

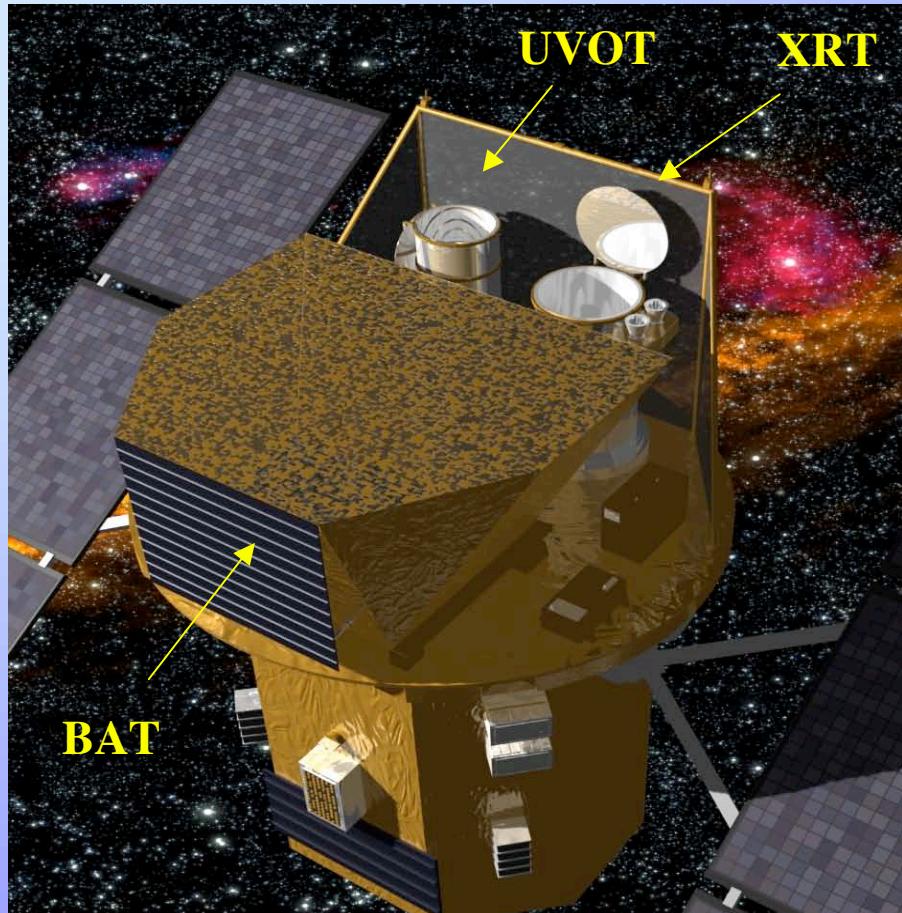


# **Swift Gamma Ray Bursts**

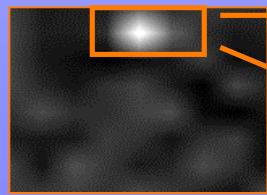
**Neil Gehrels**

**NASA-GSFC**

**July 10, 2009**  
**David Band Symposium**

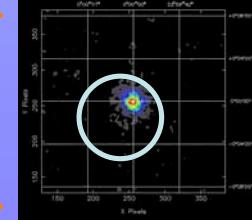


BAT Position - 2 arcmin



*T<10 sec*

XRT Position - 5 arcsec



*T<90 sec*

# Swift Mission

**3 instruments, each with:**

- lightcurves
- images
- spectra

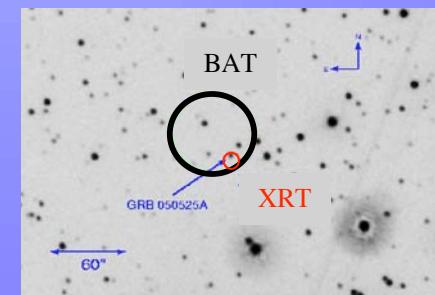
**Rapid slewing spacecraft**

**Rapid telemetry to ground**

**100 GRBs per year**

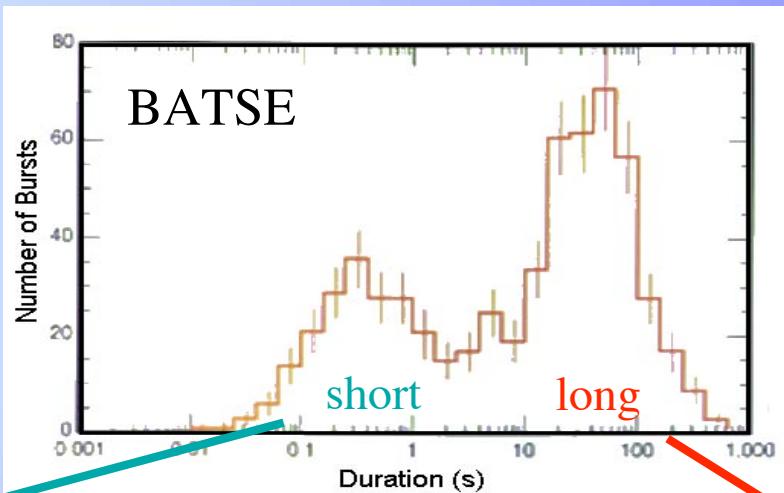
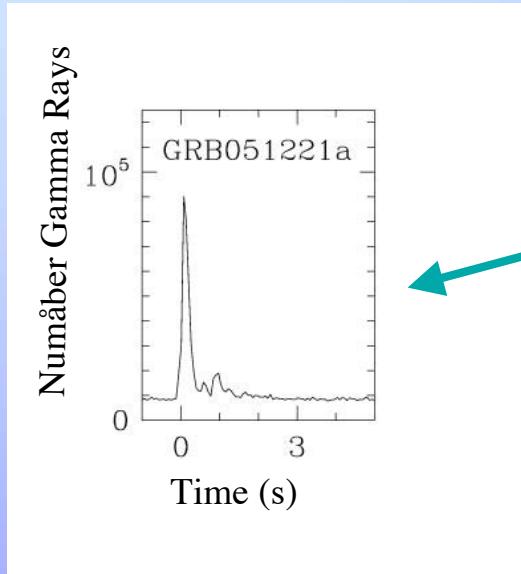


