

# GLAST Large Area Telescope Multiwavelength Planning

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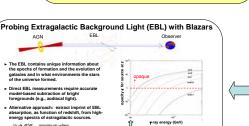


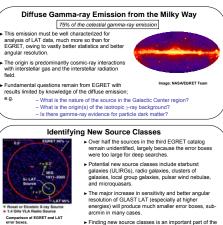
#### Abstract

Because gamma-ray astrophysics depends in many ways on multiwavelength studies, the GLAST Large Area Telescope (LAT) Collaboration has started multiwavelength planning well before the scheduled 2007 launch of the observatory. Some of the high-priority needs include: (1) radio and X-ray timing of pulsars; (2) expansion of blazar catalogs, including redshift measurements (3) improved observations of molecular clouds, especially at high galactic latitudes; (4) simultaneous broad-spectrum blazar flare measurements; (5) characterization of gamma-ray transients, including gamma ray bursts; (6) radio, optical, X-ray and TeV counterpart searches for unidentified gamma-ray sources. Work on the first three of these activities is needed before launch

The GLAST Large Area Telescope is an international effort, with U.S. funding provided by the Department of Energy and NASA

Salamon and Stecker, 1998





Finding new source classes is an important part of the discovery potential of the LAT.

#### Physics in the Extreme Environments of Pulsars

Pulsars – rotating neutron stars - are sites of interactions in extreme gravitational, electric, and magnetic fields.

YY → @\*@\*. maximum when

See. (eV) ~ ½ (1000 / E.(GeV))

- A key to deciphering these extreme conditions is having accurate, absolute timing data for many pulsars.
- With the exception of a few Xray pulsars, the radio band provides the timing information needed by observations across the spectrum. A sizeable radio timing program is beyond the scope of routine radio pulsar programs.

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(Thompson, 2004). Their diversity shows the need for a larger sample with better detail, including phase-resolved spectra at all wavelengths.

Multiwavelength Observations are Important for GLAST

## Some Goals of Multiwavelength Observations

- Source identification and population studies
- Intensive exploration of the brightest and most variable sources that will allow deep study of the source physics
- Rapid follow-up on transients (e.g. GRBs, blazar flares) GLAST mission is designed to support rapid notification for follow-up
- Understanding the high-energy diffuse emission of the Milky Way

#### SUMMARY OF SOME MULTIWAVELENGTH NEEDS AND PLANNING

#### Science Objective GLAST Provides Multiwavelength Multiwavelength Planning Requirements Activities Differential Measurement of Broadband contemporaneous measurement (vs Z) blazar spectra in simultaneous spectral Cooperate with and expand of extragalactic measurements (radio, optical hand where existing multiwavelength background light to cutoffs are X-ray, TeV) of blazar spectra. blazar and GRB campaigns Z~5.5 particularly around the expected from (e.g. WEBT, ENIGMA, GTN, $\gamma + \gamma_{ehl} \rightarrow e^* + e^$ synchrotron peak Swift) to have the broadest possible coverage during the mission Redshift and afterglow neasurements for GRB Resolve origin of All-sky monitoring narticle acceleration coverage of Radio and optical surveys of Participate with and encourage and emission blazar flares and flat-spectrum radio sources to programs to expand blazar mechanisms in Gamma Rav extend blazar catalogs, catalogs and measure systems with Bursts (GRB) redshifts for flat-spectrum radio including redshift relativistic jets, measurements sources supermassive black holes Reliable model of Mapping of Extend CO surveys to high Promote needed CO and other Milky Way diffuse cosmic ray galactic latitude; tracer observations; work with interactions with observers to reduce data and emission required for survey special directions (eg. accurate source all forms of incorporate into a model of the spiral arms, galactic center) localization and to interstellar matter diffuse gamma-ray emission with optically thin tracer (e.g. facilitate search for C18O) dark matter Search out and Large number of Counterpart searches at all Identify facilities and plan understand new source detections: other wavelengths: proposal strategies for Relatively uniform obtaining observing time classes of gamma-Population studies: rav sources sky coverage: needed to identify gamma-ray Correlated variability; Good positions, sources at other wavelengths: Multiwavelength modeling; Cooperate with existing and energy spectra Contemporary, complete planned monitoring surveys: time histories astronomical catalogs Prepare for use of the many available astronomical catalogs Understand particle Spectra and light Contemporaneous radio and Select pulsar candidates for curves resulting acceleration and X-ray pulsar timing radio timing: work with radio emission from primary observations astronomers to monitor timing mechanisms in nteractions of the of selected pulsars: plan most energetic proposals for X-ray pulsar extreme observations environments of particles rotating neutron stars

### SUMMARY

- The GLAST Large Area Telescope science will be optimized by coordinated multiwavelength observations and analysis.
- GLAST welcomes cooperative efforts from observers at all wavelengths. See http://glast.gsfc.nasa.gov/science/multi/

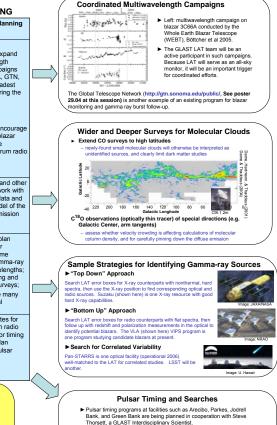
To be added to the Gamma-Ray Multiwavelength Information mailing list. please contact Dave Thompson (dit@egret.gsfc.nasa.gov).

The GLAST Guest Investigator program will have opportunities for developmental and correlative observations. See http://glast.gsfc.nasa.gov/ssc/proposals/

#### **Planning Approach**

The GLAST LAT Multiwavelength Coordination Group (GLAMCOG) has recently been formed to prioritize science-driven needs and develop an implementation plan for cooperative multiwavelength observations before and during the GLAST mission. This work will be coordinated with the GLAST Burst Monitor and GLAST Project science teams.

Some of the known multiwavelength needs are described in this poster, along with the steps being taken to meet those needs. This work is preliminary and does not represent the full range of multiwavelength activities that will be investigated.



After launch, unidentified LAT sources will provide targets for deep pulsar searches. Similar searches will be needed using X-ray telescones

