

GLAST Burst Monitor



The calibration of the GLAST Burst Monitor NaI- and BGO-detectors

by **Andreas von Kienlin¹**, Julia Wagner¹, Helmut Steinle¹, Narayana Bhat², Michael S. Briggs², Roland Diehl¹, Gerald J. Fishman³, Jochen Greiner¹, R. Marc Kippen⁴, Chryssa Kouveliotou⁵, Giselher G. Lichti¹, Charles A. Meegan³, William S. Paciesas², Robert D. Preece², Robert B. Wilson²

[1] Max-Planck-Institut für extraterrestrische Physik (MPE); [2] University of Alabama (UAH); [3] NASA/Marshall Space-Flight Center (MSFC), [4] Los Alamos National Laboratory (LANL), [5] USRA

The GBM-BGO Detectors (2 x): Introduction: The GBM-Nal(TI) Detectors (12 x): The next large NASA mission in the field of gamma-ray astronomy, GLAST, is scheduled for launch in September 2007. Aside from the main instrument LAT, a gamma-ray telescope for the energy range between ~20 MeV and ~300 GeV, a secondary instrument, the GLAST burst monitor (GBM), is foreseen. Its task is to increase ìńń the detection rate of gamma-ray bursts for the LAT and to extend the energy range of the main instrument to lower energies (from The energy range of the main mathined to lower energy range of the main mathined to lower energy range by the GBM consists of 12 thin Na1-plates, which allow the determination of the angle of incidence of the gamma radiation. These crystals are sensitive in the energy range between ~ 10 keV and ~ 1 MeV. Two additional BGO detectors, which are able to detect gamma-rays in the energy range between ~150 keV and ~30 MeV, are responsible for the overlap in energy measurement with the LAT main instrument and the NaI detectors. All flight detectors were already delivered to NASA. This poster gives an overview of the calibration measurements performed by MPE at the detector level before delivery. **Purpose of Calibration:** List of Calibration Sources: Calibration Results: provide performance verification of the GBM detectors.
provide benchmark data to compare with calculated detector The figures below summarize the main measurements performed with the flight detectors. All detectors of the same type showed nearly the same performance and all fulfilled the specified requirements. response data. provide accurate, well-characterized data for subsequent scientific analysis. The Calibration Setup: allowes the acquisition of spectra at different angles of incidence high accuracy due to laser adjustment. **Typical Spectra:** Background-subtracted NaI- and BGO-spectra, recorded at an angle where the detector effective area is maximal (0°). X-ray Transmission of the Nal Entrance Window: The installation of an 0.7 mm thick silicon pad between the The instantion of an 0.7 min tinck sincol pat between the Beryllium entrance window and the crystal was necessary so that the brittle NaI crystal is able to survive the environmental conditions during launch. This additional silicon layer caused unfortunately a degradation of the transmissivity at low energies (at 10 keV ~ 30%). 1111 400 500 600 700 800 900 **Detector Magnetic Susceptibility:** Nal Low Energy Calibration: The effect of magnetic fields on the shielded NaI-EQM module was tested from -2 G to + 2 G in 0.5 G steps at the Performed at the BESSY synchrotron in Berlin with 4 FM-NaIs: Determination of absolute efficiency at low energies (8- 60 keV) and IABG's MFSA facility. <u>Result:</u> The GBM response will not be affected by the magnetic torquers of the GLAST S/C test of response around Iodine K-Edge 40 - 25.5 keVe 11.4 keV (I K_o storage rin 100 argy (keV)

For further information, please contact: azk@mpe.mpg.de