UVOT Position - < 1 arcsec

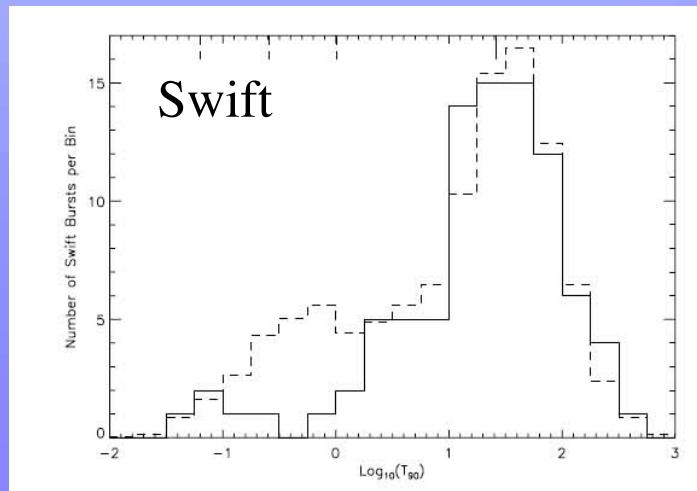
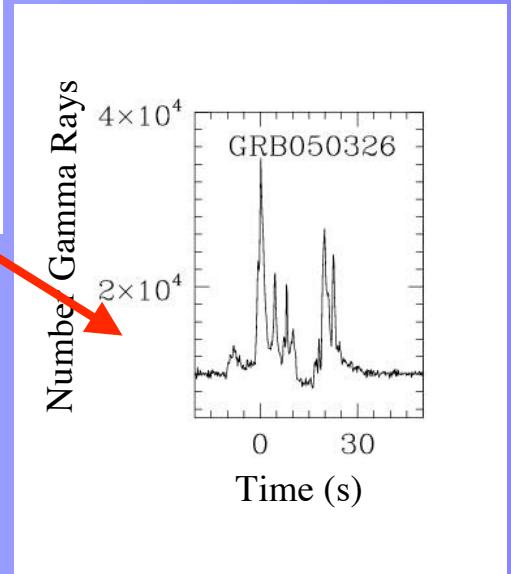


*T<2 min*

Kouveliotou et al.  
2003

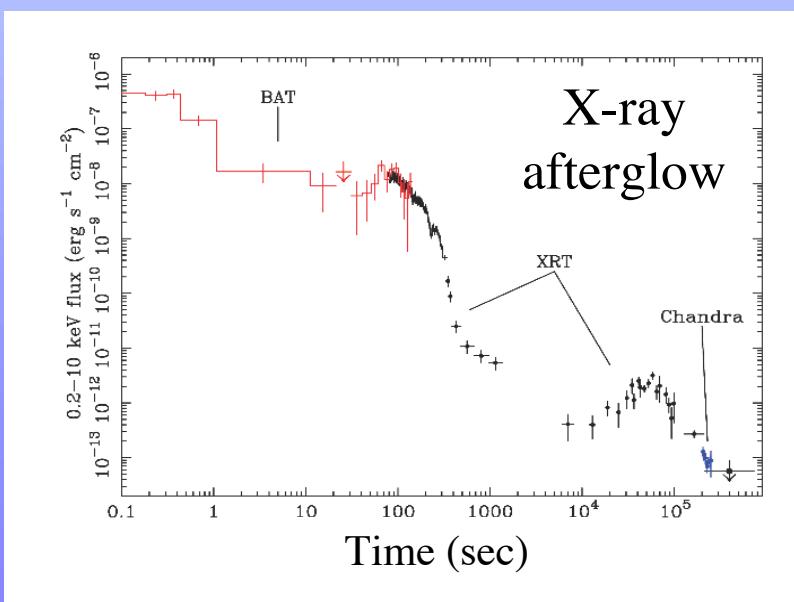
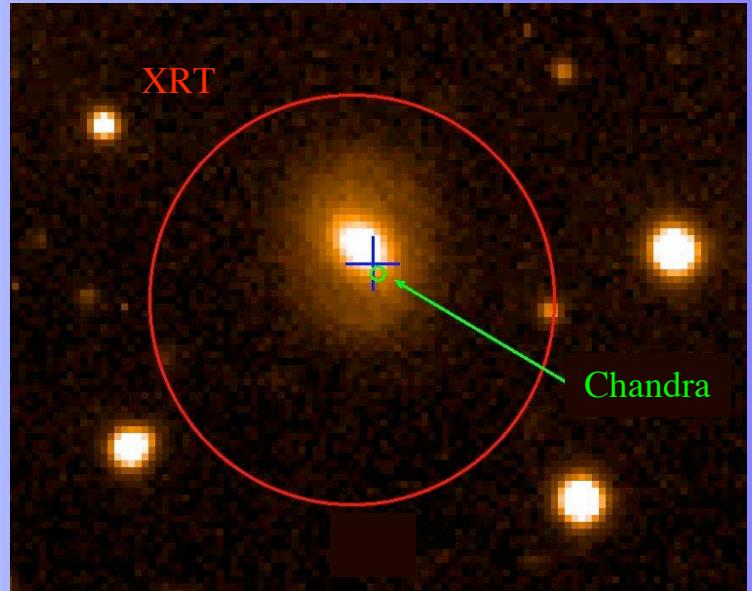
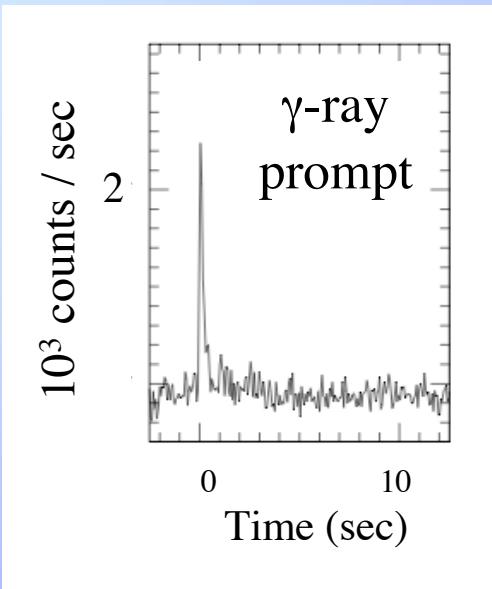


