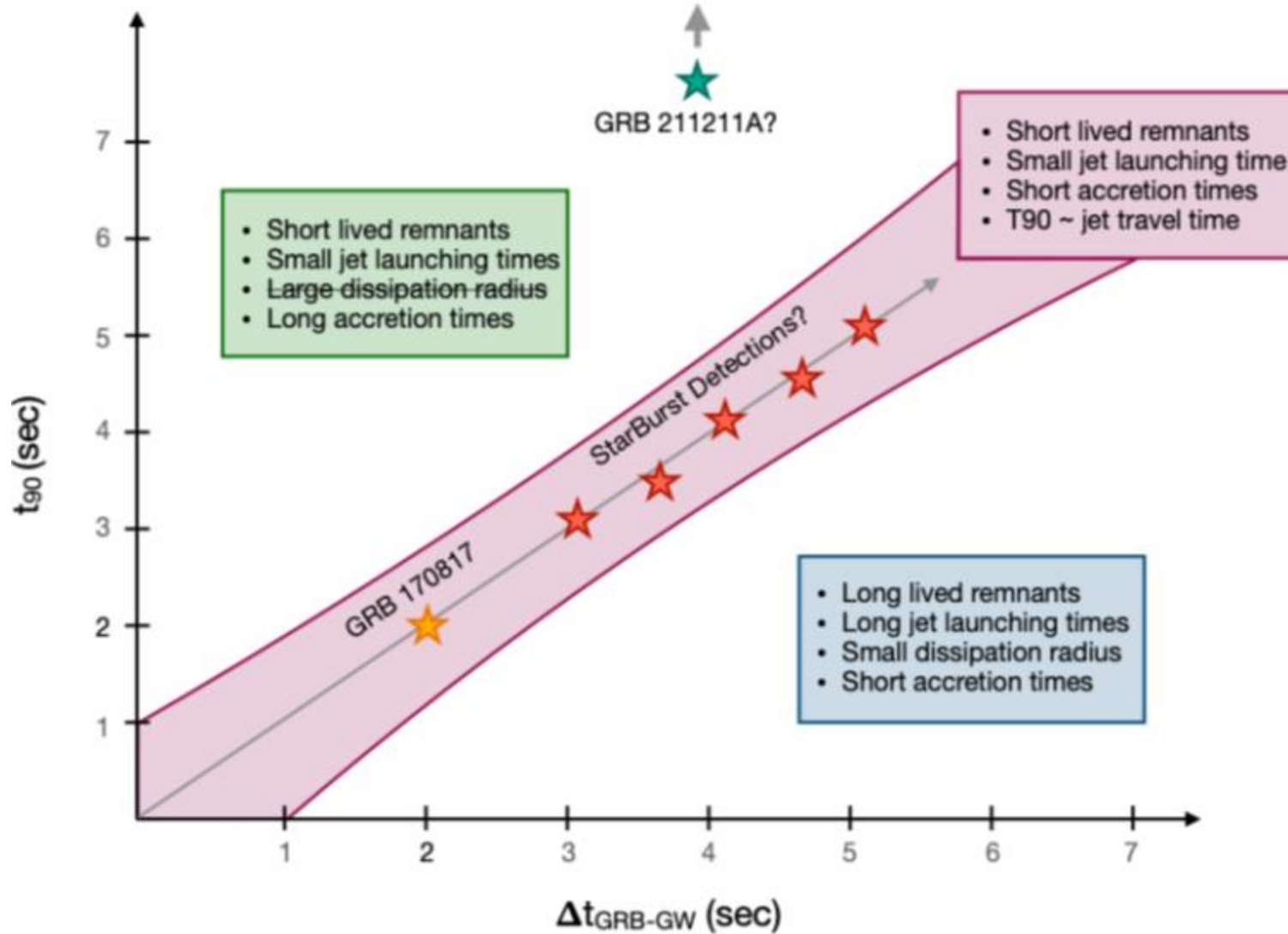
Future Science



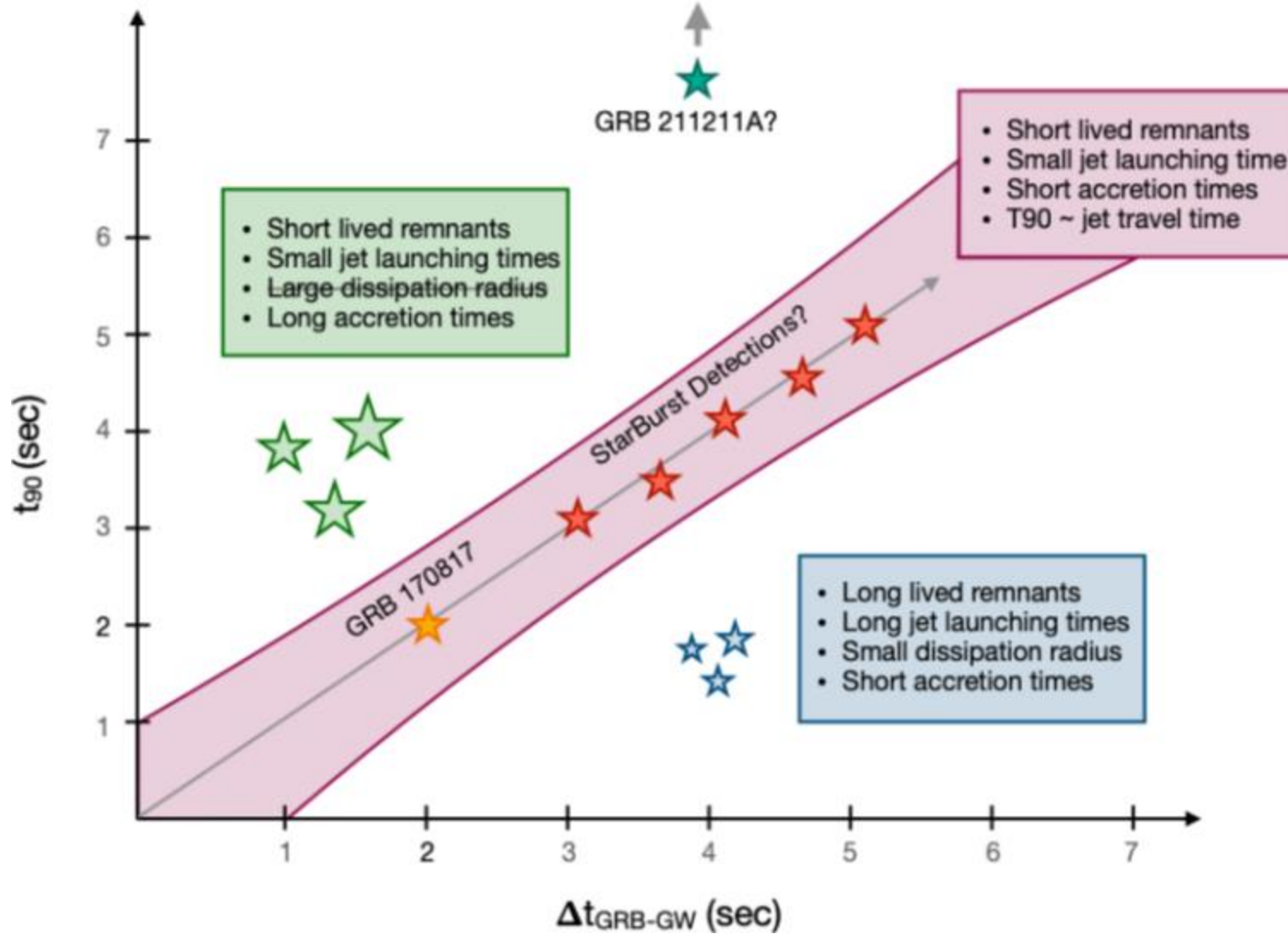
Future Science



Future Science

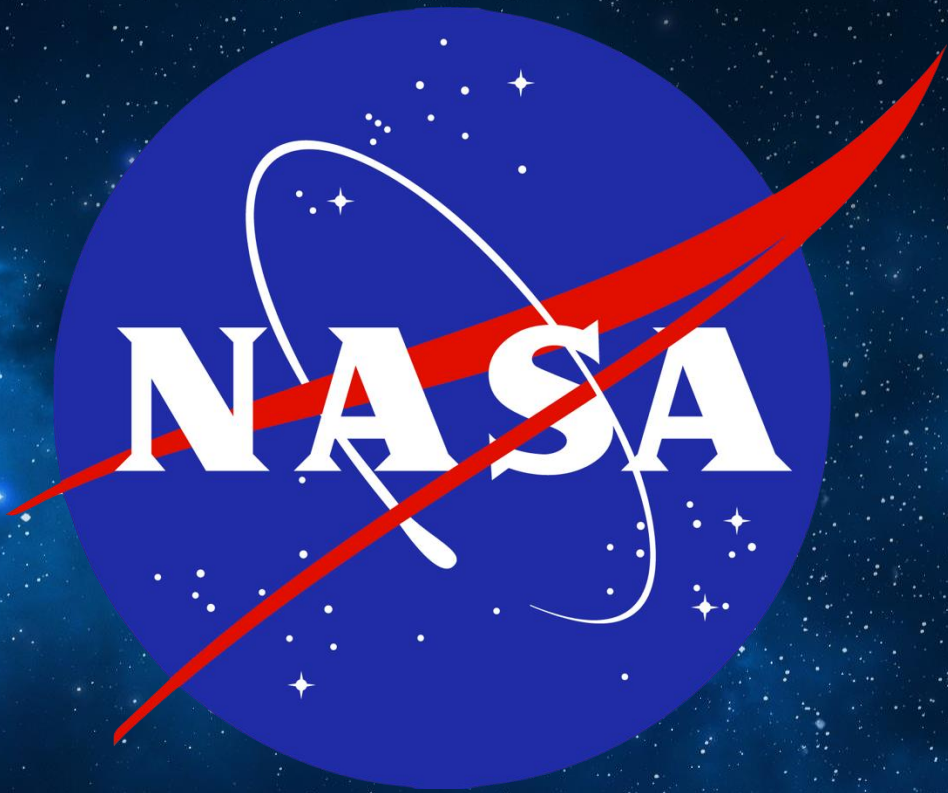


Future Science



Current Status and Schedule

- StarBurst is starting Phase D
 - Preliminary Design Review was held on on 11/16/2022
 - Critical Design Review was held on for 04/03/2024
- Flight unit fabrication and assembly beginning
 - Instrument delivery in Nov 2024 and observatory AI&T to being in Dec 2024
- Launch readiness date in mid 2026
 - Storage option allows us to wait to time the mission operations to coincide with LIGO O5 run
- Awaiting launch opportunity identification by NASA's Rideshare Office
- SpaceX announced they will begin offering mid-inclination (45 deg) orbits, which may be our ticket to LEO

