



# The variety of sources in LAT data

**Elizabeth Ferrara**

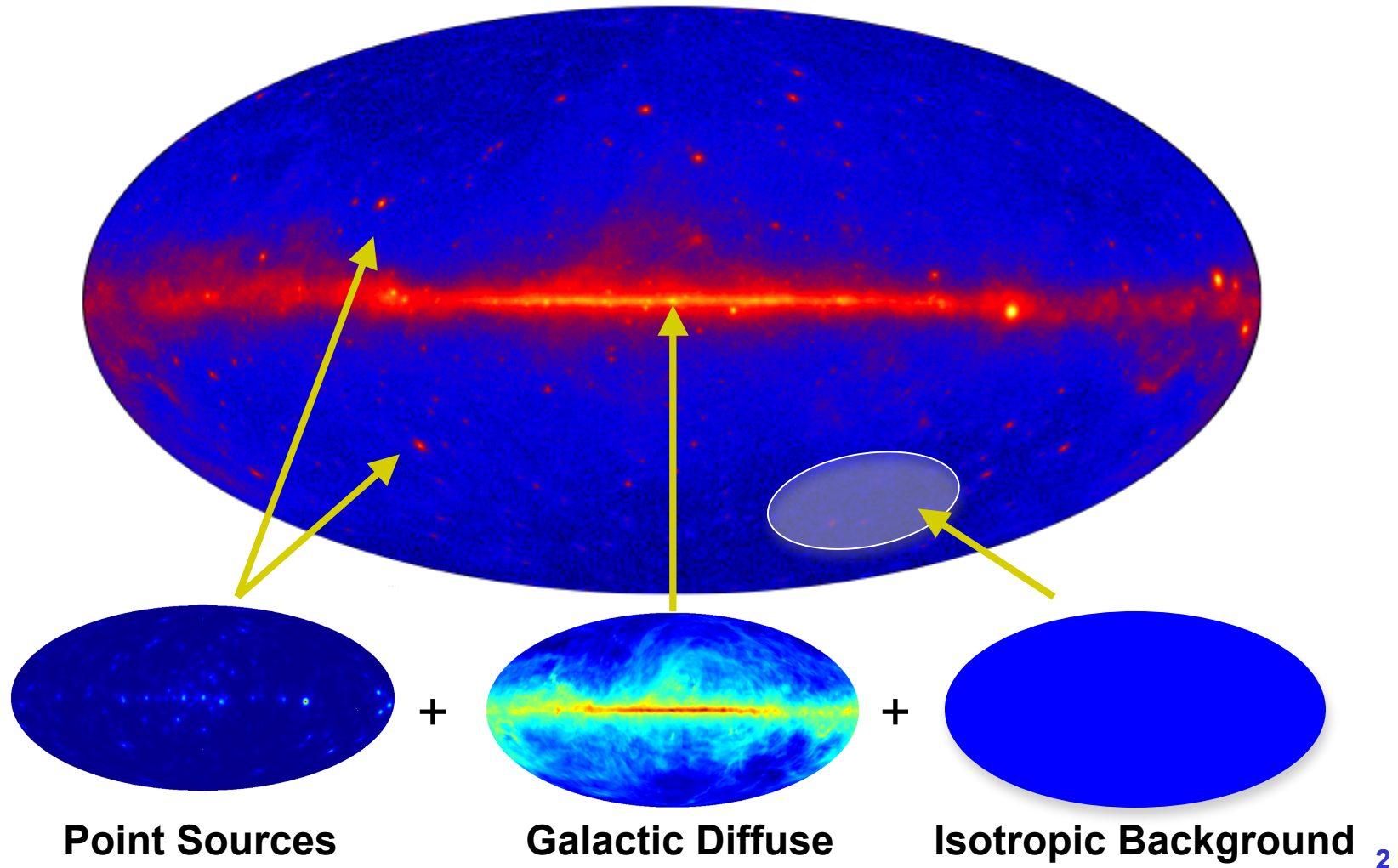
Fermi Science Support Center

**2011 Fermi Summer School**

**May 31 - June 10**

# LAT Data Set


- The LAT data consists of events from numerous different astrophysical sources + multiple sources of background



## Detection and Analysis of Point Sources

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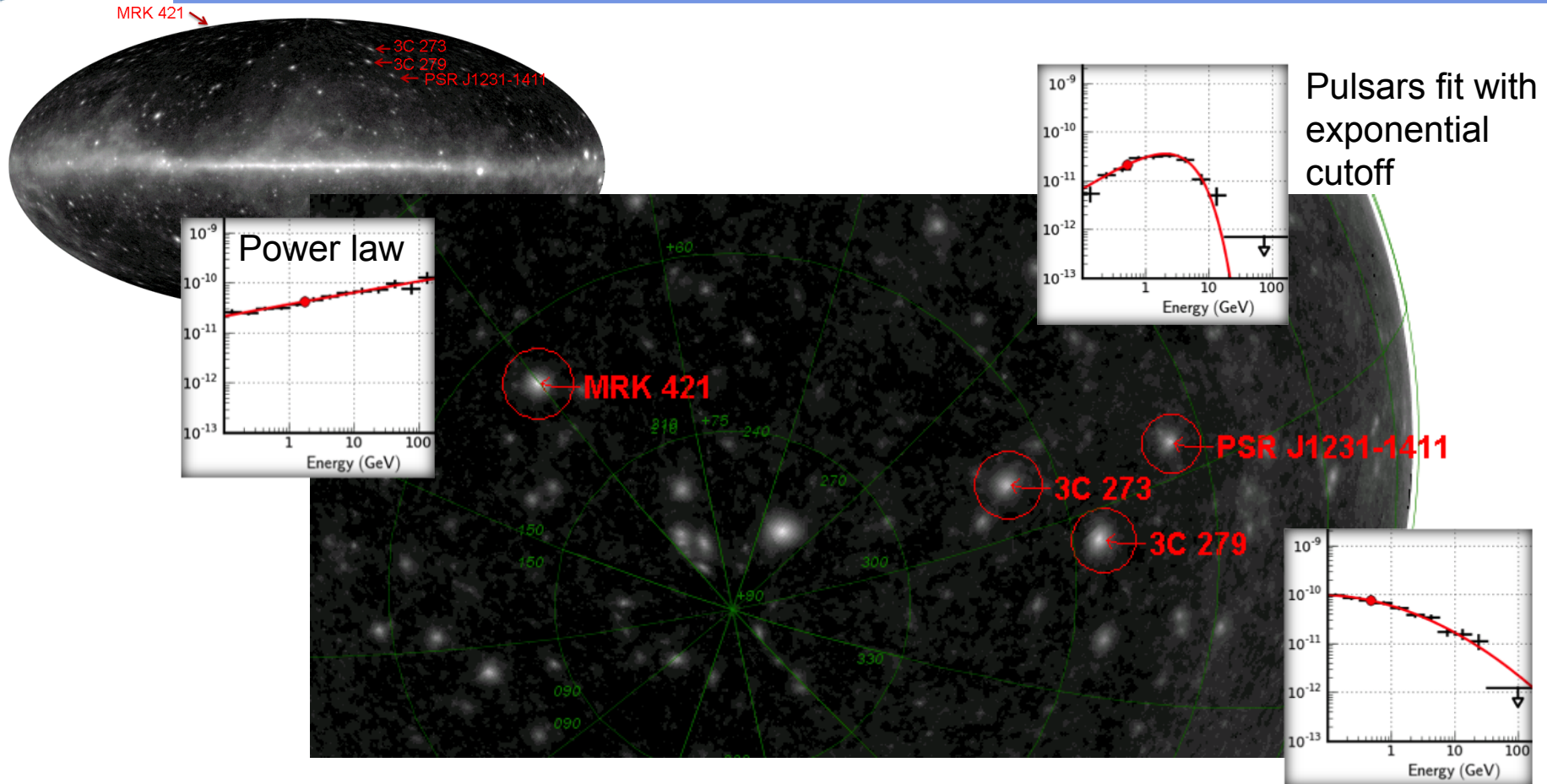
- Finding point sources in the LAT dataset is an iterative process

- 
1. Generate seed positions
  2. Simultaneously fit putative sources plus background
  3. Apply cut to remove sources that are not significant
  4. Look at residuals to find new candidate sources
  5. Iterate

- All-sky analysis is too computer-intensive
  - Requires sky be divided into “manageable” regions
  - Source extension or spectral shape can add to the complexity



# Spectral Shape



**Spectral shape is a contributing factor in whether or not a source is significantly detected**

Use log parabola or broken power law if it gives a better fit



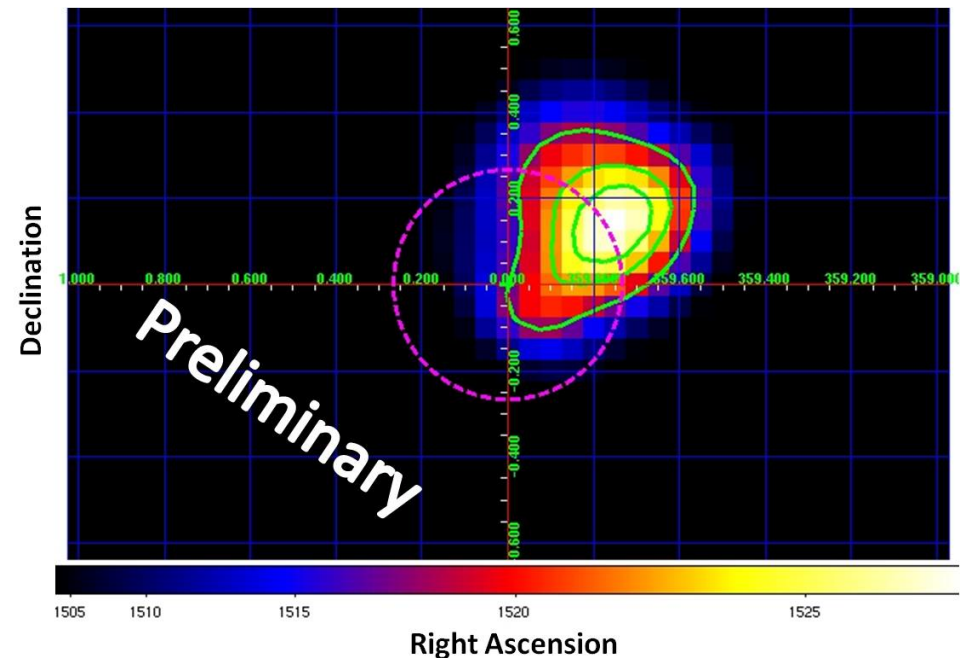
## Associations vs. Identifications

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- Sources are considered associated when their positions coincide with plausible gamma-ray-producing objects
  - Association technique is a Bayesian probability analysis based on the local density of sources from catalogs of likely objects
  - Average 95% uncertainty radius is about 7 arcmin
  - Too large in most cases to claim identification based on position
- Some LAT sources are firmly identified
  - Periodic signals, spatial morphology, or correlated variability with objects known at other wavelengths

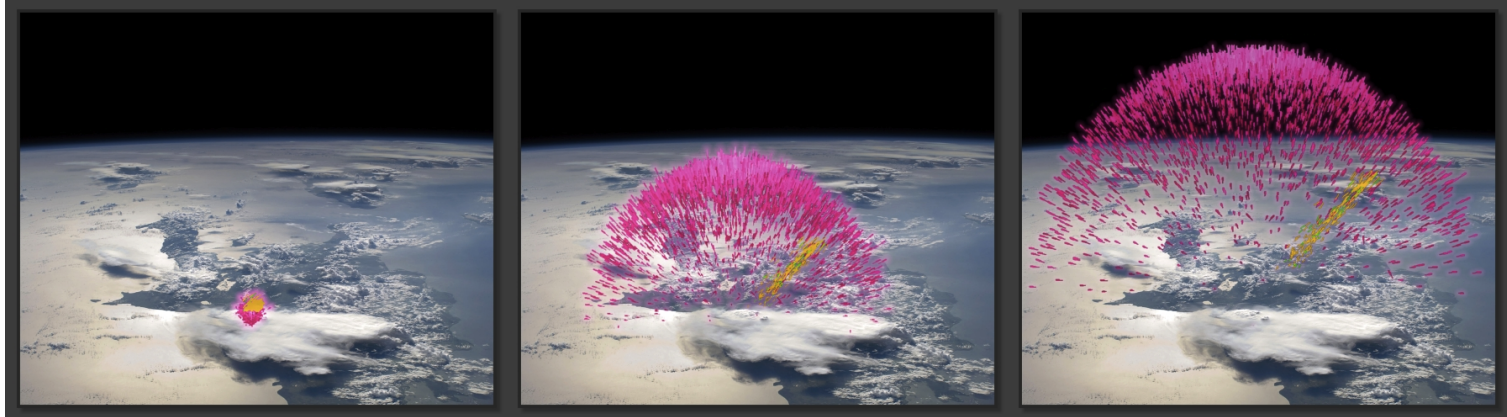
## Unlocalized Events - 1

- Some events affect the ability of the instrument to acquire data
  - In general, these events will be filtered out by the event classification process
  - Use the recommended data cuts to eliminate undesired effects
- Solar Flares
  - Seen by EGRET
  - Can last for many minutes to hours
  - Will become more prevalent as the solar cycle continues
  - As the Sun is a moving source, it can affect many other sources

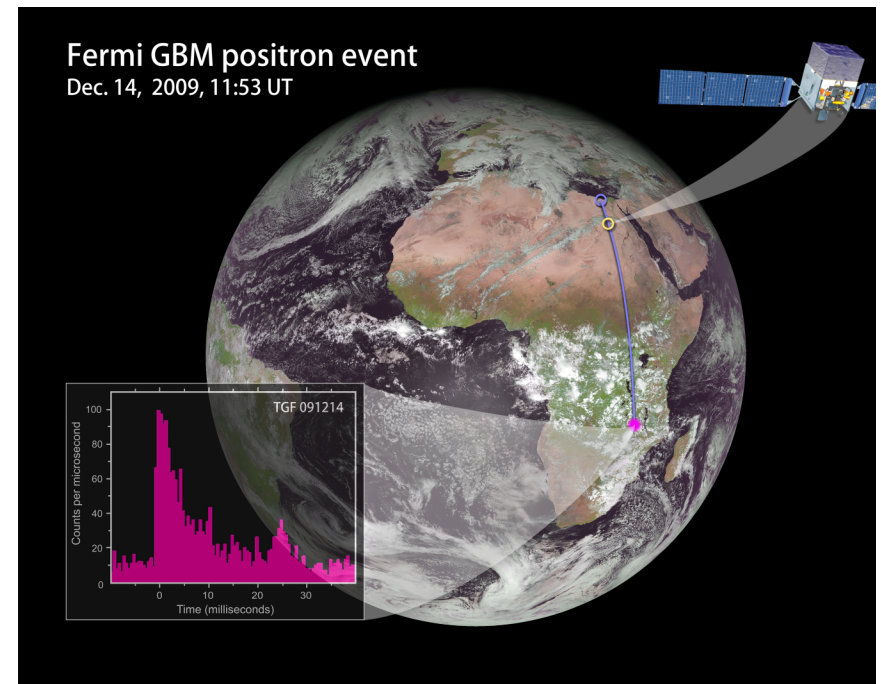


March 7, 2011 Solar Flare

## Unlocalized Events - 2

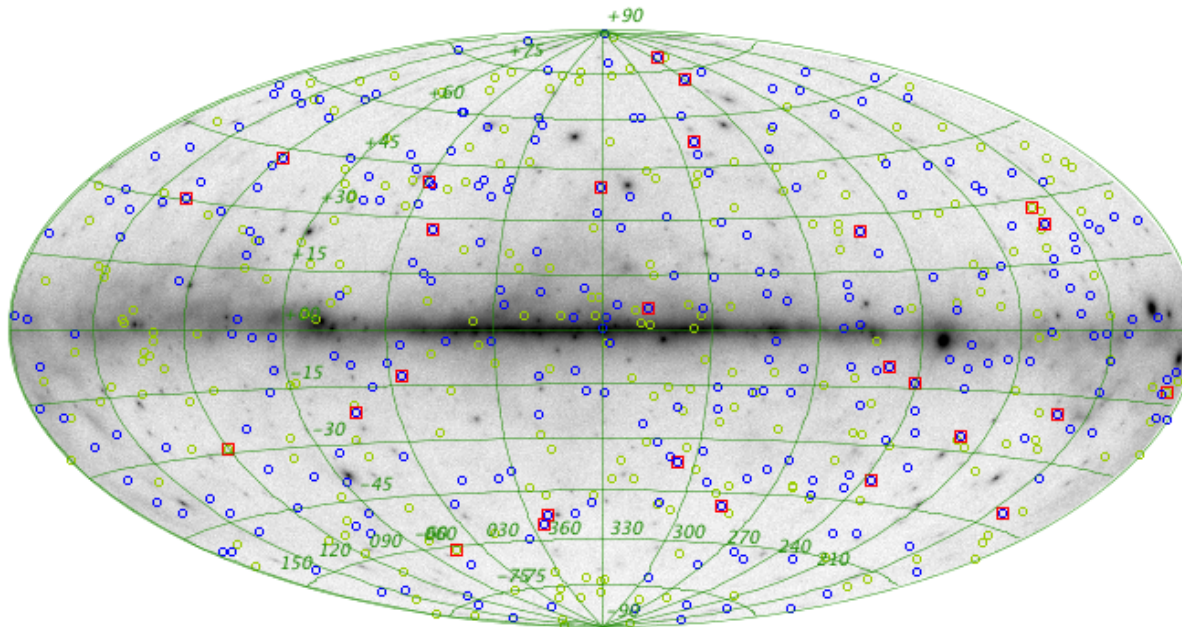


- **Terrestrial Gamma-Ray Flashes (TGFs)**
  - Several-microsecond events associated with thunderstorms on the Earth
  - Two types seen by GBM
  - Not currently detected by the LAT



