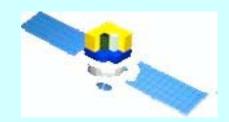


Status report on Light Simulator

Claudia Cecchi Francesca Marcucci Monica Pepe

GLAST Collaboration Meeting

GSFC Washington-DC October 22-25 2002



Why the light simulator...

Very fast simulation of the skymap seen by the GLAST experiment:

√from ~10m for the simplest case (region of 60°, with Poissonian bg,

and orbit simulation)

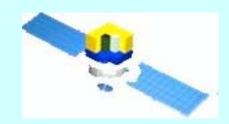
 ✓ to ~36h for the complicated one (all the sky, with the bg convoluted with the PSF and exposure time fixed)
 Generation of a photon list, with photons distributed randomly,

according to PSF, AS and ED distributions (determination of

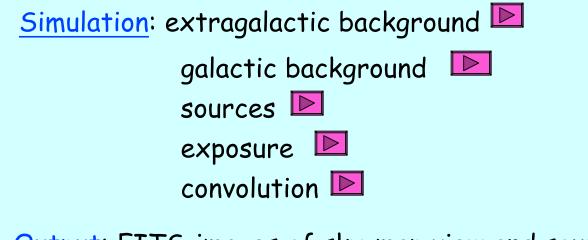
the true and measured energy, the incidence angle, the true and

the measured position of the photons)

The code (C++) is running under Windows and under Linux



<u>Input</u>: map of the galactic background (Egret or Galprop) Third Egret Catalogue with sources photon energy range (0.1 - 100 GeV) region of the sky \$20° - 90°, -180° - 180°) orbit or fix time

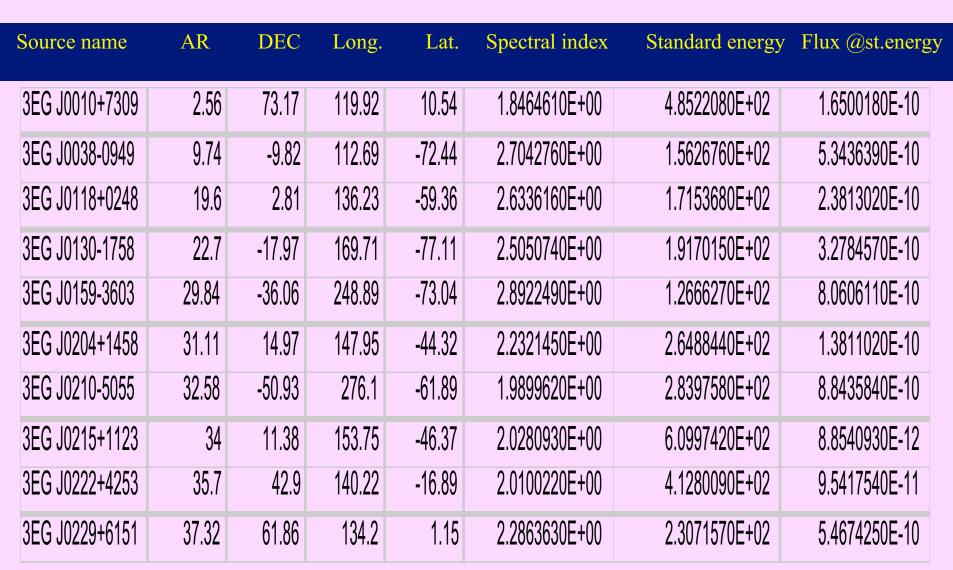


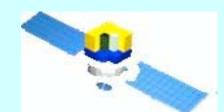


Output: FITS-images of sky-map view and sources



255 sources from the Third Egret Catalogue





The orbit

Elliptic orbit defined by the following parameters:

- a semi-major axis of the ellipse (distance of the orbit from the Earth ≅ 550 Km)
- eccentricity (is set to zero, because we assume a circular orbit)
- P epoch (time of transit to the perigee)
- **T** period (5739 s)
- $m{v}$ anomaly (the angle between the perigee and the point from where we start to calculate the orbit)
- *i* inclination of the orbit with respect to the terrestrial equator (28.5°)
- $\Omega(t)$ the orientation of the semi-major axis of the ellipse

R rocking (35°).