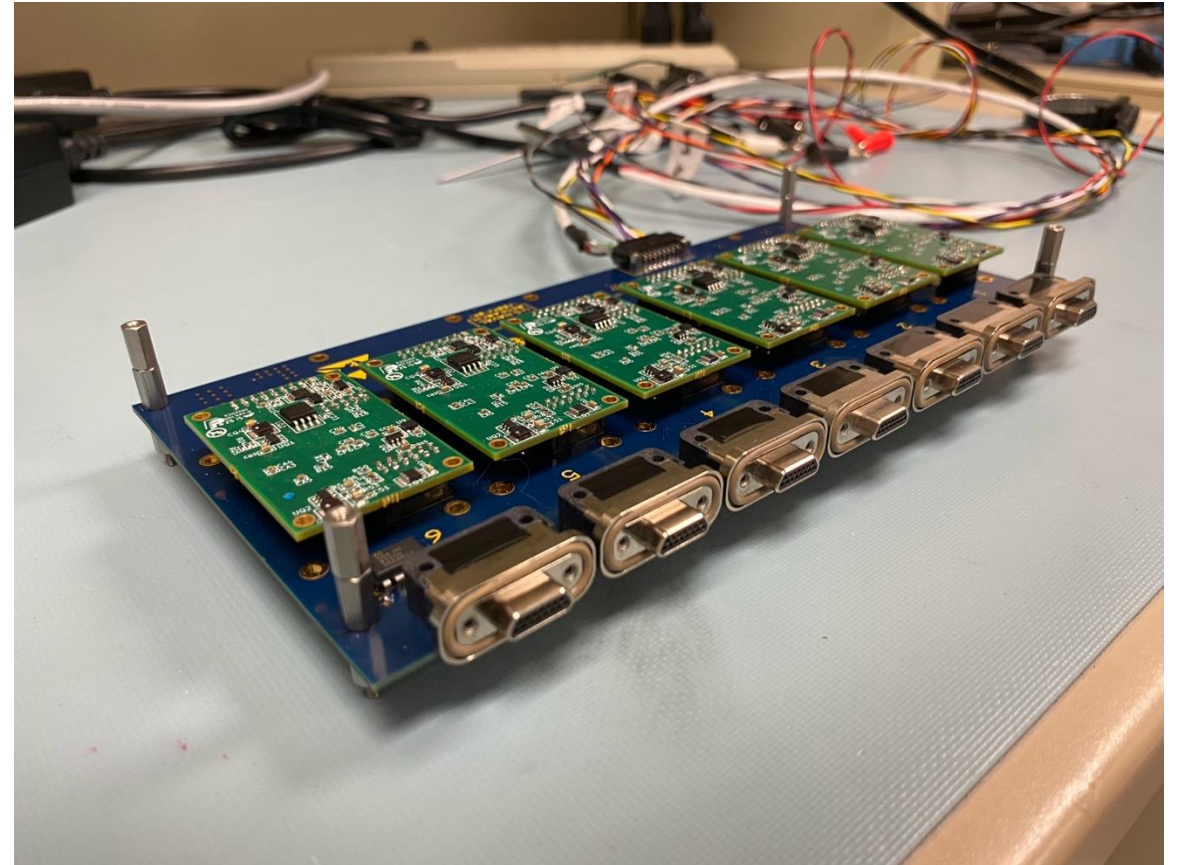