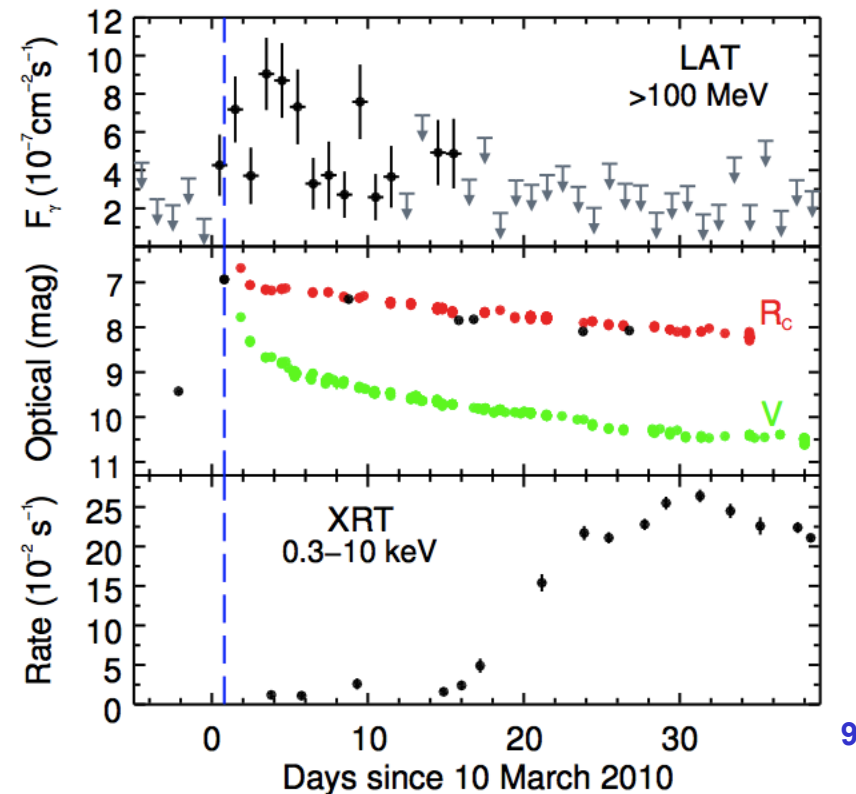
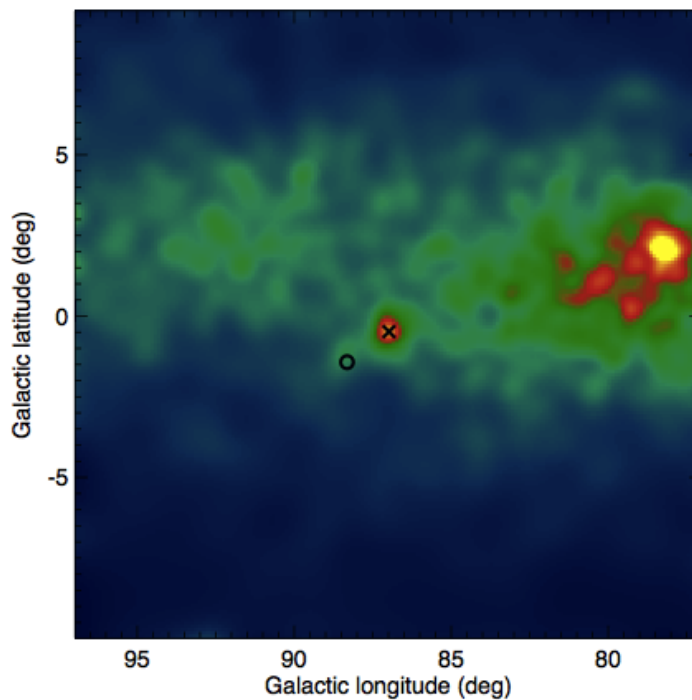


- **Extragalactic Transients: Gamma-Ray Bursts (GRBs)**
  - Rapid, very energetic events at extreme distances
  - Very short timescales (seconds to 100s of seconds)
  - Seen frequently by GBM (~550)
  - A subset of these are seen by the LAT (27)
    - *Must have a high-energy component*
    - *Must be within the LAT field of view (FOV)*
  - Can affect analysis if near the source of interest



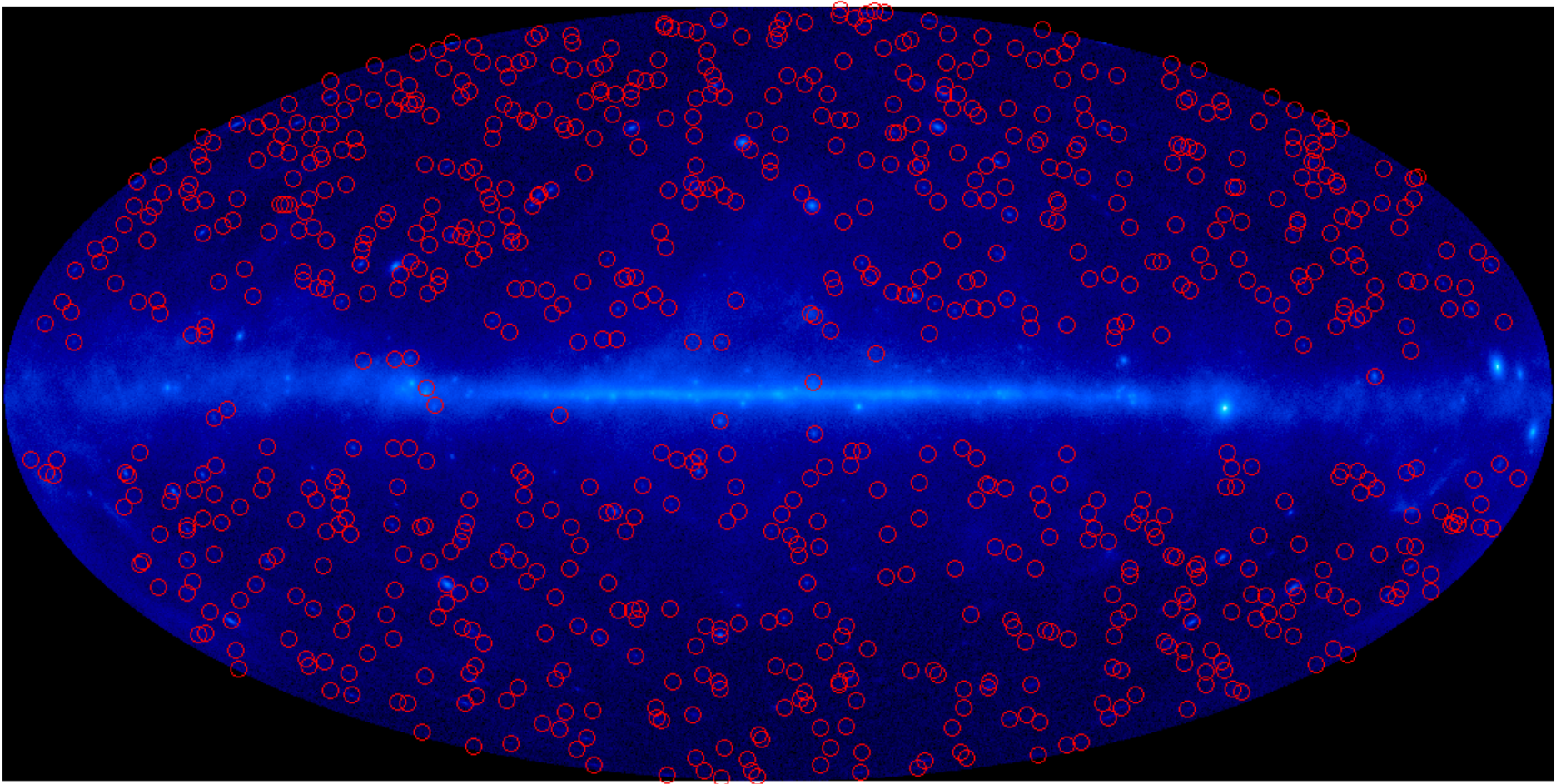
Circles:  
In Field-of-view of  
LAT ( $<70^\circ$ ): 275  
Out of the FOV  
Squares:  
LAT detections

- **Galactic Transients:**
  - Several Galactic transients have been seen by the LAT since Fermi started science operations
- **Nova V407 Cyg - symbiotic binary**
  - Explosion into surrounding medium generated a shock front that generated gammas



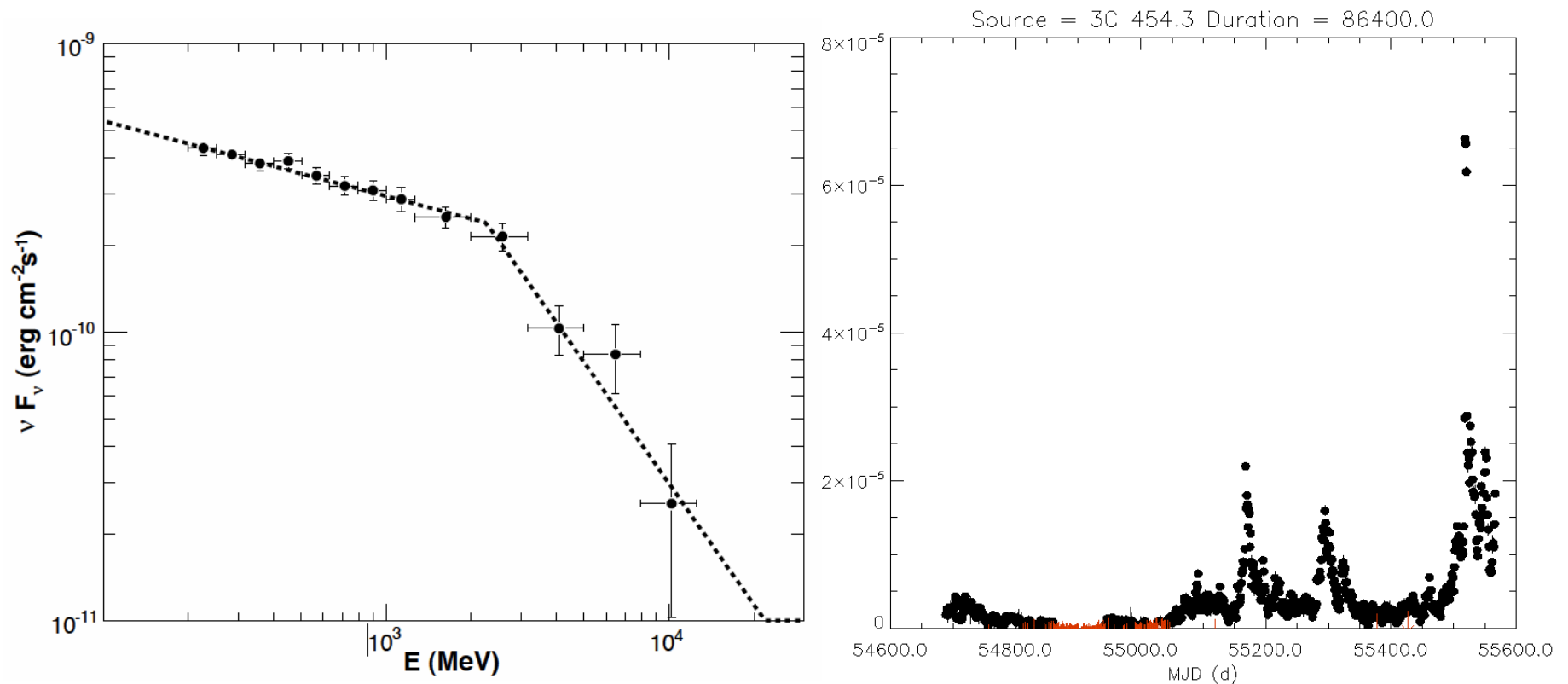
# Blazars

- **The most numerous class in the LAT data by far!**



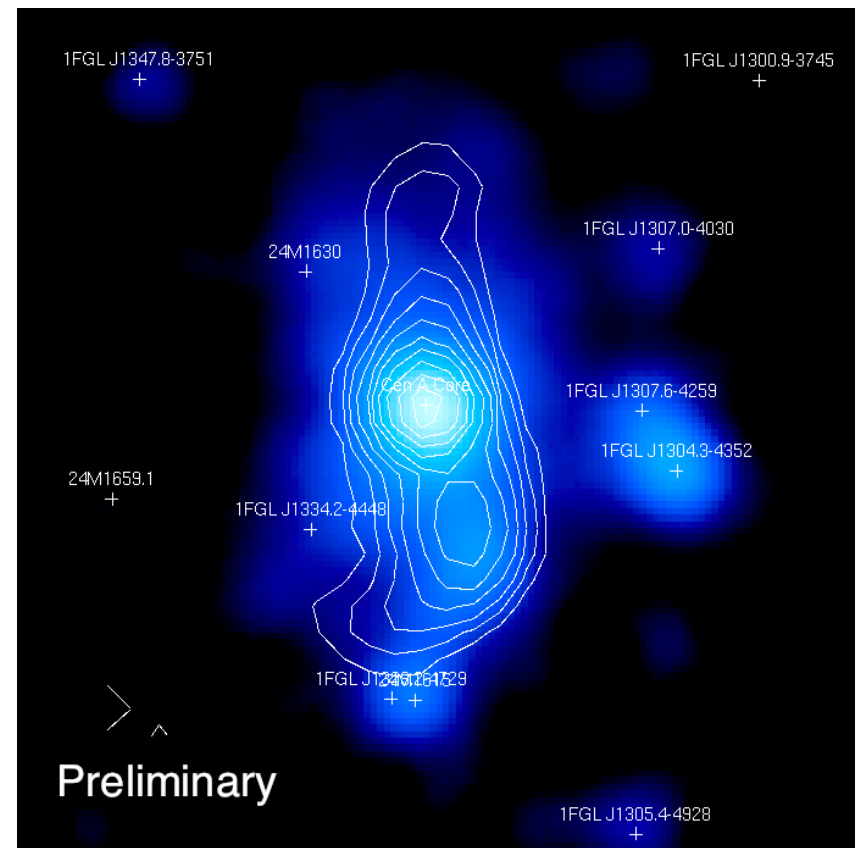


- Blazar detection is affected by both spectral and temporal characteristics
  - BL Lacs and FSRQs have spectral peaks in different energy ranges (spectral index varies widely)
  - Flux varies significantly with time