Û

Simulated with a acametrical annroach by etene of 30" (_. 2°) accumina



Diffuse extra galactic background

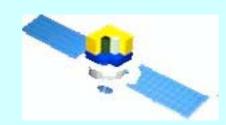
Differential flux of photons from diffuse extra galactic background described (Skreekumer, 1997) as:

$$\frac{dN}{dE} = 7.3 \cdot 10^{-6} \cdot 0.451^{2.1} \cdot E^{-2.1}$$
[1]

(photons
$$\cdot$$
 cm⁻² \cdot s⁻¹ \cdot GeV⁻¹ \cdot sr⁻²)

Total contribution of the extra galactic background obtained by the integration of (1) between E_{min} and E_{max} taken into account only if $|b| > 15^{\circ}$.





The gamma ray galactic background

Galprop simulates the gamma ray background taking into account:

- neutral pion decay;
- bremsstrahlung;
- inverse Compton scattering.

Galprop produces FITS file containing 24 images in the energy range from 0.001 to 10⁴ GeV

EGRET map measured between 0.1 and 30 GeV





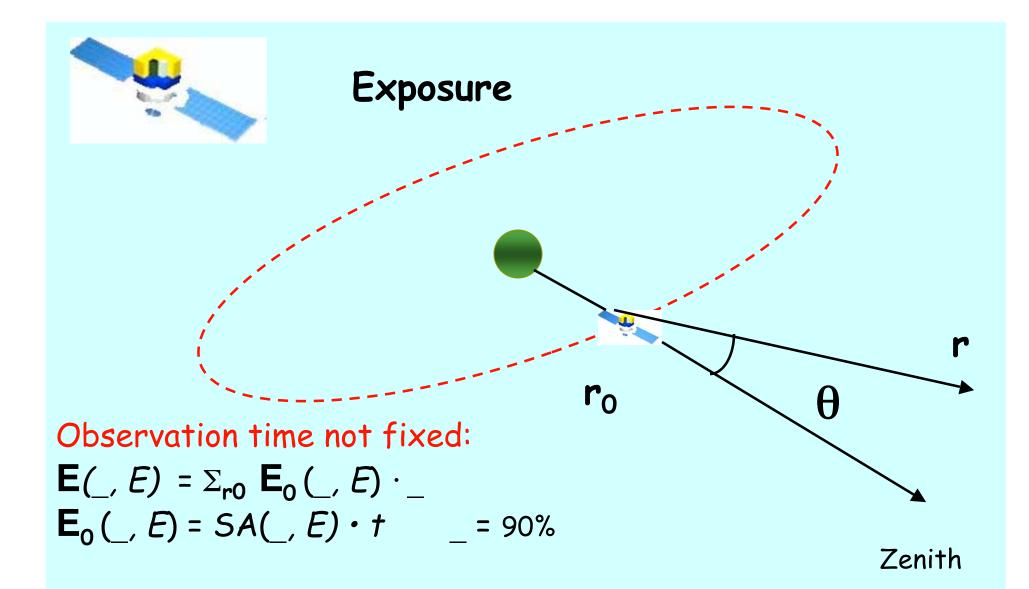
Sources

First approximation: sources point-like located at an infinite distance Intensity of photons: $I(E) = I_0 \cdot E^{-\alpha}$

(photons
$$\cdot$$
 cm $^{-2} \cdot s ^{-1} \cdot GeV ^{-1}$)

Sources information: localization, α and \mathbf{I}_0 from Third Egret Catalogue.





