

# MEV STUDIES OF HADRONIC EMISSION IN NEARBY SEYFERT GALAXIES

Marco Ajello<sup>1</sup>, Chris Karwin<sup>2</sup>, Kohta Murase<sup>3</sup>

<sup>1</sup>Clemson University, <sup>2</sup>NASA GSFC, <sup>3</sup>Penn State

*On behalf of the Fermi-LAT Collaboration*

# SEYFERT GALAXIES

- Seyferts are a type of active galaxy, characterized by their bright, compact nuclei and strong emission lines in their spectra (Carl Seyfert, 1940)
- Seyferts can be of two types: Sy1 and Sy2 depending on the visibility of the BLR

NGC 1068

10 Mpc



NGC 4945

3.6 Mpc



NGC 4151

15 Mpc



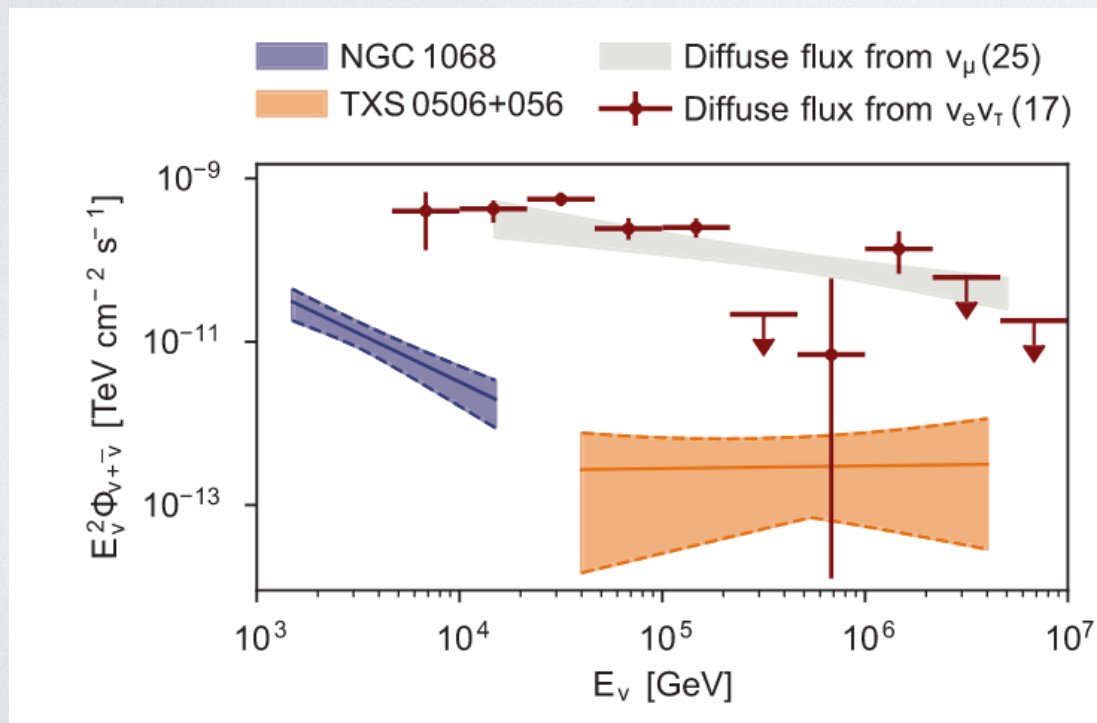
The Circinus Galaxy

4.2 Mpc

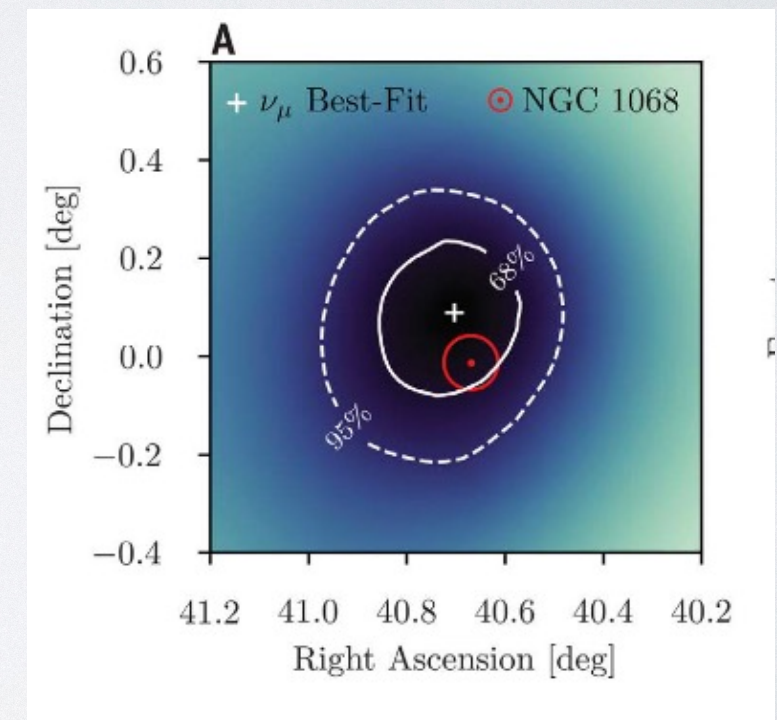


# TeV NEUTRINOS FROM NGC 1068

- IceCube has reported the detection of NGC 1068 at  $4.2 \sigma$  (IceCube Collab. 2022) at 2-20 TeV



NGC 1068



# $\gamma$ AND $\nu$ PRODUCTION

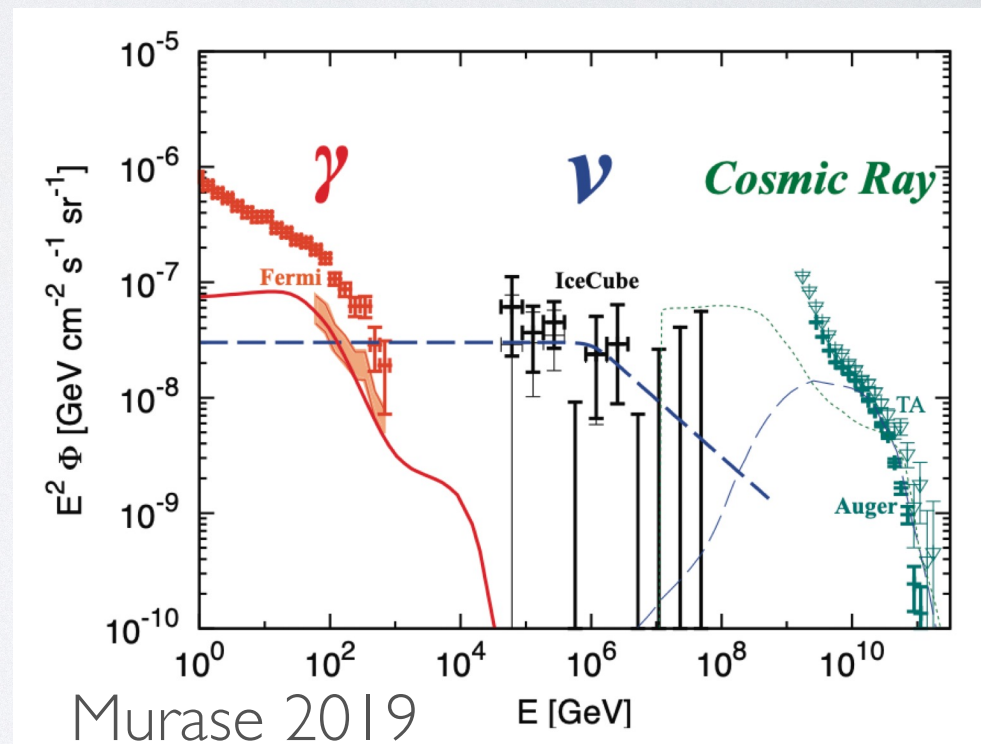
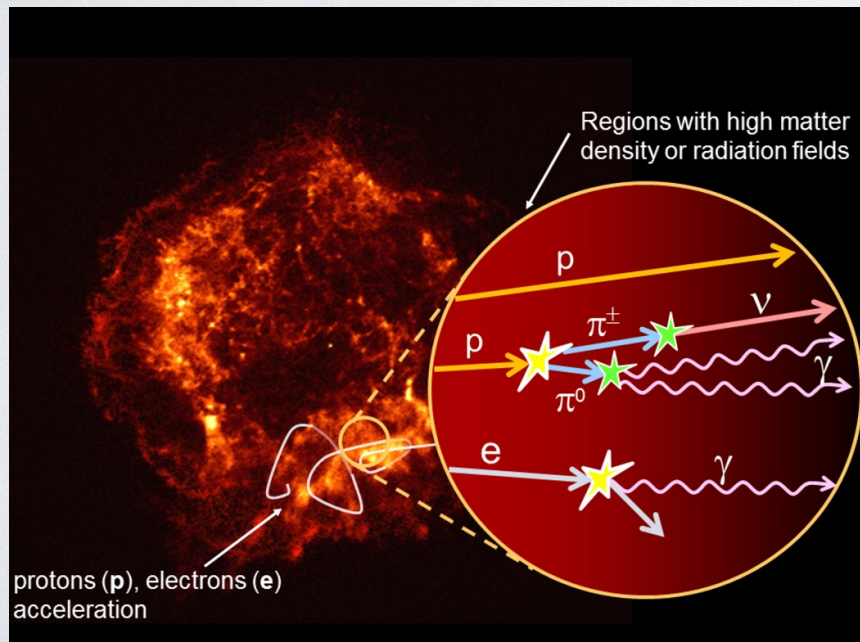
$$p + \gamma \rightarrow n + \pi^+$$