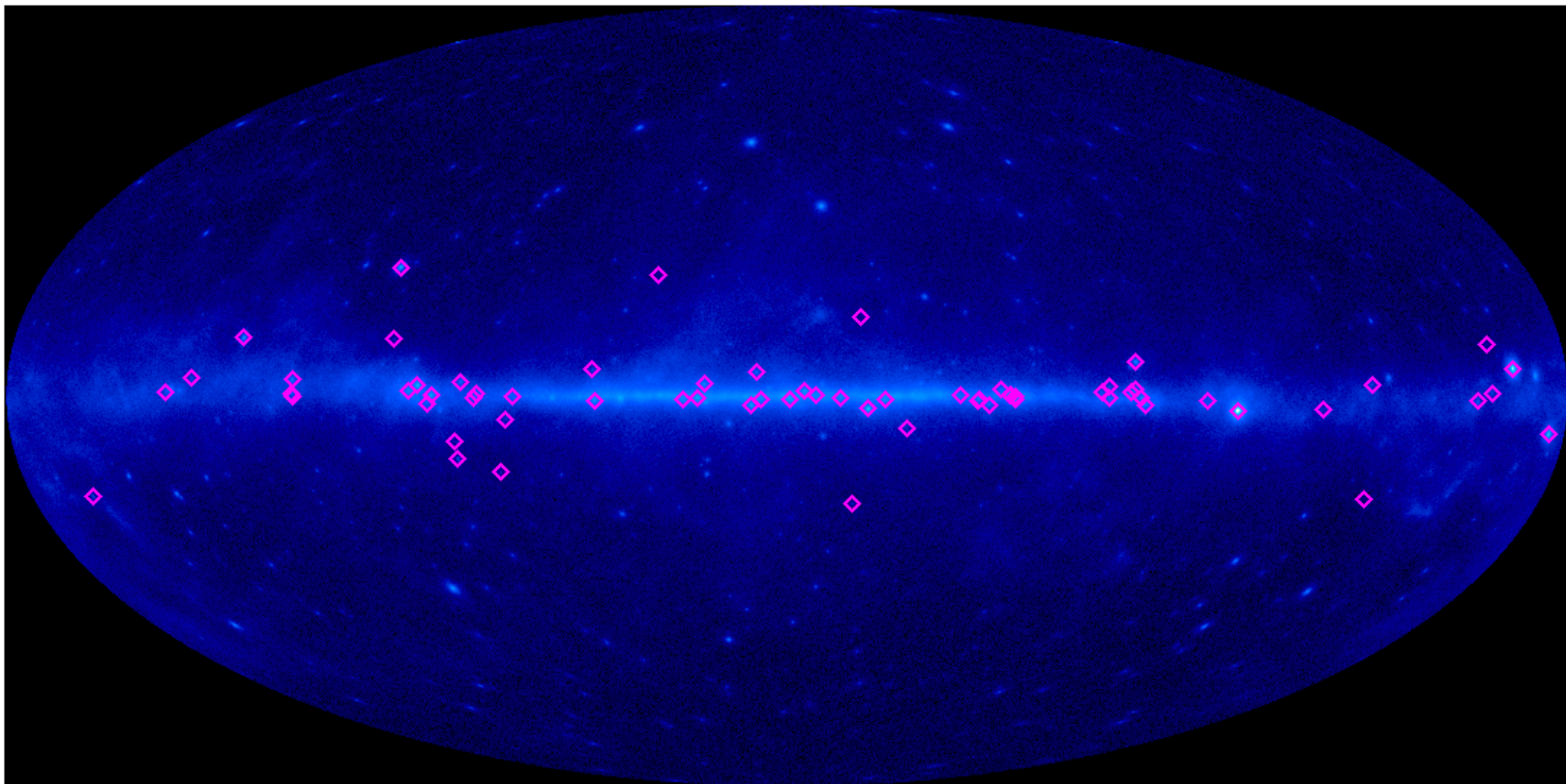
## Other AGN

- **Blazar Candidates**
  - Blazar associations use a ‘figure of merit’
  - Candidate sources have radio detections that look like blazars
  - But require additional follow-up to be able to calculate the FOM
- **Narrow-Line Seyfert 1s**
  - Three of these sources show LAT emission
- **Radio Galaxies**
  - Currently 4 of these sources are showing LAT emission
  - Centaurus A radio lobes are fully resolved by the LAT



## Young Pulsars

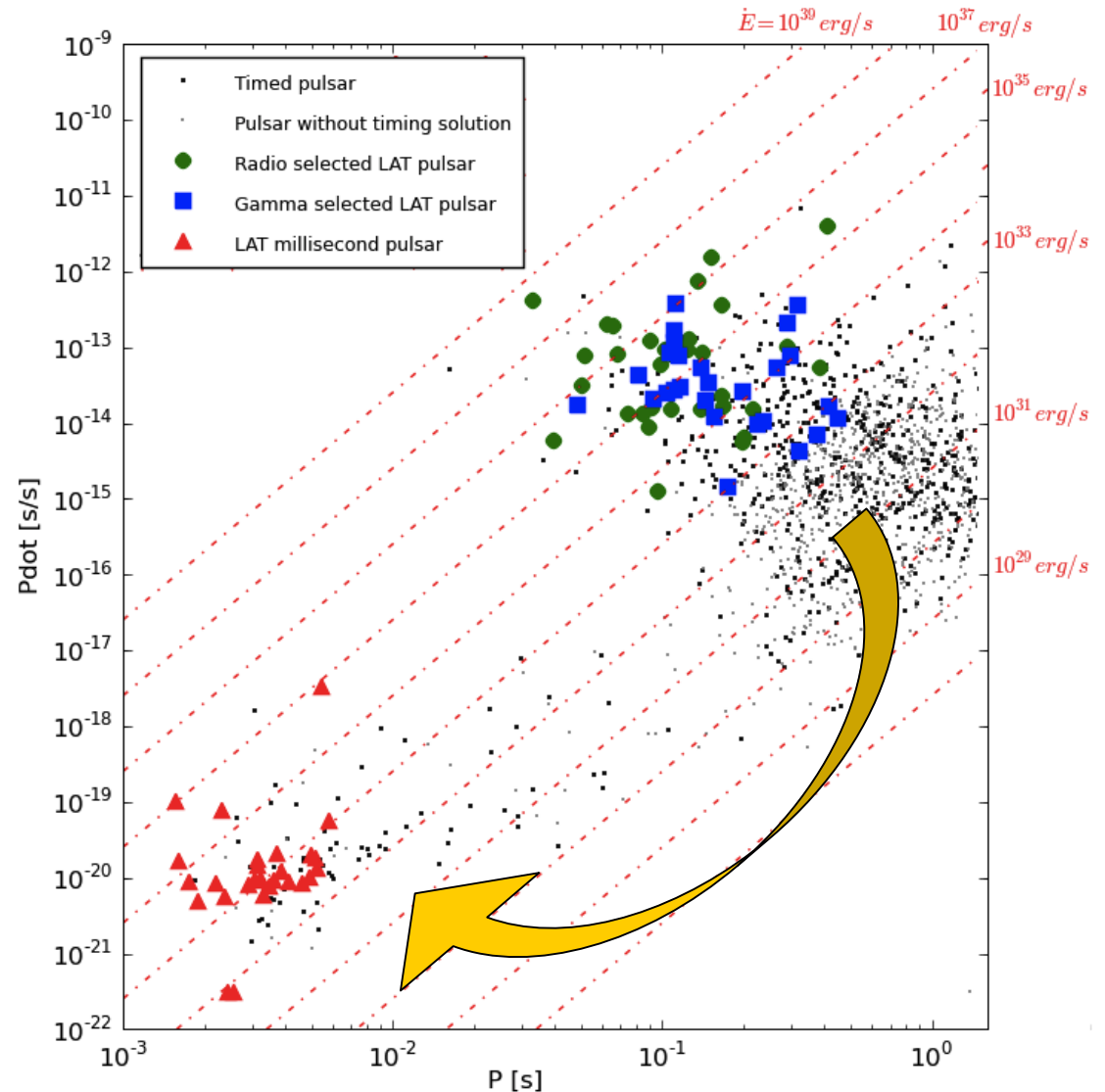
- **Highly energetic pulsars have long been known to be gamma-ray emitters**
  - Typically discovered by applying known radio ephemerides to the gamma rays
  - Can also be discovered in the gamma-ray data (26 with LAT)





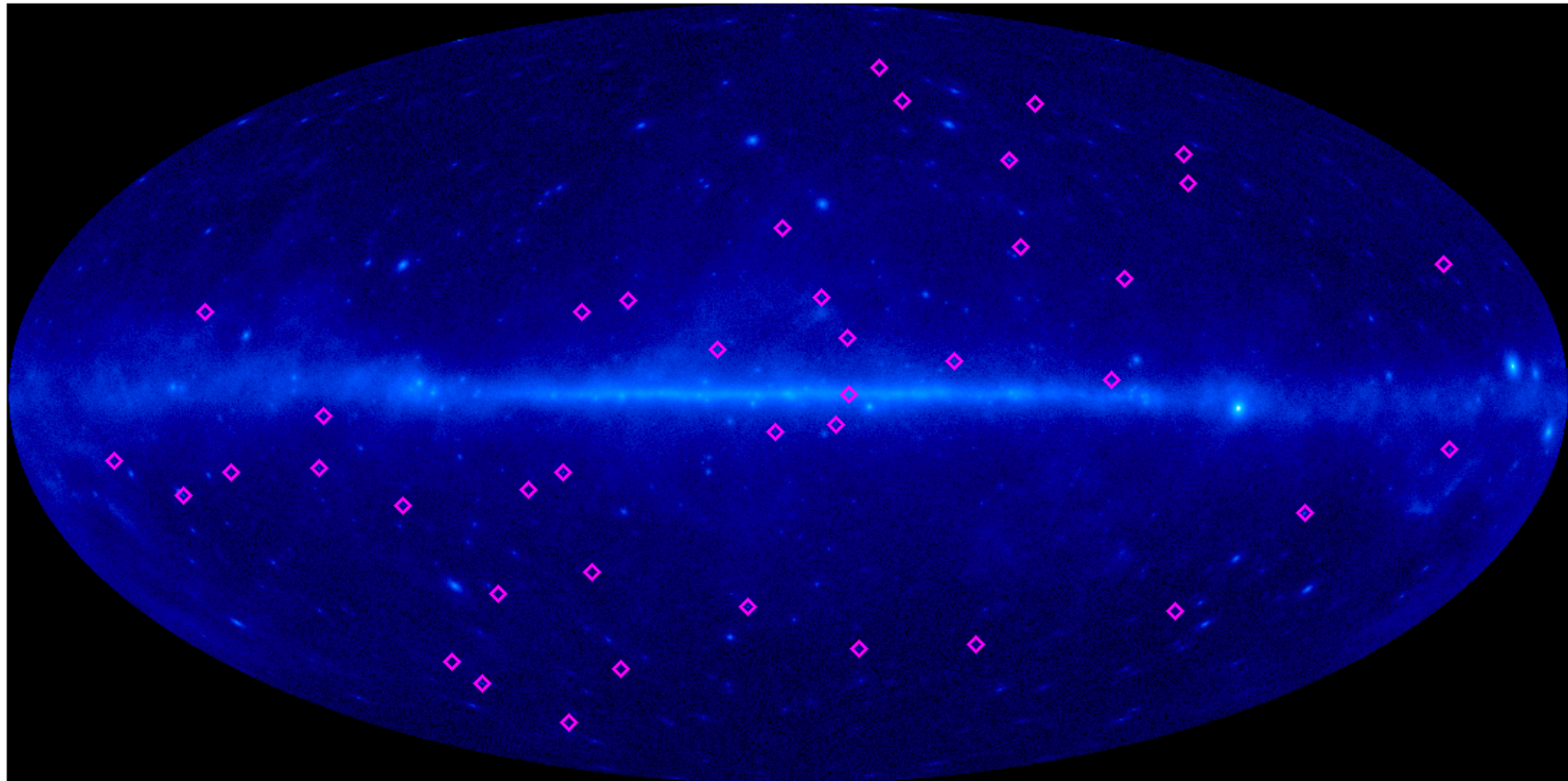
# Recycled Pulsars

- Over time young energetic pulsars slow down
  - Power for pulsations comes from rotation
  - Once energetics are no longer favorable, pulsations cease
- Pulsars in binaries can get a second life through mass transfer
  - Increase in angular momentum produces millisecond periods and high energetics

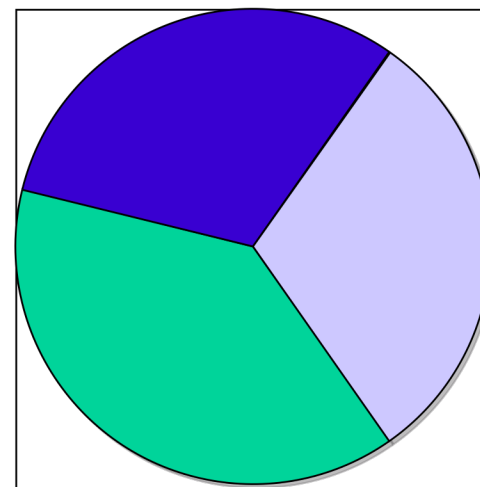


## Millisecond Pulsars

- Indeed a number of new millisecond pulsars have been found in the LAT data
  - Discovered by radio searches in LAT sources that lack counterparts
  - Significant increase in Galactic MSPs known (~60 +33 new)