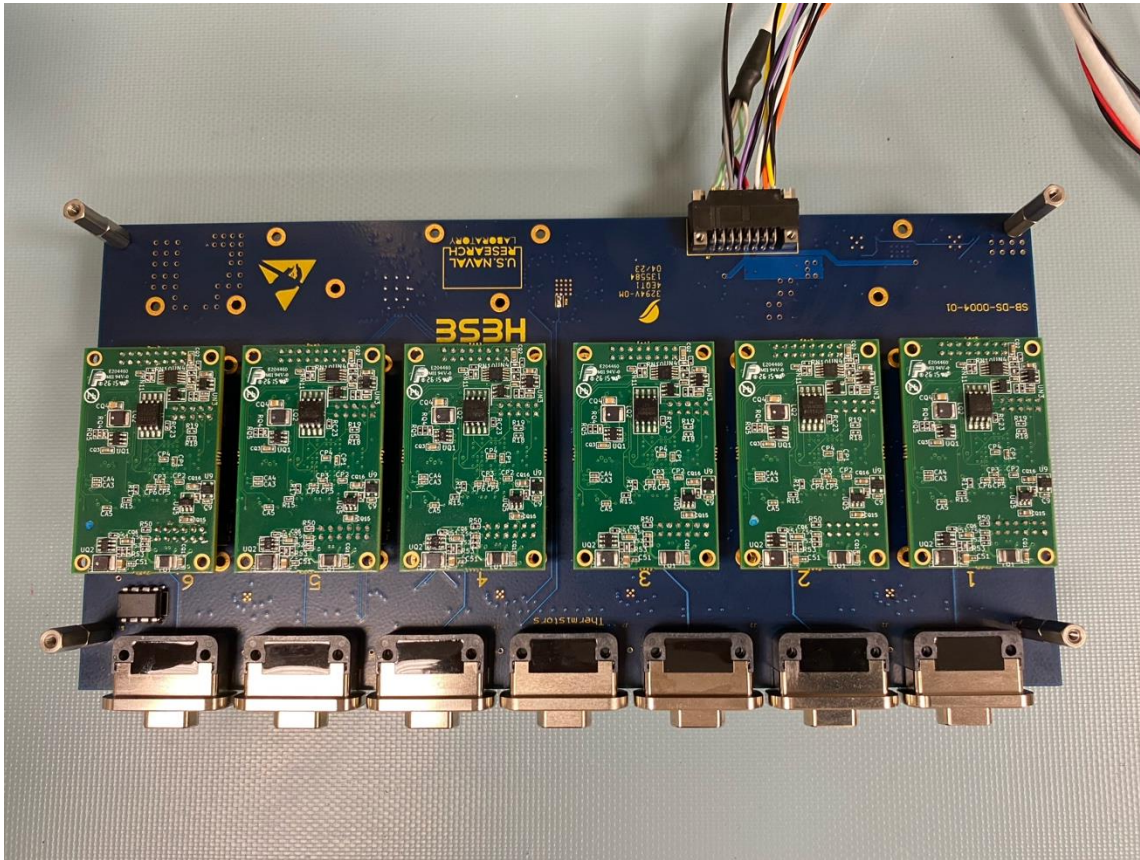