$$\pi^+ \rightarrow \mu^+ + \nu_\mu$$

$$\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$$

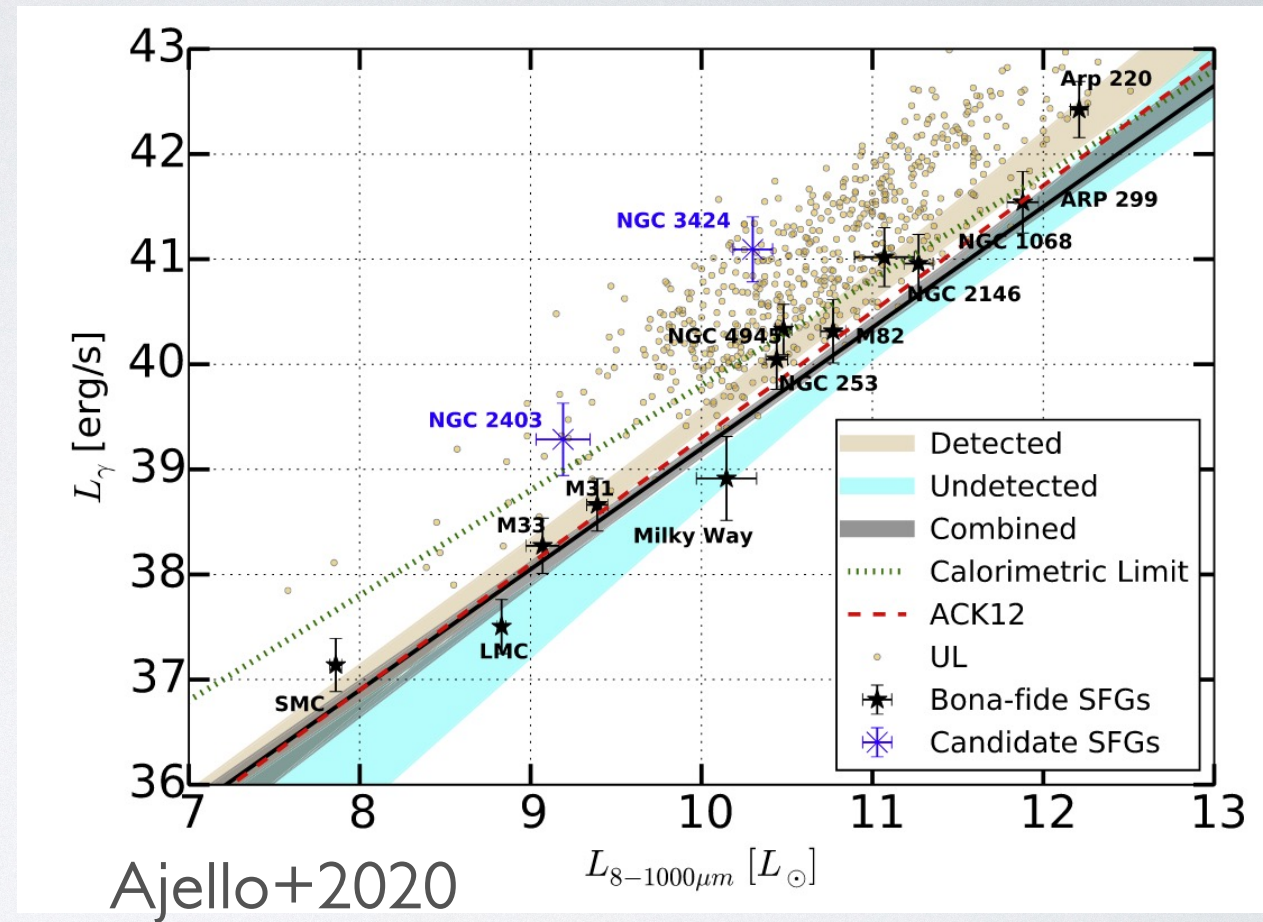
$$p + \gamma \rightarrow p + \pi^0$$

$$\pi^0 \rightarrow \gamma + \gamma$$



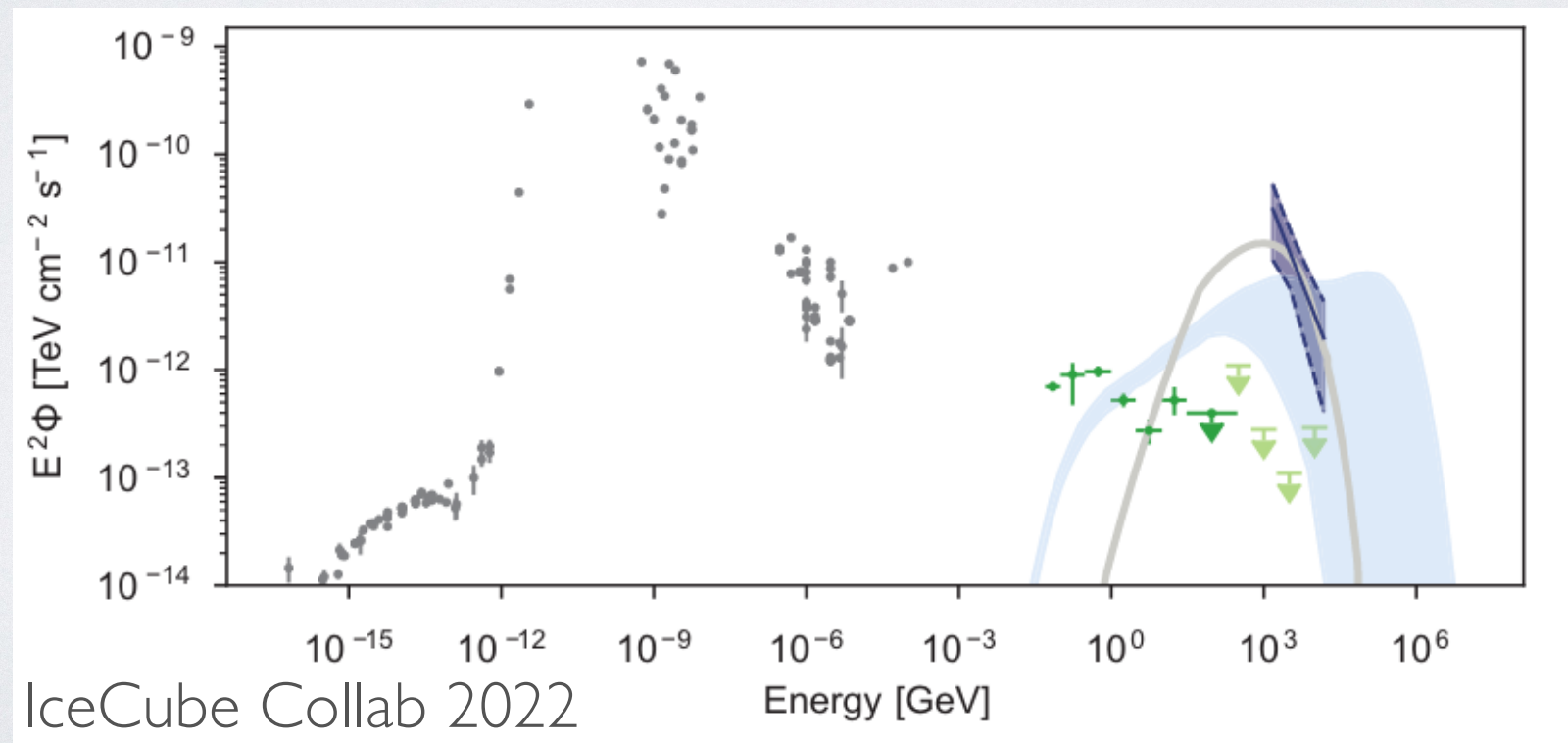
# STAR-FORMING GALAXIES

- GeV emission of star-forming galaxies correlates well with tracers of star-formation activity (IR and 1.4 GHz luminosities)
- No variability has so far been observed, consistently with the hypothesis that this emission is powered by star formation