Short  
VS  
Long



Band 2006

# GRB 050724

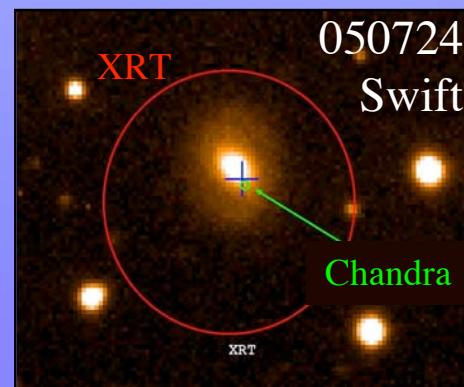
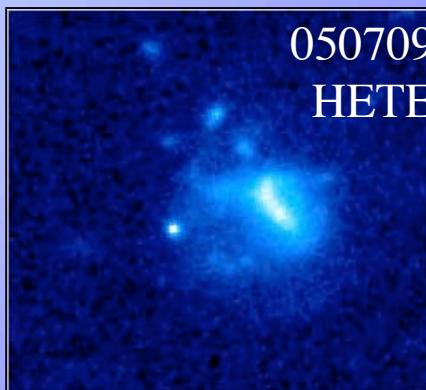
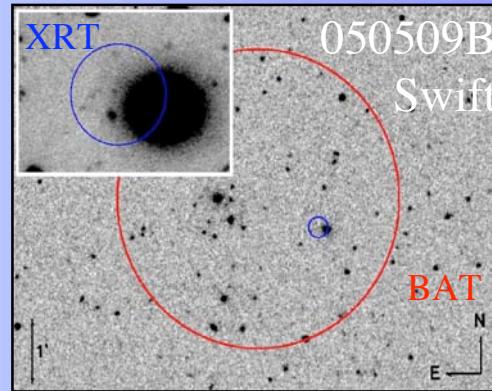
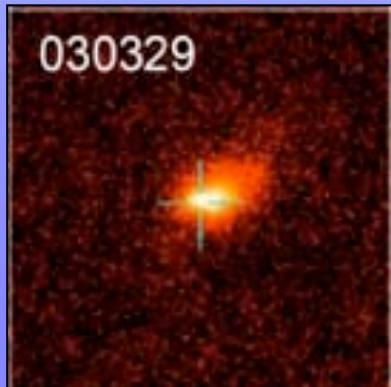
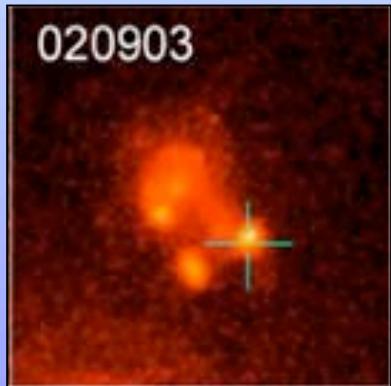
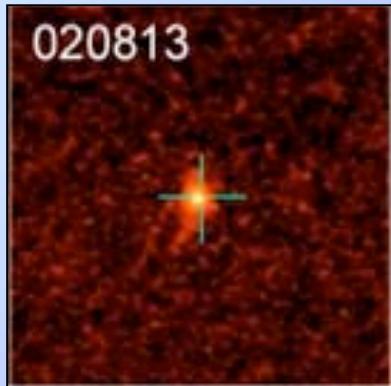


- Host:
- Elliptical
  - $z = 0.258$
  - no coincident supernova
  - $SFR < 0.02 M_{\odot} \text{ yr}^{-1}$