Observation time fixed: $E(, E) = SA(, E) \cdot t \cdot cos$



Convolution with PSF, SA and ED Background and Sources

In each pixel P_0 :

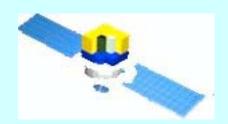
Bkg-differential flux C (P₀) (γ cm⁻² t⁻¹ Ω ⁻¹) = Σ all contributions

the sky in small regions of $0.5^{\circ} \times 0.5^{\circ}$ without detector effects

convolution of our result with the SA, PSF and ED.

$$D(E) = \int_{0-60^{\circ}} \mathbf{E}(\mathbf{E}) \cdot \mathbf{I}_{0} \cdot \mathbf{E}^{-\alpha} d\mathbf{I}_{0}$$

For Bkg: $\alpha = 2.1$ $I_0 = C (P_0) \cdot \Omega$ $\Omega = dl \cdot db \cdot cos(b)$



Generation of a photon list

 $N_{\gamma} = \int_{\Delta E} D(E)$

Each γ randomly distributed in the map with:

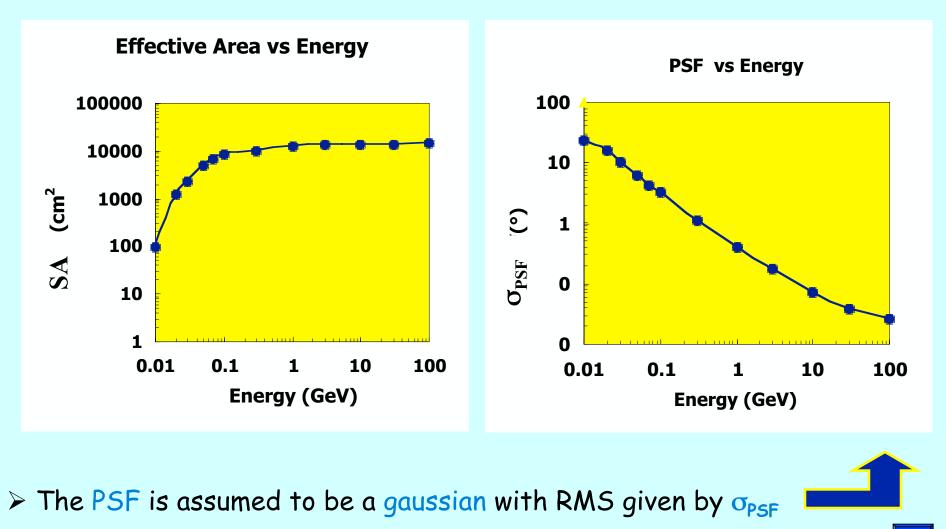
*random energy E_{true}∈[E_{min}, E_{max}] according to D(E)
*inclination angle _ distributed according to E(_, E_{true})
*E_{meas} obtained using the function ED(E_{true}, E_{meas})
*angular distance, _, from origin point P₀ = (I₀, b₀) consistent with PSF(ρ, E_{true})

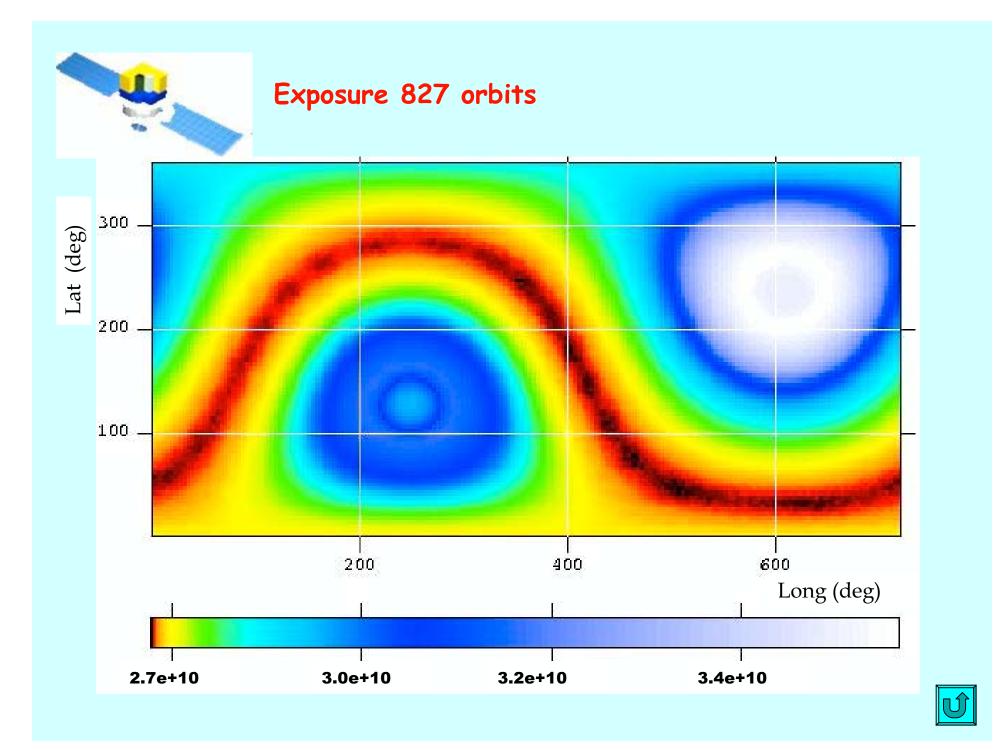
final position of each photon is $P_1 \equiv (I_1, b_1)$ obtained considering that the total angular distance from P_0 is ρ