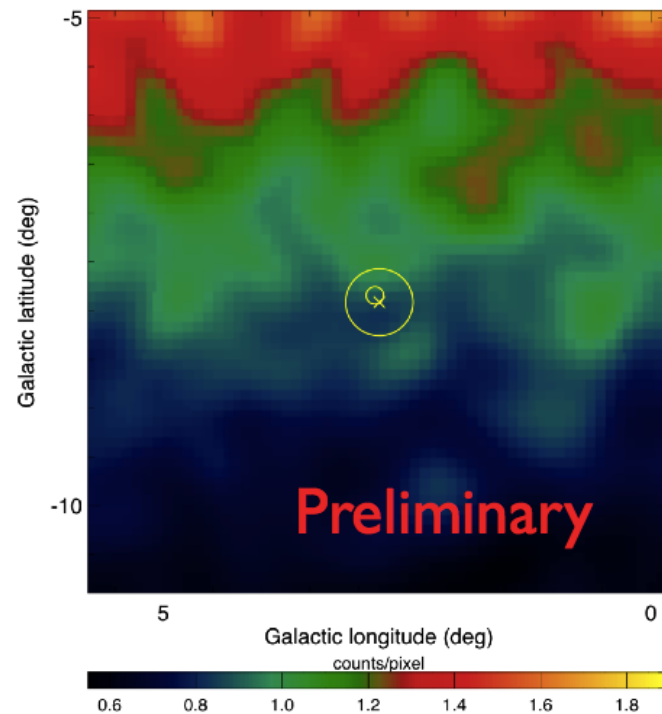
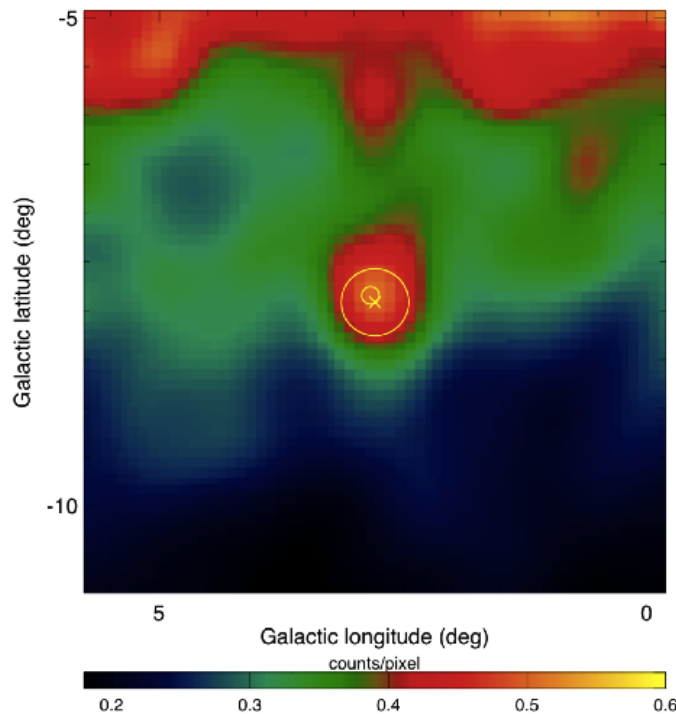
- The LAT-detected pulsars are typically referred to by their discovery method
  - Radio-selected used known radio ephemerides to find the LAT pulsations
  - Gamma-selected were discovered by folding the gamma-ray data (blind searches), and are usually radio-quiet (or very radio-faint)
  - LAT MSPs were all found using radio ephemerides
    - *Current blind search techniques are not sensitive to millisecond periods*
- Also 23 ‘radio-only’ MSPs
  - Discovered in the radio by looking at LAT sources
  - Takes 6 months to a year to get a good timing solution
  - May soon be LAT pulsars



■ Radio Pulsars  
■ Gamma Pulsars  
■ LAT MSPs

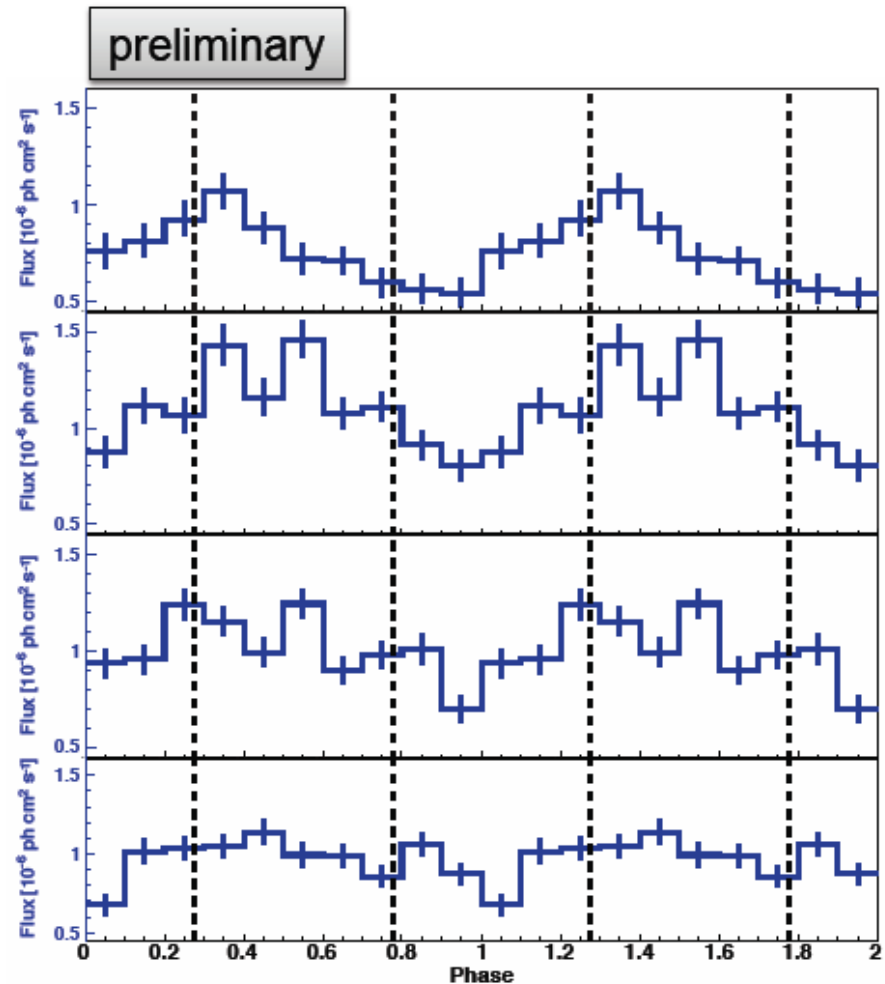
# Globular Clusters

- **Some globular clusters have long been known to contain numerous MSPs**
  - **LAT detects 11 sources coincident with globular clusters**
  - **In one instance (J1823-3021A) a single luminous gamma-ray pulsar has been found to be responsible for the entire LAT-detected emission from the cluster**





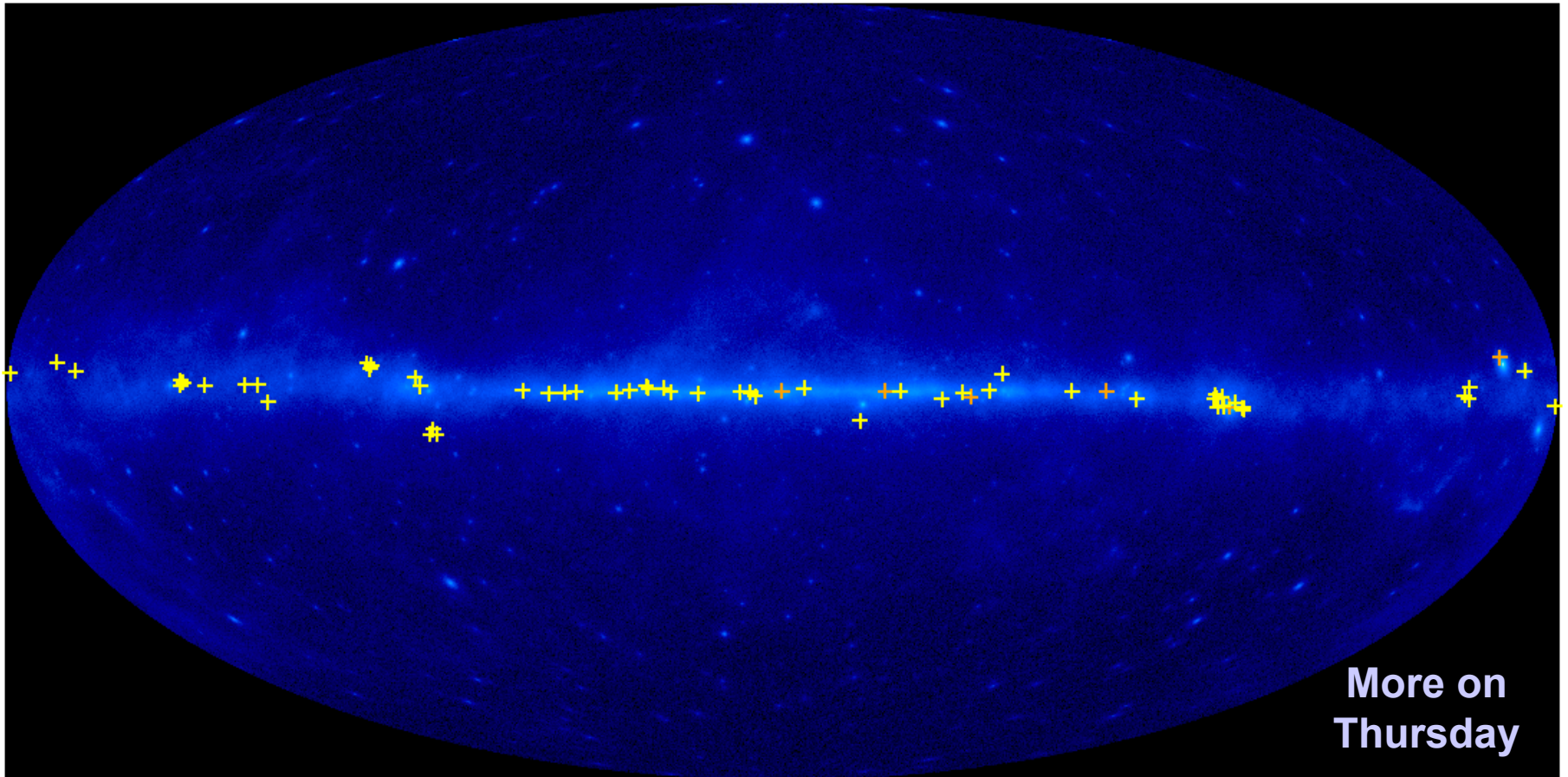
- The LAT detects periodic signals from four HMXBs:
  - Cyg X-3
  - LS 5039
  - LSI +61 303
  - 1FGL 1018.6-5856  
(new discovery!)
- LSI +61 303 orbital signal appears to have slowly disappeared since the beginning of science operations



Folded light curve in 6-month intervals

## Supernova Remnants / Pulsar Wind Nebulae

- **Third most numerous class behind blazars and pulsars**
  - **60 SNRs/PWNe in a very narrow distribution**
  - **Some positively identified by matching their extension to other wavelengths**

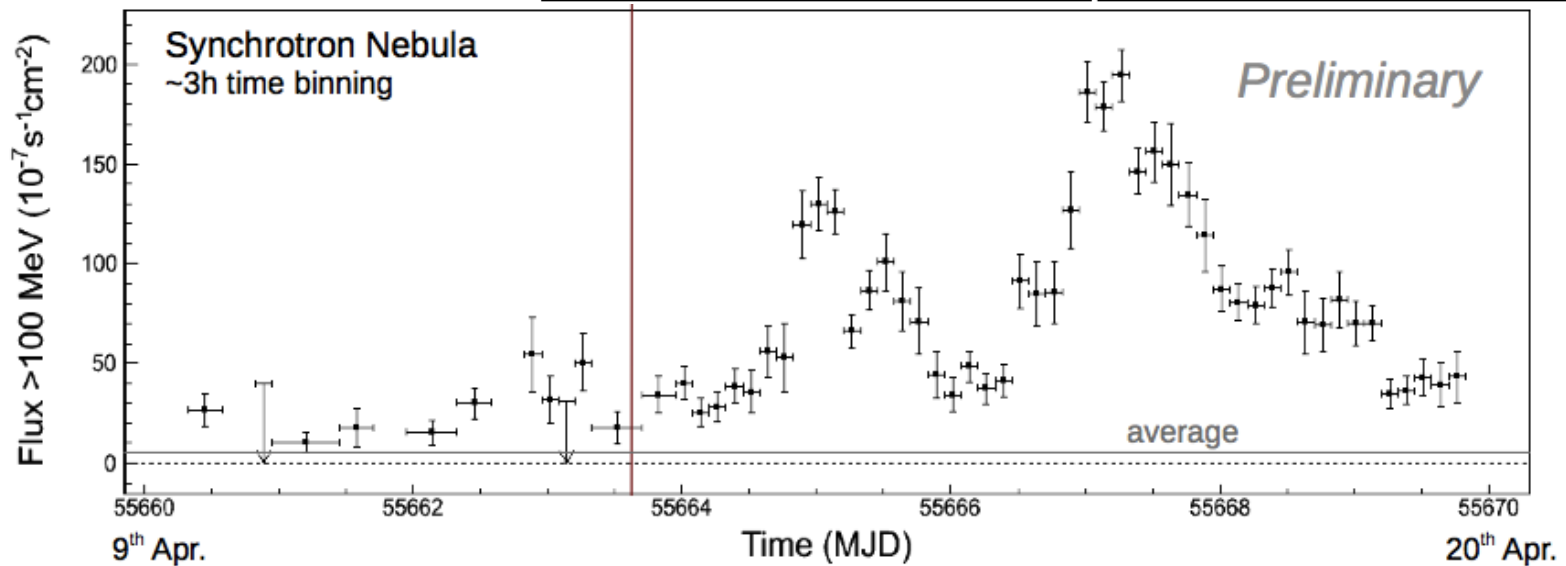
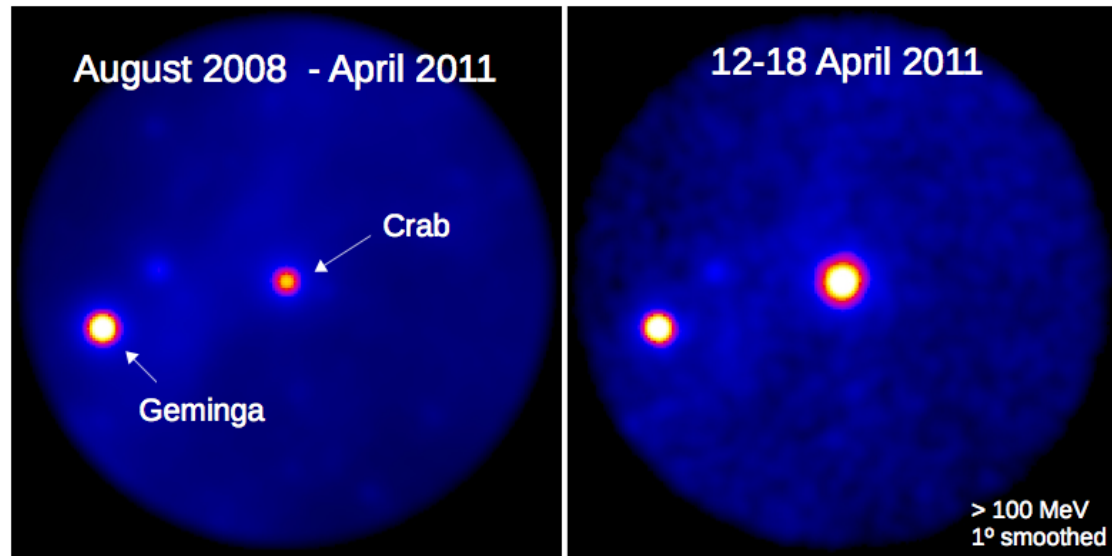


More on  
Thursday

## The Variable Crab?

- The Crab pulsar + SNR is used as a calibration source in high-energy astronomy

Recent variability from this source was quite a surprise!