## Long GRBs

## Short GRBs

star forming  
irregulars

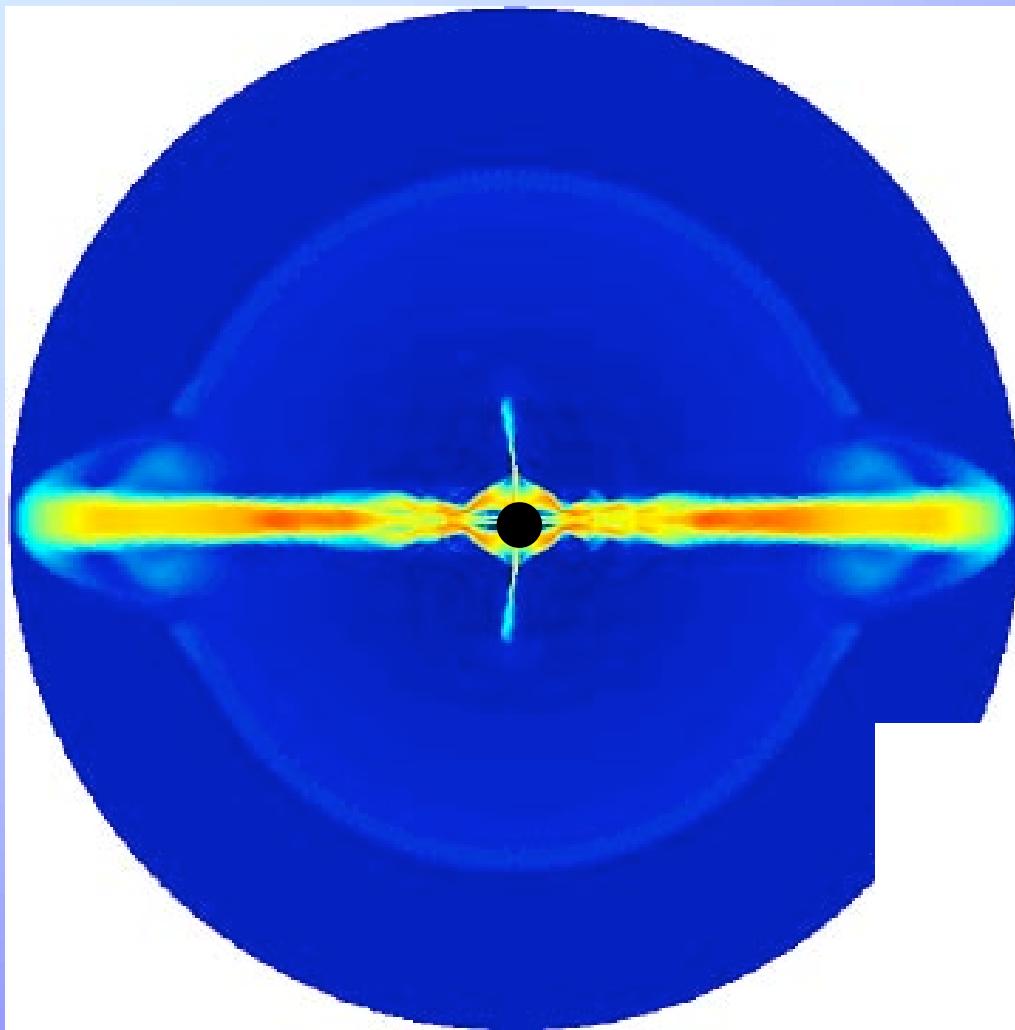


$z = 0.225$   
cD elliptical  
 $SFR < 0.2 M_O \text{ yr}^{-1}$

$z = 0.161$   
star forming galaxy  
with offset

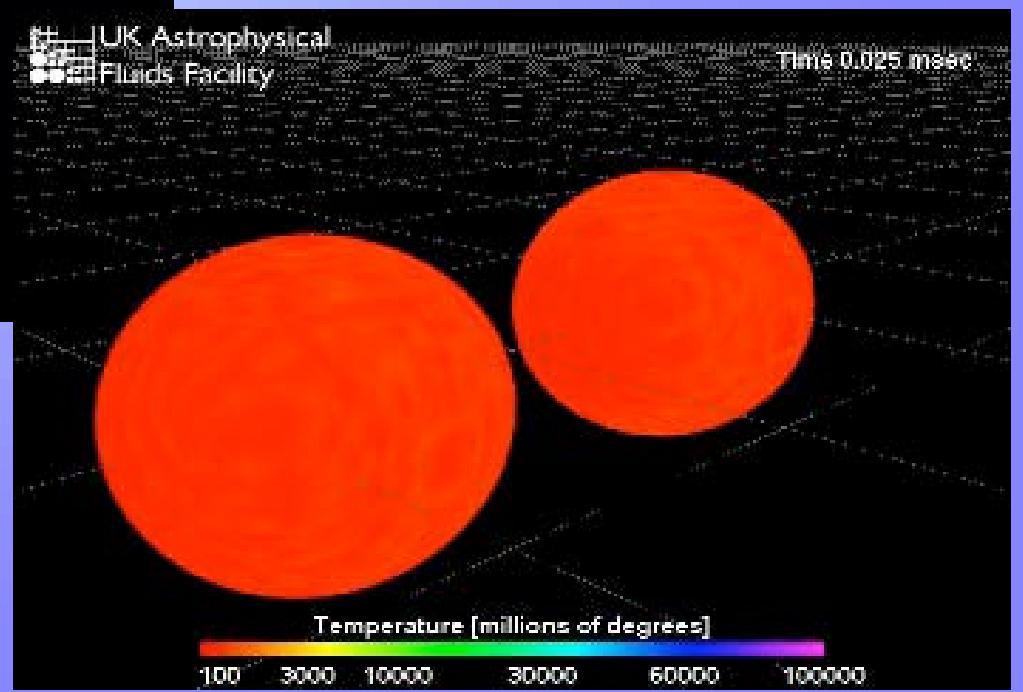
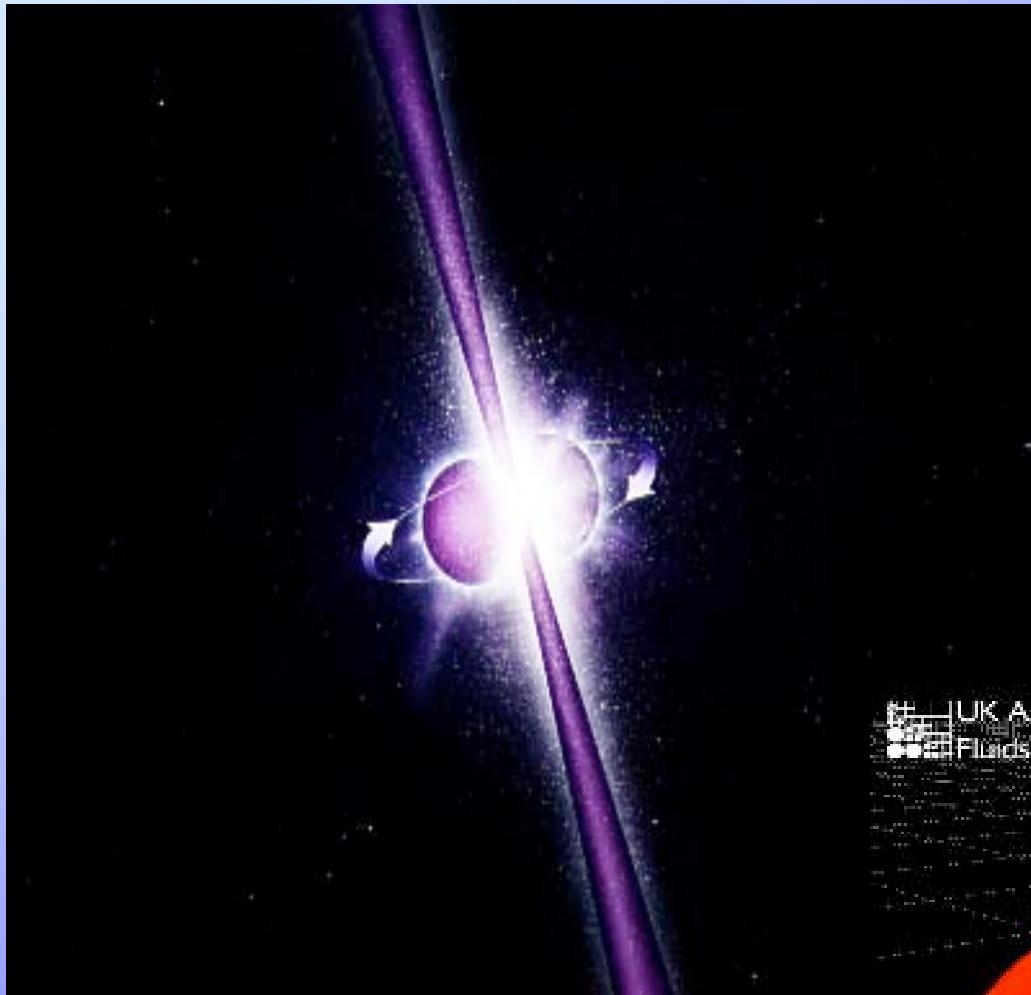
$z = 0.258$   
elliptical  
 $SFR < 0.02 M_O \text{ yr}^{-1}$

