

Telescope

Follow-up of GW with Fermi-LAT and First Results from O4

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on behalf the Fermi-LAT Collaboration

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Gamma-ray Space Telescope

LVK Observing Runs



Credit: LVK Collaboration

09/09/2024

LVK Observing Runs



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LVK Observing Runs



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LAT GW Follow-up Pipeline



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Implemented Analyses

Fixed Time Interval An independent unbinned maximum likelihood analysis is run for each pixel within 90% probability of the GW map in a fixed time window of 10 ks after the GW trigger

Adaptive Time Interval Similar to the FTI analysis, but the ATI time window is optimized for each pixel separately to get the largest exposure closer to the trigger time

LAT Low Energy Events LLE data (E < 100 MeV) are extracted for each pixel of the GW map in the LAT FoV at the trigger time and the significance of the light curves is estimated respect to the background



• PGWAVE is run over the count map to **discover candidate sources**, followed by a dedicated likelihood analysis if any of these are within the 90% probability of the GW map

LAT GW Public Table

Stanford University

Fermi-LAT Gravitational Waves Table

This page displays the outcomes of the Fermi-LAT automatic follow-up analysis pipeline used to search for electromagnetic counterparts of gravitational waves (GW). For a detailed explanation of the analysis techniques, please refer to <u>2017ApJ...841L..16V</u>. Furthermore, the Fermi-LAT Collaboration has published additional papers on GW events such as <u>GW150914</u>, <u>LVT151012 and GW151226</u>, <u>GW170104</u>, and <u>GW170817</u>

<u>All analysis results presented here should be considered preliminary</u>, unless otherwise stated. If you have any questions, please write to <u>Niccolò Di</u> <u>Lalla</u>.

Click on the following buttons to access the table associated with the corresponding observing cycle:

03

04

Stop your mouse cursor over the table headings to view a short explanation of the columns in the table or check the legend here.

* Information taken from <u>GraceDB</u> (LIGO-Virgo-KAGRA Collaboration).

Trigger Name [*]	Date [*]	Time (UTC) [*]	GraceDB [*]	FAR (Hz) [*]	Highest Probability [*]	Has NS? (%) [*]	Has Remnant? (%) [*]	Has MassGap? (%) [*]	Inst. Coverage (%)	FTI TS max	ATI TS max	Flux UB (erg/ cm2/s)	Analysis report
S240002bg	2024 00 02	14.22.06	Link	2.50.00	PPH. 100.0%			0.0		10.2	0.9	5 20 10	Link $(w02)$
5240902bq	2024-09-02	14:55:00		2.56-09	BBH: 100.0%	0.0	0.0	0.0	0.0	10.2	9.0	5.56-10	
S240830gn	2024-08-30	21:11:20	<u>Link</u>	6.3e-10	BBH: 89.1%	0.0	0.0	0.1	26.0	5.9	9.9	1.3e-09	Link (v02)
S240825ar	2024-08-25	05:51:46	<u>Link</u>	3.2e-09	BBH: 96.5%	0.0	0.0	7.1	83.2	9.2	9.5	2.8e-10	Link (v02)
S240813d	2024-08-13	04:39:13	<u>Link</u>	1.8e-18	BBH: 100.0%	0.0	0.0	0.0	27.4	12.4	11.1	1.4e-09	Link (v02)
S240813c	2024-08-13	03:45:48	<u>Link</u>	2.6e-09	BBH: 99.8%	0.0	0.0	2.8	41.8	19.5	23.4	6.9e-10	Link (v02)
S240807h	2024-08-07	21:45:59	<u>Link</u>	2.0e-11	BBH: 100.0%	0.0	0.0	28.3	39.9	37.1	35.4	5.7e-10	Link (v02)
S240716b	2024-07-16	03:49:00	<u>Link</u>	7.9e-16	BBH: 100.0%	0.0	0.0	0.0	6.6	22.5	19.6	5.7e-10	Link (v02)
S240705at	2024-07-05	05:32:15	<u>Link</u>	7.1e-16	BBH: 100.0%	0.0	0.0	0.0	2.4	7.9	8.1	6.5e-10	Link (v02)
S240703ad	2024-07-03	19:13:55	<u>Link</u>	1.2e-13	BBH: 100.0%	0.0	0.0	0.0	6.9	20.7	18.2	4.5e-10	Link (v02)
S240630t	2024-06-30	10:17:03	<u>Link</u>	1.9e-12	BBH: 100.0%	0.0	0.0	0.0	92.4	22.0	24.7	5.6e-10	Link (v02)
S240629by	2024-06-29	14:52:56	<u>Link</u>	3.2e-10	BBH: 91.5%	0.0	0.0	0.0	0.5	11.0	5.7	4.7e-10	Link (v02)

http://fermigrb.stanford.edu/GWTable/

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Analysis Report Page



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FTI/ATI TS Distribution – Significant GW



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FTI/ATI TS Distribution – Significant GW



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FTI/ATI Flux UB Distribution – Significant GW



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Considering all the GW events followed-up by the LAT



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Considering all the GW events followed-up by the LAT



Considering all the GW events followed-up by the LAT



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Considering all the GW events followed-up by the LAT



Conclusions

- Fermi-LAT successfully followed up around 50% (~1200) of all the GW alerts released during O4 looking for a gamma-ray counterpart:
 - Including all the significant GW events
 - Excluding most of the events with Terrestrial probability >~ 80-90%
- Good and timely high-energy coverage guaranteed by the LAT:
 - About 50% (75%) of the GW events reached a coverage > 90% within 1 (1.5) hour
- Analysis results ready within 12 hours of the GW trigger:
 - So far no significant counterparts beside flaring sources or spurious signals have been detected (by any EM telescopes)
 - Results released on a public webpage for significant events
 - Gamma-ray flux upper bounds (>100 MeV) ~ 10⁻¹⁰-10⁻⁹ erg/cm²/s