PSF and SA as function of the Energy

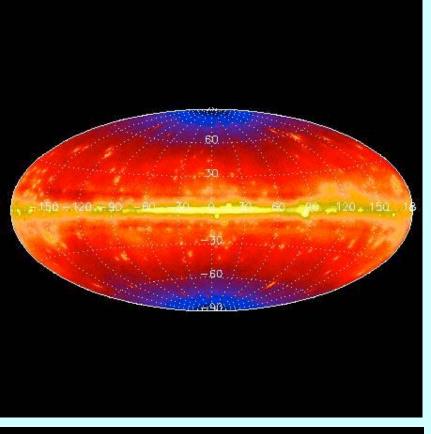




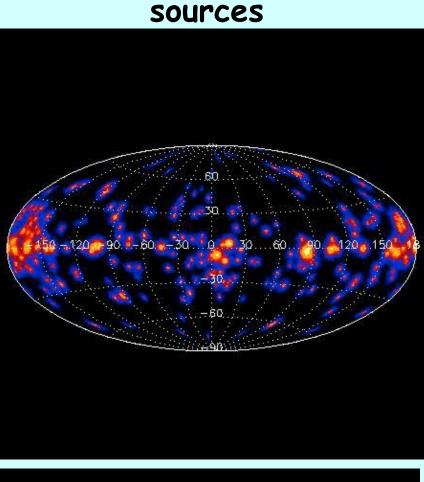


All sky + sources 0.1-30 GeV 827 orbits bg with PSF

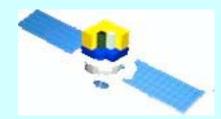
Bkg + sources



Colorbar scale 1:1								
0.E+000	2.E+003	4.E+003	6.E+003	8.E+003				



•2	Colorbar scale 1:1			
0.E+000	2.E+003	4.E+003	6.E+003	8.E+003



200

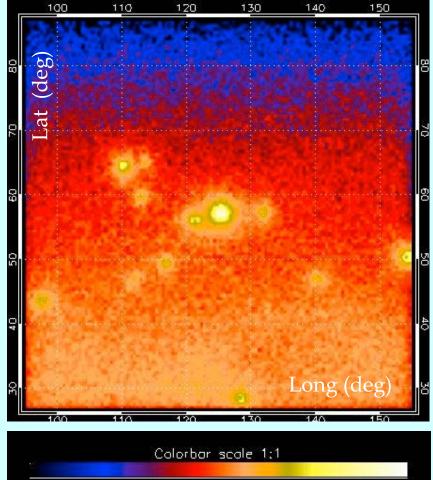
Û

400

3C279-60 sky + sources 0.1-30GeV Bg with PSF 827 orbits

sources

Bkg + sources

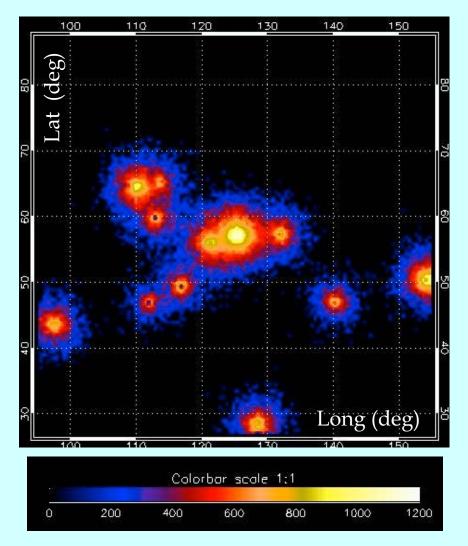


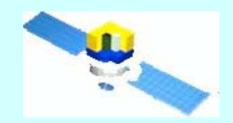
600

800

1000

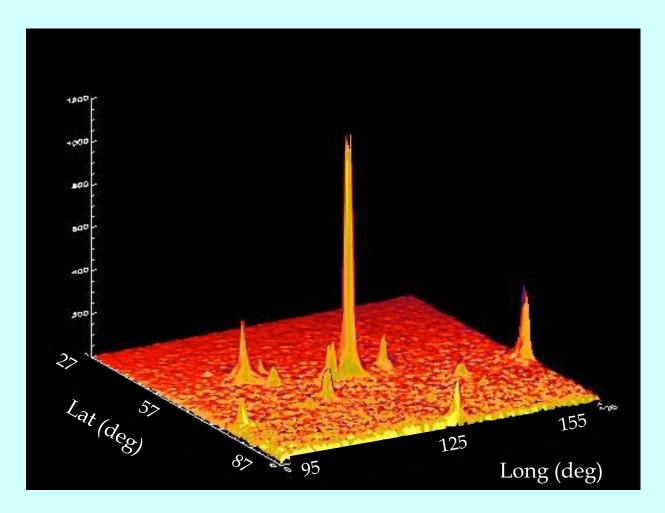
1200