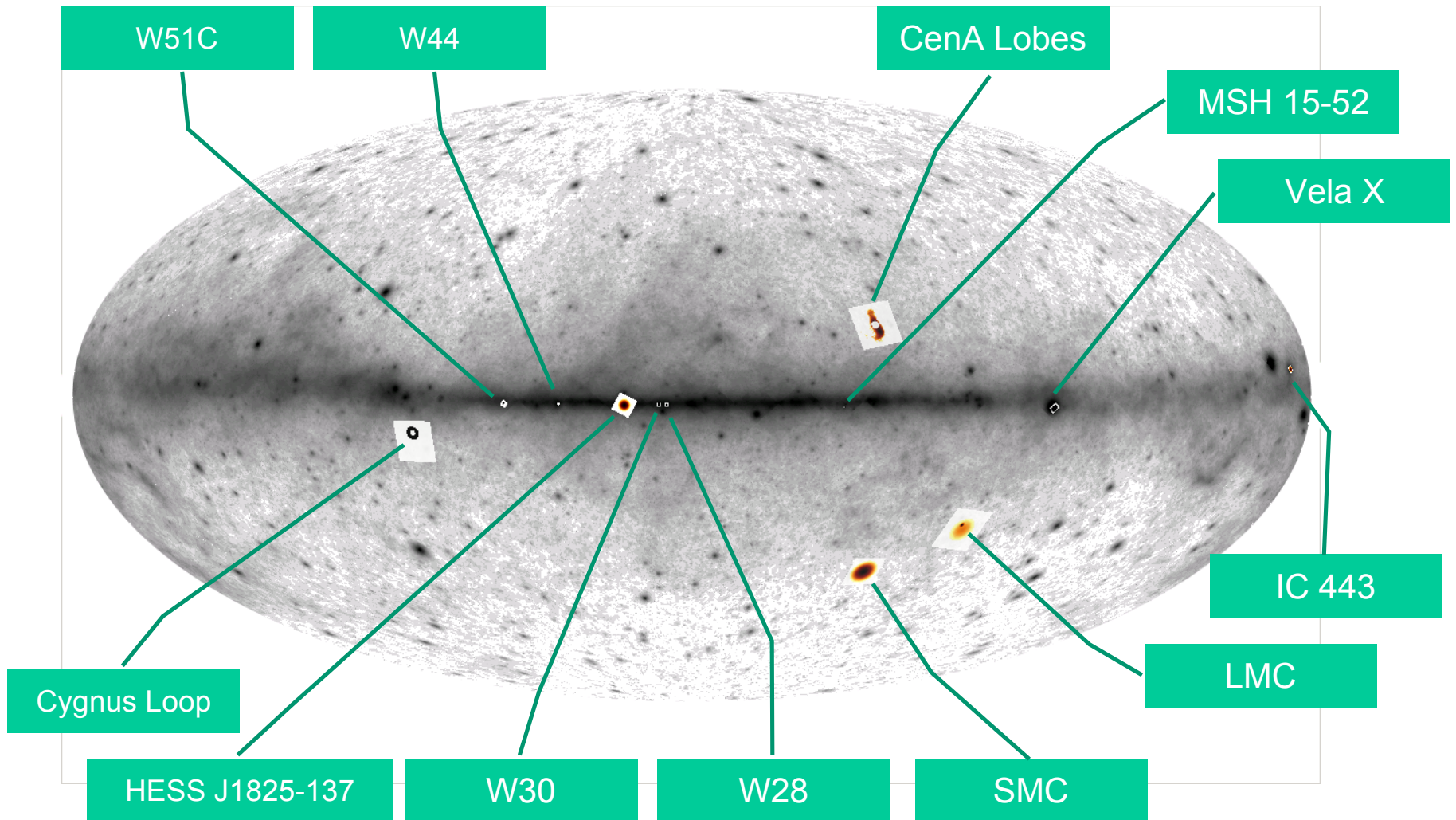


## Spatially Extended Sources

Extended Source	Spatial Form	Spectral Form
SMC	2D Gaussian	Exp Cutoff PL
LMC	2D Gaussian <sup>a</sup>	Exp Cutoff PL
IC 443	2D Gaussian	Log Parabola
Vela X	Disk	Power Law
Centaurus A (lobes)	Contour Map	Power Law
MSH 15–52	Disk	Power Law
W28	Disk	Log Parabola
W30	Disk	Log Parabola
HESS J1825–137	2D Gaussian	Power Law
W44	Ring	Log Parabola
W51C	Disk	Log Parabola
Cygnus Loop	Ring	Exp Cutoff PL

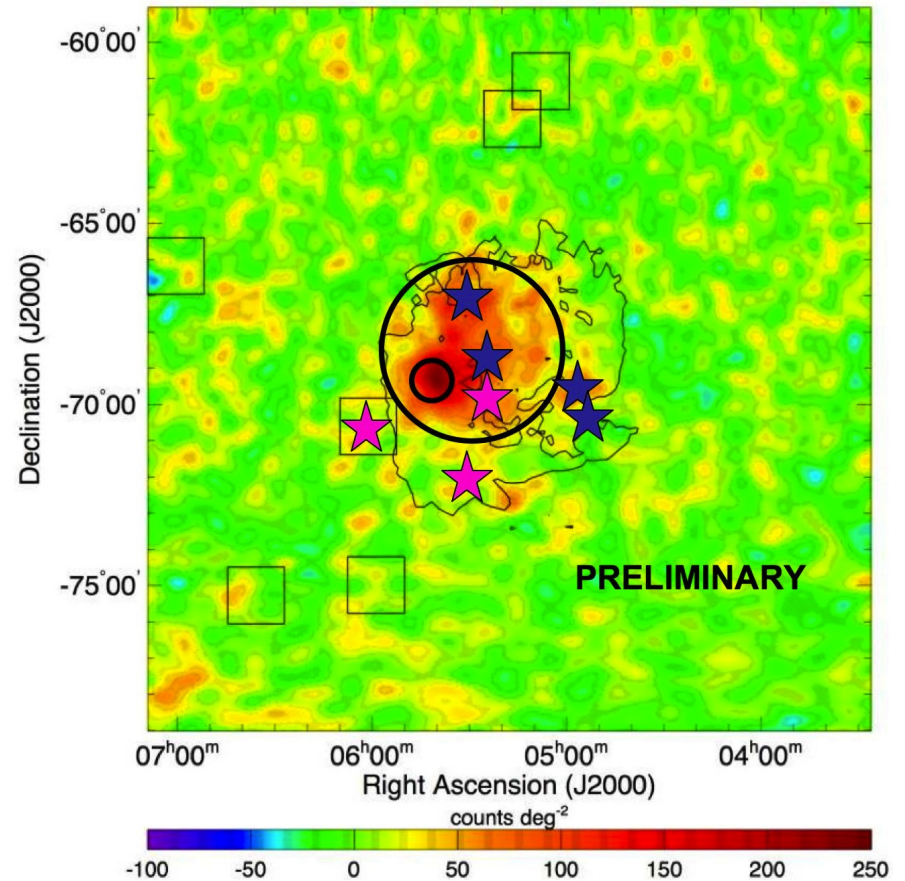


# Extended Source Templates



## Extended Source Uncertainties

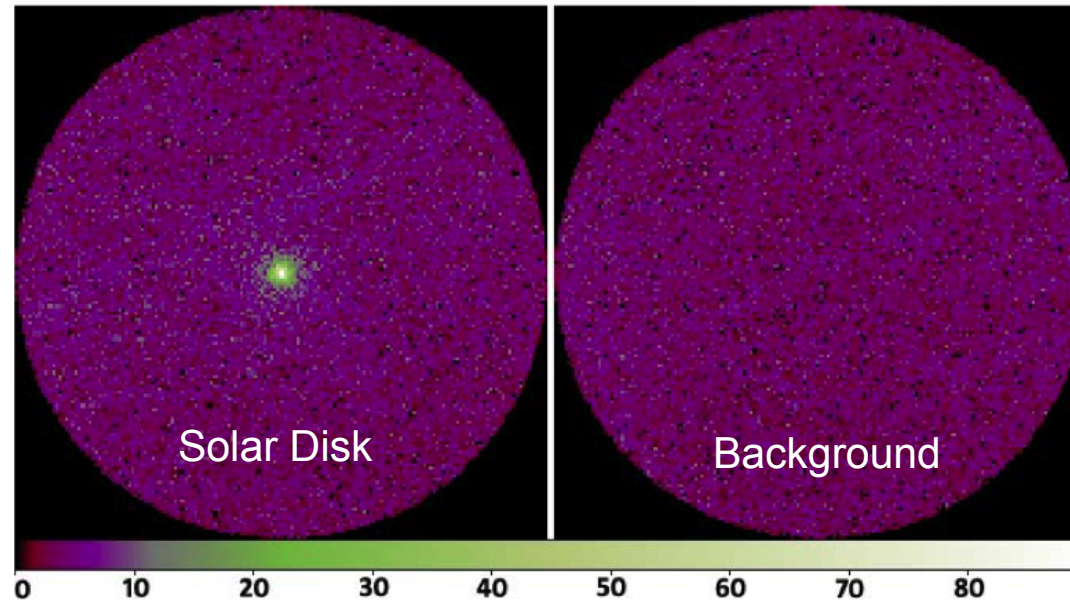
- The templates for the 12 sources that are spatially extended are approximations based on our best current knowledge
  - Analysis of regions around such sources can leave residuals that look like point sources (and in some cases may be)
  - For example, there are 7 point sources in the vicinity of the Large Magellanic Cloud. 3 of these have blazar/radio source associations, but 4 do not and could be artifacts.



LAT LMC map with overlay showing 2 component template and nearby sources

## Other Source Types...

- **Eta Carina (colliding wind binary)**
- **Starburst Galaxies**
- **Solar System bodies - Moon and quiescent Sun**

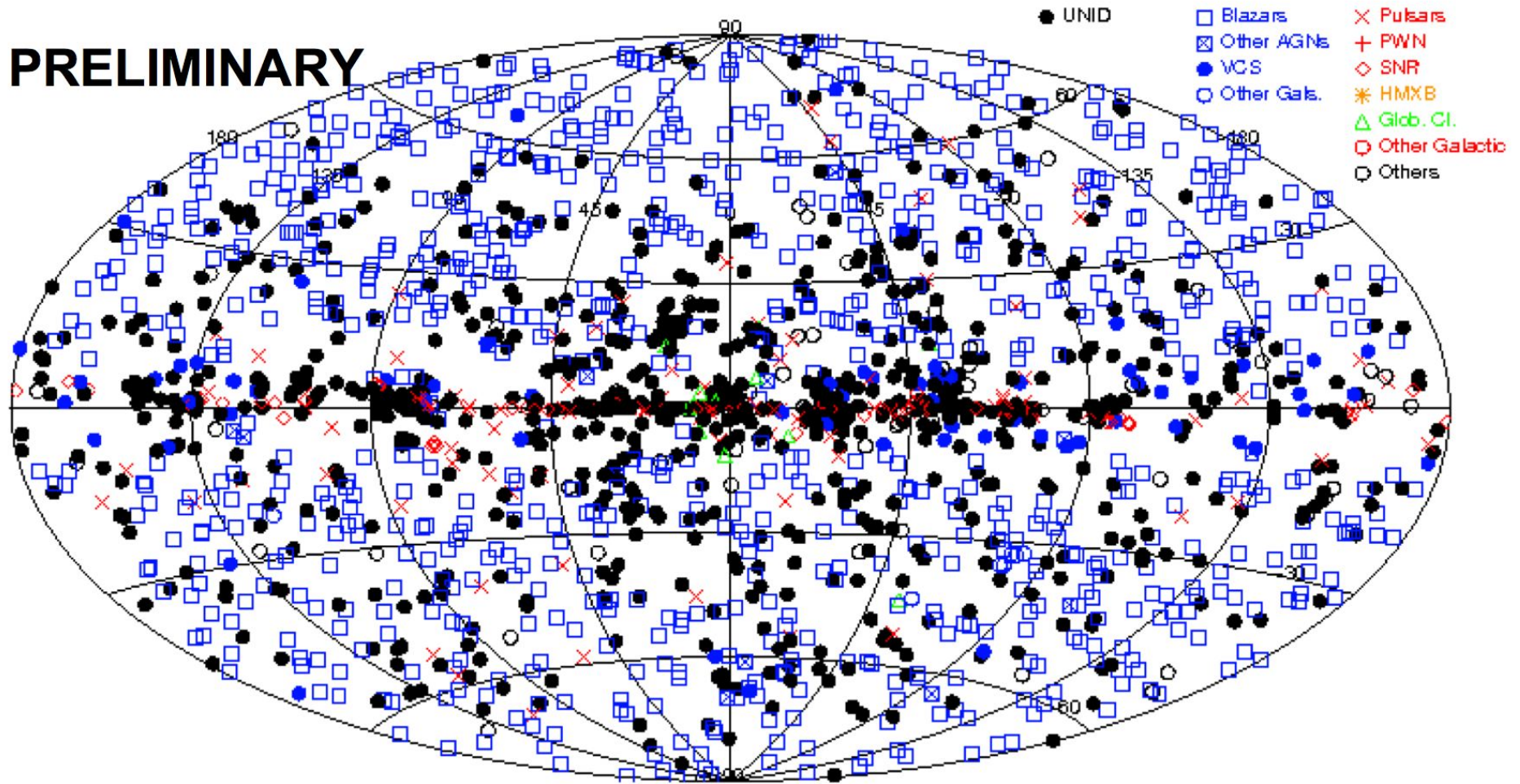


- **That's a lot of different types of sources!!**



# Classifications - 2FGL

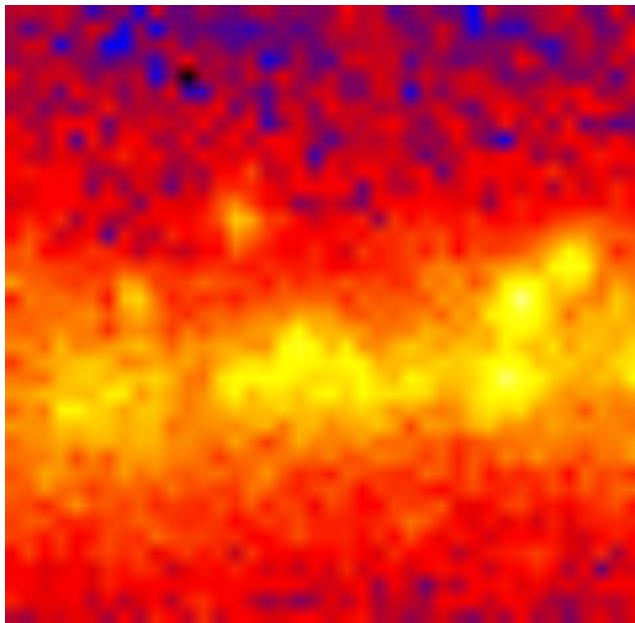
**PRELIMINARY**



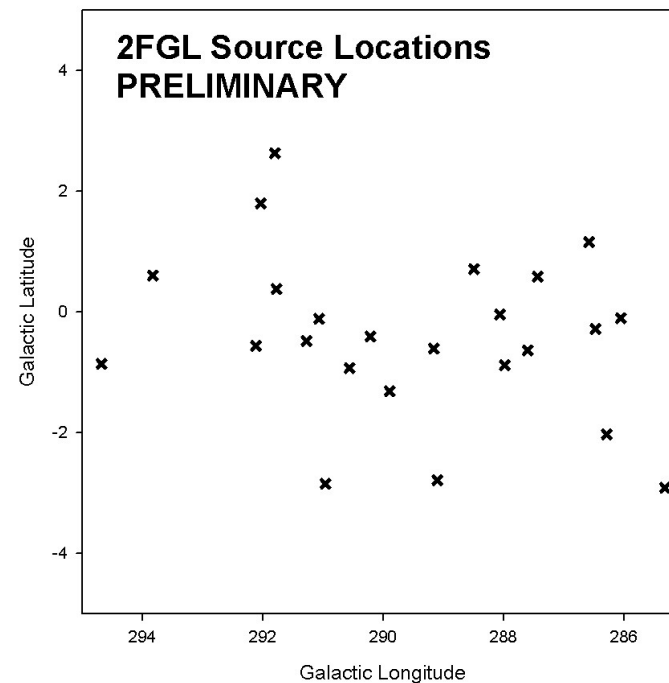


## Source Confusion

- There is a concentration of sources at low Galactic latitudes toward the inner Galaxy
  - Results in sources close enough to each other that their Point Spread Functions (PSFs) overlap
  - Particularly significant at lower energies, so affects soft sources more
- Parts of the sky away from the Galactic Plane show little impact