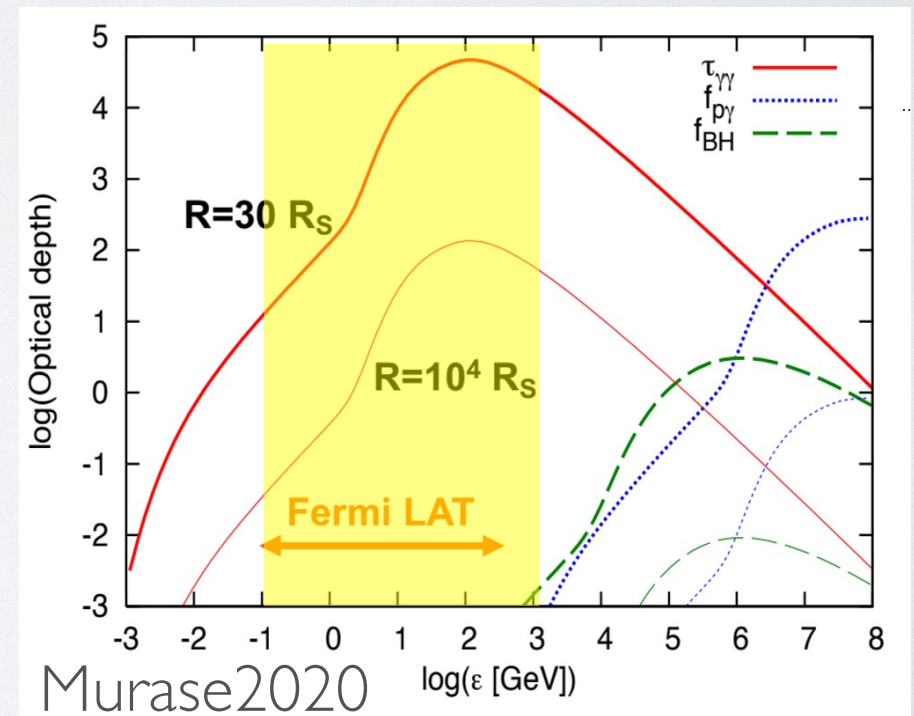


# GEV/TEV OBSERVATIONS OF NGC 1068

- At  $\sim 1$ -2 TeVs the neutrino source is  $> 10$  times brighter than the  $\gamma$ -ray one
- It implies the neutrinos are generated in a region of the galaxy which is opaque to  $\gamma$  rays

