## The Arecibo 305-m Telescope in VLBI

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## Proposing VLBI with Arecibo:

Arecibo brings uniquely high sensitivity to VLBI.
The $305-\mathrm{m}$ telescope is available for observing with;

- the VLBA
- the HSA (High Sensitivity Array)
- the GLOBAL array
- the EVN (of which Arecibo is an associate member)

Submit proposals to the usual addresses and deadlines for these Networks, giving special justification for the inclusion of Arecibo.


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Telescope Specifications:

- Diameter: 305 m (1000 ft)
- Longitude: $66^{\circ} 45^{\prime} \mathrm{W}$; Latitude: $18^{\circ} 20^{\prime} \mathrm{N}$
- Declination Range: $-1^{\circ}<$ Dec $<+38^{\circ}$
- Frequency Range: $300 \mathrm{MHz}-10 \mathrm{GHz}$
- Surface Accuracy: $\sim 2 \mathrm{~mm}$ (rms)
- Pointing Accuracy: $\sim 5 \operatorname{arcsec}$ (rms)


## Recent History:

1997-2001: 44 runs in HALCA (VSOP) mission. 2001: Received VLBA4 acquisition system allowing data rates up to 512 Mbps .
2004: Delivery of Mark-5A disc recorder permitting 1 Gbps recording.
2005: First transatlantic eVLBI (32/64 Mbps); longest eVLBI baseline yet (Arecibo -- Torun, PL). 2006: NAIC joins EVN EXPReS program, aiming at 1 Gbps recording by 2009.
2007: Improved internet connectivity (256/512 Mbps) expected by May.

Available Receivers:

- 312 - 342 MHz ( $10 \mathrm{~K} / \mathrm{Jy}$ )
- 422 - $442 \mathrm{MHz}(11 \mathrm{~K} / \mathrm{Jy})$
- 1.1 - $1.7 \mathrm{GHz}(10 \mathrm{~K} / \mathrm{Jy})$
- 1.8 - $3.1 \mathrm{GHz}(10 \mathrm{~K} / \mathrm{Jy})$
- $3.0-4.0 \mathrm{GHz}(9 \mathrm{~K} / \mathrm{Jy})$
- 3.9 - 6.1 GHz ( 8 K/Jy)
- 5.9 - $8.1 \mathrm{GHz}(6 \mathrm{~K} / \mathrm{Jy})$
- 7.8 - 10.2 GHz ( $5 \mathrm{~K} / \mathrm{Jy}$ )
(System Sensitivity in $\mathrm{K} / \mathrm{Jy}$ is given in parentheses.)
System Equivalent Flux Densities are in the range SEFD $=2.5-5 \mathrm{Jy}$ For more details, see; http://www.naic.edu/~astro


GRB 030329 at $8.4 \mathrm{GHz}, 83$ days after the burst. This used VLBA, Arecibo, phasedVLA, and phased-WSRT. The resolution is $0.5 \times 1.0$ mas, and the peak flux density 3 mJy. (Taylor, Frail, Berger \& Kulkarni.)


[^0]:    Sy-2 nucleus of LIRG. NGC 7674, observed with VLBA + Arecibo. HI absorption found towards the C, E \& NE components, but not towards J, W \& SW, reveals a rotating disk/torus associated with the central AGN. (Momjian, Romney, Carilli \& Troland)