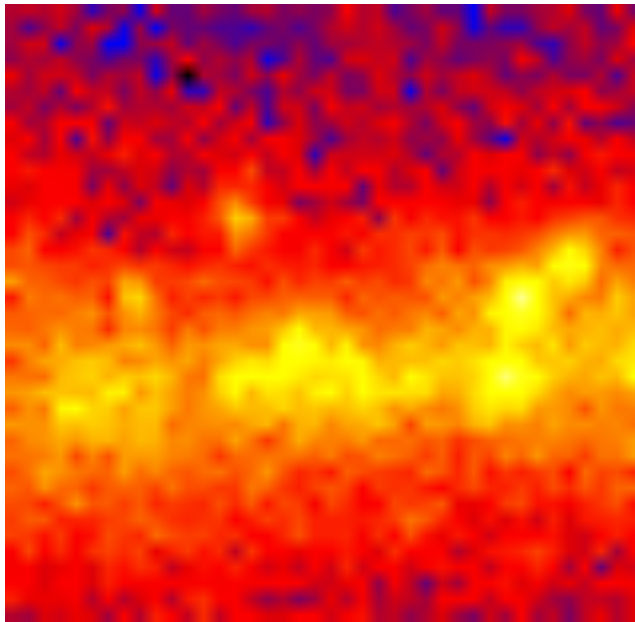


Counts map  $E > 1$  GeV

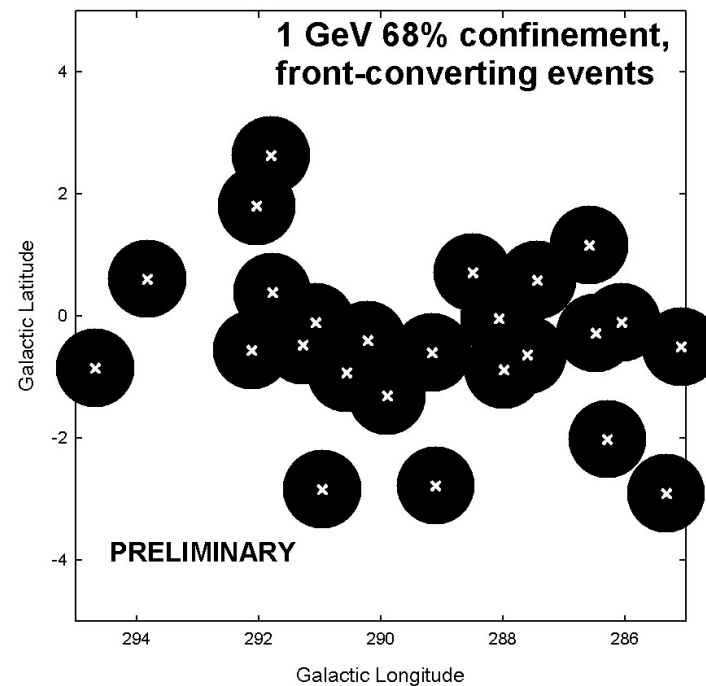


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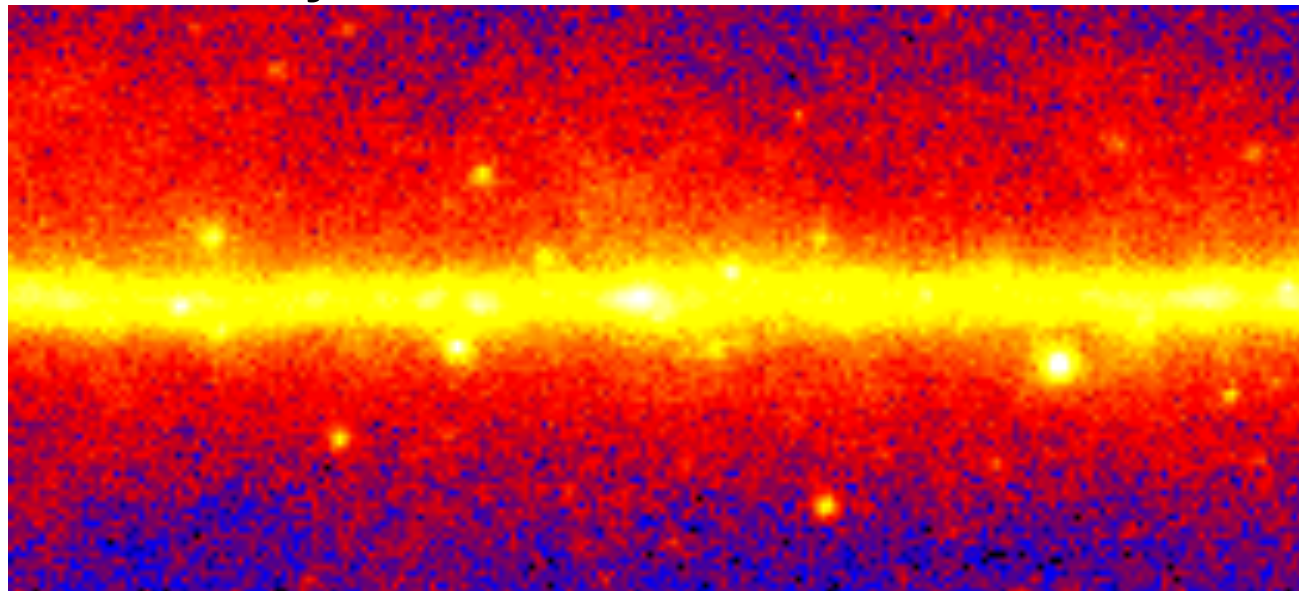
Counts map  $E > 1$  GeV



## Source Confusion

- There are many sources in the Galactic Center region
  - Many overlapping PSFs
  - Lots of soft-spectrum sources
  - The diffuse model shows some residuals compared to the large-scale diffuse emission observed in this region
- LAT catalog results for the region around the Galactic Center should be considered a good first approximation rather than a comprehensive analysis.

50° x 25°  
E>1GeV





## 32-months of LAT Sources

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~ Movie ~



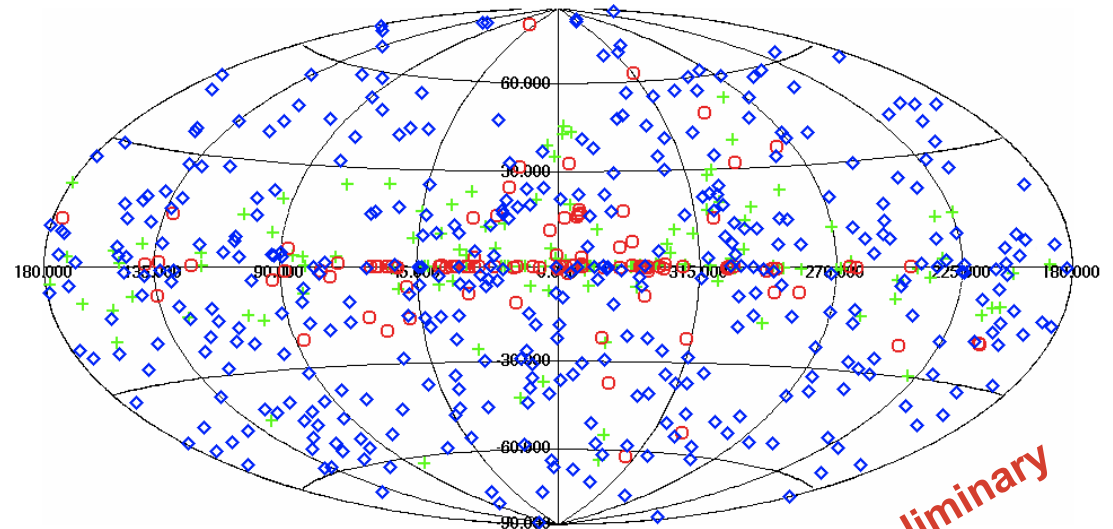
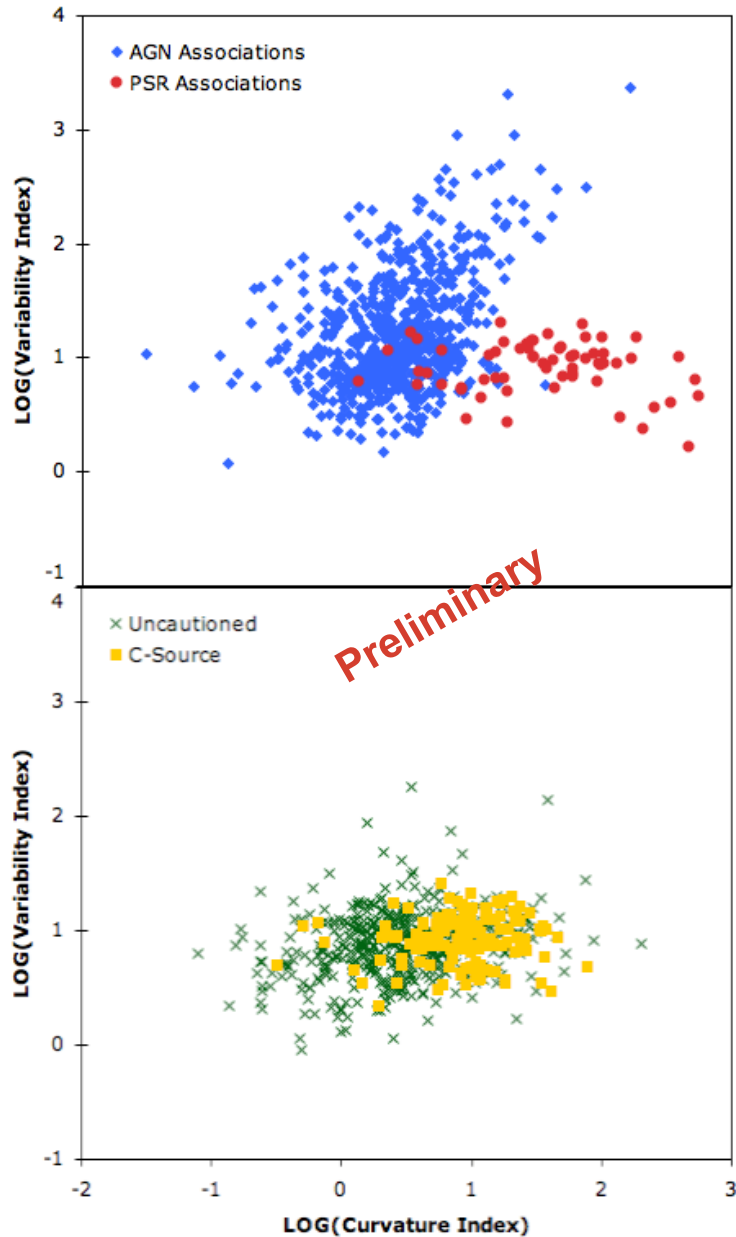
## What Else Could There Be?

- **The Upcoming 2FGL catalog has 1888 sources:**
  - **Extragalactic Sources**
    - *44% are AGN (832)*
    - *14% are 'Candidate AGN' (268)*
    - *<1% are 'Other Galaxies' (7)*
  - **Galactic Sources**
    - *6% are Pulsars (114)*
    - *3% are SNRs/PWNe (60)*
    - *<1% are Globular Clusters (11)*
    - *<1% are Binary Systems (4)*
  - **Total = 1296 sources**
- **The remaining 592 sources are still not associated!!**

**All numbers still preliminary!**

# Classifying Unassociated Sources

- Can attempt to use the gamma-ray spectral and temporal properties to determine likely source class for these 592 sources

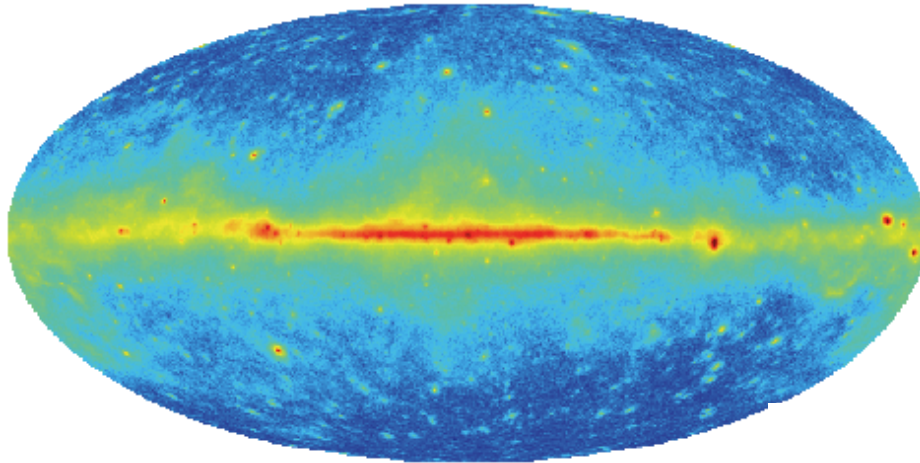


- ◆ AGN-like sources
- Pulsar-like sources
- + Unclassifiable

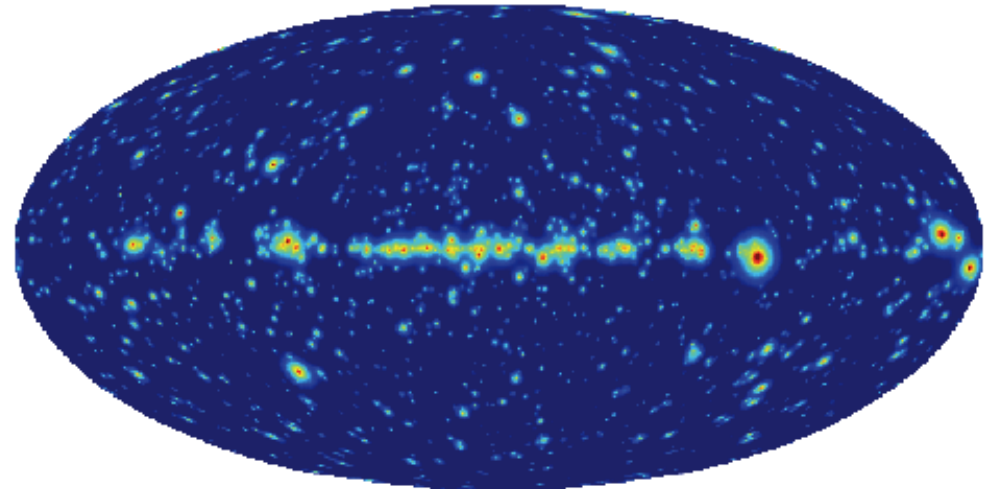
## Is that all we see?