# Long GRBs Collapsar Model



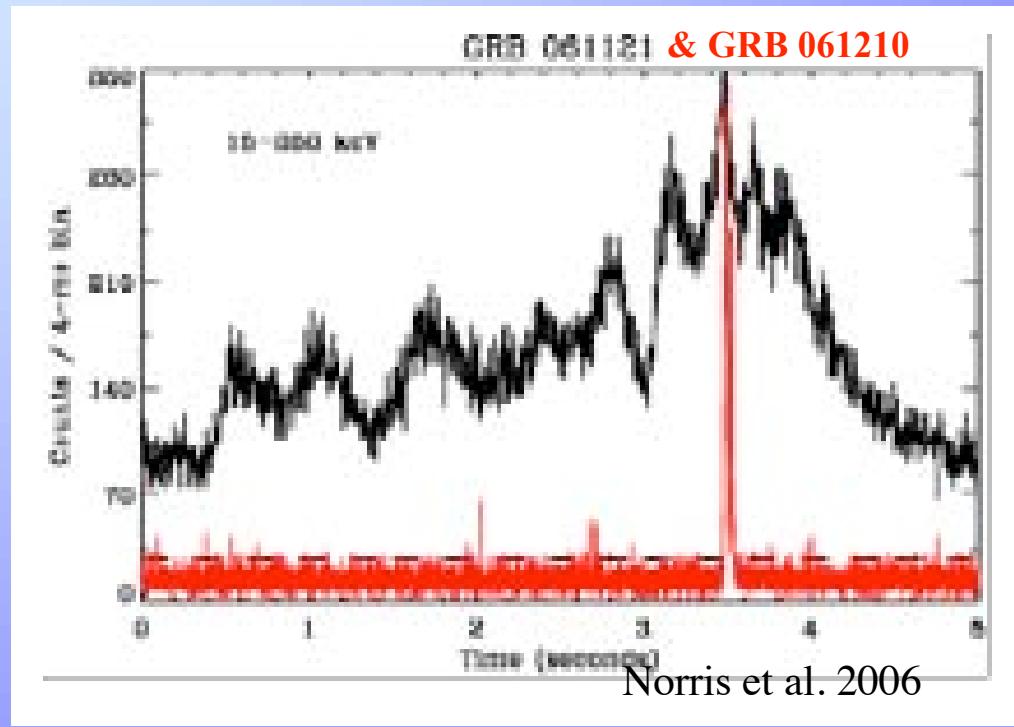
Barkov & Komissarov

# Short GRBs Merger Model



Rosswog et al.