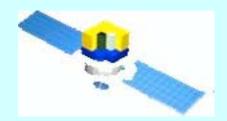




3C279-60 sky + sources 0.1-30GeV Bg with PSF 827 orbits

Surface plot





CONCLUSIONS

The simulation of the photon flux observed by GLAST works properly using sources from the Third Egret Catalogue and backgrounds either from Galprop results or Egret map.

Still to do:

 Add faint sources from Stecker and Salamon (1996)
 Include angular dependence in the PSF
 Include variability of the sources (see M. Fiorucci talk) End of November

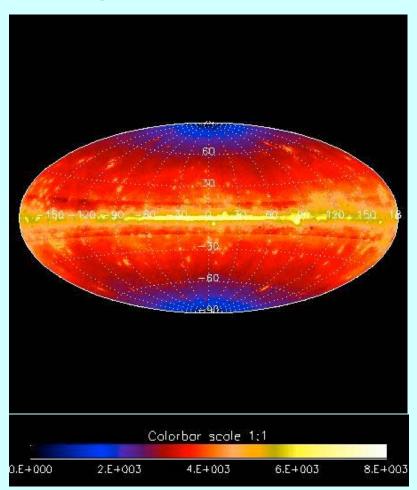


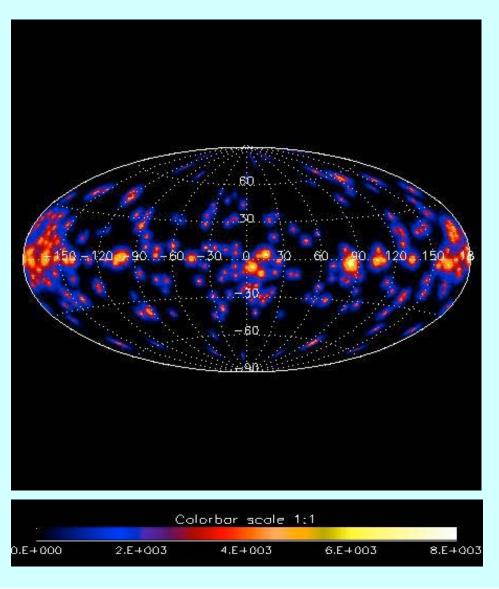


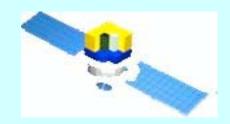
All sky + sources 0.1-30 GeV 827 orbits constant bg