- The point sources comprise only a subset of the data

*LAT counts above 300 MeV*



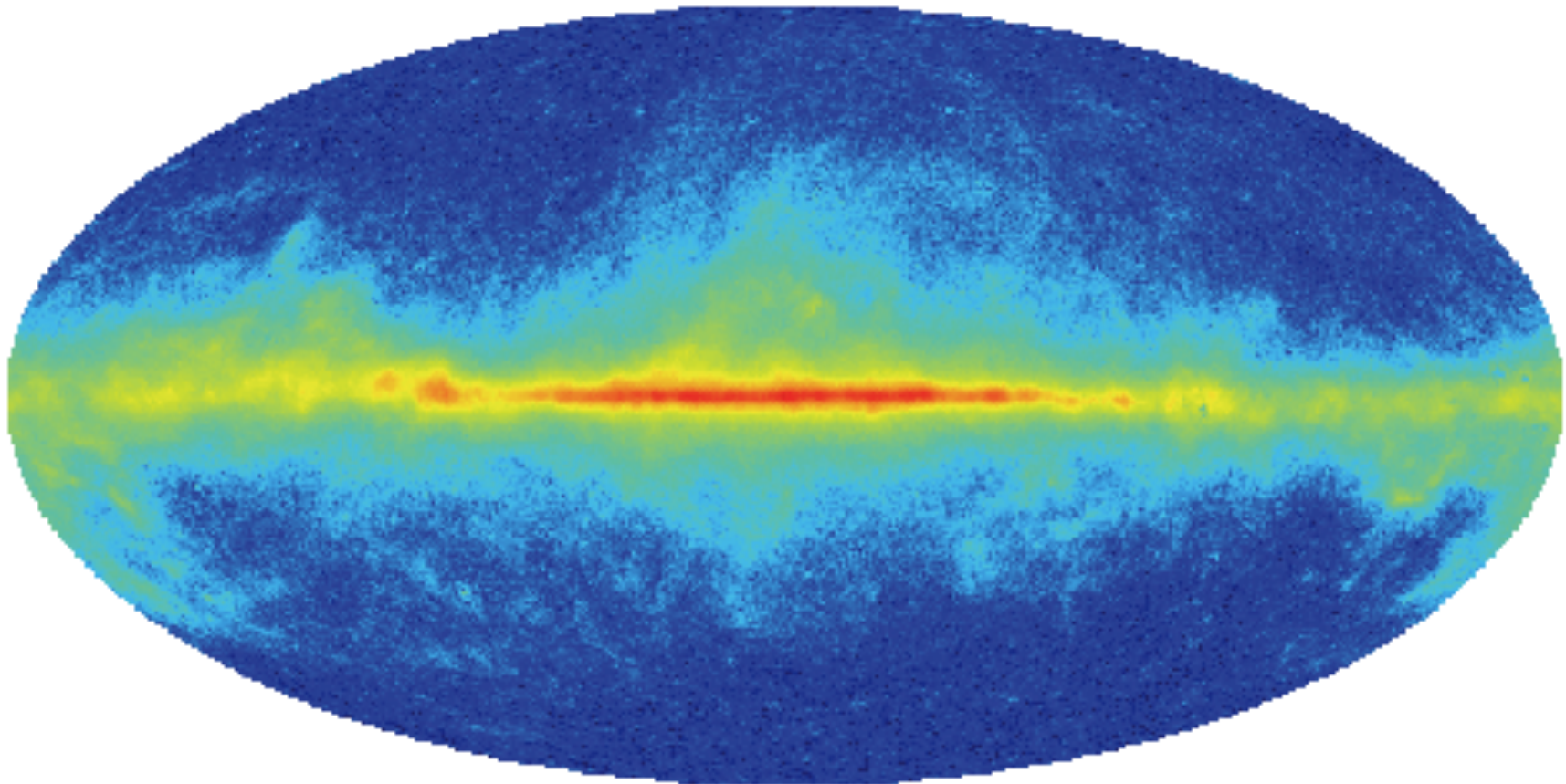
*Sources, 2FGL early version*



## What remains is 'Background'

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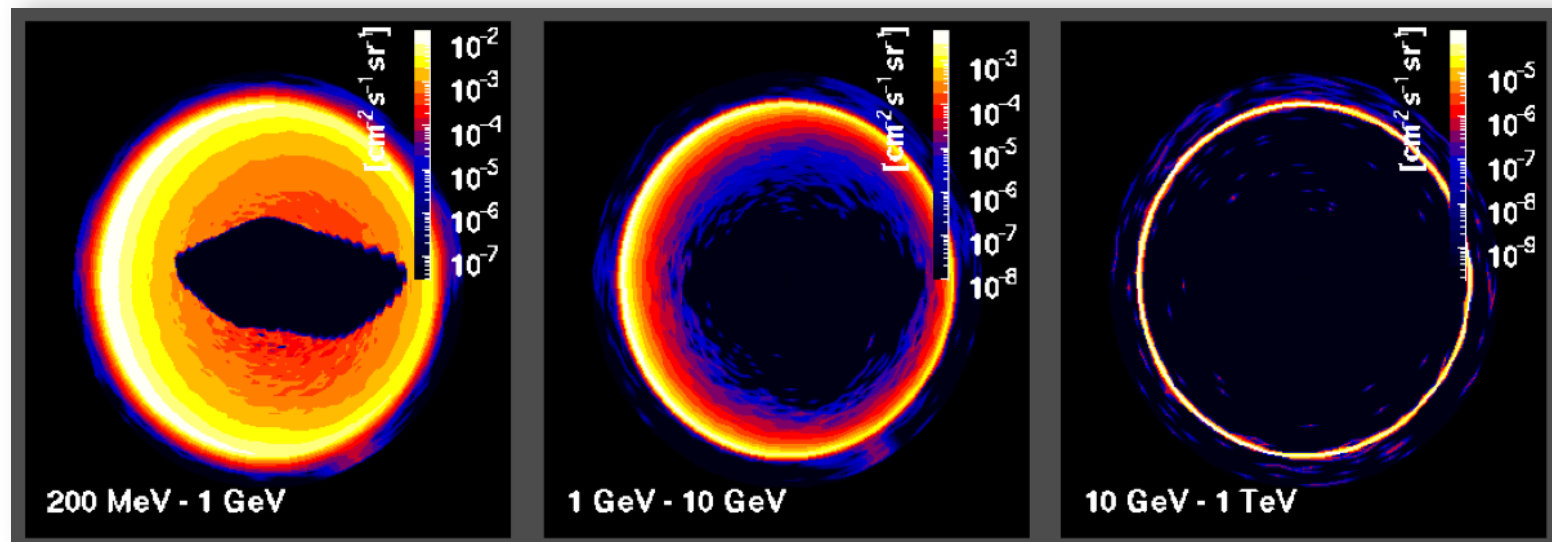
*LAT counts minus sources*





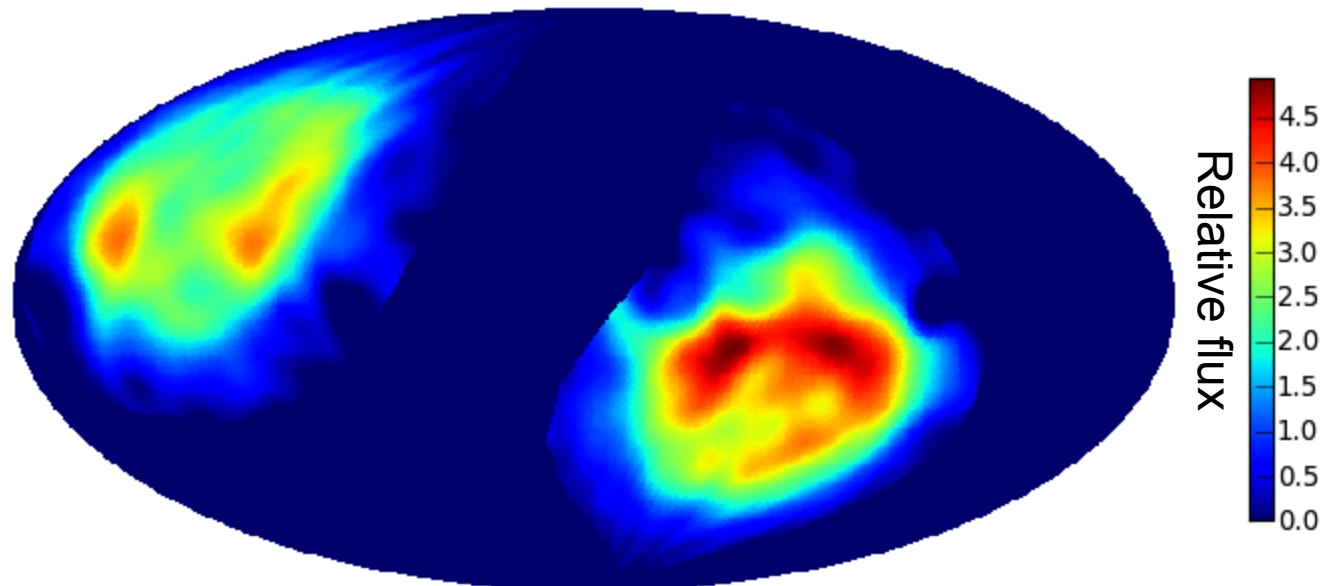
## Atmospheric gamma rays

- The Earth's limb is *bright* in gamma-rays!
  - Secondary gamma rays from cosmic ray interactions in the Earth's atmosphere
    - At Fermi's altitude, the limb is  $\sim 113$  deg from zenith
  - Far brighter than celestial sources
    - Need to remove limb gammas from analysis of celestial sources
    - Do geometry cut (standard recommendation)
      - e.g. cut on zenith angle



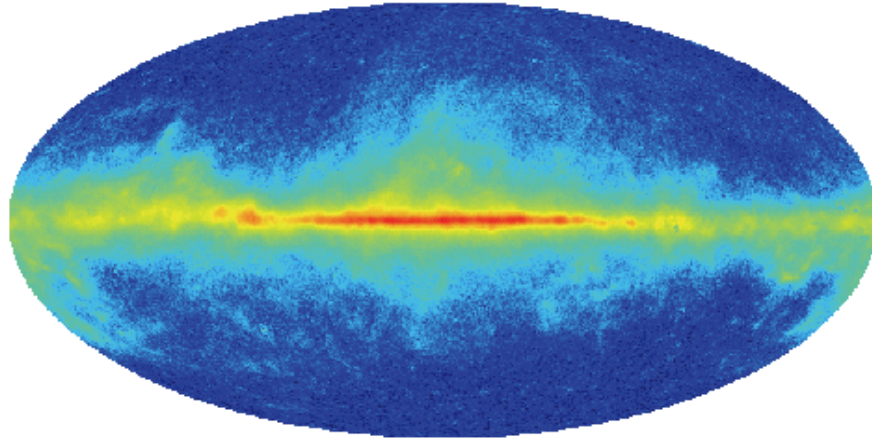
## Contribution from the Earth's limb

- **Some leakage can happen even after excluding events close to the limb**
  - Residual limb photons can be seen at the celestial poles at low energies (low energy = large PSF)
  - Can be fitted as a separate background template
  - Time-variable contribution, so average contribution may not be appropriate for time-series analysis

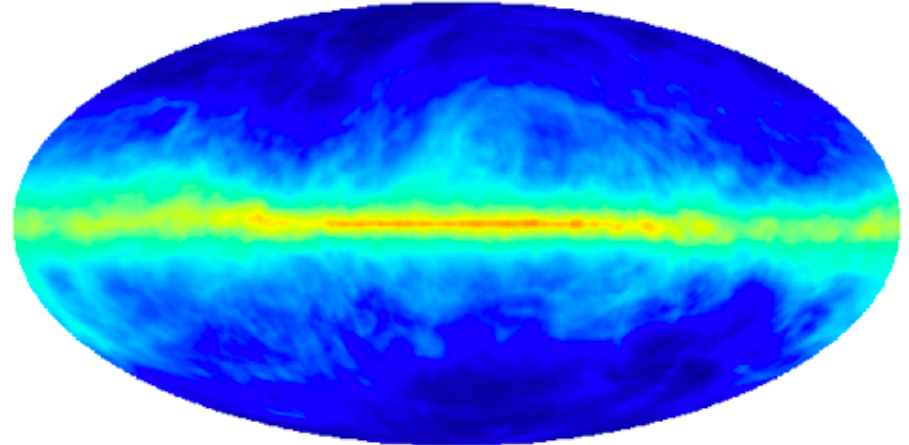


## Galactic Diffuse Component(s)

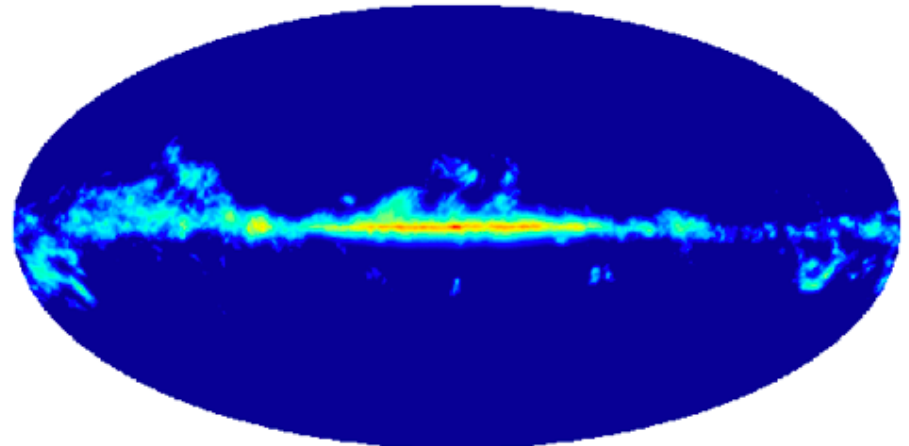
*LAT counts minus sources*



HI



CO



- **Still a lot of detail in the remaining emission**
- **Need to account for large-scale structure**
  - Requires modeling to remove
  - HI tracks ionized Hydrogen
  - CO is a tracer for neutral H<sub>2</sub>