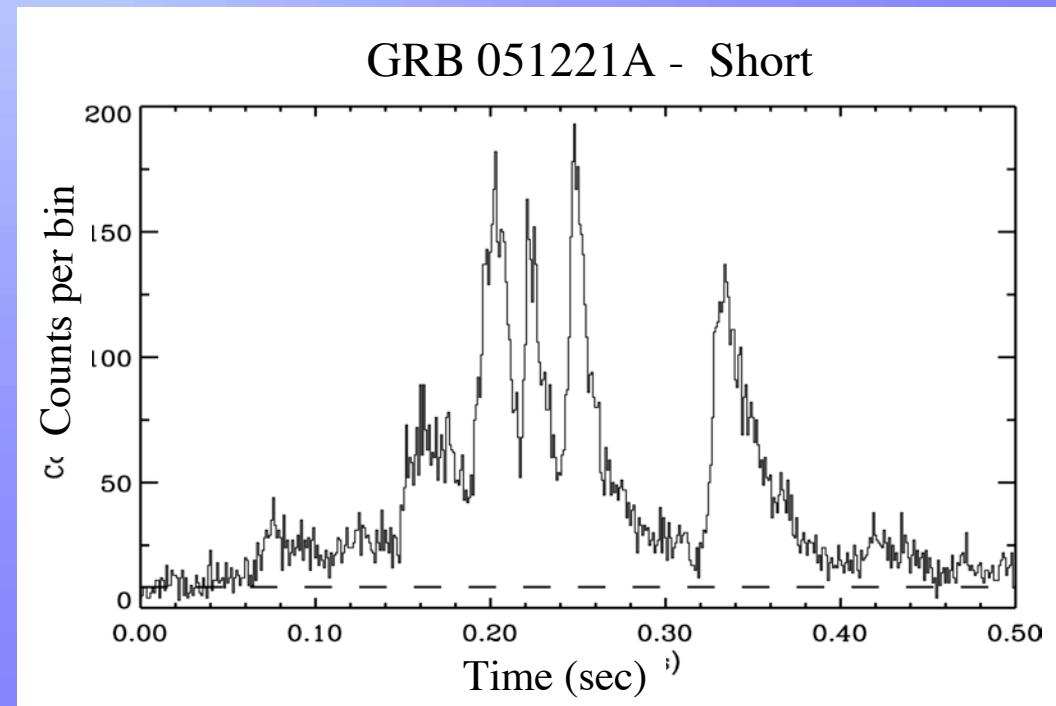
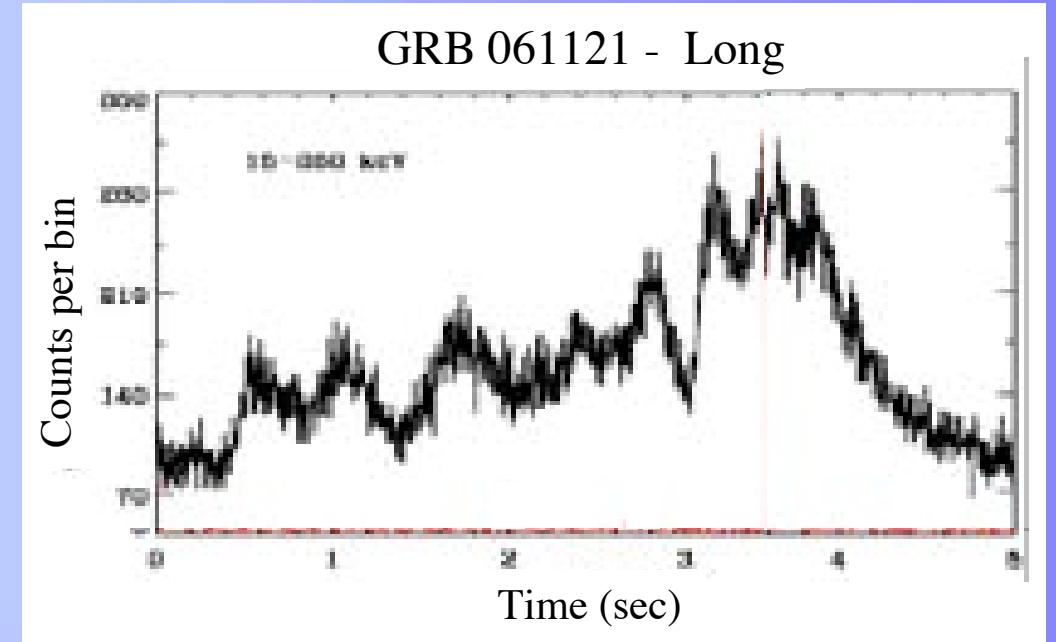
# Variability Comparing Short & Long



**GRB 061121 = brightest long GRB**  
**GRB 061210 = brightest short GRB**

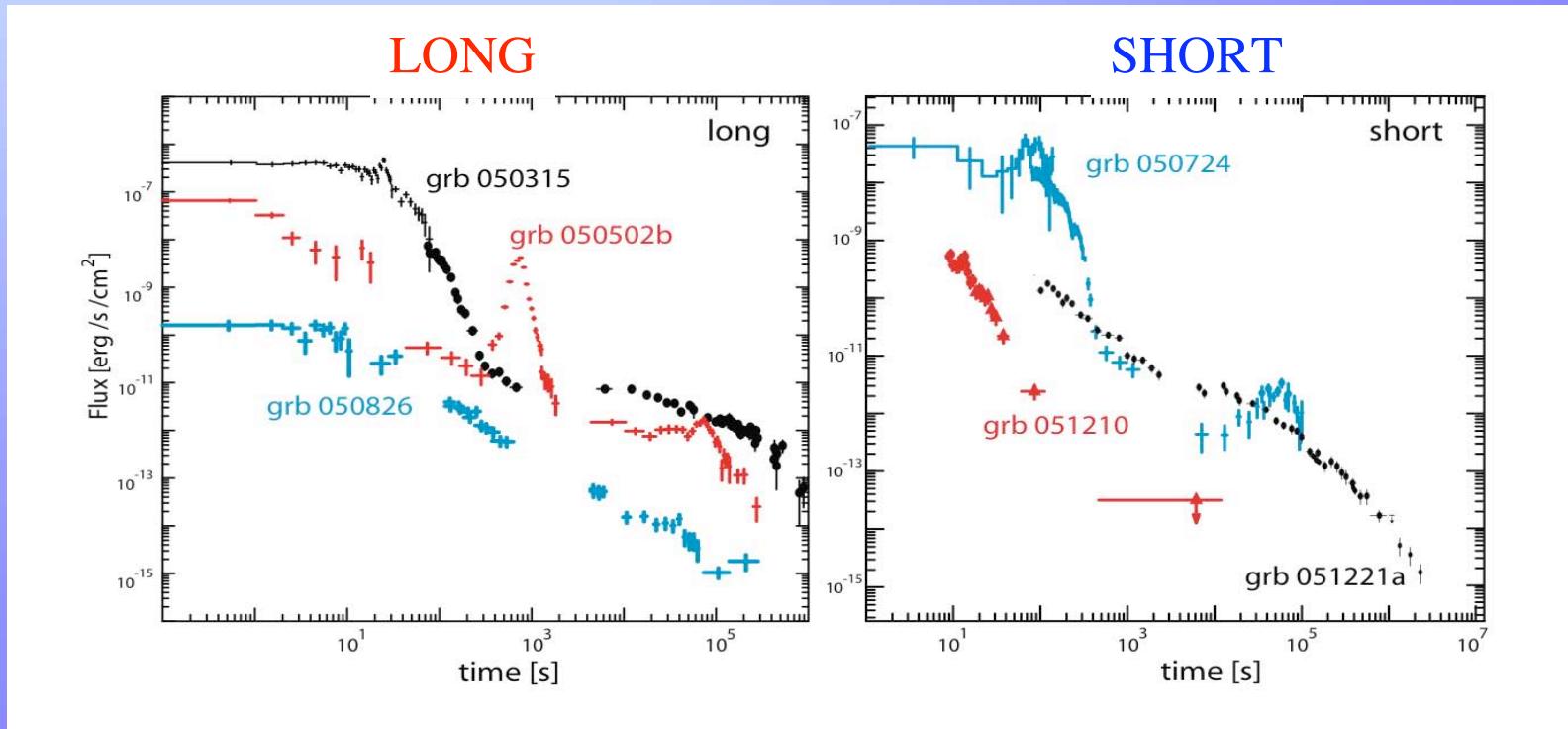
# Variability

Short & long burst  
both have highly variable  
prompt emission



# X-ray Afterglows

- There are various types of light curves for both long and short
- X-ray afterglow is weaker on average for shorts but light curves are generally similar