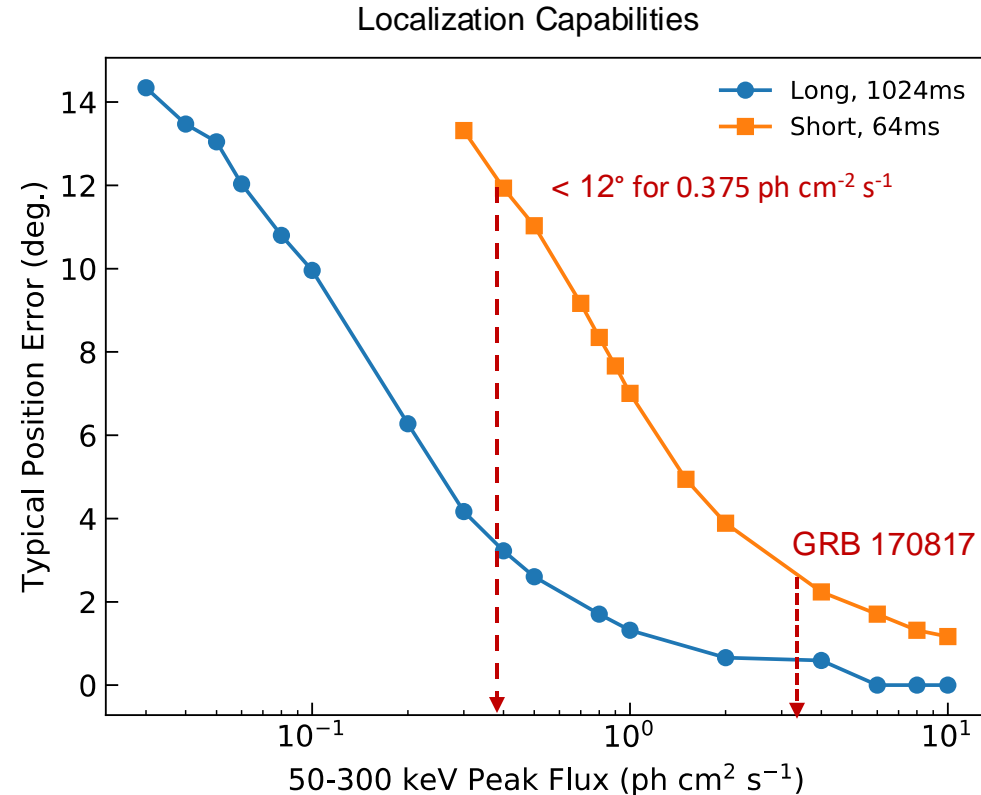
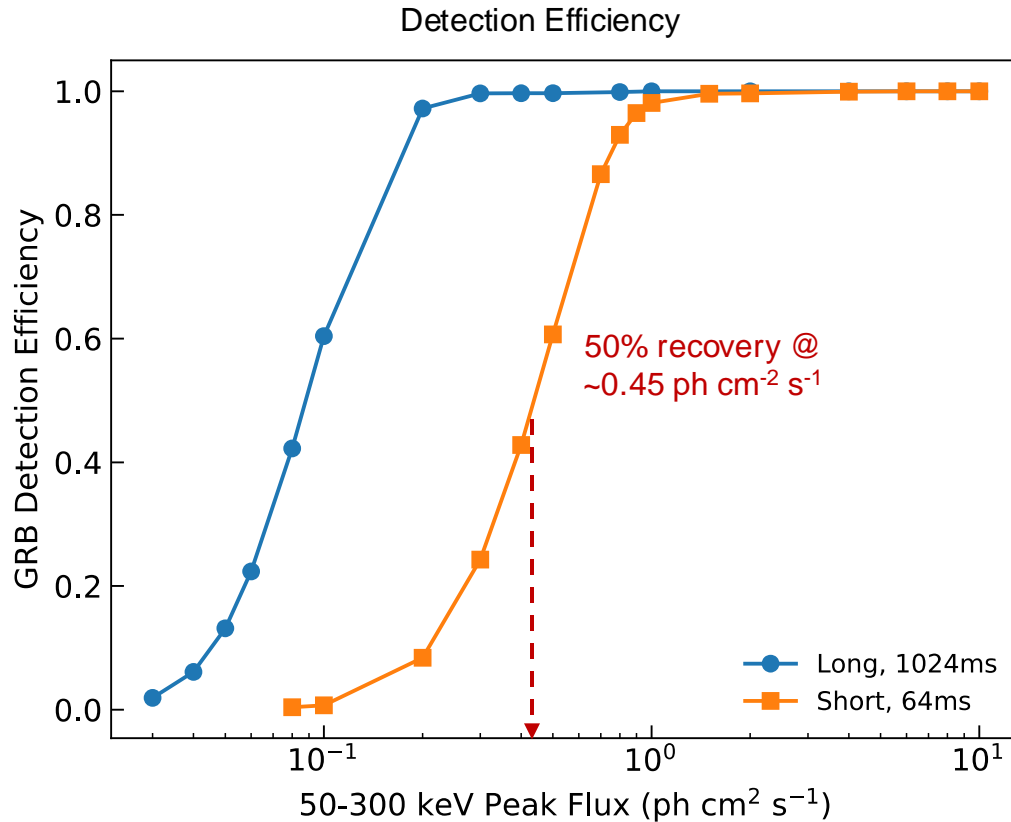
Backup Slides