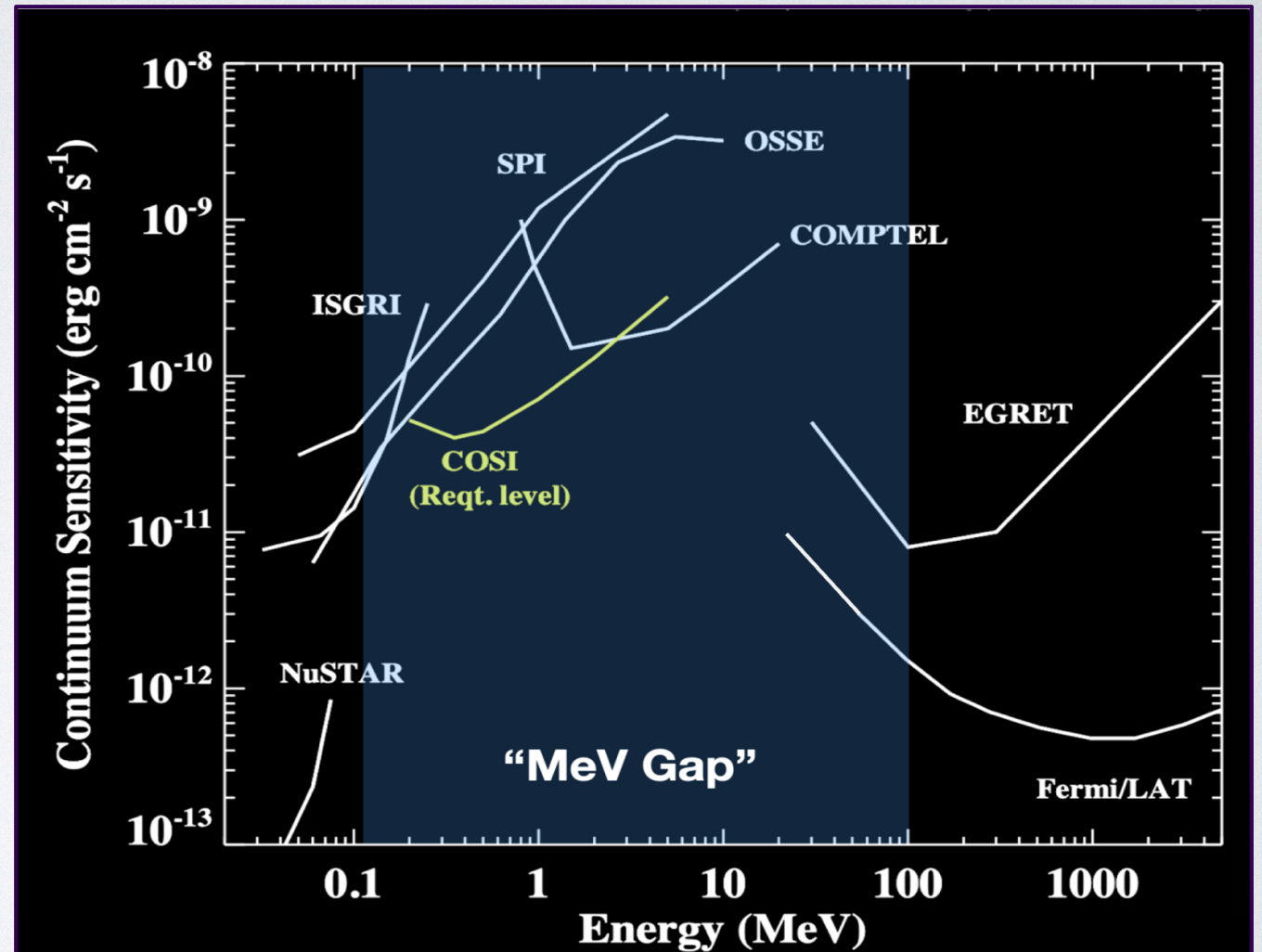


# WHERE TO HIDE A VERY BRIGHT SOURCE?



# SOLUTIONS

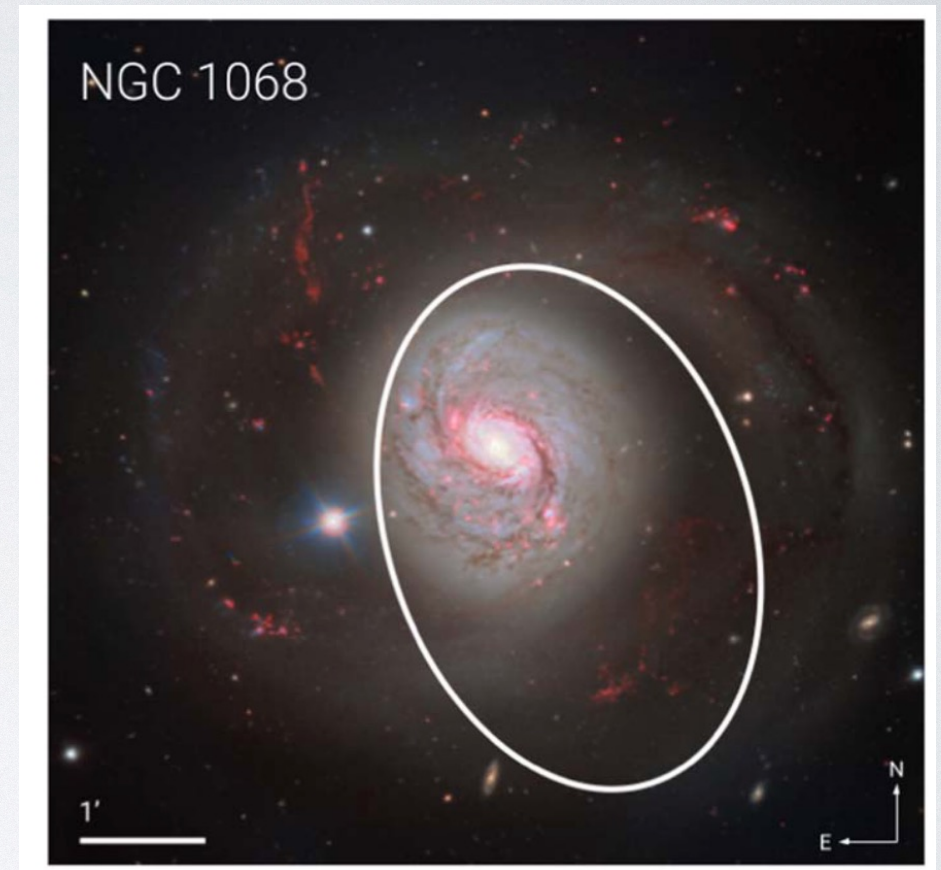
1. Wait for the launch of COSI or future MeV missions
2. Take advantage of Fermi-LAT's low-energy response





# NGC 1068: ANALYSIS

- 14.3 yr of P8R3\_Source\_V3 data
- $Z_{\text{max}} > 90$  for 50 MeV – 1 TeV
- $Z_{\text{max}} > 80$  for 20-50 MeV
- 4FGL DR3 background sources