sources

Bkg + sources

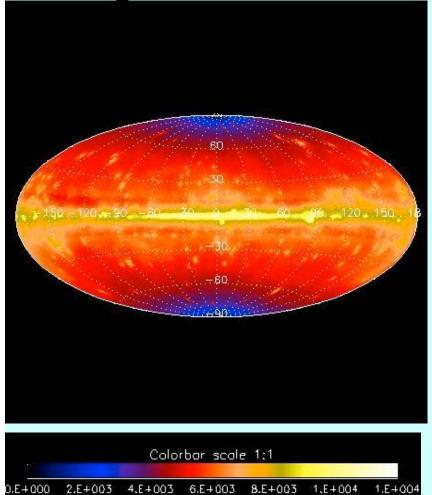




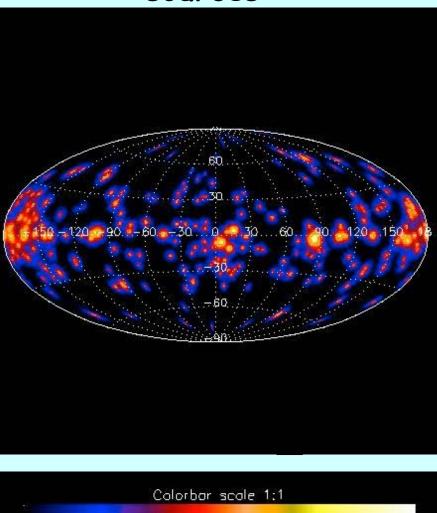


All sky + sources 0.1-30 GeV no orbit t=4E06 bg with PSF

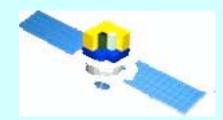
Bkg + sources



sources

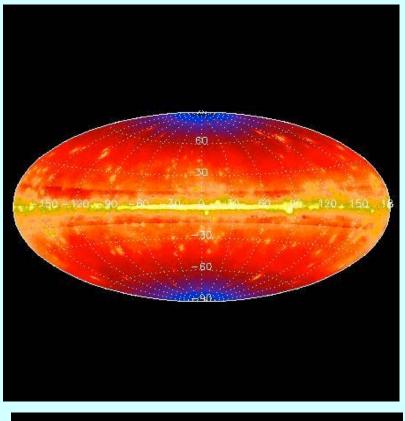


0.E+000 2.E+003 4.E+003 6.E+003 8.E+003 1.E+004 1.E+004

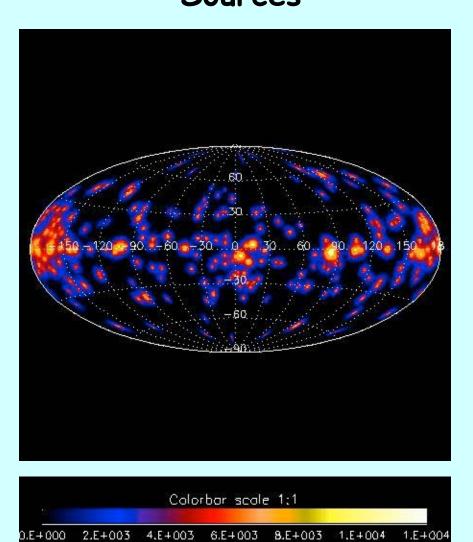


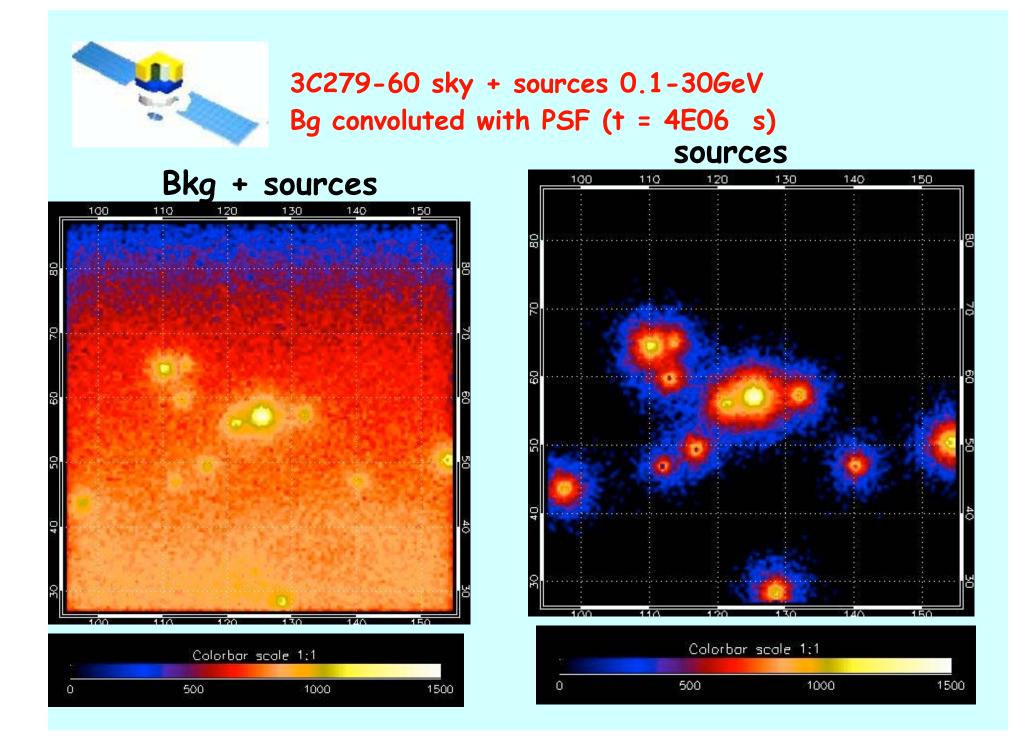