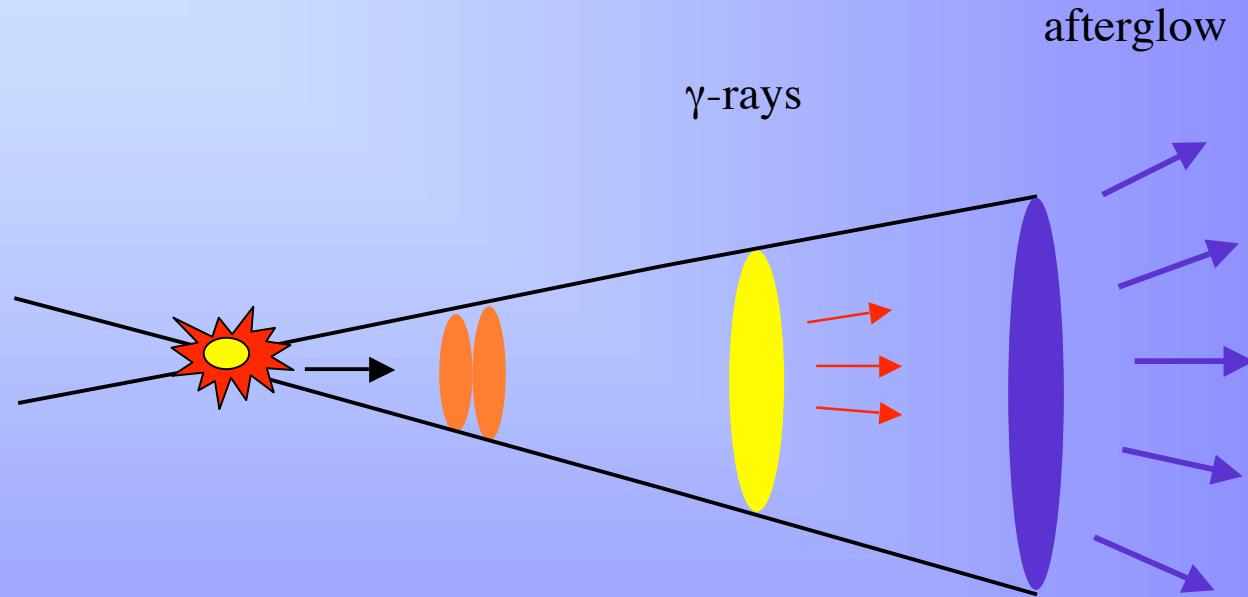


Gehrels, Ramirez-Ruiz & Fox, ARAA 2009

based on Sakamoto 2008

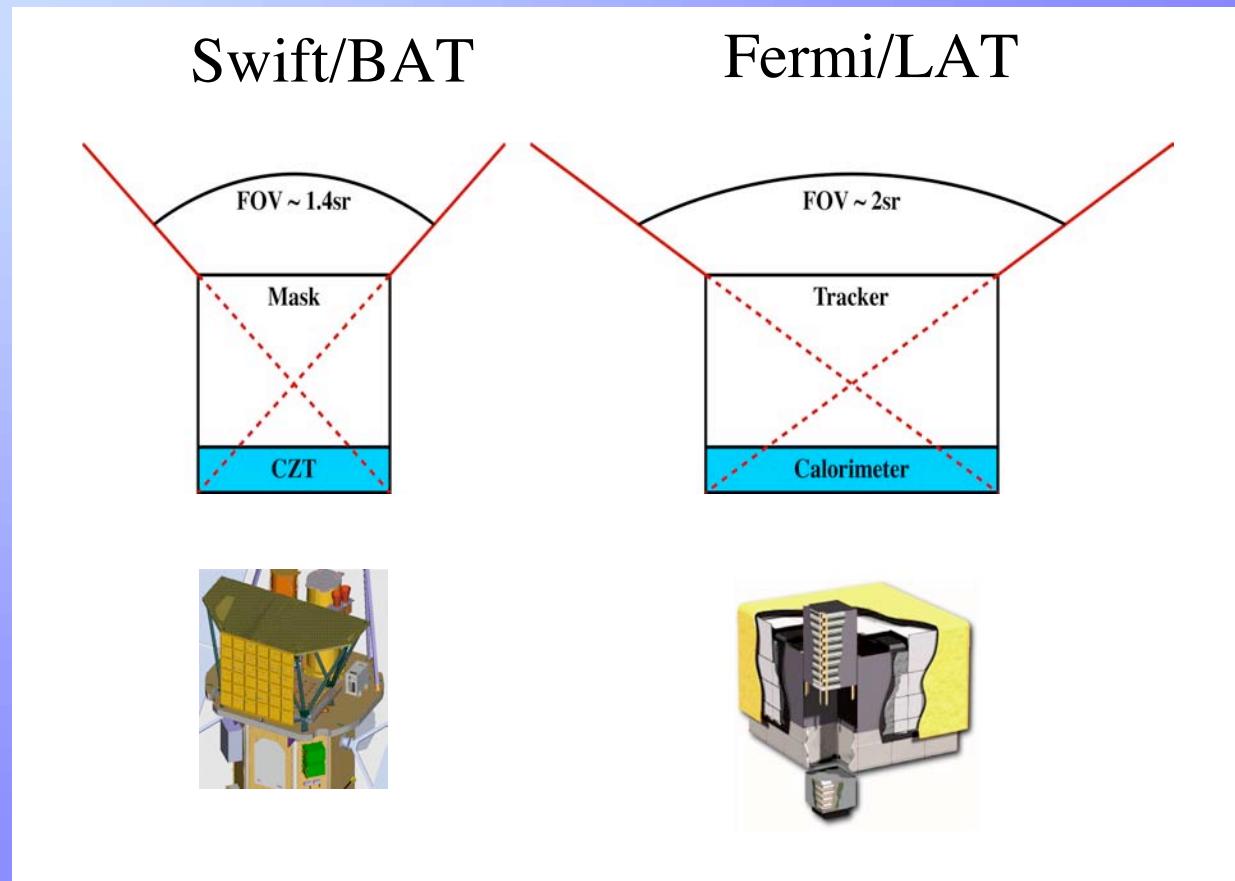
# Fireball Model

(Meszaros & Rees 1997)



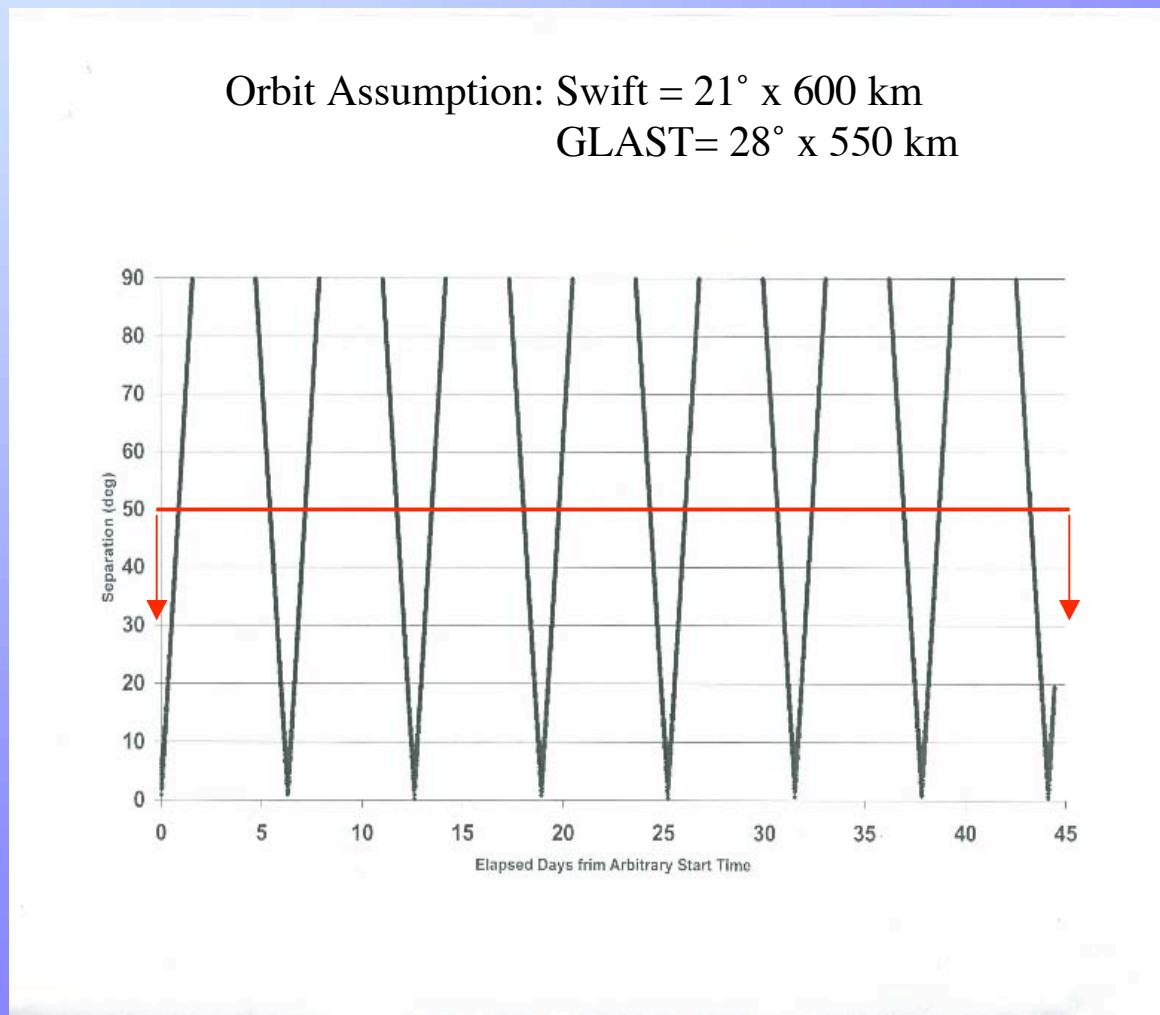
Conclusion: Given their very different origins, the observed characteristics of short & long bursts are remarkably similar.

# David's Idea: Fermi & Swift Working Together



# David's Calculation of Pointing Overlap

## BAT - LAT Joint Pointing



# David's Conclusions

DLB\_overlap\_070726.ppt

## Summary

GLAST

- **GLAST will operate in survey mode. Swift points at 5-6 different targets each orbit. Therefore, overlap will be optimized through Swift scheduling.**
- **Choosing to point Swift at/near the LAT FOV can increase the overlap by >3x! Operating constraints and Swift's other science objectives will reduce the actual overlap.**
- **Because of the large FOVs, GLAST's rocking, and Swift's pointing flexibility, the average overlap fraction is not sensitive to the relative inclination of the orbits.**
- **Therefore, we are currently developing the methodology to optimize Swift's pointing without adding a great additional burden to Swift's scheduling.**

July 26, 2007      GLAST Science Support Center      D. Band GRB Science Group EVO

Slide 2 of 5

# Swift Follow-up of Fermi GRBs