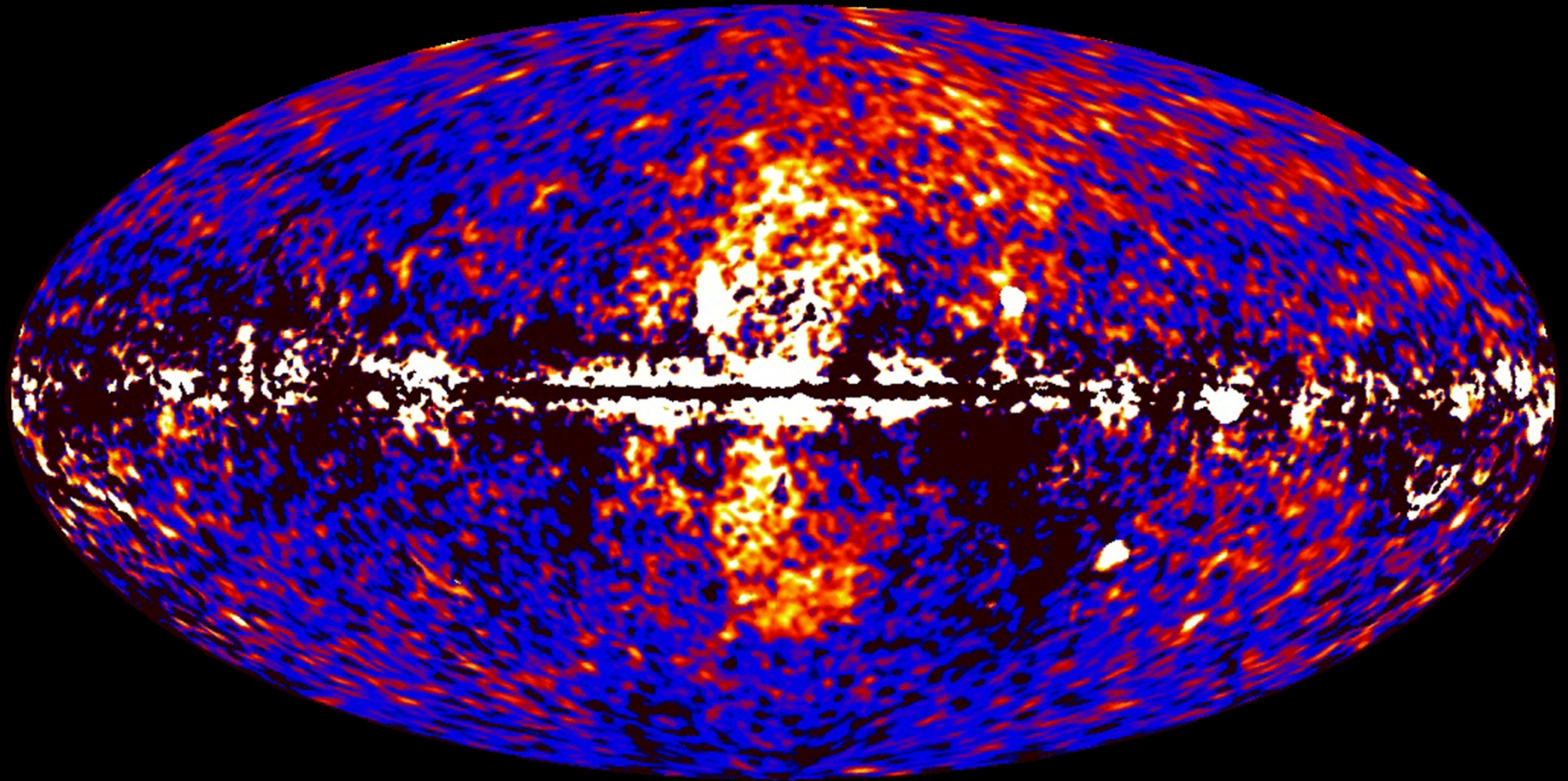
## Deconstructing the Diffuse

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~ Movie ~

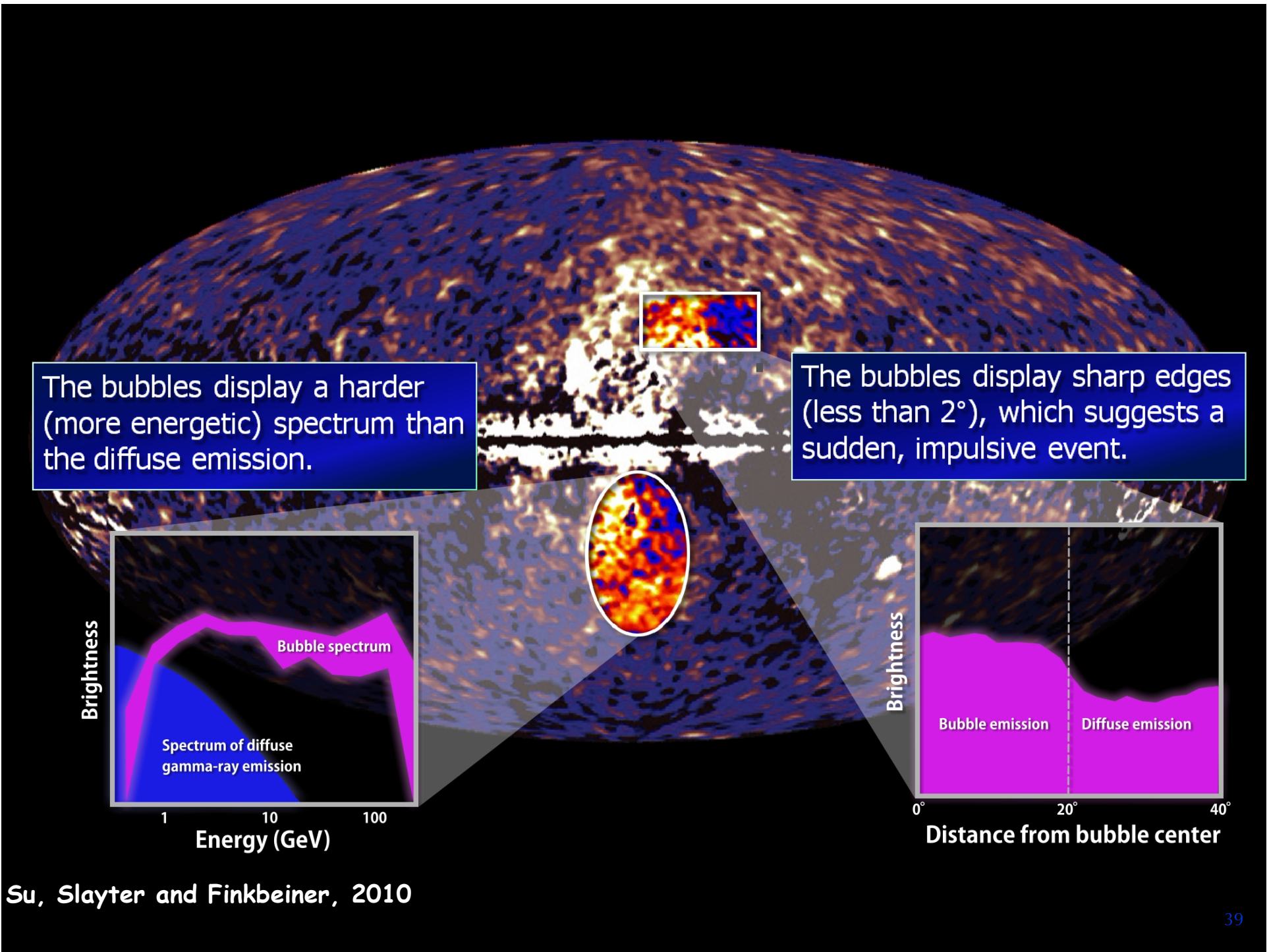


## What's Hiding in the Data



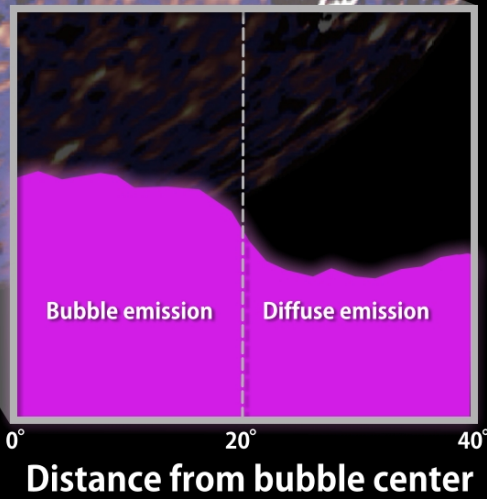
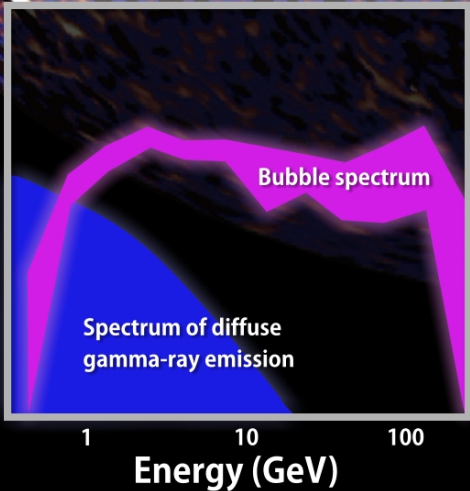
The new structure consists of enormous bubbles extending about  $50^\circ$  north and south of the galactic center.

Su, Slayter and Finkbeiner, 2010



The bubbles display a harder (more energetic) spectrum than the diffuse emission.

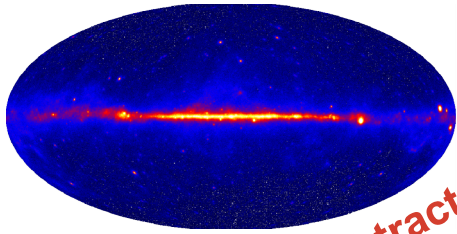
The bubbles display sharp edges (less than 2°), which suggests a sudden, impulsive event.



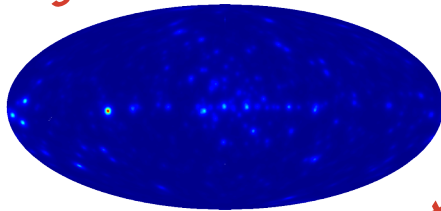


# Deriving the extragalactic diffuse spectrum

LAT Sky

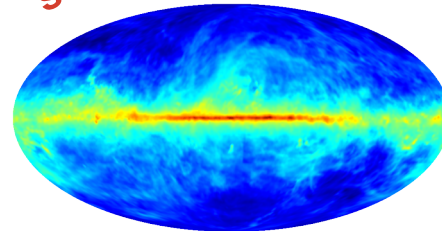


subtract



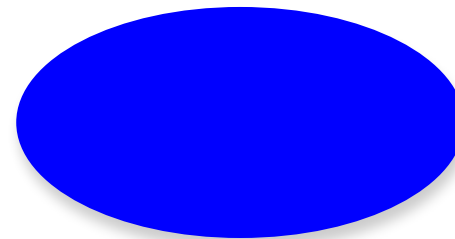
Point Sources

subtract



Galactic Diffuse Components

leaves →



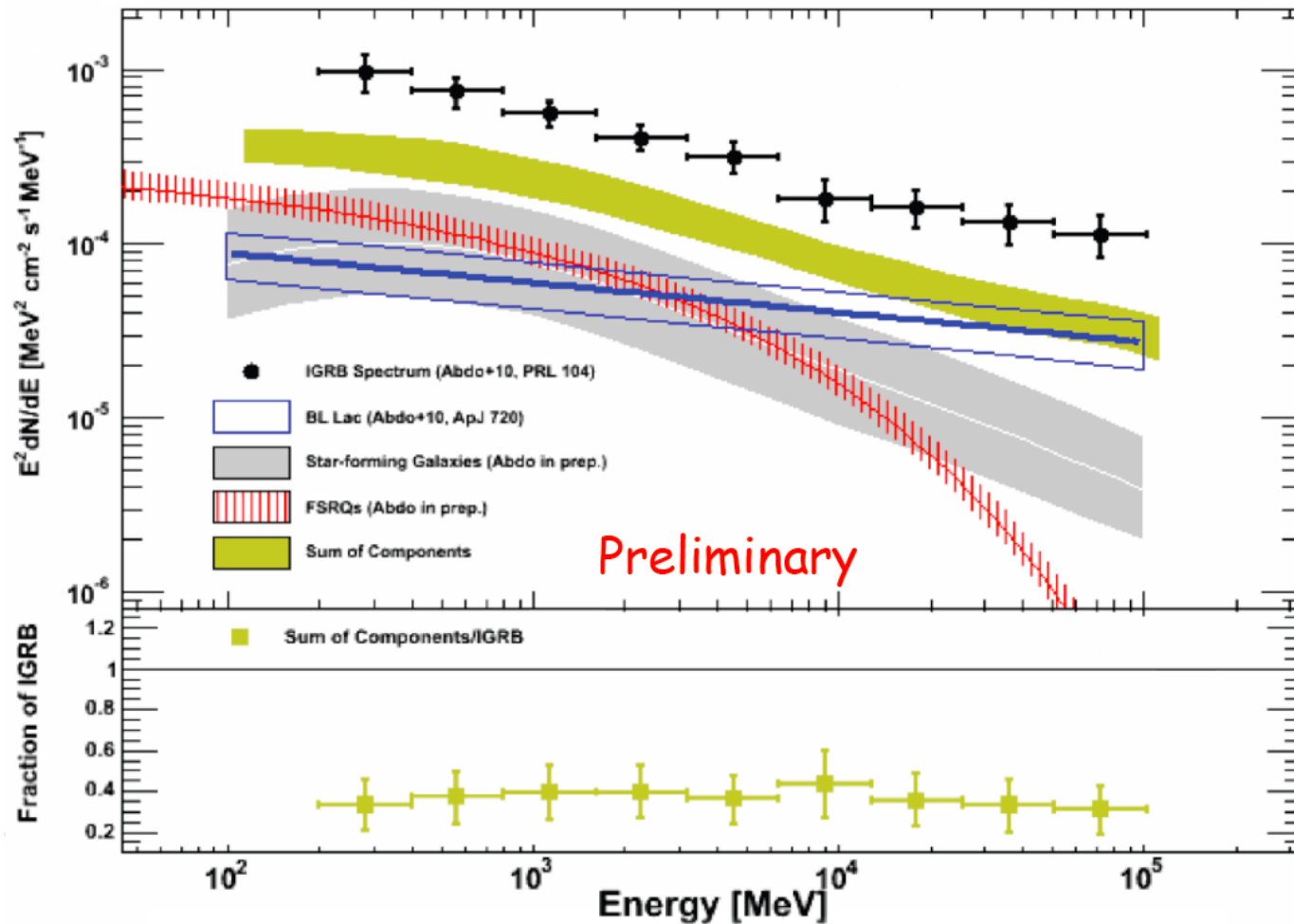
Isotropic Diffuse Emission

– Isotropic diffuse spectrum is a combination of:

- *Emission from unresolved extragalactic sources*
- *Residual charged-particle background*
- *Other signal?*

– Model the charged particles background to remove that signal

- Much, but not all, of the IGRB is accounted for
  - May provide cosmological clues





## What Doesn't the LAT See?

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- **Seyfert Galaxies**
  - They're AGN, why aren't they seen?
- **Galaxy Clusters**
  - Filled with great targets for cosmic rays, why aren't they seen?
- **Accreting X-ray Pulsars, Magnetars**
  - Seen by GBM, extreme physics, why aren't they seen?

# Dark matter searches: Nothing Yet

## Satellites

Low Galactic background and good source ID, but low statistics

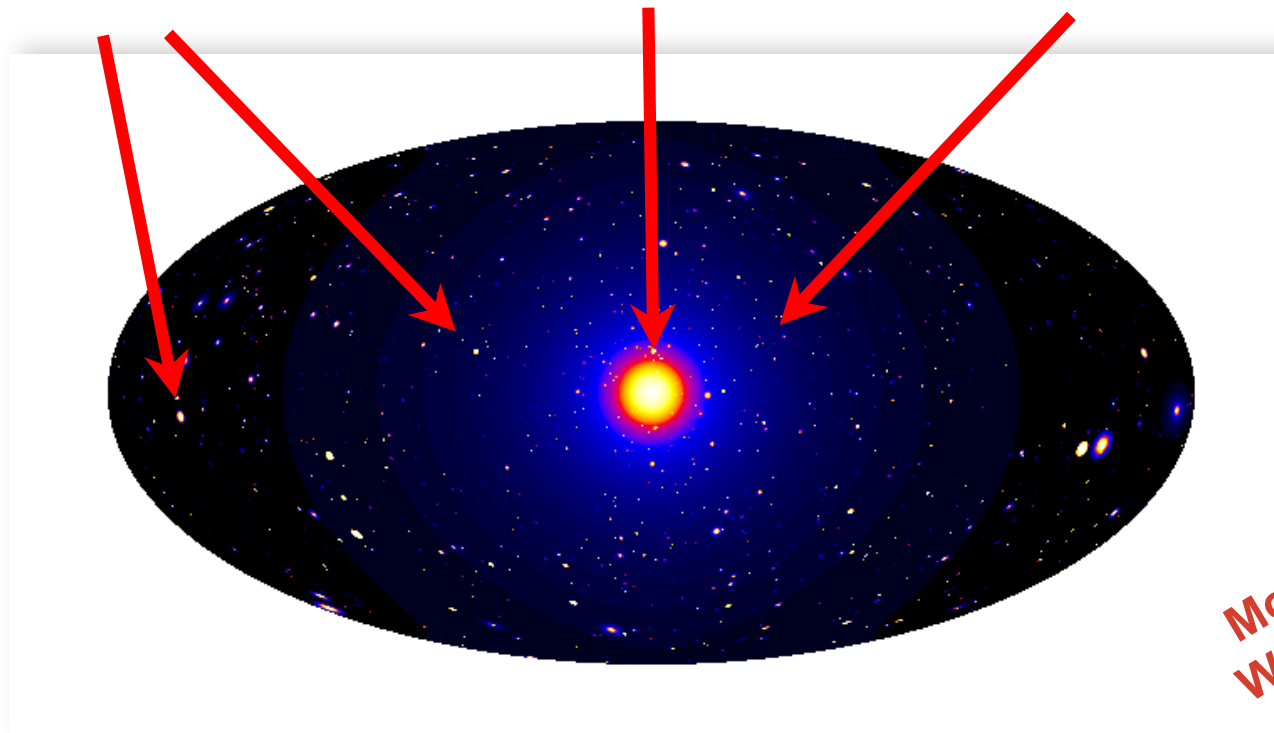
## Galactic Center

Good statistics, but source confusion and high Galactic diffuse background

## Milky Way Halo

Large statistics, but significant diffuse Galactic and extragalactic background

All-sky map of simulated gamma ray signal from DM annihilation (Pieri et al. arXiv:0908.0195)



More next  
Wednesday

**Photon statistics just beginning to reach the levels where predictions indicate DM may be detectable**

# Summary

- The LAT sees a huge variety of different source types
  - Short and long timescales
  - Variety of spectral forms
- One person's 'background' is another person's 'source of interest!'
- Separating point sources from each other and from the diffuse background is an on-going challenge
- Fermi is a significant discovery machine!
  - More of everything Gamma

**New Galactic Structures**

**Many new source types**