Detector Readout Electronics

First EM Detector Electronics (HexMorpho with 6 slimMorphos) is now complete and tested

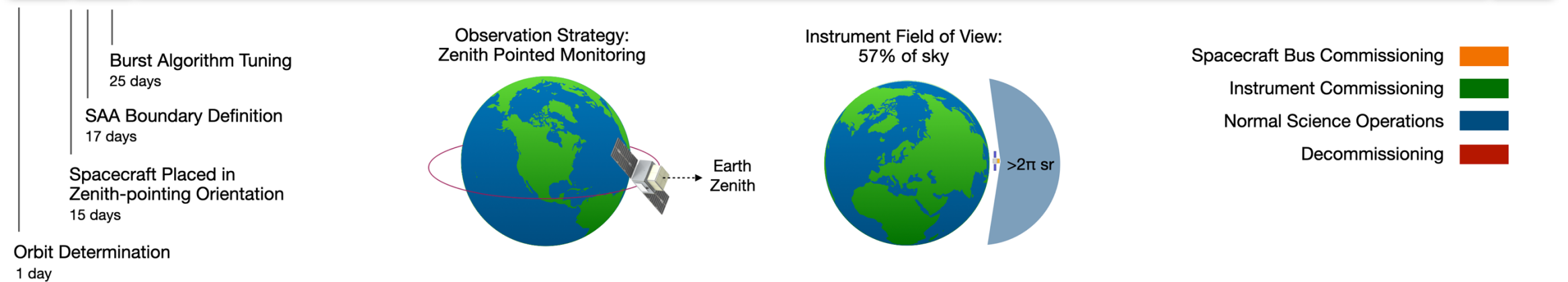
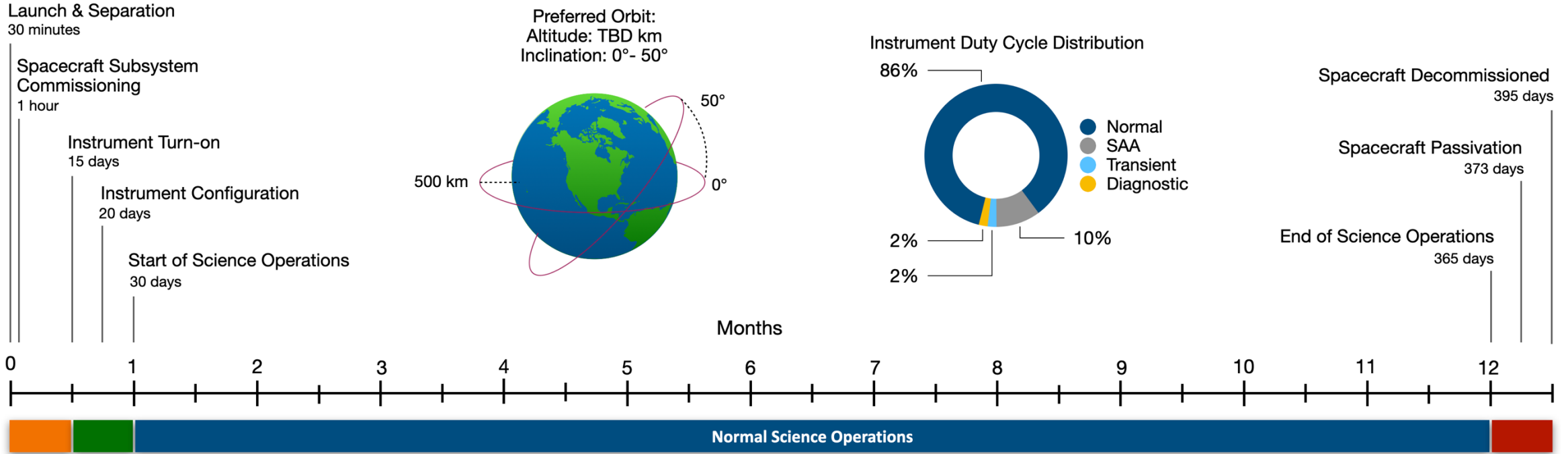