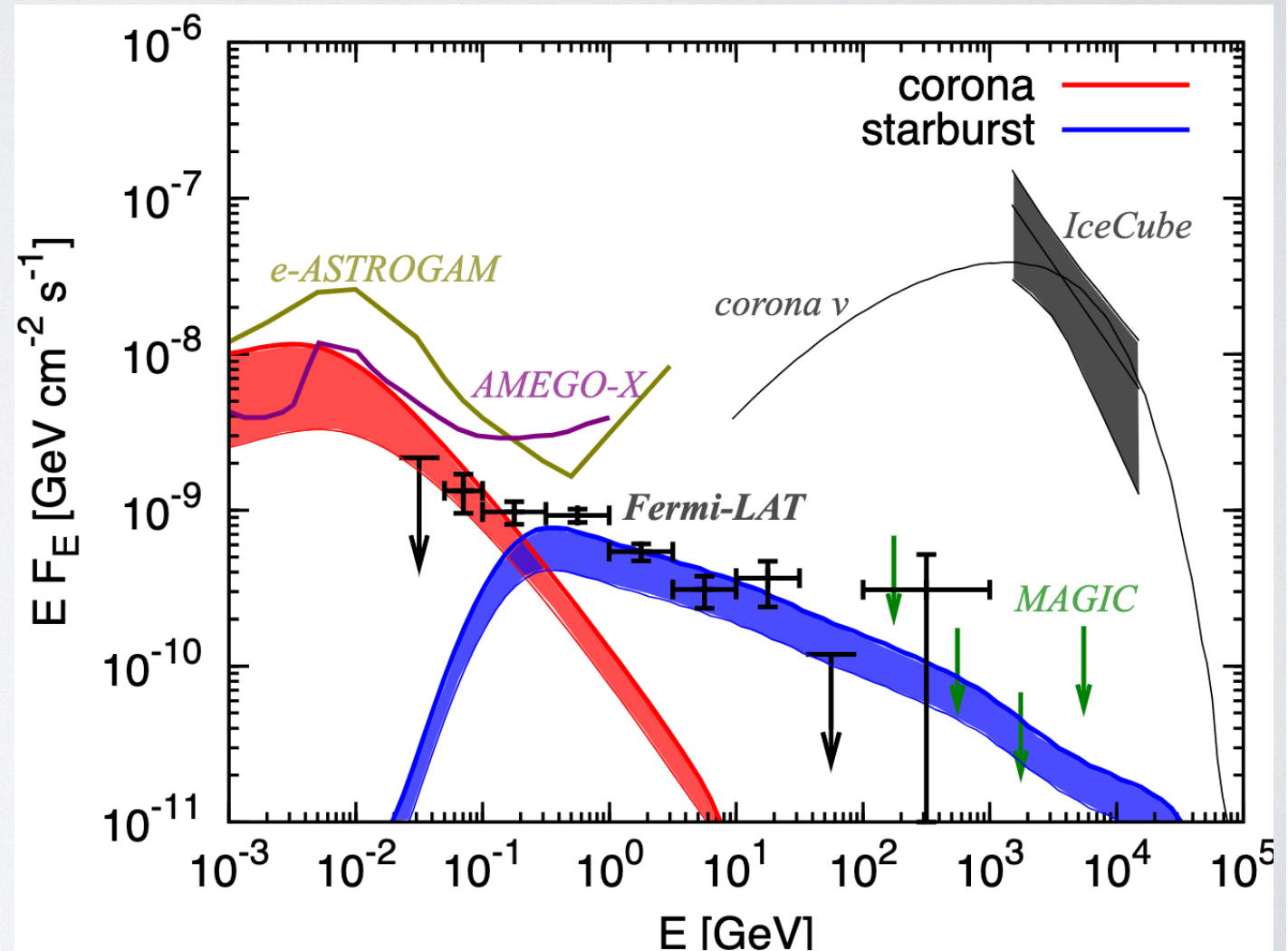
Ajello+2023

*The experience around the 2FLE (Joffre's talk) made this analysis possible*

# NGC 1068: HIDDEN AGN CORE

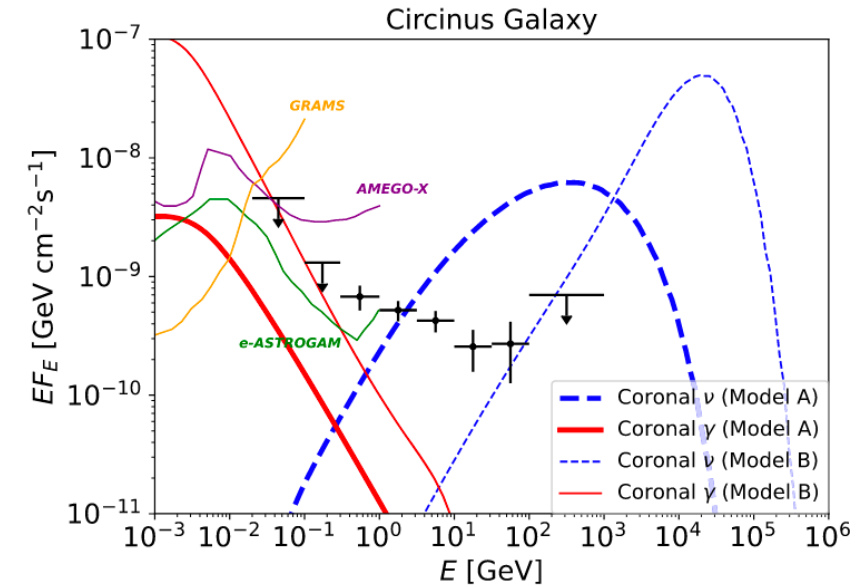
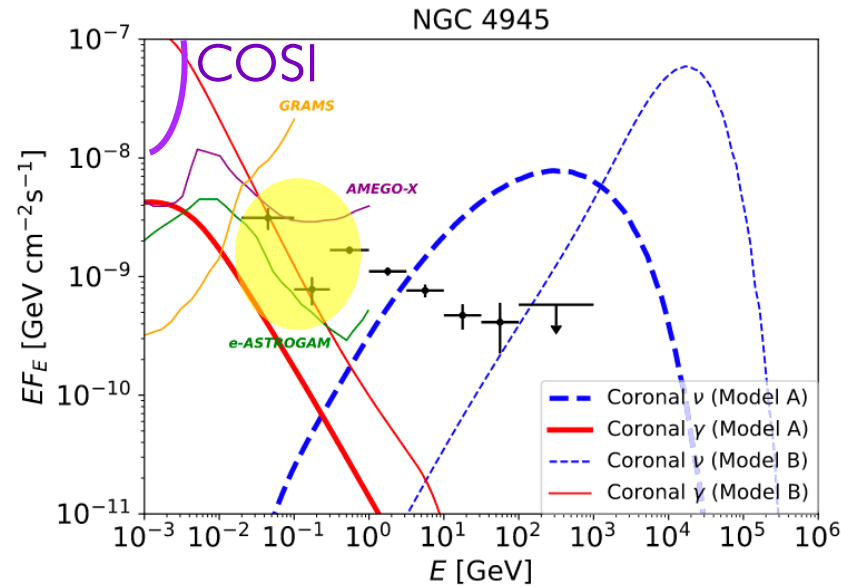
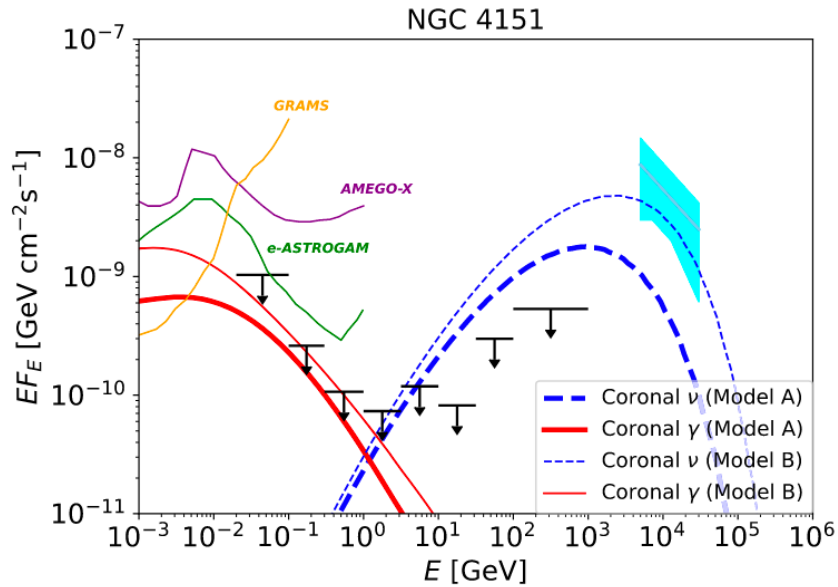
Ajello, Murase & McDaniel 2023

- The starburst component explains well the  $>300$  MeV spectrum, but cannot reproduce the Fermi-LAT low-energy data
- Hadronic emission in the AGN core reproduce the  $<300$  MeV data



# SIMILAR GALAXIES

Murase, Karwin et al. 2024



- NGC 4151, NGC 4945 and Circinus may also have low-energy data compatible with the Corona model (Murase, Karwin+24)
- NGC 4945 may be detectable by COSI

# COSI LAUNCHES IN 2027

- COSI (PI J. Tomsick) has been selected by NASA as a SMEX to launch in 2027
- 0.2-5 MeV
- Cryogenically-cooled germanium detectors to provide excellent energy resolution (0.2-1%)
- Instantaneous field of view  $>25\%$ -sky and covers the whole sky every day



**THANK YOU !**