All sky + sources map 0.1-30 GeV no orbit constant bg t=4E06 s Sources

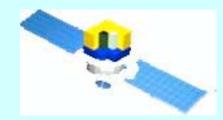
Bkg + sources





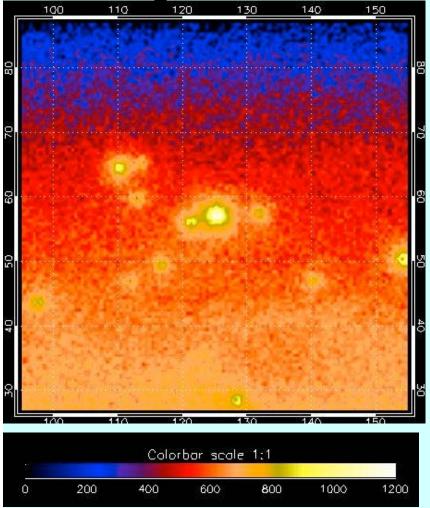


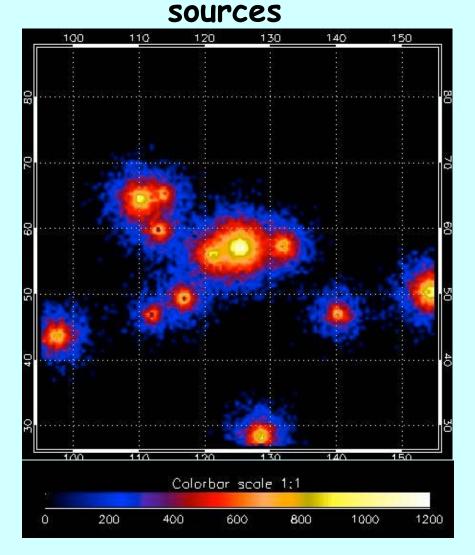


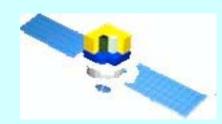


3C279-60 sky + sources 0.1-30GeV Bg const 827 orbits

Bkg + sources







3C279-60 sky + sources 0.1-30GeV constant bg (t = 4E06 s)