Instrument Sensitivity and Localization

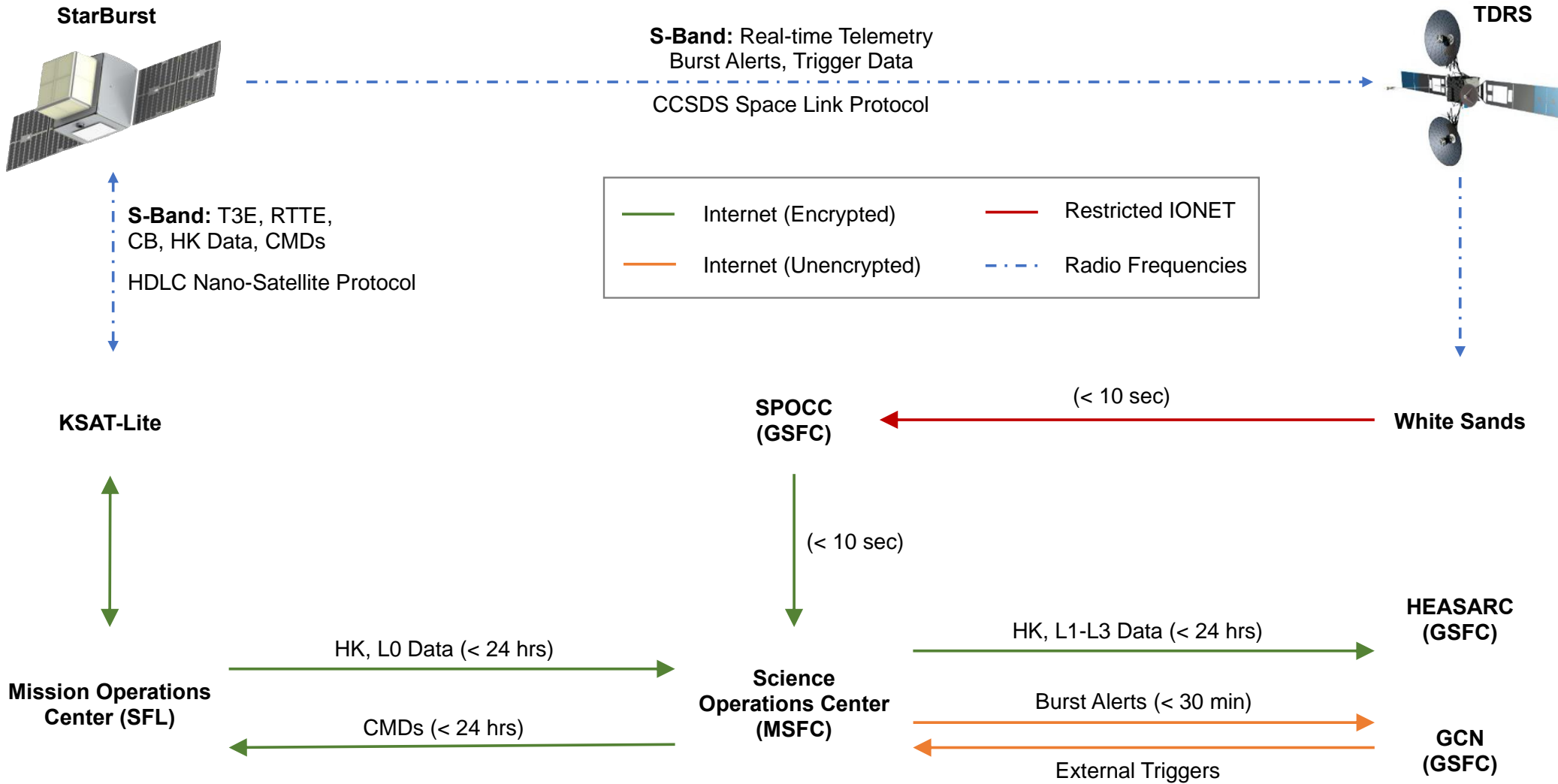


- GRB 170817A could be localized to within 3 deg, compared to ~ 10 deg with GBM
- Majority of bursts will always be near our detection threshold, so average will be same as GBM

Concept of Operations



Concept of Operations - Communications



Baseline and Threshold Mission

- Baseline Mission
 - A baseline science investigation requires detecting a sample of 140 Short Gamma-ray Bursts (SGRBs) during the A+ observing run would allow us to detect ≥ 5 joint GW-SGRB detections
- Threshold Mission
 - A threshold science investigation requires detecting a sample of 100 SGRBs during the A+ observing run would allow us to detect ≥ 4 joint GW-SGRB detections
 - Descope options are geared towards achieving the threshold mission
- Mission Success Criteria
 - A successful StarBurst mission detects at least 80 SGRB

StarBurst Organizations

- Marshall Space Flight Center
 - GRB astrophysics expertise
 - Science operations experience from BATSE and Fermi-GBM
 - Established connections with the GW community
 - Management and engineering support
- Naval Research Laboratory
 - Extensive space instrumentation expertise
 - Experience building & flying SiPM-based scintillation detectors
- University of Alabama Huntsville
 - Onboard science analysis (flight software) expertise
- University Space Research Association
 - Science operations experience from BATSE and Fermi-GBM
- Space Flight Laboratory
 - Small satellite expertise



Science Team



- The StarBurst science team leverages expertise from several lead institutions in the study of gamma-ray and multi-messenger astrophysics
- The team consists of a diverse group of researchers, including ten early career scientists, with a wide variety of backgrounds
 - Includes both observers and theorists with GRBs, GWs, and kilonova expertise
 - Early career scientists are mentored by team members with decades of experience in building and operating space-based gamma-ray telescopes
- Several science team members are also members of the LIGO collaboration and will facilitate collaboration between the two groups
- Several science team members are long-standing members of the multimessenger GW-GRB working group that led the joint work on GW170817